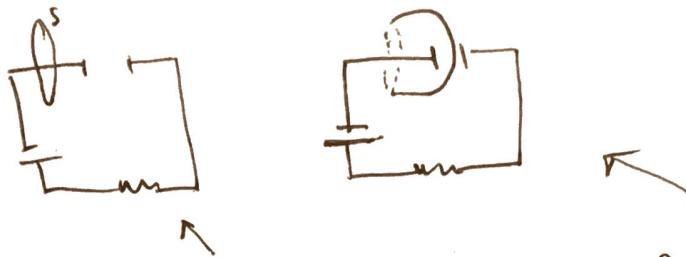


Displacement Current

Inconsistency



$$\int_{\partial S} \bar{B} \cdot d\bar{l} = \mu_0 \int_S \bar{T} \cdot d\bar{s} + \mu_0 e_0 \frac{d}{dt} \int_S \bar{E} \cdot d\bar{s}$$

$$\mu_0 e_0 \frac{d}{dt} \int_S \bar{E} \cdot d\bar{s}$$

So what? Field equations become coupled.

Boring

→

Not Boring

$$\nabla \cdot \bar{B} = \nabla \cdot \bar{E} = 0 = \nabla \cdot \bar{B}$$

$$\nabla \cdot \bar{B} = \nabla \cdot \bar{E} = 0$$

$$\nabla \times \bar{E} = -\frac{\partial \bar{B}}{\partial t}$$

$$\nabla \times \bar{E} = -\frac{\partial \bar{B}}{\partial t}$$

DRAMA

$$\nabla \times \bar{B} = \mu_0 e_0 \frac{\partial \bar{E}}{\partial t}$$

We cannot solve the equations on the right for \bar{E}

$$\frac{\partial}{\partial t} \left[\mu_0 e_0 \frac{\partial \bar{E}}{\partial t} = \nabla \times \bar{B} \right]$$

$$\mu_0 e_0 \frac{\partial^2 \bar{E}}{\partial t^2} = \nabla \times \frac{\partial \bar{B}}{\partial t} \quad \rightarrow \quad \frac{\partial \bar{B}}{\partial t} = -\nabla \times \bar{E}$$

$$= -\nabla \times (\nabla \times \bar{E}) = -\nabla (\nabla \cdot \bar{E}) + \nabla^2 \bar{E}$$

$$\rightarrow \frac{1}{c^2} \frac{\partial^2 \bar{E}}{\partial t^2} = \nabla^2 \bar{E} \quad \left\{ \text{Where } c = \sqrt{\frac{1}{\mu_0 e_0}} \approx 3 \times 10^8 \text{ m/s} \right.$$

$$\text{Same for } \bar{B} \rightarrow \frac{1}{c^2} \frac{\partial^2 \bar{B}}{\partial t^2} = \nabla^2 \bar{B}$$

["The velocity of the transverse" -]
 = "Oh shit." J. C. Maxwell, paraphrased by me.

Greece - fur on Amber 400 BC	Optics 400 BC
Leoo BC ηλεκτρον	Lenses & Euclid

[One of the first great unifications in physics.]

Plane Wave Solutions to Wave Eqs / Maxwell's Eqs:

$$\bar{E} = \bar{E}_0 e^{i(\bar{k} \cdot \bar{x} - \omega t)} \quad \bar{B} = \bar{B}_0 e^{i(\bar{k} \cdot \bar{x} - \omega t)}$$

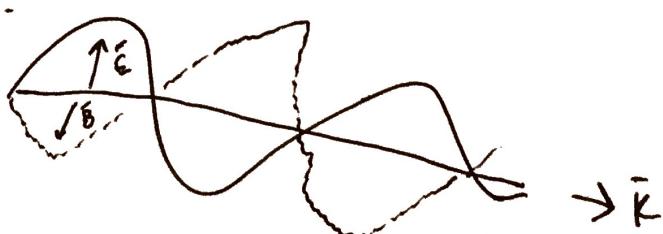
are solutions when $\bar{k} = k \hat{k}$, $\omega = ck$, and

$$\bar{k} \cdot \bar{E}_0 = \bar{k} \cdot \bar{B}_0 = 0, \quad \bar{k} \times \bar{E}_0 = k E_0 \hat{B} = \omega \bar{B}_0$$

[Writing \bar{E} and \bar{B} as complex seems strange, but they satisfy Maxwell's Eqs and are easier to manipulate than sines & cosines. Just take the real part before doing any ∇^2 or shenanigans.]

