Quantum Field Theory

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1 Quantum Electrodynamics (QED)

5.1 Classical Electrodynamics and Maxwell's equations

We have the gauge potential $A^{\mu}=(A^0, \mathbf{A})=(\phi, \mathbf{A})$ (or $A_{\mu}=(A^0, -\mathbf{A})=(\phi, -\mathbf{A})$) and the field strength tensor $F_{\mu\nu}=\partial_{\mu}A_{\nu}-\partial_{\nu}A_{\mu}$.

Then

• electric field is $E_i = F_{0i} = \partial_0 A_i - \partial_i A_0 \rightarrow \mathbf{E} = -\dot{\mathbf{A}} - \nabla \phi$

• magnetic field is $B^i = -\frac{1}{2} \epsilon^{ijk} F_{jk} \rightarrow \boldsymbol{B} = \nabla \times \boldsymbol{A}$

Lagrangian density $\mathcal{L}_{EM} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} = -\frac{1}{2}(\boldsymbol{E}\cdot\boldsymbol{E}-\boldsymbol{B}\cdot\boldsymbol{B})$. The field equation $\partial_{\mu}\left(\frac{\partial\mathcal{L}}{\partial(\partial_{\mu}A_{\nu})}\right) - \frac{\partial\mathcal{L}}{\partial A_{\nu}} = 0$ leads to

$$\partial_{\mu}F^{\mu\nu} = 0 \tag{5.1.1}$$

It is half of Maxwell's equations (in vacuum).

The other half is Bianchi identities following from the definition of $F_{\mu\nu}$:

$$\begin{split} \partial_{\lambda}F_{\mu\nu} + \partial_{\mu}F_{\nu\lambda} + \partial_{\nu}F_{\lambda\mu} &= 0 \iff \epsilon^{\sigma\lambda\mu\nu}\partial_{\lambda}F_{\mu\nu} = 0 \\ \text{or} \quad \partial_{\lambda}\tilde{F}^{\sigma\lambda} &= 0, \ \tilde{F}^{\sigma\lambda} &= \frac{1}{2}\epsilon^{\sigma\lambda\mu\nu}F_{\mu\nu} \end{split}$$

In terms of **E** and **B**:

$$\nabla \cdot \mathbf{E} = 0, \ \dot{\mathbf{E}} = \nabla \times \mathbf{B}$$
 dynamical equations $\nabla \cdot \mathbf{B} = 0, \ \dot{\mathbf{B}} = -\nabla \times \mathbf{E}$ Bianchi identities

Remarks

• Lagrangian density does not depend on \dot{A}_0 , since A_0 is not really dynamical.

$$\nabla \cdot \boldsymbol{E} = 0 \rightarrow \nabla^2 A_0 + \nabla \cdot \dot{\boldsymbol{A}} = 0$$

Solve this <u>Poisson</u> equation for $A_0(\mathbf{x}, t) = \frac{1}{4\pi} \int \mathrm{d}^3 y \frac{\nabla \cdot \dot{A}(\mathbf{y}, t)}{|\mathbf{y} - \mathbf{x}|}$. Thus A_0 is given in terms of the other components of A.

• Gauge invariance ensures field strength tensor invariant under the transformation $A_{\mu} \mapsto A_{\mu} - \partial_{\mu} X$ due to commuting derivatives. This leads to gauge invariance of Maxwell equations. Choose X to satisfy $\partial_{\mu}\partial^{\mu}X = \partial^{2}X = \partial_{\mu}A^{\mu}$ allows us to demand the condition (Lorenz condition)

$$\partial_{\mu}A^{\mu} = 0 \tag{5.1.2}$$

such that A_{μ} belongs to the "Lorenz gauge" and reduces the degrees of freedom from 4 to 3.

• Further freedom is eliminated by adding any X with $\partial^2 X = 0$, e.g. $\partial_t X = A_0$. Then we get the Coulomb or radiation gauge

$$A_0 = 0, \ \nabla \cdot \mathbf{A} = 0 \tag{5.1.3}$$

Note: vice versa imposing $\nabla \cdot \mathbf{A} = 0$ first, yields $A_0 = 0$. In Coulomb gauge:

$$\begin{aligned}
\mathbf{E} &= -\dot{\mathbf{A}}, \ \mathbf{B} = \nabla \times \mathbf{A}, \ \nabla \times \mathbf{A} = 0 \\
-\ddot{\mathbf{A}} &= \dot{\mathbf{E}} \stackrel{\text{Maxwell}}{=} \nabla \times \mathbf{B} = \nabla \times (\nabla \times \mathbf{A}) = \nabla (\underbrace{\nabla \cdot \mathbf{A}}_{=0}) - \nabla^2 \mathbf{A} \\
&\Rightarrow \partial^2 \mathbf{A} = 0
\end{aligned}$$

This wave equation is massless KG equation for each spatial component.

