Quantum Field Theory

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Contents

1 Classical field theory

1.1 Field theory in continuum

Euler-Lagrange-equation

$$\partial_{\mu} \left(\frac{\partial \mathcal{L}}{\partial (\partial_{\mu} \phi)} \right) - \frac{\partial \mathcal{L}}{\partial \phi} = 0 \tag{1.1.1}$$

momentum density

$$\pi(x) = \frac{\partial \mathcal{L}}{\partial \dot{\phi}(\mathbf{x})} \tag{1.1.2}$$

Hamiltonian density

$$\mathcal{H}(\phi(\mathbf{x}), \pi(\mathbf{x})) = \pi(\mathbf{x})\dot{\phi}(\mathbf{x}) - \mathcal{L}(\phi, \partial_{\mu}\phi)$$
(1.1.3)

1.2 Noether Theorem

If a Lagrangian field theory has an infinitisimal symmetry, then there is an associated current j^{μ} , which is conserved.

$$\partial_{\mu}j^{\mu} = 0 \tag{1.2.1}$$

$$j^{\mu} = \frac{\partial \mathcal{L}}{\partial(\partial_{\mu}\phi)} \Delta \phi - X^{m} u \tag{1.2.2}$$

Energy-momentum tensor (stress-energy tensor)

Asymmetric version

$$\Theta_{\nu}^{\mu} = \frac{\partial \mathcal{L}}{\partial(\partial_{\mu}\phi)} \partial_{\nu}\phi - \delta_{\nu}^{\mu}\mathcal{L}$$
 (1.2.3)

General version

$$T^{\mu\nu} = \Theta^{\mu\nu} + \partial_{\lambda} f^{\mu\nu\lambda} \tag{1.2.4}$$

with $f^{\lambda\mu\nu} = -f^{\mu\lambda\nu}$ or $\partial_{\mu}\partial\nu f^{\lambda\mu\nu} = 0$

2 Klein-Gordon theory

(Real) Lagrangian density

$$\mathcal{L} = \frac{1}{2} \partial_{\mu} \phi \partial^{\mu} \phi - \frac{m^2}{2} \phi^2 \tag{2.0.1}$$

Quantization

$$[\phi(\mathbf{x}), \phi(\mathbf{x}')] = [\pi(\mathbf{x}), \pi(\mathbf{x}')] = 0$$

$$[\phi(\mathbf{x}), \pi(\mathbf{x}')] = i\delta^{(3)}(\mathbf{x} - \mathbf{x}')$$
 (2.0.2)

Decomposition into Fourier modes

$$\phi(\mathbf{x}) = \int \frac{\mathrm{d}^3 p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} \left(a_p e^{i\mathbf{p}\cdot\mathbf{x}} + a_p^{\dagger} e^{-i\mathbf{p}\cdot\mathbf{x}} \right)$$
(2.0.3)

$$\pi(\mathbf{x}) = \int \frac{\mathrm{d}^3 p}{(2\pi)^3} (-i) \sqrt{\frac{E_p}{2}} \left(a_{\mathbf{p}} e^{i\mathbf{p}\cdot\mathbf{x}} - a_{\mathbf{p}}^{\dagger} e^{-i\mathbf{p}\cdot\mathbf{x}} \right)$$
(2.0.4)

thus the commutation relations for ladder operators:

$$\left[a_{\boldsymbol{p}}, a_{\boldsymbol{p}'}\right] = \left[a_{\boldsymbol{p}}^{\dagger}, a_{\boldsymbol{p}'}^{\dagger}\right] = 0 \tag{2.0.5}$$

$$\left[a_{\mathbf{p}}, a_{\mathbf{p'}}^{\dagger}\right] = (2\pi)^{3} \delta^{(3)}(\mathbf{p} - \mathbf{p'})$$
(2.0.6)

Hamiltonian in terms of ladder operator

$$H = \int \frac{\mathrm{d}^3 p}{(2\pi)^3} E_p \left(a_{\mathbf{p}} a_{\mathbf{p}}^{\dagger} + \frac{1}{2} \left[a_{\mathbf{p}}, a_{\mathbf{p}}^{\dagger} \right] \right) \tag{2.0.7}$$

