Recap:

Each EM mode behaves like a harmonic oscillator, where we know how to quantize each mode. Specifically,

$$ec{A}(ec{r},t) = rac{1}{L^{3/2}} \left(\sum_{ec{k},lpha} rac{2\pi \hbar c^2}{\omega_k}
ight)^2 [a^\dagger_{ec{k},lpha} e^{-iec{k}\cdotec{r}+i\omega_k t} + a_{ec{k},lpha} e^{iec{k}\cdotec{r}-i\omega_{kt}}] \hat{e}_{ec{k},lpha}$$

with the algebra of the operator a^{\dagger} , a as:

$$egin{aligned} [a^\dagger_{k,lpha},a^\dagger_{k',lpha'}] &= 0 \ [a_{k,lpha},a_{k',lpha'}] &= 0 \ [a_{k,lpha},a^\dagger_{k',lpha'}] &= \delta_{ec{k},ec{k}'}\delta_{lpha,lpha'} \end{aligned}$$

with equation of \vec{E}, \vec{B} as:

$$ec{E}(ec{r},t) = -rac{1}{c}rac{\partialec{A}}{\partial t} \ ec{B}(ec{r},t) = ec{
abla} imesec{A}$$

Where, after doing some algebra, we get

$$\mathcal{H} = \sum_{ec{k}, lpha} \hbar \omega_k \left[a_{ec{k}, lpha}^\dagger a_{ec{k}, lpha} + rac{1}{2}
ight]$$
 (1)

the free space Hamiltonian. Note that this come from the classical result:

$${f E} = rac{1}{8\pi} \int [|ec E|^2 + |ec B|^2] \, d^3 r$$

For the QM case, we can interpret our operators as:

- $a_{ec{k},lpha}^{\dagger}$: creation of a photon of energy $\hbar\omega_k$ with mode $ec{k},lpha$
- $a_{\vec{k},\alpha}$: destruction of a photon with energy ...
- $a_{\vec{k},\alpha}^{\dagger}a_{\vec{k},\alpha}^{}$: the number of photons in mode \vec{k},α .

Which is just like the harmonic oscillator case!

Note that because (1) is true, we have a infinite amount of modes, and there exist a lowest energy state $\frac{1}{2}\hbar\omega_k$, this implies that the sum of (1) $\to \infty$, which implies a infinite vacuum energy.

While this is fine for many problems, as the zero point energy is trivial in many questions- especially experimentally as we always measure $\Delta \mathbf{E}$, but there is one very important physical interaction that cares about the ground state energy- gravity.

Now, the good news(not for the gravity case) is that sometimes we can use analytical extension to make \mathcal{H} converge (e.g. $\Gamma(-1) = -\frac{1}{12}$)

Trivial Exercise:

Integrate gravity into quantum mechanics.

We must note that a photon is a boson. In fact, the operator relation is the defining feature of the boson. as the commutation relation is symmetric. Specifically:

$$egin{aligned} a^{\dagger}_{k,lpha}a^{\dagger}_{k',lpha'}|0
angle &= |ec{k}lpha,ec{k}'lpha'
angle \ a^{\dagger}_{k'lpha'}a^{\dagger}_{klpha}|0
angle &= |ec{k}'lpha',ec{k}lpha
angle \end{aligned}$$

Which indicates that the photon is a boson.

Spontaneous Emission:

Spontaneous emission means we start from an excited state $|i\rangle$, and the particle spontaneously goes to the final state $|f\rangle$ (ground state) and emit a photon.

This means that there is no EM emission to begin with, but we get a photon in the final state.

For a hydrogen atom transitioning from $|2,l,m\rangle \to |1,0,0\rangle$, we can write down the unperturbed Hamiltonian

$$egin{aligned} \mathcal{H} &= \mathcal{H}_0 + \mathcal{H}_1 \ &= rac{p^2}{2m} - rac{e^2}{r} + \sum_{ec{k}\,lpha} \hbar \omega_k (a_{k,lpha}^\dagger a_{k,lpha}) + \mathcal{H}_1 \end{aligned}$$

with:

 $|i
angle=|2,l,m
angle\otimes|0
angle$ where the second ket is the ket of the EM ground state, $|f
angle=|1,0,0
angle\otimes|1
angle_{ec{k}lpha}$ in some mode. $\implies E_f-E_i=(E_{100}+\hbar\omega_k)-E_{2lm}$

We can use Fermi's Golden Rule:

$$\Gamma_{i
ightarrow f} = rac{2\pi}{\hbar} |\langle f|H_1|i
angle|^2 \delta(E_f-E_i)$$

Where H_1 is how the hydrogen atom couples with the EM field:

$$H_1 = rac{e}{mc} ec{A} \cdot ec{p}$$

(go back and check the dipole interaction Hamiltonian! Please check the commutation relation!)

Note that the algebra between \hat{p} and a^{\dagger} , a is an abelian one, as a^{\dagger} , a is in the subspace of the photon- **not** the hydrogen wave function.

Matrix Element:

$$\langle f|H_1|i
angle = \langle 100|\langle ec{k},lpha|ec{A}\cdotec{p}|0
angle|2lm
angle$$

We focus on the matrix element of the photon space

$$\langle ec{k} lpha | ec{A} | 0
angle$$

Note that we can write \vec{A} to the sum of modes, and because $|0\rangle$ is on the right, only the creation operator a^{\dagger} part of \vec{A} will contribute. It is then easy to see that:

$$\langle ec{k}lpha |ec{A}|0
angle = rac{1}{L^{3/2}}igg(rac{2\pi\hbar c^2}{\omega_k}igg)^{1/2}e^{-iec{k}\cdotec{r}}\hat{e}_{ec{k},lpha}$$

in which we can make the dipole approximation such that $e^{-i\vec{k}\cdot\vec{r}}\sim 1$ because $\lambda=rac{2\pi}{k}\gg a_0.$

We still have to compute $ec{p}$ part of the matrix $\langle 100 | ec{p} | 2lm
angle \cdot ec{e}_{ec{k},\alpha}$

We learned earlier by looking in $[\vec{R}, H_0]$ that:

$$\langle f_0 | ec{p} | i_0
angle = rac{m}{i\hbar} (E_i^0 - E_f^0) \langle f_0 | ec{R} | i_0
angle$$

where $E_i^0-E_f^0=\hbar\omega.$ This gives:

$$egin{aligned} \langle 100 | ec{p} | 2lm
angle \cdot ec{e}_{ec{k},lpha} \ &= im\omega ec{e} \cdot \langle 100 | ec{R} | 2lm
angle \ &= im\omega ec{e} \cdot \int \psi_{100}(ec{r}) ec{r} \psi_{2lm}(ec{r}) \, d^3r \end{aligned}$$

where we can use dipole selection rule and get:

 $egin{aligned} l &= 0 \ m &= 0, \pm 1 \end{aligned}$