Then the solutions are obvious: $\mathbf{A} = \epsilon e^{-ik \cdot x}$ with $k^2 = 0$ and $\epsilon \cdot \mathbf{k} = 0$. The polarization vector ϵ is transverse to \mathbf{k} .

Can write the lagrangian in Coulmb gauge

$$\mathcal{L}_{\rm EM} = \frac{1}{2}\dot{\boldsymbol{A}}\dot{\boldsymbol{A}} - \frac{1}{2}\boldsymbol{B}\cdot\boldsymbol{B}$$

Then the conjugate momentum to \mathbf{A} is $\mathbf{\Pi} = \frac{\partial \mathcal{L}}{\partial \dot{A}} = \dot{\mathbf{A}} = -\mathbf{E}$. It has only 3 components, there is no conjugate momentum to A_0 !. Because of Coulomb gauge $\mathbf{\Pi}$ is subject to the constraint $\nabla \cdot \mathbf{\Pi} = 0$

Hamiltonian

$$H_{\rm EM} = \int d^3x \left(\frac{1}{2} \mathbf{\Pi} \cdot \mathbf{\Pi} + \frac{1}{2} \mathbf{B} \cdot \mathbf{B} \right)$$

5.2 Quantizing the Maxwell field

We would like to impose canonical commutation relations, à la

$$[A_i(\mathbf{x}), A_j(\mathbf{y})] = [\Pi_i(\mathbf{x}), \Pi_j(\mathbf{y})] = 0$$
$$[A_i(\mathbf{x}), \Pi_j(\mathbf{y})] = i\delta_{ij}\delta^{(3)}(\mathbf{x} - \mathbf{y})$$

However this cannot be true. Take either derivative of the last equation and it needs to vanish due to $\nabla \cdot \mathbf{A} = \nabla \cdot \mathbf{\Pi} = 0$. But

$$[\partial^i A_i(\boldsymbol{x}), \Pi_k(\boldsymbol{y})] = i\delta_{ij}\partial^i \delta^{(3)}(\boldsymbol{x} - \boldsymbol{y})$$

here the derivative is taken with respect to \mathbf{x} , i.e. $\partial^i = \frac{\partial}{\partial x_i}$. Replace δ_{ij} by Δ_{ij}

$$[\partial^{i} A_{i}(\mathbf{x}), \Pi_{j}(\mathbf{y})] = i\Delta_{ij}\partial^{i} \frac{1}{(2\pi)^{3}} \int d^{3}k e^{i\mathbf{k}\cdot(\mathbf{x}-\mathbf{y})}$$
$$= -\frac{1}{(2\pi)^{3}} \int d^{3}k (k^{i}\Delta_{ij}) e^{i\mathbf{k}\cdot(\mathbf{x}-\mathbf{y})} \stackrel{!}{=} 0$$

It works for $\Delta_{ij} = \delta_{ij} - \frac{k_i k_j}{k^2}$ in momentum space or $\Delta_{ij} = \delta_{ij} - \nabla^{-2} \partial_i \partial_j$ in position space.

$$[A_i(\mathbf{x}), \Pi_i(\mathbf{y})] = i \left(\delta_{ij} - \nabla^{-2} \partial_i \partial_j \right) \delta^{(3)}(\mathbf{x} - \mathbf{y})$$
(5.2.1)