Normlisation it's also lorentz-invariante

$$\langle p|q\rangle = 2E_p(2\pi)^3 \delta^{(3)}(\boldsymbol{p} - \boldsymbol{q}) \tag{2.0.8}$$

2.1 Heisenberg-picture fields

Heisenberg-picture

$$|\psi_H\rangle = e^{iHt}|\psi_s(t)\rangle$$
 (2.1.1)

$$O_H(t) = e^{iHt} O_S e^{-iHt} (2.1.2)$$

Field operator

$$\phi(x) = \phi(\mathbf{x}, t) = \int \frac{\mathrm{d}^3 p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} \left(a_p e^{ipx} + a_p^{\dagger} e^{-ipx} \right)$$
 (2.1.3)

2.2 Commutations and propogators

Commutations

$$[\phi(x), \phi(y)] = D(x - y) - D(y - x) \begin{cases} = 0 & \text{if } (x - y) \text{ is space-like} \\ \neq 0 & \text{otherwise} \end{cases}$$
 (2.2.1)

$$D(x-y) = \int \frac{\mathrm{d}^3 p}{(2\pi)^3} \frac{1}{2E_p} e^{-ip(x-y)}$$
 (2.2.2)

Propogator

$$\langle 0|\phi(x)\phi(y)|0\rangle = D(x-y) \tag{2.2.3}$$

Feynman propagator

$$D_F(x - y) = \langle 0|T\phi(x)\phi(y)|0\rangle$$

= $\Theta(x^0 - y^0)D(x - y) + \Theta(y^0 - x^0)D(y - x)$ (2.2.4)

$$D_F(x-y) = \int \frac{d^4p}{(2\pi)^4} \frac{i}{p^2 - m^2 + i\epsilon} e^{-ip(x-y)}$$
 (2.2.5)

3 Quantization of the Dirac field

3.1 Dirac equation

$$\left(i\gamma^{\mu}\partial_{\mu}-m\right)\phi(x)=0\tag{3.1.1}$$

Standard representation (Dirac's)

$$\gamma_0 = \begin{pmatrix} \mathbb{1}_2 & 0 \\ 0 & -\mathbb{1}_2 \end{pmatrix} \quad \boldsymbol{\gamma} = \begin{pmatrix} 0 & \boldsymbol{\sigma} \\ -\boldsymbol{\sigma} & 0 \end{pmatrix}$$
 (3.1.2)

Lorentz transformation

$$\Lambda = \exp\left(\frac{1}{2}\omega_{\mu\nu}M^{\mu\nu}\right) \tag{3.1.3}$$

with ω set of parameters and M the generator of Lie algebra.

Spinor representation

$$S^{\rho\sigma} = \frac{1}{4} \left[\gamma^{\rho}, \gamma^{\sigma} \right] = \frac{1}{2i} \sigma^{\rho\sigma} \tag{3.1.4}$$

(3.1.5)

Spinor transformation

$$S(\Lambda) = \exp\left(\frac{1}{2}\omega_{\mu\nu}S^{\mu\nu}\right) \tag{3.1.6}$$

$$\psi_a'(x) = S_{ab}(\Lambda)\psi_b(\Lambda^{-1}x) \tag{3.1.7}$$

adjoint spinor

$$\bar{\psi} = \psi^{\dagger} \gamma^0 \tag{3.1.8}$$

Fifth gamma matrix

$$\gamma^5 := i\gamma^0 \gamma^1 \gamma^2 \gamma^3 \tag{3.1.9}$$

$$\left\{\gamma^{\mu}, \gamma^{5}\right\} = 0 \tag{3.1.10}$$

$$(\gamma^5)^2 = \mathbb{1}_4 \tag{3.1.11}$$

Plane wavesolutions

$$\psi(x) = \begin{cases} u(p)e^{-ipx} & \text{positive frequency} \\ v(p)e^{ipx} & \text{negative frequency} \end{cases}$$
(3.1.12)

$$u_s(p) = \sqrt{E_p + m} \left(\frac{\chi_s}{\frac{u \cdot p}{E_p + m} \chi_s} \right) e^{-ipx} v_s(p) = \sqrt{E_p + m} \left(\frac{\frac{u \cdot p}{E_p + m} \tilde{\chi}_s}{\tilde{\chi}_s} \right) e^{ipx}$$
(3.1.13)

with

$$\chi_{\frac{1}{2}} = \begin{pmatrix} 1 \\ 0 \end{pmatrix} \quad x_{-\frac{1}{2}} = \begin{pmatrix} 0 \\ 1 \end{pmatrix}$$
$$s = \pm \frac{1}{2} \quad \tilde{\chi}_s = \chi_{-s}$$

