$7 + k\alpha 7 + (k\alpha)^{2} 7 + \cdots = \cos k(x-ct) e$   $\begin{cases} (expend in Taylor and since ka is still small) \end{cases}$ = cos k(x-ct) (1+ ka7,+(ka)2+...) Equate terms of order 1, o(ka), o(ka)... (this is what we got from  $2 = \cos k(x-ct)$  $n_2 = n_1 \cos h(x-ct) = \cos^2 h(x-ct) = \frac{1}{2} + \frac{1}{2} \cos 2h(x-ct)$  $\frac{7}{3} = \frac{7}{2} \cos k(x-ct) = \cdots$ Thus, 7 - 1/2 ka2 = a cus k(x-ct) + 1/2 (ka) a cus 2 k(x-ct)+... This is equivalent to the Kinematic Surface Condition (which implied that fluid element at Surface remains there) Choose origin so that mean Value of of is Zero:

(Reality check:)  $2 = a \cos k(x-ct) + \frac{1}{2}(ka) a \cos 2k(x-ct) + \cdots$ 

Is this egn. Consistent with Dynamic F5 Condition?

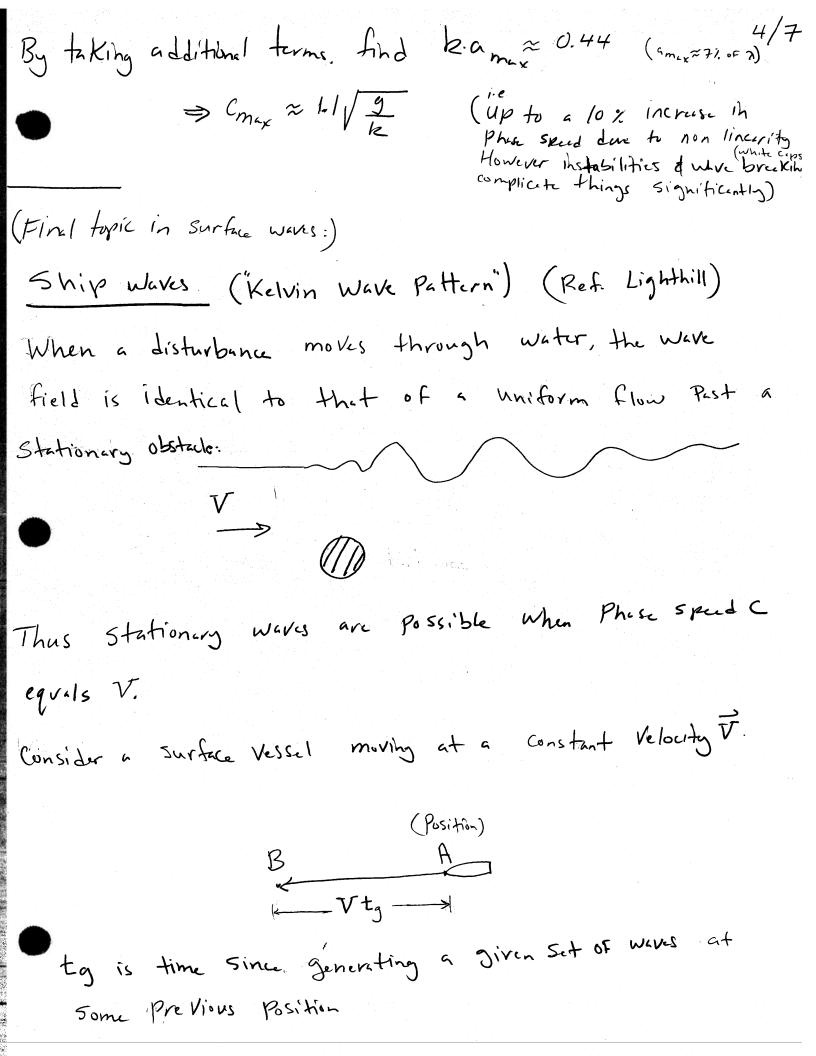
(Since at Serface.) P= Po (atmospheric passure) ...
(substituting back that the unsteady Bernoulli equi)

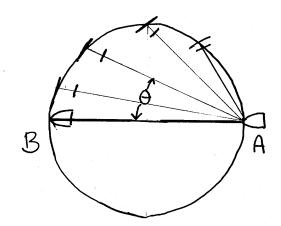
at y= y (New origin Constant =  $97 + \phi_t + \frac{1}{2}(\phi_x^2 + \phi_y^2)$ (Since our origin (Substitute in the 7:)

New Constant = 9 (a cos k(x-ct) + ...)