As before we have the mode expansion

$$\begin{aligned} \boldsymbol{A}(\boldsymbol{x}) &= \int \frac{\mathrm{d}^3 k}{(2\pi)^3 \sqrt{2|\boldsymbol{k}|}} \left(\boldsymbol{a}_{\boldsymbol{k}} e^{i\boldsymbol{k}\cdot\boldsymbol{x}} + \boldsymbol{a}_{\boldsymbol{k}}^{\dagger} e^{-i\boldsymbol{k}\cdot\boldsymbol{x}} \right) \\ \boldsymbol{\Pi}(\boldsymbol{x}) &= \int \frac{\mathrm{d}^3 k}{(2\pi)^3} (-i) \sqrt{\frac{|\boldsymbol{k}|}{2}} \left(\boldsymbol{a}_{\boldsymbol{k}} e^{i\boldsymbol{k}\cdot\boldsymbol{x}} - \boldsymbol{a}_{\boldsymbol{k}}^{\dagger} e^{-i\boldsymbol{k}\cdot\boldsymbol{x}} \right) \end{aligned}$$

with $\mathbf{k} \cdot \mathbf{a_k} = \mathbf{k} \cdot \mathbf{a_k^{\dagger}} = 0$.

Introduce two orthogonal polarization vectors $\boldsymbol{\epsilon}^{(1)}(\boldsymbol{k})$ and $\boldsymbol{\epsilon}^{(2)}(\boldsymbol{k})$ for each \boldsymbol{k} .

$$a_{\mathbf{k}} = a_{\mathbf{k}}^{(1)} \boldsymbol{\epsilon}^{(1)} + a_{\mathbf{k}}^{(2)} \boldsymbol{\epsilon}^{(2)} = \sum_{\lambda=1}^{2} a_{\mathbf{k}}^{(\lambda)} \boldsymbol{\epsilon}^{(\lambda)}(\mathbf{k})$$
with $\mathbf{k} \cdot \boldsymbol{\epsilon}^{(1)}(\mathbf{k}) = \mathbf{k} \cdot \boldsymbol{\epsilon}^{(2)}(\mathbf{k}) = 0$, $\boldsymbol{\epsilon}^{(\lambda)} \cdot \boldsymbol{\epsilon}^{(\lambda;)} = \delta_{\lambda \lambda'}$

Creation and annihilation operator have the standard commutation relations

$$[a_{\mathbf{k}}^{(\lambda)}, a_{\mathbf{k}'}^{(\lambda')\dagger}] = (2\pi)^3 \delta_{\lambda\lambda'} \delta^{(3)}(\mathbf{k} - \mathbf{k}')$$

$$(5.2.2)$$

and all other commutators vanish. Geometrically, still possible to write including the unphysical longitudinal components:

$$[\boldsymbol{a_k}, \boldsymbol{a_l}] = [\boldsymbol{a_k^{\dagger}}, \boldsymbol{a_l^{\dagger}}] = 0$$
$$[\boldsymbol{a_k^i}, \boldsymbol{a_l^{j\dagger}}] = (2\pi)^3 \left(\delta^{ij} - \frac{k^i k^j}{k^2}\right) \delta^{(3)}(\boldsymbol{k} - \boldsymbol{l})$$

 $a_{\pmb{k}}^{(\lambda)}$ and $a_{\pmb{k}}^{(\lambda)\dagger}$ create and destroy photons of momentum \pmb{k} , energy $|\pmb{k}|$ and (electric) polarization along $\pmb{\epsilon}^{(\lambda)}(\pmb{k})$.

Next steps are analogous to KG theory.

Hamiltonian

$$H = \frac{1}{2} \int d^3x \left(\mathbf{E}^2 + \mathbf{B}^2 \right) = \frac{1}{2} \int d^3x \left(\dot{\mathbf{A}}^2 + (\nabla \times \mathbf{A}) \cdot (\nabla \times \mathbf{A}) \right)$$

using identity $\mathbf{A} \cdot (\mathbf{B} \times \mathbf{C}) = \mathbf{B} \cdot (\mathbf{C} \times \mathbf{A})$

$$= \frac{1}{2} \int d^3x \left(\dot{\boldsymbol{A}}^2 + \boldsymbol{A} \cdot \boldsymbol{\nabla} \times (\boldsymbol{\nabla} \times \boldsymbol{A}) \right)$$

using the identity $\nabla \times (\nabla \times \mathbf{A}) = \nabla (\nabla \cdot \mathbf{A} - \nabla^2 \mathbf{A})$

$$= \frac{1}{2} \int d^3x \left(\dot{\boldsymbol{A}}^2 - \boldsymbol{A} \cdot \nabla^2 \boldsymbol{A} + \boldsymbol{A} \cdot \nabla (\nabla \cdot \boldsymbol{A}) \right)$$

using coulomb gauge condition

$$= \frac{1}{2} \int d^3x \left(\dot{\boldsymbol{A}}^2 - \boldsymbol{A} \cdot \nabla^2 \boldsymbol{A} \right)$$

the first term vanishes and use normal ordering

$$= \int \frac{\mathrm{d}^3 k}{(2\pi)^3} |\mathbf{k}| \mathbf{a}_{\mathbf{k}}^{\dagger} \cdot \mathbf{a}_{\mathbf{k}} = \sum_{k=1}^2 \int \frac{\mathrm{d}^3 k}{(2\pi)^3} |\mathbf{k}| a_{\mathbf{k}}^{(\lambda \dagger)} a_{\mathbf{k}}^{\lambda}$$

