Practice Science Motivation (1 page)

Table 1: Summary of Subfields Supported by TDE Science

Astronomical Subfield	Decadal Subsec-	How TDE Observations are Useful	
	tion		
TDE Emission	B-Q2a	_	
Transients Powered by	B-Q2a	Since TDEs are powered by Black Holes,	
Black Holes		they are directly applicable.	
Multimessenger as-	B-DA8	There has been some evidence of neutrino	
tronomy		detections from nearby TDEs.	
Galaxy Nuclei	D-Q4d	Radio observations of TDEs provide circum-	
		nuclear $(r < 1 \text{ pc})$ contents and density pro-	
	D 00 D 04	files.	
Black Holes	B-Q2a, B-Q4b,	TDEs provide measurement of the mass of a	
	D-Q3b,	central SMBH independent of galaxy prop-	
		erties and scaling relations (e.g. $M-\sigma$). They	
		also provide a means to discover and study the elusive intermediate mass black hole.	
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Shock Physics	B-Q2b	Radio observations of TDEs reveal the prop-	
		erties of the shock front as it propagates through the circumnuclear medium.	
Jet Physics	B-Q2a, B-Q3a	Prompt radio observations of TDEs will cap-	
Jet I Hysics	D-Q2a, D-Q3a	ture the early-time jet evolution minutes to	
		days after the initial launch.	
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A primary science goal of the 2023 Astronomy Decadal, and a personal interest of mine, is the continued prompt multiwavelength followup of transient events. These violent and explosive transient events provide a unique probe of high energy phenomenon and environments that are impossible to reproduce on the surface of the Earth. Of particular interest are tidal disruption events (TDEs). TDEs occur when a star approaches within the tidal radius of a supermassive black hole (SMBH) and is torn apart, releasing emission across the electromagnetic spectrum. Understanding the emission from TDEs on their own and how it relates to other transient emission is an open question in astronomy and a key focus of the 2023 Astronomy Decadal. The relevant Decadal subsections are listed in the first part of Table 1.

Studying TDEs will not only further our understanding of them but also many other areas of astronomy. For instance, studying TDE emission provides a probe the SMBH properties, the sub-parsec galactic nuclei environment, jet evolution and fundamental physics, and shock physics. For instance, using radio observations of TDEs we are able to study the density and particle makeup of the previously quiescent galactic nuclei. A summary of areas of astronomy probed by TDE observations, as well as the relevant Atronomy 2020 Decadel (sub)sections corresponding to each of these areas, is given in Table 1. Overall, this makes multiwavelength observations of TDEs necessary to further our fundamental understanding of high energy, dense astronomical environments.

In particular, as described in Table 1, prompt radio/millimeter observations of TDEs allow us to study the early evolution of jets. As the jets evolve outwards, they shock the previously quiescent circumnuclear medium, producing synchrotron emission visible in the radio and millimeter. Fitting a complete spectral energy distribution of TDEs allows us to extract properties of the jet, shock front, and ambient circumnuclear medium. Since this SED normally peaks around 1-10 GHz (depending on the time of the observation), observations with sensitive telescopes like the Very Large Array (VLA) and Atacama Large Millimeter Array (ALMA) necessary to study TDE radio/millimeter emission and further our understanding of jet evolution. A list of recently radio bright TDEs is given in Table 2 as examples of objects that require followup. Although, for a true study of this nature, we would need "Target of Opportunity" (ToO) observations within days after the initial optical or X-ray detection.

Table 2: Sample of Radio Bright TDEs for Observations

IAU Name	RA [hours]	Declination [degree]	Redshift	Approximate Flux Density at 6GHz $[\mu Jy]$
AT2022dyt AT2022wtn AT2021sdu AT2020zso AT2018cqh	10:00:08.008 23:23:23.778 01:11:23.924 22:22:17.130 02:33:46.930	+26:27:38.57 $+10:41:07.99$ $+50:34:29.67$ $-07:15:59.08$ $-01:01:28.38$	0.072 0.049 0.059 0.061 1	75 300 75 150 10 ⁵
AT2018hyz	10:06:50.871	+01:41:34.08	0.04573	10^{5}

¹No redshift publicly available for AT2018cqh on the Transient Name Server (TNS).