Orthogonality of spinor

$$\bar{u}_s(p)u_{s'}(p) = -\bar{v}_s(p)v_{s'}(p) = 2m\delta_{ss'}$$
(3.1.14)

$$\bar{u}_s(p)v_{s'}(p) = 0 (3.1.15)$$

Spin sums

$$\sum_{s} u_{s}(p)\bar{u}_{s}(p) = p + m \tag{3.1.16}$$

$$\sum_{s} v_{s}(p)\bar{v}_{s}(p) = p - m \tag{3.1.17}$$

3.2 Dirac Lagrangian and quantization

$$\mathcal{L} = \bar{\psi}(x)(i\partial \!\!\!/ - m)\psi(x) \tag{3.2.1}$$

Quantization

$$\left\{\psi_a(\mathbf{x}), \psi_b^{\dagger}(\mathbf{x}')\right\} = \delta_{ab}\delta^{(3)}(\mathbf{x} - \mathbf{x}') \tag{3.2.2}$$

$$\{\psi_a(\mathbf{x}), \psi_b(\mathbf{x}')\} = \{\psi_a^{\dagger}(\mathbf{x}), \psi_b^{\dagger}(\mathbf{x}')\} = 0$$
(3.2.3)

Field operators

$$\psi(\mathbf{x}) = \int \frac{\mathrm{d}^3 p}{(2\pi)^3 \sqrt{2E_p}} \sum_s (a_{\mathbf{p}}^s u_s(\mathbf{p}) e^{i\mathbf{p}\cdot\mathbf{x}} + b_{\mathbf{p}}^{s\dagger} v_s(\mathbf{p}) e^{-i\mathbf{p}\cdot\mathbf{x}})$$
(3.2.4)

thus the anticommutations of ladder operators:

$$\begin{cases}
a_{\mathbf{p}}^{s}, a_{\mathbf{p'}}^{s'\dagger} \\
\end{cases} = \begin{cases}
b_{\mathbf{p}}^{s}, b_{\mathbf{p'}}^{s'\dagger} \\
\end{cases} = (2\pi)^{3} \delta_{ss'} \delta^{(3)}(\mathbf{p} - \mathbf{p'})$$

$$\{a, a\} = \begin{cases}
a^{\dagger}, a^{\dagger} \\
\end{cases} = \dots = 0$$

Hamiltonian in terms of Fourier modes (with normal ordering)

$$H = \int \frac{\mathrm{d}^{3} p}{(2\pi)^{3}} \sum_{s} E_{p} (a_{\mathbf{p}}^{s\dagger} a_{\mathbf{p}}^{s} - b_{\mathbf{p}}^{s\dagger} b_{\mathbf{p}}^{s})$$
 (3.2.5)

3.3 Particles and antiparticles

$$Q = e \int d^3x \psi^{\dagger}(x)\psi(x)$$
 (3.3.1)

$$: Q := e \int \frac{\mathrm{d}^{3} p}{(2\pi)^{3}} \sum_{s} (a_{\mathbf{p}}^{s\dagger} a_{\mathbf{p}}^{s} - b_{\mathbf{p}}^{s\dagger} b_{\mathbf{p}}^{s})$$
 (3.3.2)

3.4 Dirac propagator and anticommutators

$$S_{ab}(x - y) = \{ \psi_a(x), \bar{\psi}_b(y) \}$$

= $(i\partial + m) [D(x - y) - D(y - x)]$ (3.4.1)

Time ordering of Dirac fields

$$T(\phi_a(x)\bar{\psi}_b(y)) = \Theta(x^0 - y^0)\psi_a(x)\bar{\psi}_b(y) - \Theta(y^0 - x^0)\bar{\psi}_b(y)\psi_a(x)$$
(3.4.2)