Heisenberg field

$$\boldsymbol{A}(\boldsymbol{x},t) = \int \frac{\mathrm{d}^3 k}{(2\pi)^3} \frac{1}{\sqrt{2|\boldsymbol{k}|}} \left(\boldsymbol{a}_{\boldsymbol{k}} e^{-ik \cdot x} + \boldsymbol{a}_{\boldsymbol{k}}^{\dagger} e^{ik \cdot x} \right)$$

Photon propagator

$$\langle 0|TA_{i}(x)A_{j}(y)|0\rangle =: D_{ij}^{tr}(x-y) = \int \frac{\mathrm{d}^{4}k}{(2\pi)^{4}} \frac{i}{k^{2}+i\epsilon} \left(\delta_{ij} - \frac{k_{i}k_{j}}{|\mathbf{k}|^{2}}\right) e^{-ik\cdot(x-y)}$$
(5.2.3)

tr stands for transverse: photon polarization perpendicular to its momentum. This is **NOT** the final version of the photon propagator!

5.3 Inclusion of matter - QED

$$\mathcal{L}_{\text{QED}} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \bar{\psi} (i\not\!\!D - m) \psi \tag{5.3.1}$$

where $D_{\mu} = \partial_{\mu} + ieA_{\mu}$ is the (gauge) covariante derivative

$$= \mathcal{L}_{EM} + \mathcal{L}_D - e \underbrace{\bar{\psi} \gamma^{\mu} \psi A_{\mu}}_{i^{\mu}}$$
 (5.3.2)

Field equations would be

$$\partial_{\mu}F^{\mu\nu} = ej^{\nu}$$
 $(iD - m)\psi = 0$

where ej^{ν} is the electromagnetic 4-current.

Gauge invariance under the transformation

$$\begin{cases} \psi(x) \longmapsto \psi'(x) = e^{ie\chi(x)} \psi \\ A_{\mu}(x) \longmapsto A'_{\mu}(x) = A_{\mu}(x) - \partial_{\mu}\chi(x) \end{cases}$$

To check the consistence: cavariant derivative transforms like $D_{\mu} \longmapsto D'_{\mu} \psi'(x) = e^{ie\chi(x)} D_{\mu} \psi(x)$. Since the adjoint spinor transforms like $\bar{\psi}(x) \longmapsto \bar{\psi}'(x) = \bar{\psi}(x) e^{-ie\chi(x)}$, the Lagrangian and field equations are gauge invariant.

Again we choose Coulomb gauge $\nabla \cdot \mathbf{A} = 0$, then equation for A^0 :

$$\partial_{i}F^{i0} = ej^{0}$$

$$\Rightarrow -\nabla^{2}A^{0} = ej^{0} = e\bar{\psi}\gamma^{0}\psi$$

$$= e\bar{\psi}\gamma^{0}\psi = e\psi^{\dagger}\psi$$

$$= e\rho(x)$$

$$A^{0}(\mathbf{x}, t) = e\int d^{3}y \frac{\rho(\mathbf{y}, t)}{4\pi |\mathbf{x} - \mathbf{y}|}$$
(5.3.3)

We want to derive the interaction Hamiltonian. Note

$$\int d^3x \frac{1}{2} \boldsymbol{E}^2 = \int d^3x \frac{1}{2} \left(\dot{\boldsymbol{A}} + \boldsymbol{\nabla} A^0 \right)^2$$

cross terms vanish after integration by parts due to $\nabla \cdot \dot{A} = 0$

$$= \int d^3x \frac{1}{2} \left(\dot{\mathbf{A}}^2 + (\nabla A^0)^2 \right)$$

$$= \int d^3x \frac{1}{2} \left(\dot{\mathbf{A}}^0 - A^0 \nabla^2 A^0 \right)$$

$$= \int d^3x \frac{1}{2} \dot{\mathbf{A}}^2 + \underbrace{\frac{e^2}{2} \int d^3x d^3y \frac{\rho(\mathbf{x})\rho(\mathbf{y})}{4\pi |\mathbf{x} - \mathbf{y}|}}_{=\frac{e^2}{2} j^0 A_0}$$