[I'm sure you all have seen this before, but we talk about EM waves as propagating in \hat{k} with perp \bar{E} and \bar{B} . When we talk about Polarization, we are just referring to the direction \hat{E} is pointing.]

Here's the picture:



Any old plant will tell you that EM waves carry energy.⁴

Starting with

$$U = \int_V d^3x \left(\frac{\epsilon_0}{2} \bar{E} \cdot \bar{E} + \frac{1}{2\mu_0} \bar{B} \cdot \bar{B} \right) \quad \text{Energy of Fields in } V$$

Calculate

$$\begin{aligned} \frac{dU}{dt} &= \int_V d^3x \left(\epsilon_0 \bar{E} \cdot \frac{\partial \bar{E}}{\partial t} + \frac{1}{\mu_0} \bar{B} \cdot \frac{\partial \bar{B}}{\partial t} \right) \\ &= \int_V d^3x \left(\frac{1}{\mu_0} \bar{E} \cdot (\nabla \times \bar{B}) - \bar{E} \cdot \bar{J} - \frac{1}{\mu_0} \bar{B} \cdot (\nabla \times \bar{E}) \right) \end{aligned}$$

use $\bar{E} \cdot (\nabla \times \bar{B}) - \bar{B} \cdot (\nabla \times \bar{E}) = -\nabla \cdot (\bar{E} \times \bar{B})$ + div thm.

$$\rightarrow \frac{dU}{dt} = - \int_V d^3x (\bar{J} \cdot \bar{E}) - \frac{1}{\mu_0} \int_{\partial V} (\bar{E} \times \bar{B}) \cdot d\bar{S}$$



work done on \bar{J}

by \bar{E}

(don't care, not radiation)



energy radiated away

Poynting Vector $\bar{S} = \frac{1}{\mu_0} \bar{E} \times \bar{B}$ \approx Energy fluence

- energy loss is

$$\oint_S \bar{S} \cdot d\bar{S}$$

Thomson Scattering: n 30 years after Maxwell, pre QM
 People are beginning to ask how matter is affected by EM Waves.

Harmor Formula:

Given: for $r \gg \lambda$ the magnetic field is well approximated by $\bar{B}(t, \bar{x}) = -\frac{\mu_0}{4\pi r c} \hat{x} \times \ddot{\vec{p}}(t - r/c)$

{ "electric dipole approximation" } 

where $\bar{p} = \int_V d^3x p(\bar{x}) \bar{x}$ - "center of charge"

[So we can have some arbitrary charge distribution far away doing crazy things and describe \bar{B} from it]

In this case we can write the Poynting vector

$$\bar{S} = \frac{1}{\mu_0} \bar{E} \times \bar{B} = \frac{c}{\mu_0} |\bar{B}|^2 \hat{x} = \frac{\mu_0}{(4\pi r^2 c)} |\hat{x} \times \ddot{\vec{p}}|^2 \hat{x}$$

So for oscillations in the \hat{z} direction

$$\bar{S} = \frac{\mu_0}{(4\pi r^2 c)} |\ddot{\vec{p}}|^2 \sin^2 \theta \hat{z} \quad \text{take } \ddot{\vec{p}} = |\ddot{\vec{p}}| \hat{z} \\ \rightarrow |\hat{x} \times \ddot{\vec{p}}| = |\ddot{\vec{p}}| \sin \theta$$

Integrate over sphere

$$P = \frac{\mu_0}{(4\pi c)} |\ddot{\vec{p}}|^2 - \bar{P} = \frac{\mu_0}{8\pi c} |\ddot{\vec{p}}_{\text{max}}|^2$$

