## Introduction into CAMB

to CAMB-Version 10/2013

all texts within "" are extracted out of the CAMB-code or from http://camb.info/readme.html!

Martina Schwind
Institute for Theoretical Physics Heidelberg

## Overview

- Running CAMB
- Fortran: program, procedure, module
- The program-files of CAMB:
  - main-program-files
    - can be used to run different tests
  - useful program-files (independent of cosmology)
  - program-files with cosmology inside
    - steering modules
    - Recombination, Reionization, Initial Powerspectrum, Bispectrum, Halofit, Lensing
    - modules, equations
- Example: inidriver.f90: program driver

## Running CAMB

Compile with the makefile: *make* 

includes makefile\_main

choose FFLAGS-Command inside the makefile fitting to Your technical environment

Run with ./camb <param-filename>

Compile again with: make clean && make

## Fortran: program, procedure, module

## program:

- consists of a sequence of statements: these statements are written on lines that may contain from 0 to 132 characters.
- two kinds of procedures:
  - subroutine: may be used to perform any computation and is invoked by executing a *call* statements.
  - function: looks much like a Fortran program, except that it begins with the keyword *function* instead of the keyword *program*.

## module:

- provide another way of sharing constants, variables, and type definitions.
- also provide a way of sharing procedures, as well as data.
- is a program unit that is not executed directly, but contains data specifications and procedures that may be utilized by other program units via the use-statement.

## All program-files of CAMB

bessels.f90		camb.f90	cmbmain.f90	constants.f90
cosmorec.F90		equations.f90	halofit.f90	hyrec.F90
inidriver.F90		inifile.f90	lensing.f90	Matrix_utils.F90
modules.f90		power_tilt.f90	recfast.f90	reionization.f90
SeparableBispectru .F90	m	sigma8.f90	subroutines.f90	tester.f90
utils.F90		writefits.f90		

22 fortran-program-files

## main-program-files of CAMB:

drivers: can be used to run different tests

bessels.f90	camb.f90	cmbmain.f90	constants.f90
cosmorec.F90	equations.f90	halofit.f90	hyrec.F90
inidriver.F90	inifile.f90	lensing.f90	Matrix_utils.F90
modules.f90	power_tilt.f90	recfast.f90	reionization.f90
SeparableBispectrum .F90	sigma8.f90	subroutines.f90	tester.f90
utils.F90	writefits.f90		

• 1 testdriver, 2 examples for calling CAMB

## main-program-files of CAMB:

drivers: can be used to run different tests



sigma8.f90

program GetSigma8

tester.f90

program tester

- Inidriver: "Reads in parameters from a file of name/value pairs and calls CAMB.
   Modify this file to generate grids of models, change the parameterization, etc."
- **sigma8, tester**: "Sample programs ... are supplied showing how to use CAMB from your own programs."

# "Useful" program-files of CAMB "without cosmological logic"

bessels.f90	camb.f90	cmbmain.f90	constants.f90
cosmorec.F90	equations.f90	halofit.f90	hyrec.F90
inidriver.F90	inifile.f90	lensing.f90	Matrix_utils.F90
modules.f90	power_tilt.f90	recfast.f90	reionization.f90
SeparableBispectrum .F90	sigma8.f90	subroutines.f90	tester.f90
utils.F90	writefits.f90		

• 6 program-files with useful routines: only use them - never change them!

# "Useful" program-files of CAMB "without cosmological logic"

### bessels.f90

## module SpherBessels

bessels.f90: "Module to calculate spherical and hyperspherical Bessel functions. Hyper-spherical functions generated by use of either the recursion relation or Kosowsky's WKB approximation. Based on Arthur Kosowsky's "hyperjl.c"."

## inifile.f90

## module IniFile

"Module to read in name/value pairs from a file, with each line of the form line 'name = value' Should correctly interpret FITS headers"

## utils.F90

module Ranges
module Lists
module AMLutils
MODULE Ziggurat
module Random
01/27/14

## writefits.f90

"subroutine to export Cls in FITS format for HEALPix 1.2"

writefits.f90: "Subroutine WriteFitsCls that uses HEALPIX routines to output power spectrum in FITS format."

## constants.f90

module Precision module constants module Errors

Matrix\_utils.F90

module MatrixUtils

## subroutines.f90

"General numerical routines and global accuracy. Includes modified dverk for CAMB."

