

P7502 HW3

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Problem 1

1. We have the equation

$$E_n^1 = \langle n^0 | H^1 | n^0 \rangle \quad (1)$$

where n^0 is just the n th state of the unperturbed oscillator. Note that

$$\begin{aligned} E_n^1 &= \langle n^0 | \lambda x^4 | n^0 \rangle \\ &= \lambda \langle n^0 | \left(\sqrt{\frac{\hbar}{2m\omega}} (a^\dagger + a) \right)^4 | n^0 \rangle \\ &= \lambda \left(\frac{\hbar}{2m\omega} \right)^2 \langle n^0 | (a^2 a^{\dagger 2} + a^\dagger a^2 a^\dagger + a a^\dagger a a^\dagger + a^\dagger a a^\dagger a + a a^{\dagger 2} a + a^{\dagger 2} a^2) | n^0 \rangle \end{aligned}$$

By exploiting commutating relation $[a, a^\dagger] = 1$

$$\begin{aligned} &= \lambda \left(\frac{\hbar}{2m\omega} \right)^2 (6(n+1)(n+2) - 12(n+1) + 3) \\ &= \frac{3\lambda\hbar^2}{4m^2\omega^2} (2n^2 + 2n + 1) \checkmark \end{aligned}$$

2. Because the first perturbation term E_n^1 scales with n^2 , $\forall \lambda \in \mathbb{C}, \exists \lambda$ such that $E_n^1 > \Delta E_n$.

Physically, because our Hamiltonian $H = \frac{p^2}{2m} + \frac{1}{2}m\omega x^2 + \lambda x^4$ will approach $H' = \frac{p^2}{2m} + \lambda x^4$ for n sufficiently large regardless how small a fixed λ . This means our potential term will be dominated by the perturbation.

Problem 2

We have that $H = H^0 + H^1 = -\gamma S_z B_0 - \gamma S_x B$

Because the main magnetic field B_0 is in \hat{z} direction, we recognize that the spin up state is going to be the ground state as it has the lowest energy.

Therefore,

$$\begin{aligned} E_+^1 &= \langle + | H^1 | + \rangle \\ &\doteq \frac{-\hbar\gamma B}{2} \begin{pmatrix} 1 & 0 \end{pmatrix} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \begin{pmatrix} 1 \\ 0 \end{pmatrix} = 0 \end{aligned}$$

$$\begin{aligned} E_+^2 &= \frac{|\langle + | H^1 | - \rangle|^2}{E_+^0 - E_-^0} \\ &= \frac{\gamma^2 B^2 \hbar^2}{4} \frac{1}{-\gamma B_0 \hbar} = -\frac{B^2 \hbar \gamma}{4B_0} \end{aligned}$$

$$\begin{aligned} | +^1 \rangle &= | +^0 \rangle + \frac{\langle - | H_1 | + \rangle}{E_+^0 - E_-^0} \\ &= | + \rangle + \frac{B}{2B_0} | - \rangle \end{aligned}$$

To compare the expanded exact solution, we first note that to get the exact answer, we just need to diagonalize

$$H \doteq -\gamma \frac{\hbar}{2} \begin{pmatrix} B_0 & B \\ B & -B_0 \end{pmatrix} \text{ Because we only care about the ground state, we}$$

only select negative eigenvalue and its eigenvector. This gives $E = \frac{-\sqrt{B^2 + B_0^2} \hbar \gamma}{2}$ and $|n\rangle = |+\rangle - \frac{B}{-B_0 - \sqrt{B^2 + B_0^2}} |-\rangle$ We can expand our E

$$\begin{aligned} E &= \frac{-\sqrt{B^2 + B_0^2} \hbar \gamma}{2} \\ &= -\frac{\hbar \gamma}{2} B_0 \sqrt{1 + \frac{B^2}{B_0^2}} \\ &\approx -\frac{\hbar \gamma}{2} (B_0 + \frac{B^2}{2B_0}) \end{aligned}$$

This agrees with our perturbation parameters where $E^0 = -\frac{\hbar \gamma B_0}{2}$, $E^1 = 0$, and $E^2 = -\frac{B^2 \hbar \gamma}{4B_0}$ We can expand our $|n\rangle$

$$\begin{aligned} |n\rangle &= |+\rangle - \frac{B}{-B_0 - \sqrt{B^2 + B_0^2}} |-\rangle \\ &\approx |+\rangle + \frac{B}{2B_0 + \frac{B^2}{B_0}} |-\rangle \\ &\approx |+\rangle + \frac{B}{2B_0} |-\rangle \end{aligned}$$