For $\bar{p} = \cancel{\partial} Q d\hat{z} e^{i\omega t}$, $\ddot{p} = -Q d\omega^2 \hat{z} e^{i\omega t}$

$$\boxed{\bar{P}_{\text{radiated}} = \frac{\mu_0 Q^2 d^2 \omega^4}{12\pi c} \quad - \text{Larmor Formula}}$$

Thomson Scattering - First application of radiation formulas to scattering

~30 yrs after Maxwell, but pre QM

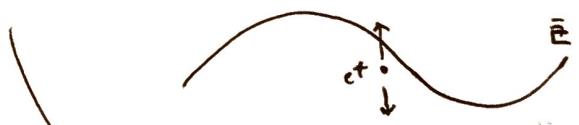
Suppose we have a free charged particle sitting in an ~~stationary~~ oscillating electric field, and it doesn't ever move far from the origin. ~~oscillates~~

i.e. long wavelength.

Strated :-

Newton: $m\ddot{x} = q\bar{E}$

$$\ddot{x}(t) = -\frac{q E_0}{m \omega^2} \sin(\omega t)$$



$$\frac{q E_0}{m \omega^2} \ll \frac{c}{\omega} \Rightarrow \frac{q E_0}{m \omega c} \ll \frac{e}{c}$$

↑
v_{max}

According to Larmor :

$$\bar{P}_{\text{radiated}} = \frac{\mu_0 q^2 (q^2 E_0^2) \omega^4}{12 \pi m^2 c^4}$$

$$\bar{S}_{\text{incident}} = \frac{1}{\mu_0} \bar{E} \times \bar{B} = \frac{c E_0^2}{\mu_0} \hat{k} \sin^2(\hat{k} \cdot \hat{x} - \omega t)$$

What's going to happen to energy in

$$\text{Stress} = \frac{\omega^2}{2\pi} \int_0^{2\pi/\omega} \frac{cE_0^2}{\mu_0} \sin^2(kx - \omega t) dt$$

$$= \frac{cE_0^2 \omega}{2\pi\mu_0} \cdot \frac{1}{2} \int_0^{2\pi/\omega} [1 - \cos(2kx/\lambda - 2\omega t)] dt$$

$$= \frac{cE_0^2}{2\pi\mu_0}$$

$$\frac{P_{\text{radiated}}}{S_{\text{incident}}} = \frac{\mu_0 g^4 E_0^2}{12\pi m c} \cdot \frac{2\mu_0}{cE_0^2} = \boxed{\frac{\mu_0 g^4}{16\pi m^2 c} = \sigma_{\text{Th}}}$$

$$\approx 10^{-30} \text{ m}^2$$

Rayleigh Scattering:

We can imagine doing the same thing for a neutral atom like hydrogen. When you put hydrogen in an oscillating field, the com does not change, but the constituents will want to move.

Given: $\alpha \approx \frac{g^2}{\omega_0^2 m}$, $\bar{p} = \alpha \bar{E}$ good assumption

$$\omega_0 = \frac{g^2}{m}$$

ω_0 is the resonant frequency of the atom

$$\alpha = \frac{g^2}{\omega_0^2 m}$$

$$\omega_0 = \frac{g^2}{4\pi q^3 \epsilon_0 m}$$

$$\bar{P}_{\text{rad}} = \frac{\mu_0 p^2 \omega^4}{12\pi c} = \frac{\mu_0 \alpha^2 E_0^2 \omega^4}{12\pi c} = \frac{\mu_0 g^2 E_0^2 \omega^4}{12\pi c \omega_0^4 m^2}$$

$$\bar{S}_{\text{inc}} = \frac{c E_0^2}{2 \mu_0} \rightarrow \boxed{O_{\text{Ray}} = \frac{\mu_0 g^4}{6\pi m^2 c^2} \left(\frac{\omega}{\omega_0}\right)^4}$$

Depends on atomic structure ω_0 .