Feynman propogator for the Dirac field

$$S_F(x-y) = \langle 0|T\psi(x)\bar{\psi}(y)|0\rangle = \int \frac{d^4p}{(2\pi)^4} \frac{i(\not p+m)}{p^2 - m^2 + i\epsilon} e^{-ip\cdot(x-y)}$$
(3.4.3)

3.5 Discrete symmetries of the Dirac Field

	orientation perserving	orientation not perserving
(ortho)chronous	$\mathcal{L}_{+}^{\uparrow}$	$\mathcal{L}_{-}^{\uparrow}=\mathcal{P}\mathcal{L}_{+}^{\uparrow}$
non-orthochronous	$\mathcal{L}_{-}^{\downarrow}=\mathcal{T}\mathcal{L}_{+}^{\uparrow}$	$\mathcal{L}_{+}^{\downarrow} = \mathcal{PTL}_{+}^{\uparrow}$

4 Interacting QFT

4.1 Introduction and examples

Theories discussed so far are Klein-Gordon theory with spin 0

$$\mathcal{L}_{KG} = \frac{1}{2} \partial_{\mu} \phi \partial^{\mu} \phi - \frac{m^2}{2} \phi^2$$

and Dirac theory spin $\frac{1}{2}$

$$\mathcal{L}_D = \bar{\psi}(i\partial \!\!\!/ - m)\psi$$

There is also $\mathcal{L}_{EM} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu}$ with $F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}$ for a massless vector filed. Its quantiasation gives photon.

One thing they have in common is quadratic in the fields. As result:

- linear field equations
- exact quantisation
- multi-particle states without scattering or interaction
- linear fourier decompositions, no mementum changes

To have an interacting theory with scattering, need higher powers in the field in the Lagrangians. A few examples are following

Scalar ϕ^4 theory

$$\mathcal{L} = \mathcal{L}_{KG} + \frac{\lambda}{4!} \phi^4$$

need positive sign $\lambda > 0$ for a stable theory, otherwise classical energy can be arbitarily negative. Equation of motions

$$(\partial^2 + m^2)\phi = -\frac{\lambda}{3!}\phi^3$$

is nonlinear, cannot be solved by Fourier decomposition.

Yukawa-theory

$$\mathcal{L} = \mathcal{L}_{KG} + \mathcal{L}_D - g\bar{\psi}\psi\phi$$

It is originally developed as a theory for nuclear forces with ψ nucleon, ϕ pion. In the Standard Model it is similar to interactions in Higgs mechanism.

Quantum Electrodynamics (QED)

$$\mathcal{L} = \mathcal{L}_{EM} + \mathcal{L}_D - eA_\mu \bar{\psi} \gamma^\mu \psi$$

descreibes electrons, their antiparticles positrons and photons.

Yang-Mills theory generalises \mathcal{L}_{EM} with terms like A^4 or $A^2 \partial A$

Scalar QED descreibes pions and photons

$$\mathcal{L} = \mathcal{L}_{EM} + D_{\mu}\phi D^{\mu}\phi^* - m^2|\phi|^2$$

$$= \mathcal{L}_{EM} + \partial_{\mu}\phi\partial^{\mu}\phi^* - m^2\phi\phi^* + ieA_{\mu}(\phi\partial^{\mu}\phi^* - \phi^*\partial^{\mu}\phi) + e^2A_{\mu}A^{\mu}\phi\phi^*$$

Remarks

- 1. Interaction terms in $H_{\text{int}} = \int d^3x \mathcal{H}_{\text{int}} = -\int d^3x \mathcal{L}_{\text{int}}$ always involves products of fields at the same point \boldsymbol{x} . It ensures causality, no "instant at a distance".
- 2. There are no derivative interactions. These may complicate quantisation as

$$\pi(\mathbf{x}) = \frac{\partial \mathcal{L}}{\partial (\partial_0 \phi(\mathbf{x}))}$$