Combined Hamiltonian

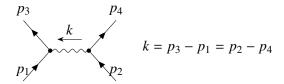
$$H = \int d^3x \left\{ \frac{1}{2} \mathbf{\Pi} \cdot \mathbf{\Pi} + \frac{1}{2} \mathbf{B} \cdot \mathbf{B} + i \bar{\psi} \mathbf{\gamma} \cdot \nabla \psi + m \bar{\psi} \psi \right\}$$
 free photon and fermion

$$+ \frac{e^2}{2} \int d^3x d^3y \frac{\rho(\mathbf{x})\rho(\mathbf{y})}{4\pi |\mathbf{x} - \mathbf{y}|} - e \int d^3x \mathbf{j} \cdot \mathbf{A}$$
 interactions

where $\rho = \psi^{\dagger} \psi = \bar{\psi} \gamma^0 \psi$, $\mathbf{j} = \bar{\psi} \boldsymbol{\gamma} \psi$ for 2 types of interactions.

5.4 Lorentz-invariant propagator

Consider e^-e^- scattering at $O(e^2)$



We expect this to involve

- spinors for external fermions
- $-ie\gamma^{\mu}$
- Photon propagator $D_{\mu\nu}(x-y)$

What we have found in Coulomb gauge is actually

- vertices $ie\gamma^i$, transverse propagator $D_{\mu\nu}^{\rm tr}(x-y)$
- vertices $\pm ie\gamma^0$, instantaneous Coumlomb interaction $\frac{1}{4\pi|\mathbf{x}-\mathbf{y}|}\delta(x^0-y^0)$

Effectively combine these propagators terms into $D_{\mu\nu}^{\text{Coul}}(x-y)$, where the $D_{00}^{\text{Coul}}(x-y) = \frac{1}{4\pi |\mathbf{x}-\mathbf{y}|} \delta(x^0 - y^0)$. This component in momentum space is simply

$$\int \frac{\mathrm{d}^3 k}{(2\pi)^3} \frac{e^{i\boldsymbol{k}\cdot\boldsymbol{r}}}{|\boldsymbol{k}|^2} = \frac{1}{4\pi|\boldsymbol{r}|}$$

Therefore Coumlomb propagator in momentum space:

$$D_{\mu\nu}^{\text{Coul}}(k) = \begin{cases} \frac{i}{k^2 + i\epsilon} \left(\delta_{ij} - \frac{k_i k_j}{|\mathbf{k}|^2} \right) & \mu = i, \nu = j \\ \frac{i}{|\mathbf{k}|^2} & \mu = \nu = 0 \\ 0 & \text{otherwise} \end{cases}$$

Consider contraction to scattering amplitude from vertex at x:

$$\sim e\bar{u}(p_3)\gamma^{\mu}u(p_1)e^{i(p_3-p_1)x}$$

current conservation $\partial_{\mu}j^{\mu}=0$ written in momentum space

$$\underbrace{(p_3-p_1)_\mu}_{k_\mu}\bar{u}(p_3)\gamma^\mu u(p_1)=0$$

so in the complete diagram $D_{\mu\nu}^{\mathrm{Coul}}$ occurs in a form

$$a^{\mu}D_{\mu\nu}^{\text{Coul}}(k)b^{\nu}$$

$$=a^{0}\frac{i}{|\mathbf{k}|^{2}}b^{0}+a^{i}\left[\frac{i}{k^{2}+i\epsilon}\left(\delta_{ij}-\frac{k_{i}k_{j}}{|\mathbf{k}|^{2}}\right)\right]b^{j}$$

where $k^{\mu}a^{\mu} = 0, k_{\mu}a^{\mu} = 0$

$$=i\left[\frac{\boldsymbol{a}\cdot\boldsymbol{b}}{k^{2}}\underbrace{-\frac{k_{0}^{2}a_{0}b_{0}}{k^{2}|\boldsymbol{k}|^{2}}+\frac{a_{0}b_{0}}{|\boldsymbol{k}|^{2}}}_{=\frac{-k_{0}^{2}a_{0}b_{0}+a_{0}b_{0}(k_{0}^{2}-|\boldsymbol{k}|^{2})}{k^{2}|\boldsymbol{k}|}^{2}}\right]$$

$$=\frac{i}{k^2}(\boldsymbol{a}\cdot\boldsymbol{b}-a_0b_0)=-\frac{i}{k}a_\mu b^\mu$$