**subroutines.f90**: "Various subroutines for interpolation, and modified Runge-Kutta dverk for parallelized evolution."

# program-files of CAMB with cosmology inside

camb.f90 cmbmain.f90 bessels.f90 constants.f90 equations.f90 halofit.f90 cosmorec.F90 hyrec.F90 lensing.f90 inidriver.F90 inifile.f90 Matrix\_utils.F90 modules.f90 power\_tilt.f90 recfast.f90 reionization.f90 SeparableBispectrum subroutines.f90 tester.f90 sigma8.f90 .F90 utils.F90 writefits.f90

12 program-files with cosmology inside

## program-files of CAMB: cosmology inside: **steering modules**

 cmbmain.f90:"The main subroutine that does integrations, etc. Encompasses CMBFAST's cmbflat and cmbopen."

## camb.f90

module CAMB

SUBROUTINES/ FUNCTIONS

#### **PUBLIC**

CAMB\_SetDefParams,
CAMB\_GetResults,
CAMB\_cleanup,
CAMB\_ValidateParams,
CAMB\_GetTransfers,
CAMB\_GetCls,
CAMB\_GetAge,
CAMB\_GetZreFromTau,
CAMB\_InitCAMBdata,
CAMB\_FreeCAMBdata,
CAMB\_TransfersToPowers

cmbmain.f90

module CAMBmain

SUBROUTINES/ FUNCTIONS

#### **PUBLIC**

→cmbmain, ClTransferToCl, InitVars

camb.f90: "Main wrapper routines for running CAMB in your programs.

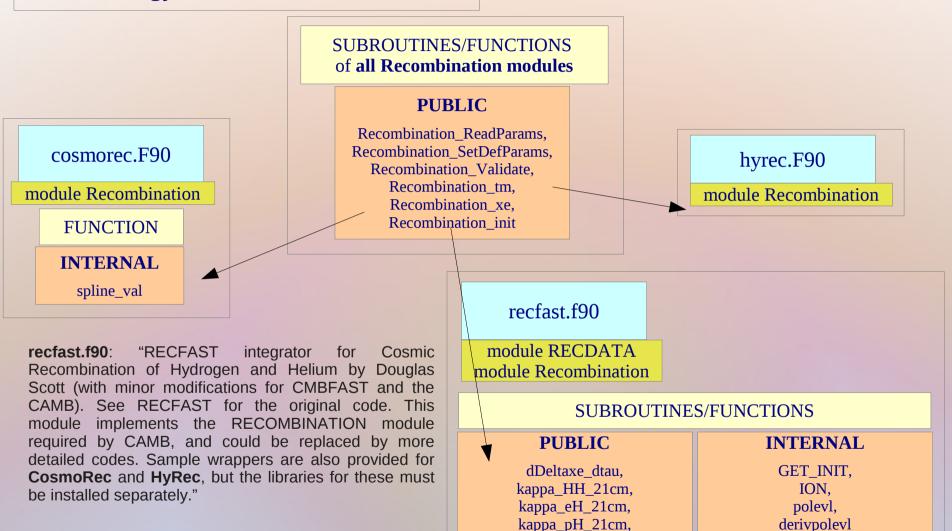
- Add "use camb" to your programs and call CAMB\_GetResults to generate output from a set of model parameters (specified in the CAMBparams type defined at the top of modules.f90).
- You can **call CAMB\_ValidateParams(P)** to check that the parameter set is valid.
- use CAMB\_GetAge to compute the age of a model in gigayears, and CAMB\_GetCls to retrieve the computed Cls.
- The results can also be accessed directly using the arrays in the ModelData module (defined in modules.f90)."

#### **INTERNAL**

CalcLimberScalCls.

GetLimberTransfers. SourceToTransfers, InitTransfer. DoSourcek. GetSourceMem. FreeSourceMem. SetkValuesForSources. SetClosedkValuesFromArr. CalcScalarSources. CalcTensorSources. CalcVectorSources. TransferOut. GetTransfer. MakeNonlinearSources. InitSourceInterpolation, SetkValuesForInt. InterpolateSources, IntegrationVars\_Init, DoSourceIntegration, DoFlatIntegration, IntegrateSourcesBessels, DoRangeInt, DoRangeIntTensor, GetInitPowerArrayVec, GetInitPowerArrayTens. CalcScalCls. (CalcScalCls2?), CalcTensCls. CalcVecCls. **InterpolateCls** 