3. Why are we taking the examples above? There must be zillions of theories (Lagrangians)? We have the criterion of **renormalizability**. Note the mass dimensions of fields;

$$[S] = 1 \text{ so } [\mathcal{L}] = [M]^4 \Rightarrow [\phi] = [M], [\psi] = [M]^{\frac{3}{2}}, [A_{\mu}] = [M]$$

So in all the interaction terms indicated above, the coupling constant λ , e, g are all **dimensionless!** Can add $-\frac{\mu}{3!}\phi^3$ to the ϕ^4 theory. This leads to $[\mu] = [M]$ and all these generate renormalisable interactions

All higher interaction terms require coupling constants of **negative** mass dimension, e.g. $G\bar{\psi}\psi\bar{\psi}\psi$ and then $[G] = [M]^{-2}$. These are nonrenormalisable and create trouble when performing higher-order calculation in perturbation theory. (with energy cutoff; corrections $G\Lambda^2$, $\Lambda \to \infty$)

4. we haven't quantised the photon yet. The reason is that its is a vector field, i.e. 4 degrees of freedom, but photon has just 2 physical polarisaion states. It is linked to gauge symmetry and complicates quantisation somewhat.

4.2 The interaction picture

Consider the ϕ^4 theory,

$$\mathcal{L}_{int} = -\frac{\lambda}{4!}\phi(x)^4 \tag{4.2.1}$$

Hamiltonian $H = H_0 + H_{int}$ with

$$H_0 = \int d^3x \left\{ \frac{1}{2} \pi^2(x) + \frac{1}{2} (\nabla \phi)^2 + \frac{1}{2} m^2 \phi^2 \right\}$$
 (4.2.2)

$$H_{int} = -\int d^3x \mathcal{L}_{int} = \frac{\lambda}{4!} \int d^3x \phi^4$$
 (4.2.3)

Interaction picture means that operators evolve in time using H_0 (only), in particular

$$\phi_I(t, \mathbf{x}) = e^{iH_0 t} \phi(\mathbf{x}) e^{-iH_0 t}$$
(4.2.4)

Time-dependence of the free field, obeys classical equation of motion $(\partial^2 + m^2)\phi_I(t, \mathbf{x}) = 0$. Solution in terms if fourier modes as before:

$$\phi_I = \int \frac{\mathrm{d}^3 p}{(2\pi)^3 \sqrt{2E_p}} (a_p^I e^{-ipx} + a_p^{I\dagger} e^{+ipx})$$
 (4.2.5)

as in the free theory with standard commutation relations $[a_{\bf p}^I,a_{\bf p}^{I\dagger}]=(2\pi)^3\delta^{(3)}({\bf p}-{\bf p}')$. The state satisfing $a_p^I|0\rangle=0$ is the vacuum of the free, noninteracting theory.

Relation between interaction and Schrödinger picure states:

$$|\phi_I(t)\rangle = e^{iH_0t}|\psi_S(t)\rangle \tag{4.2.6}$$

Schrödinger equation becomes:

$$i\frac{\partial}{\partial t}|\psi_{S}\rangle = (H_{0} + H_{\text{int}})|\psi_{S}\rangle$$

$$LHS = i\frac{\partial}{\partial t}\left(e^{-iH_{0}t}|\phi_{I}\rangle\right) = H_{0}e^{-iH_{0}t}|\phi_{I}\rangle + e^{-iH_{0}t}i\frac{\partial}{\partial t}|\phi_{I}\rangle$$

$$RHS = (H_{0} + H_{\text{int}})e^{-iH_{0}t}|\phi_{I}\rangle$$

$$\Rightarrow i\frac{\partial}{\partial t}|\phi_{I}\rangle = e^{iH_{0}t}H_{int}e^{-iH_{0}t} = H_{I}(t)|\phi_{I}\rangle$$
(4.2.7)

with H_I interaction Hamiltonian in the interaction picture. Clearly

$$H_I = \frac{\lambda}{4!} \int \mathrm{d}^3 x \phi_I^4(x)$$

What is the solution of $\ref{eq:property}$ for the time evolution of $|\phi_I(t)\rangle$? Define time-evolution operator in the interaction picture.

$$|\phi_I(t)\rangle = U(t, t_0) |\phi_I(t_0)\rangle \tag{4.2.8}$$

where
$$U(t, t_0) = e^{iH_0(t-t_0)}e^{-iH(t-t_0)}$$
 (4.2.9)