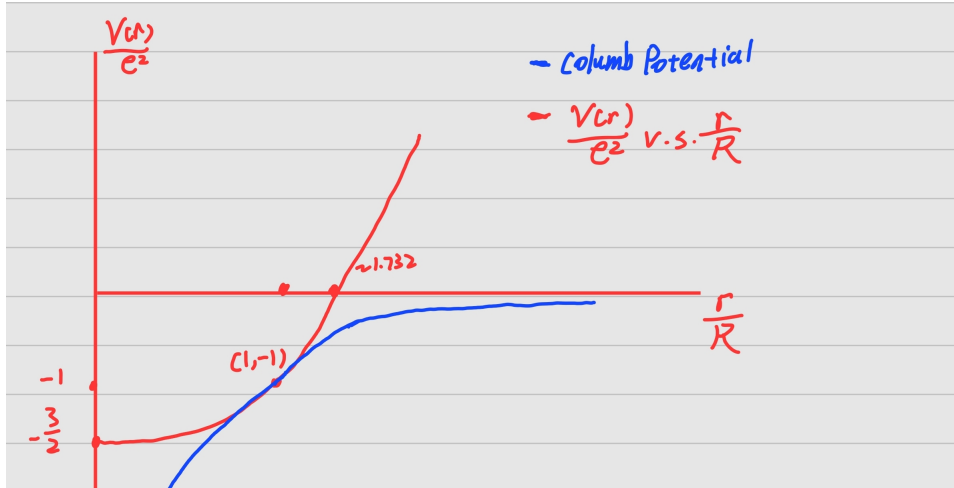
Which is exactly the same as $| +^1 \rangle$ the first perturbed eigenstate.

Problem 3

for $r > R$, the problem returns to be Columb potential problem. Thus we conly consider the case where $r < R$ that corresponds to $V(r) = \frac{3e^2}{2R} + \frac{e^2 r^2}{2R^3}$

$$\text{We have } H^1 = V(r) - V_{\text{Columb}}(r) = -\frac{3e^2}{2R} + \frac{e^2 r^2}{2R^3} + \frac{e^2}{r}$$

$$\begin{aligned} E_1^1 &= \langle 100 | H_1 | 100 \rangle \\ &= \int_0^{2\pi} \int_0^\pi \int_0^R e^{-2r/\alpha_0} \frac{1}{\pi a_0^3} \left(-\frac{3e^2}{2R} + \frac{e^2 r^2}{2R^3} + \frac{e^2}{r} \right) r^2 \sin \theta dr d\theta d\phi \\ &= \frac{4}{a_0^3} \left(\frac{1}{2} e^2 R^2 - \frac{1}{2} e^2 R^2 + \frac{1}{10} e^2 R^2 \right) \\ &= \frac{2e^2 R^2}{5a_0^3} \end{aligned}$$



Notice that in our graphics of $V(r)$, the Columb potential is only similiar to our potential near $r = R$. I think this is fine, because for $r > R$, we define our potential to be the columb potential, and for $r \ll R$, the angular effective potential will likely dominate anyway.

Now for Numerical Estimation, we bring back the missing $\frac{1}{4\pi\epsilon_0}$

$$\begin{aligned} E_1^1 &\approx \frac{2}{5} e^2 (1E - 9\mu\text{m})^2 / (5.3E - 5\mu\text{m})^3 \frac{1\text{eV}}{4\pi \cdot 55e^2 \mu\text{m}} \\ &= 4 \cdot 10^{-9} \text{eV} \end{aligned}$$

To figure out what change in mass would result in change in perterbation result, we first realize that the radius a_0 must change when we increase the mass of the

particle. Using classical mech and Bhor's QUantization conditiion

$$\begin{aligned}\frac{e^2}{4\pi\epsilon_0 r^2} &= m_\mu \frac{v^2}{r} \\ \Rightarrow r &= \frac{4\pi\epsilon_0 \hbar^2}{m_\mu e^2} = \frac{1}{200} a_0\end{aligned}$$

Once we realized that our new radius is $\frac{1}{200}a_0$, we realize that our new perturbation term must be $200^3 = 8000000$ times stronger. Using previous result, this will grant us 0.032eV.

Note that Technically speaking, our eigenenergy E_n should also increase by a factor of 200. This still makes our perturbation term $200^2 = 40000$ times larger with respect to E^0 .

Problem 4

1) We have

$$W = \frac{e^2}{R^3} [\mathbf{r}_A \cdot \mathbf{r}_B - 3(\mathbf{r}_A \cdot \hat{\mathbf{n}})(\mathbf{r}_B \cdot \hat{\mathbf{n}})] \quad (2)$$

A small parameter that would fit the problem would be $\lambda = \frac{a_0^3}{R^3} \propto \frac{W}{H^0}$. Because we have $R \gg \{r_a, r_B\} \propto a_0$, we must have that $\frac{a_0^3}{R^3}$ is a small parameter that compares W with H^0 .