# program-files of CAMB: cosmology inside: **Recombination**



# program-files of CAMB cosmology inside: **InitialPower**, **Reionization**

power\_tilt.f90: "This file defines a module called InitialPower that returns the initial power spectra. Change this file to use your own initial power spectrum, change how the spectra are parameterized, or to change how the Cls are normalized. Comments in the code explain this further."

power\_tilt.f90

module InitialPower
SUBROUTINES/
FUNCTIONS

## **PUBLIC**

lens\_Cls,
SetDefPowerParams,
InitializePowers,
ScalarPower,
TensorPower,
Power\_Descript,
InitialPower\_ReadParams

reionization.f90: "This file a module called defines Reionization that parameterizes the reionization history and supplies function Reionization xe that gives xe as a function of redshift. Optical depth input parameters are mapped into zre (defined as where xe is half its maximum (ex second He reionization)) using a binary search. See the CAMB discussion. notes for This module should be easily modifiable for alternative reionization models."

## reionization.f90

module Reionization

SUBROUTINES/ FUNCTIONS

#### INTERNAL

Reionization\_xe,
Reionization\_timesteps,
Reionization\_ReadParams,
Reionization\_SetParamsForZre,
Reionization\_Init,
Reionization\_SetDefParams,
Reionization\_Validate,
Reionization\_doptdepth\_dz,
Reionization\_GetOptDepth,
Reionization\_zreFromOptDepth,
Reionization\_SetFromOptDepth

## program-files of CAMB cosmology inside: NonLinear, Bispectrum, lensing

### SUBROUTINES/FUNCTIONS

### **INTERNAL**

NonLinear GetNonLinRatios, halofit. wint, omega m, omega v

#### Not within the module

NonLinear GetRatios. NonLinear GetRatios all

halofit.f90

module NonLinear

halofit.f90: "Implements the NonLinear module, to calculate non linear scalings of the matter power spectrum as a function of redshift. This module can be replaced to use a different non-linear fitting method if desired."

and a lensing power spectrum."

## lensing.f90

module lensing SUBROUTINES/ **FUNCTIONS** 

## INTERNAL

CorrFuncFullSky, CorrFuncFullSkyImpl, CorrFuncFlatSky, BadHarmonic, GetBessels. bessi, BESSIO. **BESSI1** 

## **PUBLIC**

lens Cls, **BESSI** 

lensing.f90: "Lensing module for computing the lensed CMB power spectra from the unlensed spectra

## SeparableBispectrum .F90

module Bispectrum

SUBROUTINES/ **FUNCTIONS** 

## **INTERNAL**

InitBesselDerivs, NonGauss\_l\_r\_localOpt, NonGauss 1 r. GetBispectrum, NonGauss\_deriv\_l\_r, Bispectrum\_SetDefParams, Bispectrum ReadParams

### SeparableBispectrum.f90:

"Implements calculation simple separable primordial bispectra, specifically the local constant fNL model, and the CMB lensing bispectrum due to the linear temperature and polarization cross-correlation with the lensing potential."

# program-files of CAMB cosmology inside: equations.f90

## module LambdaGeneral

### SUBROUTINES/FUNCTIONS

#### **INTERNAL**

subroutine DarkEnergy\_ReadParams

## **INTERNAL**

subroutine init\_background, function dtauda(a)

## equations.f90

module LambdaGeneral module GaugeInterface

module LambdaGeneral is **used** in camb.f90 cmbmain.f90 halofit.f90 inidriver.f90 module GaugeInterface is **used** in camb.f90 cmbmain.f90

**equations.f90**: "Files containing background and perturbation evolution equations. The perturbations equations used are derived in the covariant approach, fixing to the CDM (zero acceleration) frame, which are essentially equivalent to the synchronous gauge equations.

The file defines a module called "GaugeInterface" which provides the necessary perturbation calculation routines for "cmbmain".

The subroutine **dtauda**(a) returns dt/da and is used wherever the background evolution is needed. It can be modified for different backgrounds. You may also need to change the GetOmegak routine if you add additional components, and can edit the init background routine to do additional initialization.

**outtransf** writes out the matter transfer functions.

The "output" subroutine computes the scalar sources at a given time for a given wavenumber. These are the temperature, E polarization and (if doing lensing) the lensing source. By editing the equation for the lensing source it should be straightforward to compute power spectra for other matter tracers, e.g. for cross-correlation with the CMB. The lensing power spectrum is automatically computed if DoLensing=T."