With ?? and ??:

$$i\frac{\partial}{\partial t}U(t,t_0) = H_I(t)U(t,t_0) \tag{4.2.10}$$

To solve with boundary conditions: $U(t_0, t_0) = 1$. The formal solution:

$$U(t, t_0) = 1 - i \int_{t_0}^t dt' H_I(t') U(t', t_0)$$

Substitute back in and we get:

$$U(t,t_0) = 1 - i \int_{t_0}^t dt' H_I(t') + (-i)^2 \int_{t_0}^t dt' \int_{t_0}^{t'} dt'' H_I(t') H_I(t'') + \dots$$

 H_I inside the integral is automatically time-ordered. Ranges of integration is not.

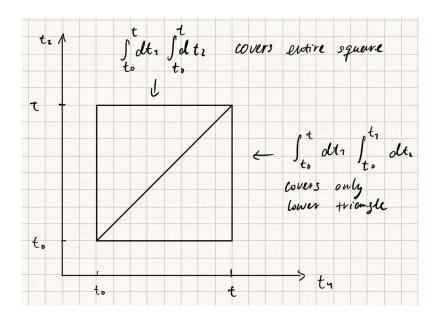


Figure 4.1: Time ordering

Upper triangle has the wrong time order. We are going to "repair" it by hand.

$$U(t,t_0) = 1 - i \int_{t_0}^{t} dt' H_I(t') + \frac{(-i)^2}{2} \int_{t_0}^{t} dt' \int_{t_0}^{t} dt'' T(H_I(t')H_I(t'')) + \dots$$

$$= \sum_{n=0}^{\infty} \frac{(-i)^n}{n!} \int_{t_0}^{t} dt_1 \cdots \int_{t_0}^{t} dt_n T(H_I(t_1) \dots H_I(t_n))$$

$$= T \exp\left\{-i \int_{t_0}^{t} dt' H_I(t')\right\}$$
(4.2.11)

It is interesting for scattering to transition into asymptotic state for $t \to \infty$

$$S = \lim_{t \to \infty} U(t, -t) = T \exp\left\{-i \int_{-\infty}^{\infty} dt H_I(t)\right\}$$

$$\stackrel{\phi^4}{=} T \exp\left\{-i \int d^4 x \frac{\lambda}{4!} \phi_I^4(x)\right\}$$
(4.2.12)

Both U and S are formally unitary

Composition law for time evolution operator:

$$U(t_2, t_0) = U(t_2, t_1)U(t, t_0) = U(t_2, t_1)U(t_0, t_1)^{\dagger}$$
(4.2.13)

4.2.1 Scattering amplitudes and the S-matrix

Take $|i\rangle$ the initial (multi-particle) state and $|f\rangle$ the final (multi-particle) state. Time evolution of $|i\rangle$ then is

$$\lim(t \to \infty)U(t, -\infty)|i\rangle = S|i\rangle$$

Probability that $|i\rangle$ evolves into $|f\rangle$ is proportional to the squared "S-matrix element"

$$|\langle f, t \to \infty | i, t \to -\infty \rangle|^2 = |\langle f | S | i \rangle|^2 = |S_{fi}|^2$$
(4.2.14)

The nontrivial part of the S-matrix is the T-matrix:

$$S_{fi} := \delta_{fi} + iT_{fi} \tag{4.2.15}$$

Use momentum conservation (from translation invariance) to define matrix element

$$S_{fi} = \delta_{fi} + i(2\pi)^4 \delta^{(4)}(p_f - p_i) M_{fi}$$
 (4.2.16)

 M_{fi} measures "genuine scattering" from $|i\rangle$ to $|f\rangle$.

How are we going to calculate correlation functions in the interacting theory:

$$\langle \Omega | T \phi(x) \phi(y) | \Omega \rangle$$
 (4.2.17)

or more generally $\langle \Omega | T \phi(x_1) \phi(x_2) \dots | \Omega \rangle$, where $| \Omega \rangle$ is the vaccum/ground state of the interacting theory and $\phi(x)$ the Heisenberg operators?