Conclusion in this diagram (and in fact, in general), we may replace the $D_{\mu\nu}^{\text{Coul}}(k)$ by the manifestly Lorentz covariant propagator

$$D_{\mu\nu}(k) = -\frac{ig_{\mu\nu}}{k + i\epsilon} \tag{5.4.1}$$

This can be generalised to

$$D_{\mu\nu}(k) = -\frac{i}{k^2 + i\epsilon} \left(g_{\mu\nu} - (1 - \alpha) \frac{k_{\mu} k_{\nu}}{k^2} \right)$$
 (5.4.2)

as, by current consevation, additional term doesn't contribute.

Feynman gauge $\alpha = 1$; Landau gauge $\alpha = 0$.

Remark one can also try to quantise photons in a manifestly covariant way, imposing Lorentz gauge $\partial_{\mu}A^{\mu}=0$

$$[A_{\mu}(\boldsymbol{x}), \Pi_{\nu}(\boldsymbol{y})] = ig_{\mu\nu}\delta^{(3)}(\boldsymbol{x} - \boldsymbol{y})$$

This is trouble since $\Pi^0 = \frac{\partial \mathcal{L}}{\partial \dot{A}_0} = 0$. This cannot hold!

We thus change the Lagrangian

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{2\alpha}(\partial_{\mu}A^{\mu})^2$$

with "gauge fixing term". The equation of motion from this is

$$\partial^2 A^{\mu} - \left(1 - \frac{1}{\alpha}\right) \partial^{\mu} (\partial_{\lambda} A^{\lambda}) = 0$$

e.g. $\alpha = 1$ is the Feynman gauge.

With this Lagrangian we can the 0th component of conjugate momentum

$$\Pi^0 = -\frac{1}{\alpha} \partial_\mu A^\mu$$

but this seems as bad as before!

We cannot impose Coulomb gauge condition $\partial_{\mu}A^{\mu}=0$ as an operator identity. Instead demand a weaker condition $\langle \text{out}|\partial_{\mu}A^{\mu}|\text{in}\rangle=0$ for all physical states.

This in turn tells us which states are actually physical. The 4 polarisation states consist of physical, timelike(scalar) and longitudinal states. The negative-norm states cancel each ther out (Gopta-Bleuler formalism).

Feynman rules for QED diagrams constucted from electron (positron) — and photon , and photon , rules for fermions are valid as before.

In addition

• vertex
$$\mu = -ie\gamma^{\mu};$$

- photon propagator $\begin{picture}(1,0) \put(0,0){\line(1,0)} \put($
- external photons

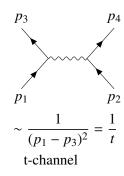
$$\mu \stackrel{\kappa_{\text{in}}}{\longleftarrow} \nu = \epsilon_{\mu}$$

$$\mu \xrightarrow{k_{\text{out}}} \nu = \epsilon_{\nu}^{*}$$

 ϵ_{μ} polarisation vector of in/out photon and ϵ_{μ}^* for out photon required for complex (circular) polarisation.

5.5 QED process at tree level

Example $e^-e^-
ightarrow e^-e^-$



$$p_{1}$$

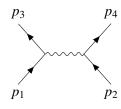
$$p_{1}$$

$$p_{2}$$

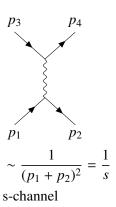
$$\frac{1}{(p_{1} - p_{4})^{2}} = \frac{1}{u}$$

$$\text{u-channel}$$

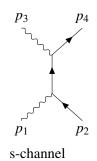
Example $e^-e^+ \rightarrow e^-e^+$

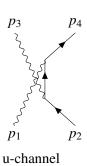


$$\sim \frac{1}{(p_1 - p_3)^2} = \frac{1}{t}$$
t-channel

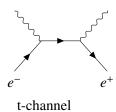


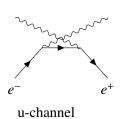
Compton scattering $\gamma e^- \to \gamma e^-$





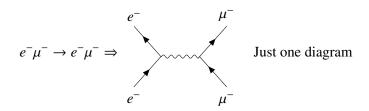
Example $e^+e^- o \gamma\gamma$

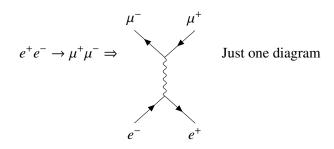




These are important for lifetime of positronium.