 $R(a Cus k(x-ct) + \cdots)$ 3/7 - kac² cosk(x-ct) e 2 ka los k (x-ct) +---0 + 1 k2a2 e + HOT Newer = { ga-kac² + k²a²c² zka} cosk(x-ct) + Constant (Since LHIS is constant and con not depend on x or t) Egn is Satisfied only if { = 0, i.e. 0 = ga - kac2 + k2a2c2 2 ka  $C^{2} = \frac{9}{k \sqrt{1 - k^{2}a^{2}}} \Rightarrow C^{2} = \frac{9}{k} \left(1 + k^{2}a^{2} + \cdots\right) \leq \ln a \cdot ka < c$ No longer just  $C=\frac{9}{12}$ , but also a Small correction. So, the dispersion relation now depends on amplitude. Using C= W => (W= VOR (1+ 1 ka2+--) The correction to of makes crests sharper & stratches out the troughs.

By a local solution can show angle, at now a sharp crist => 120° as a -> a max (see Lamb or Yih)

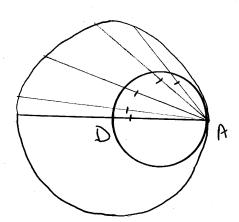




Circle AB would contain the waves traveling with the Ship, different wave lengths going at different speeds, the longest wavelength I max corresponding to  $V=C=\sqrt{\frac{9}{R}}$  traveling in the direction of the Ship, all others moving at

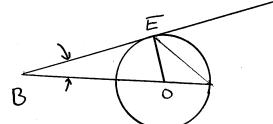
C= V cos 0, i.e. Shortest whees moving laterally.

But we know that the waves associated with a Ship ave of gravity type  $(c=\sqrt{\frac{q}{R}})$  and a Packet of gravity waves moves at group velocity  $Cg=\frac{1}{2}c$ . Thus, Ship waves can only  $g_0 \pm as$  far as depicted above.



Therefore, Waves traveling with the Ship that wen generated when Ship was at A will be contained in Circle AD

so, when the Snip. gets to B, its waves must be contained inside a triangle:



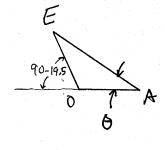
So the "Kelvin Ship-Wave" wedge has a half angle oF:

$$Sin'\left(\frac{1}{3}\right) = 19.5^{\circ}$$

The waves traveling at the Center of the wedge have a Phose Skeed  $C = V = \sqrt{\frac{9}{k}} = \sqrt{\frac{92}{2\pi}}$ 

> 2 mx 27 V2. Waves at the boundary of wedge (E)

are traveling at angle 0



 $\theta = \frac{1}{2} (90 - 19.5^{\circ}) = 35.3^{\circ}$ 

Thus, their Phase speed  $c = V \cos \theta = 0.816 V$ .
and their  $\lambda = \frac{2}{3} \lambda_{max}$ 

Overall Wave Pattern: T.P. Wote

T.P.) FIDTI Wares,)

Recent Observations. (SAR images) TIP Visual
Munk, Scully-Power & Zachariasen Proc. Royal Soc. Lundon A 412
231-254 (1987) Narrow V & Wedges.

## Fig 70 Light hill

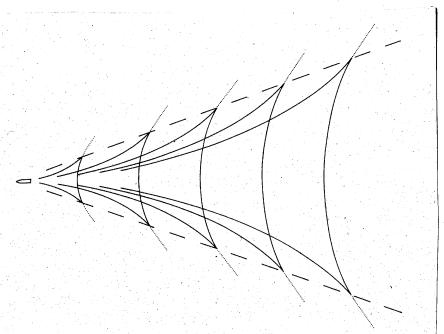


Figure 70. Plain lines: Kelvin ship-wave pattern. Broken lines: boundary of Kelvin wedge. Dotted lines: extension of waves beyond the Kelvin wedge indicated by the theory of sections 4.11 and 4.12.

## Fig.71 Lighthill (P. 278)

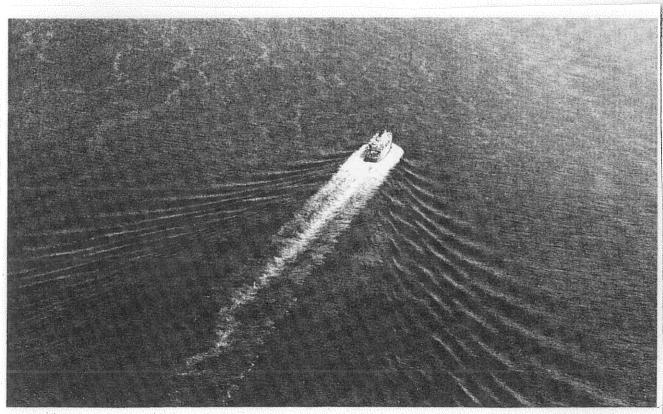


Figure 71. An observed ship-wave pattern. [Courtesy of Aerofilms Ltd.]