2) WLOG, we can choose \hat{n} to be \hat{z} . This gives a specific equation representation for W in Cartesian Coordinates

$$\begin{aligned}W &= \frac{e^2}{R^3} [\mathbf{r}_A \cdot \mathbf{r}_B - 3(\mathbf{r}_A \cdot \hat{\mathbf{n}})(\mathbf{r}_B \cdot \hat{\mathbf{n}})] \\ &= \frac{e^2}{R^3} [x_A x_B + y_A y_B + z_A z_B - 3z_A z_B] \\ &= \frac{e^2}{R^3} [x_A x_B + y_A y_B - 2z_A z_B]\end{aligned}$$

3) We can use the fact that our $|100_A; 100_B\rangle$ is spherically symmetric. and realize $W = H^1$ has two positive contribution in X, Y and 2 times negative contribution in Z . The spherical symmetric nature of the wave function means the contribution from all directions are the same. Therefore, the contribution of $X + Y - 2Z$ should be 0.

Alternatively, we just realize for each wavefunction,

$$\int_0^{2\pi} \int_0^\pi d\theta d\phi \cos \phi \sin \theta + \sin \phi \sin \theta - 2 \cos \theta = 0 \quad (3)$$

This makes physical sense because the first degree multiple expansion is between E -field and charges. However, the net charge of each Hydrogen atom is neutral.

4)

For our second order energy perturbation term E^2 , we have the following equation

$$E_n^2 = \sum_m \frac{|\langle n^0 | H^1 | m^0 \rangle|^2}{E_n^0 - E_m^0} \quad (4)$$

First, we realize that our denominator is negative since $E_n^0 < E_m^0 \forall n \neq m$ as we set n to be the ground state.

Second, we realize that our numerator will at least have the factor $|\frac{e^2}{R^3}|^2$.
Combining these two facts, we can say that $E_n^2 \propto -\frac{c}{R^6}$ for some $c > 0$.

5) As we have two particles in H ,

$$\begin{aligned} E_n^2 &= \sum_m \frac{|\langle n^0 | H^1 | m^0 \rangle|^2}{2E_n^0 - 2E_m^0} \\ E_n^2 &\approx \sum_m \frac{|\langle n^0 | H^1 | m^0 \rangle|^2}{2E_1} \text{ by using identity} \\ &\approx \frac{|\langle n^0 | H^1 H^1 | n^0 \rangle|}{2E_1} \\ &\approx -\frac{e^4}{2E_1 R^6} \langle 100; 100 | (X_A X_B + Y_A Y_B - 2Z_A Z_B)^2 | 100; 100 \rangle \\ &\Rightarrow c \propto \frac{e^2}{2|E_1|} \langle 100; 100 | (X_A X_B + Y_A Y_B - 2Z_A Z_B)^2 | 100; 100 \rangle \end{aligned}$$

6) Explicitly, we can integrate $X_A Y_A$ explicitly. Note that our radial part of the integration does not involve θ, ϕ , we can simply integrate our wavefunction over θ, ϕ

$$\begin{aligned} &\int_0^\pi \int_0^{2\pi} d\phi d\theta Y_0^{02} \sin \theta \sin \theta \sin \phi \sin \theta \cos \phi \\ &= \frac{1}{4\pi} \int_0^\pi \int_0^{2\pi} d\phi d\theta \sin^3 \theta \sin \phi \cos \phi \\ &= \frac{1}{4\pi} \int_0^{2\pi} d\phi \sin \phi \cos \phi \cdot \int_0^\pi d\theta \sin^3 \theta \end{aligned}$$

Where the $d\phi$ integral is 0.

7) We have explicitly

$$\begin{aligned}
& \langle 100; 100 | R^2 | 100; 100 \rangle \\
&= \frac{1}{\pi a_0^3} \int_0^\infty \int_0^\pi \int_0^{2\pi} r^4 \sin \theta d\phi d\theta dr \\
&= 3a_0^2 \\
& \langle 100; 100 | Z^2 | 100; 100 \rangle \\
&= \frac{1}{\pi a_0^3} \int_0^\infty \int_0^\pi \int_0^{2\pi} r^4 \sin \theta \cos^2 \theta d\phi d\theta dr \\
&= a_0^2 \\
& \langle 100; 100 | X^2 | 100; 100 \rangle \\
&= \frac{1}{\pi a_0^3} \int_0^\infty \int_0^\pi \int_0^{2\pi} r^4 \sin \theta \sin^2 \theta \cos^2 \phi d\phi d\theta dr \\
&= a_0^2 \\
& \langle 100; 100 | Y^2 | 100; 100 \rangle \\
&= \frac{1}{\pi a_0^3} \int_0^\infty \int_0^\pi \int_0^{2\pi} r^4 \sin \theta \sin^2 \theta \sin^2 \phi d\phi d\theta dr \\
&= a_0^2
\end{aligned}$$

Note that the expectation value of each direction squared is just $\frac{1}{3}\langle R^2 \rangle$

8) In Gaussian unit, energy $\propto e^2/r$. Therefore, for our unit to be in the unit of energy, we must have $p = 2$. As we have R^6 in the denominator and we want r^{-1} , we must have $q - 6 = -1 \rightarrow q = 5$