Goes as ω_0 or ω^4 — short wavelengths scatter more. Blue light appears to come from all regions of the sky.

Not really scattering
Relativistic Larmor Formula & Bremsstrahlung

The derivation of the Relativistic Larmor formula is significantly more complicated,

but follows the same steps as the previous derivation. Same thing, but

$$\frac{dP}{d\Omega} = \frac{g^2}{(4\pi r^2 \epsilon_0 c^3)} \frac{[\hat{R} \times [(\hat{R} - \bar{v}/c) \times \bar{a}]]^2}{k^5}$$

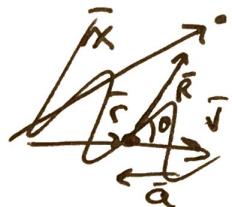
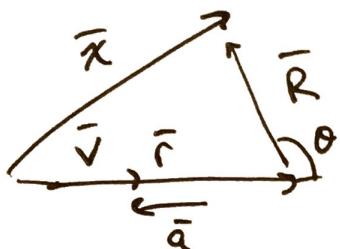
1. not far away
2. account for light travel time.
relativistic

$$\text{where } K = 1 - \frac{\hat{R} \cdot \bar{v}}{c}, \text{ and } \bar{R}(t) = \bar{x} - \bar{r}(t)$$

In the case of Bransstetling, we assume linear acceleration in the opposite direction as \vec{v}

$$\hat{\vec{a}} = -\hat{\vec{v}}$$

$$\hat{\vec{R}} \cdot \vec{v} = v \cos \theta$$



$$\frac{dP}{d\Omega} = \frac{q^2}{16\pi^2 \epsilon_0 c} \left[\frac{\hat{\vec{R}} \times (\hat{\vec{R}} \times \frac{\vec{a}}{mc^2})}{k^5} \right]^2 = \frac{q^2}{16\pi^2 \epsilon_0 c} \left[\hat{\vec{R}} \times (-a \sin \theta \hat{\vec{R}} \times \hat{\vec{a}}) \right]^2$$

Perp so
we get
the whole
thing
dotted
into itself.

$$= \frac{q^2}{16\pi^2 \epsilon_0 c} \left[\hat{\vec{R}} \times (-a \sin \theta (\hat{\vec{R}} \times \hat{\vec{a}})) \right]^2$$

$$= \frac{q^2 a^2}{16\pi^2 \epsilon_0 c} \frac{\sin^2 \theta}{(1 - \frac{v}{c} \cos \theta)^5}$$

$$\text{For } v \ll c \rightarrow \frac{q^2 a^2}{16\pi^2 \epsilon_0 c} \sin^2 \theta, \quad \text{max @ } \theta/2$$

- (a) relativistic speeds - two lobes in the forward direction.

$$\int d\Omega \left(\frac{dP}{d\Omega} \right) = P = \frac{q^2 \gamma^6 a^2}{6\pi \epsilon_0 c^3}$$

Ch. 7. Gamma and X-ray Interactions

Introductions - Name, Undergrad, favorite movie

How's the class going? Any questions abt. HW?

Outline: - You will not be tested on today's - feel free to leave
Friday - Scattering in Classical Electrodynamics:

Review, Thompson scattering, Rayleigh scattering,
Bremsstrahlung

Tuesday - "Modern" physics Interactions:

Compton, Photoelectric, Pair prod., Photoneuclear,

Summary ~~as it relates to us.~~

Maxwell's Equations:

*Gauss' Law $\nabla \cdot \bar{E} = \frac{\rho}{\epsilon_0}$

No Magnetic Monopoles $\nabla \cdot \bar{B} = 0$

Faraday's Law $\nabla \times \bar{E} = -\frac{\partial \bar{B}}{\partial t}$

Ampere's Law $\nabla \times \bar{B} = \mu_0 (\bar{J} + \epsilon_0 \frac{\partial \bar{E}}{\partial t})$

What did Maxwell actually contribute? Why
are these "Maxwell's Equations"?