## module GaugeInterface

### SUBROUTINES/FUNCTIONS

#### **INTERNAL**

(cm) subroutine GetNumEqns,

(cm) subroutine GaugeInterface\_Init, subroutine GaugeInterface\_ScalEv,

(cm) subroutine initial,

(cm) subroutine GaugeInterface\_EvolveScal,

(cm) subroutine derivs,

(cm) subroutine output,

(cm) subroutine outtransf

(cm) subroutine initialy,

(cm) subroutine derivsv (with dverk),

(cm) subroutine outputy,

(cm) subroutine initialt,

(cm) subroutine GaugeInterface\_EvolveTens, subroutine derivst.

(cm) subroutine outputt,

(cm) subroutine initial,

(cm) subroutine GaugeInterface\_EvolveScal,

(cm) subroutine outtransf

function DeltaTimeMaxed, function next\_nu\_nq,

subroutine SetupScalarArrayIndices, subroutine CopyScalarVariableArray, subroutine SetupTensorArrayIndices, subroutine CopyTensorVariableArray,

subroutine SwitchToMassiveNuApprox,
subroutine MassiveNuVars,,
subroutine MassiveNuVarsOut,
subroutine Nu\_Integrate\_L012,
subroutine Nu\_pinudot,
function Nu\_pi,
subroutine Nu\_Intvsq

# program-files of CAMB cosmology inside: equations\_quint.f90

### module LambdaGeneral

### SUBROUTINES/FUNCTIONS

### **INTERNAL**

subroutine DarkEnergy\_ReadParams

## equations\_quint.f90

module LambdaGeneral module GaugeInterface *additional:* module Quint

## module Quint

### SUBROUTINES/FUNCTIONS

## **INTERNAL**

function Vofphi, subroutine EvolveBackground, function Quint\_GetOmegaFromInitial, function Quint\_phidot\_start, subroutine Quint\_init\_background, subroutine Quint\_ValsAta **equations\_quint.f90**: "There is also a more general quintessence module that lets you specific a single scalar field potential. It will need customization for different potentials, and changes to the way initial conditions are set if trying to use a tracker model. You can download the modified equations module for CAMB. It is not very well tested. "

**INTERNAL** 

subroutine init background,

function dtauda(a)

module Quint is **used** only in equations.f90

module LambdaGeneral
is **used** in
camb.f90
cmbmain.f90
halofit.f90
inidriver.f90

module GaugeInterface is **used** in camb.f90 cmbmain.f90

## module GaugeInterface

### SUBROUTINES/FUNCTIONS

#### INTERNAL

(cm) subroutine GetNumEqns,

(cm) subroutine GaugeInterface\_Init

(cm) subroutine initial,

(cm) subroutine GaugeInterface\_EvolveScal,

(cm) subroutine derivs,

(cm) subroutine output,

(cm) subroutine outtransf

(cm) subroutine initialy, (cm) subroutine derivsy (with dverk), (cm) subroutine outputy,

(cm) subroutine initialt,
(cm) subroutine GaugeInterface\_EvolveTens,
subroutine derivst,
(cm) subroutine outputt,

(cm) subroutine initial, (cm) subroutine GaugeInterface\_EvolveScal, (cm) subroutine outtransf

> subroutine GaugeInterface\_ScalEv, function DeltaTimeMaxed, function next\_nu\_nq,

subroutine SetupScalarArrayIndices, subroutine CopyScalarVariableArray, subroutine SetupTensorArrayIndices, subroutine CopyTensorVariableArray,

subroutine SwitchToMassiveNuApprox, subroutine MassiveNuVars,, subroutine MassiveNuVarsOut, subroutine Nu\_Integrate\_L012, subroutine Nu\_pinudot, function Nu\_pi, subroutine Nu\_Intvsq

# program-files of CAMB cosmology inside: part 3

## module LambdaGeneral

### SUBROUTINES/FUNCTIONS

### **INTERNAL**

subroutine DarkEnergy\_ReadParams

### **INTERNAL**

subroutine init\_background, function dtauda(a)

## equations.f90

module LambdaGeneral module GaugeInterface

**equations\_ppf.f90**: " is an alternative module that allows evolving dark energy crossing w=-1."

equations\_quint.f90: "There is also a more general quintessence module that lets you specific a single scalar field potential. It will need customization for different potentials, and changes to the way initial conditions are set if trying to use a tracker model. You can download the modified equations module for CAMB. It is not very well tested. "

**equations.f90**: "Files containing background and perturbation evolution equations. The perturbations equations used are derived in the covariant approach, fixing to the CDM (zero acceleration) frame, which are essentially equivalent to the synchronous gauge equations.