All these amplitudes are $O(e^2)$, $\alpha = \frac{e^2}{4\pi} = \frac{1}{137.036}$ the fine structure constant. Muons μ^{\pm} , like electrons, just ca. 200 times heavier.





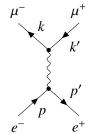
 μ^{\pm} decay into e^{\pm} and neutrinos in weak interactions.

For tree level diagrams the photon propagator does not need to have the $i\epsilon$ in the denominator, since we will never be able to see a singularity/pole.

5.5.1 Some hints and tricks for cross section calculations

See application in exercises!

Example $e^+e^- \rightarrow \mu^+\mu^-$



$$\begin{split} iM &= \bar{v}_e^{s'}(-ie\gamma^\mu)u_e^s(p) \left. \frac{-ig_{\mu\nu}}{s} \right|_{s=q^2} \bar{u}_\mu^r(k)(-ie\gamma^\nu)v_\mu^{r'}(k') \\ &= \frac{ie^2}{s} \left(\bar{v}_e(p')\gamma^\mu u_e(p) \right) \left(\bar{u}_\mu(k)\gamma_\mu v_\mu(k') \right) \end{split}$$

See section ??, $|M|^2$ is needed for cross section. M^* involves things like

$$(\bar{\nu}\gamma^{\mu}u)^* = (\bar{\nu}\gamma^{\mu}u)^{\dagger} = u^{\dagger}\gamma^{\mu\dagger}\gamma_0^{\dagger}\nu$$
$$= u^{\dagger}\gamma_0\gamma^{\mu}\gamma_0\gamma_0\nu = \bar{u}\gamma^{\mu}\nu$$

So

$$|M|^2 = \frac{e^4}{s^2} \left[\bar{v}(p') \gamma^{\mu} u(p) \bar{u}(p) \gamma^{\nu} v(p') \right]_{e^{\pm}} \cdot \left[\bar{u}(k) \gamma_{\mu} v(p) \bar{v}(k') \gamma_{\nu} u(k) \right]_{\mu^{\pm}}$$

Unpolarized scattering= $\frac{1}{4} \sum_{r,s,r',s'} |M|^2$.

Now $\bar{v}\gamma^{\mu}u$, $\bar{u}\gamma^{\nu}v$ etc. are scalars in Dirac/spinor space:

$$\sum_{s,s'} \bar{v}_{s'} p' \gamma^{\mu} u_{s}(p) \bar{u}_{s}(p) \gamma^{\nu} v_{s'}(p')$$
(taking trace of scalar)
$$= \sum_{s,s'} \operatorname{Tr} \left(\bar{v}_{s'} p' \gamma^{\mu} u_{s}(p) \bar{u}_{s}(p) \gamma^{\nu} v_{s'}(p') \right)$$

$$= \sum_{s,s'} \operatorname{Tr} \left(v_{s'}(p') \bar{v}_{s'}(p') \gamma^{\mu} u_{s}(p) \bar{u}_{s}(p) \gamma^{\nu} \right)$$

using spin sums

$$= \operatorname{Tr} \left((p' - m) \gamma^{\mu} (p + m) \gamma^{\nu} \right)$$

Trace technology

- remember $Tr\gamma_{\mu} = 0$
- $\operatorname{Tr}(\gamma^{\mu}\gamma^{\nu}) = 4g_{\mu\nu}$
- Tr(odd number of γ) = 0
- $\operatorname{Tr}(\gamma_{\mu}\gamma_{\nu}\gamma_{\alpha}\gamma_{\beta}) = 4\left(g_{\mu\nu}g_{\alpha\beta} + g_{\mu\beta}g_{\nu\alpha} g_{\mu\alpha}g_{\nu\beta}\right)$
- more rules involving γ_5 (weak interactions!)