The file defines a module called "GaugeInterface" which provides the necessary perturbation calculation routines for "cmbmain".

The subroutine **dtauda**(a) returns dt/da and is used wherever the background evolution is needed. It can be modified for different backgrounds. You may also need to change the GetOmegak routine if you add additional components, and can edit the init background routine to do additional initialization.

outtransf writes out the matter transfer functions.

The "output" subroutine computes the scalar sources at a given time for a given wavenumber. These are the temperature, E polarization and (if doing lensing) the lensing source. By editing the equation for the lensing source it should be straightforward to compute power spectra for other matter tracers, e.g. for cross-correlation with the CMB. The lensing power spectrum is automatically computed if DoLensing=T."

## module GaugeInterface

### SUBROUTINES/FUNCTIONS

#### **INTERNAL**

subroutine GaugeInterface\_ScalEv, function next\_nu\_nq, subroutine GaugeInterface\_EvolveTens, function DeltaTimeMaxed, subroutine GaugeInterface\_Init,

subroutine SetupScalarArrayIndices, subroutine CopyScalarVariableArray, subroutine SetupTensorArrayIndices, subroutine CopyTensorVariableArray,

subroutine GetNumEqns,

subroutine SwitchToMassiveNuApprox, subroutine MassiveNuVars,, subroutine MassiveNuVarsOut, subroutine Nu\_Integrate\_L012, subroutine Nu\_pinudot, function Nu\_pi, subroutine Nu\_Intvsq

> subroutine initial, subroutine initialy, subroutine initialt.

subroutine derivs, subroutine derivst, subroutine derivst,

subroutine output, subroutine outputv, subroutine outputt,

subroutine outtransf

## program-files of CAMB cosmology inside: part 3

**equations\_ppf.f90**: " is an alternative module that allows evolving dark energy crossing w=-1."

**equations\_quint.f90:** "There is also a more general quintessence module that lets you specific a single scalar field potential. It will need customization for different potentials, and changes to the way initial conditions are set if trying to use a tracker model. You can download the modified equations module for CAMB. It is not very well tested. "

## equations.f90

module LambdaGeneral module GaugeInterface

**equations.f90:** "Files containing background and perturbation evolution equations. The perturbations equations used are derived in the covariant approach, fixing to the CDM (zero acceleration) frame, which are essentially equivalent to the synchronous gauge equations.

The file defines a module called "GaugeInterface" which provides the necessary perturbation calculation routines for "cmbmain".

The subroutine **dtauda**(a) returns dt/da and is used wherever the background evolution is needed. It can be modified for different backgrounds. You may also need to change the GetOmegak routine if you add additional components, and can edit the init\_background routine to do additional initialization.

outtransf writes out the matter transfer functions.

The "output" subroutine computes the scalar sources at a given time for a given wavenumber. These are the temperature, E polarization and (if doing lensing) the lensing source. By editing the equation for the lensing source it should be straightforward to compute power spectra for other matter tracers, e.g. for cross-correlation with the CMB. The lensing power spectrum is automatically computed if DoLensing=T."

## modules.f90

module ModelParams module lvalues module ModelData module MassiveNu module Transfer module ThermoData

**modules.f90:** "Various modules used by the other parts of the program, Module "ModelParams" contains most of the model parameters. Boolean vars flat, open and closed determine the model type."

## SUBROUTINES/FUNCTIONS

PUBLIC	INTERNAL
d,	G,
k,	I,
k,	р,
,	dl

## EXAMPLE: **inidriver.f90**: program driver

camb.f90 module CAMB SUBROUTINES/ **FUNCTIONS PUBLIC** CAMB SetDefParams, CAMB GetResults, CAMB\_cleanup, CAMB ValidateParams, CAMB GetTransfers, CAMB\_GetCls, CAMB GetAge, CAMB GetZreFromTau, CAMB InitCAMBdata, CAMB FreeCAMBdata. CAMB\_TransfersToPowers

cmbmain.f90 **INTERNAL** CalcLimberScalCls. module CAMBmain GetLimberTransfers. SourceToTransfers, SUBROUTINES/ InitTransfer. **FUNCTIONS →** DoSourcek, **PUBLIC** GetSourceMem, FreeSourceMem. →cmbmain, SetkValuesForSources, ClTransferToCl. SetClosedkValuesFromArr. InitVars CalcScalarSources. CalcTensorSources. CalcVectorSources, TransferOut. GetTransfer. MakeNonlinearSources, InitSourceInterpolation, SetkValuesForInt. InterpolateSources, IntegrationVars\_Init, DoSourceIntegration, DoFlatIntegration, IntegrateSourcesBessels, DoRangeInt, DoRangeIntTensor, GetInitPowerArravVec, GetInitPowerArrayTens, CalcScalCls. (CalcScalCls2?), CalcTensCls.

CalcVecCls, InterpolateCls

01/27/14

inidriver.F90

program driver

**SUBROUTINE** 

**INTERNAL** 

SetIdle

## EXAMPLE: **inidriver.f90**: program driver

program driver

modules used: IniFile, CAMB, LambdaGeneral, Lensing, AMLUtils, Transfer,

constants, Bispectrum, CAMBmain,

#ifdef NAGF95 F90\_UNIX #endif

data structures used:

Type(CAMBparams) P

DefIni% ???, CP%, BispectrumParams

#### process:

#### Getting started

check param-file error → write&stop

call Ini\_Open(InputFile) file inifile.f90

module IniFile

error → write&stop error → write&stop

check **outroot** error → write&stop
call **CAMB\_SetDefParams**(P) file **CAMB**.f90

module **CAMB**• get parameters of param file and check them error → write&stop

#### Read initial parameters

• call DarkEnergy\_ReadParams(DefIni) file equations.f90
module LambdaGeneral

· if non-linear and lensing:

call **Transfer\_SetForNonlinearLensing**(P%Transfer) file **modules**.f90
end-if module **Transfer**call **Transfer\_SortAndIndexRedshifts**(P%Transfer) file **modules**.f90
module **Transfer** 

call Reionization\_ReadParams(P%Reion, DefIni) file reionization.f90
module Reionization

call InitialPower ReadParams(P%InitPower, DefIni, P%WantTensors)

file power\_tilt.f90 module InitialPower

• call **Recombination\_ReadParams**(P%Recomb, DefIni) file **cosmorec**.f90 file **hyrec**.f90 file **recfast**.f90

module Recombination

check parameters
 error → write&stop

call Bispectrum\_ReadParams(BispectrumParams, DefIni, outroot)

file SeparableBispectrum.f90 module Reionization

set specific for Scalars(& Transfer), Vectors, Tensors

#### R????? initial

• if (version\_check == ") then
call TnameValueList\_Add(DefIni%ReadValues,...)
end if

file inifile.f90
module IniFile

if (outroot /= ") then

if (InputFile /= trim(outroot) //'params.ini') then

call Ini\_SaveReadValues(trim(outroot) //'params.ini',1) file inifile.f90 end if module IniFile

end n

call **Ini\_Close** file inifile.f90
module **IniFile** 

if .not. CAMB\_ValidateParams(P)
 error → write&stop

#ifdef RUNIDLE

call **SetIdle** file **inidriver**.f90 module **NONE** 

#endif

#### Calculating the results

• if (global\_error\_flag==0)
call CAMB\_GetResults(P)

file CAMB.f90
module CAMB

else error → write&stop

end-if

#### Write output

if (P%PK\_WantTransfer) then

call **Transfer\_SaveToFiles**(MT,TransferFileNames) file **modules**.f90 module **Transfer** 

call **Transfer\_SaveMatterPower**(MT,MatterPowerFileNames)

file modules.f90 module Transfer

call **Transfer\_output\_sig8**(MT) file modules.f90 end-if module **Transfer** 

if (P%WantCls) then

call output\_cl\_files(...) file modules.f90
module ModelData

call output\_lens\_pot\_files(LensPotentialFileName, output\_factor)

file modules.f90 module ModelData

if (P%WantVectors) then

call **output\_veccl\_files**(VectorFileName, output\_factor) file **modules**.f90 end if module **ModelData** 

#ifdef WRITE\_FITS, if (FITSfilename /= ")

call **WriteFitsCls**(FITSfilename, CP%Max\_l) file **writefits**.f90 end-if, #endif module **NONE** 

end if

#### Cleanup

call CAMB\_cleanup

file CAMB.f90

module CAMB