So

$$\operatorname{Tr}\left((p'-m)\gamma^{\mu}(p+m)\gamma^{\nu}\right) = 4\left(p'^{\mu}p^{\nu} + p'^{\nu}p^{\mu} - g^{\mu\nu}(p\cdot p' + m^2)\right)$$

Mandelstam variables with 4 equal masses, center-of-mass system (CMS):

$$p = (E, \mathbf{p}), \quad p' = (E, -\mathbf{p}), \quad k = (E, \mathbf{k}), \quad \theta = \measuredangle(\mathbf{p}, \mathbf{k})$$

$$s \stackrel{\text{CMS}}{=} (p + p')^2 = 4E^2$$
 (5.5.1)

$$t = (p - k)^{2} = -(\mathbf{p} - \mathbf{k})^{2} = -2|\mathbf{p}|^{2}(1 - \cos\theta)$$
 (5.5.2)

$$u = (p' - k)^2 = -2|\mathbf{p}|^2(1 + \cos\theta)$$
 (5.5.3)

$$|\mathbf{p}|^2 = E^2 - m^2 = \frac{s}{4} - m^2 \tag{5.5.4}$$

Only 2 Mandelstam variables are independent.

$$s + t + u = (p_1 + p_2)^2 + (p_1 - p_3)^2 + (p_1 - p_4)^2$$

$$= p_1^2 + p_2^2 + p_3^2 + p_4^2 + 2p_1 \underbrace{(p_1 + p_2 - p_3 - p_4)}_{=0}$$

$$= \sum_i m_i^2 = \text{const}$$

photon polarisation sums Analogy to fermion spin sums before, Feynman rules for external photons involve $\epsilon^{(*)}_{\mu}$; e.g. Compton amplitude of the form

$$M \sim \epsilon_{\mu}^*(p_3)\epsilon_{\nu}(p_1)T^{\mu\nu}$$

Thus

$$\sum_{\text{spin, pol.}} |M|^2 = \sum_{\text{spin, pol.}} \epsilon_{\mu}^*(p_3) \epsilon_{\alpha}(p_3) \epsilon_{\beta}^*(p_1) \epsilon_{\nu}(p_1) T^{\mu\nu} T^{\alpha\beta^*}$$

How can we simplify $\sum_{pol} \epsilon_{\mu}^*(k) \epsilon_{\nu}(k)$? Again we have only 2 physical polarisation states, but want to do it in a covariant form.

Assume a simpler process (than Compton) with a single external photon, $\epsilon_{\mu}^{*}(k)M^{\mu}$. Choose

$$k^{\mu} = (k, 0, 0, k), \quad \epsilon_{(1)}^{\mu} = (0, 1, 0, 0), \quad \epsilon_{(2)}^{\mu} = (0, 0, 1, 0)$$

so
$$\sum_{\text{pol}} |\epsilon_{\mu}^*(k)M^{\mu}|^2 = |M_1|^2 + |M_2|^2$$

Remember that photon coupled source j^{μ} , current sonvervation $\partial_{\mu}j^{\mu}=0$. We will see (next term) this holds in general as Ward identity

$$k_{\mu}M^{\mu} = 0 \tag{5.5.5}$$

In exercises, show $p_{3\mu}T^{\mu\nu} = 0 = p_{1\nu}T^{\mu\nu}$ for Compton Here $kM^0 - kM^3 = 0 \Rightarrow M^0 = M^3$ and we can rewrite

$$\sum_{pol} \epsilon_{\mu}^{*} \epsilon_{\nu} M^{\mu} M^{*\nu} = |M_{1}|^{2} + |M_{2}|^{2} + \underbrace{|M_{3}|^{2} - |M_{0}|^{2}}_{=0} = -g_{\mu\nu} M^{\mu} M^{*\nu}$$

so effectively

$$\sum_{pol} \epsilon_{\mu}^*(k)\epsilon_{\nu}(k) = -g_{\mu\nu} \tag{5.5.6}$$

side remark

- KG propagator $\frac{i}{p^2 M^2 + i\epsilon}$
- Dirac propagator $\frac{i(\not p+m)}{p^2-m^2+i\epsilon} = \frac{i\sum_s u_s(p)\bar{u}_s(p)}{p^2-m^2+i\epsilon}$
- Photon propagator $\frac{-ig_{\mu\nu}}{p^2+i\epsilon} = \frac{i\sum_{pol} \epsilon_{\mu}^*(p)\epsilon_{\nu}(p)}{p^2+i\epsilon}$