Ignore $|\Omega\rangle \neq |0\rangle$ for the moment other than saying: we want to study the time evolution from the vacuum at $t \to -\infty$ to $t \to +\infty$. So rewriting in terms $\phi_I(x)$, assuming $x^0 > y^0$ for now:

$$\langle 0|U(\infty, x^{0})\phi_{I}(x^{0})U(x^{0}, y^{0})\phi_{I}(y^{0})U(y^{0}, -\infty)|0\rangle = \langle 0|T(\phi_{I}(x)\phi_{I}(y)S)|0\rangle \tag{4.2.18}$$

still holds if $x^0 < y^0$ because of T.

Now $|\Omega\rangle \neq |0\rangle$: this can be taken care of by dividing out the time evolution of the (free) vacuum $\langle 0|S|0\rangle$, so

$$\langle \Omega | T(\phi(x)\phi(y)) | \Omega \rangle$$

$$= \frac{\langle 0 | T(\phi_I(x)\phi_I(y)S) | 0 \rangle}{\langle 0 | S | 0 \rangle}$$

$$\stackrel{\phi^4}{=} \frac{\langle 0 | T\phi_I(x)\phi_I(y) \exp\{-i \int d^4 x' \frac{\lambda}{4!} \phi^4(x')\} | 0 \rangle}{\langle 0 | T \exp\{-i \int d^4 x' \frac{\lambda}{4!} \phi^4(x')\} | 0 \rangle}$$

$$(4.2.19)$$

Proof can be found in Peskin. It will also be illustrated parctically later ("vacuum bubbles").

Perturbation theory is viable when λ (or some other coupling) is "small" and then expands $U(t, t_0)$ or S in powers of λ .

4.3 Wick's theorem

From now on drop the subscript for interaction pictire fields $\phi_I(x) \to \phi(x)$.

Want to calculate stuff like $\langle 0|T\phi(x_1)\dots\phi(x_n)S|0\rangle$ in perturbation theory; so e.g. at order λ^n . So

$$\frac{1}{n!} \left(-i\frac{\lambda}{4!} \right)^n \int d^4 y_1 \dots d^4 y_n \langle 0 | T\phi(x_1) \dots \phi(x_n) \phi^4(y_1) \dots \phi^4(y_n) | 0 \rangle$$

$$\tag{4.3.1}$$

is tough!

We know $\langle 0|T\phi(x_1)\phi(x_2)|0\rangle$ is the Feynman propagator!

Recall **normal ordering** with $\phi(x) = \phi^{+}(x) + \phi^{-}(x)$

$$: \phi^{+}\phi^{-} :=: \phi^{-}\phi^{+} := \phi^{-}\phi^{+} \tag{4.3.2}$$

where

$$\phi^{+} = \int \frac{\mathrm{d}^{3} p}{(2\pi)^{3}} \frac{1}{\sqrt{2E_{\boldsymbol{p}}}} a_{\boldsymbol{p}} e^{-ip \cdot x}$$
$$\phi^{-} = \int \frac{\mathrm{d}^{3} p}{(2\pi)^{3}} \frac{1}{\sqrt{2E_{\boldsymbol{p}}}} a_{\boldsymbol{p}}^{\dagger} e^{+ip \cdot x}$$

Wick's therem expresses time-ordered products in terms of normal-ordered ones. Then it is easy to take vacuum expectation values, as $\langle 0|: \phi(x_1) \dots \phi(x_n): |0\rangle = 0$

Take two fields and $x^0 > y^0$:

$$T\phi(x)\phi(y) = \phi(x)\phi(y) = (\phi^{+}(x) + \phi^{-}(x))(\phi^{+}(y) + \phi^{-}(y))$$

$$= \phi^{+}(x)\phi^{+}(y) + \phi^{-}(x)\phi^{-}(y) + \phi^{-}(x)\phi^{+}(y) + \phi^{-}(x)\phi^{+}(y) + [\phi^{+}(x), \phi^{-}(y)]$$

$$=: \phi(x)\phi(y) : +[\phi^{+}(x), \phi^{-}(y)]$$

Particularly for $y^0 > x^0$:

$$T\phi(x)\phi(y) =: \phi(x)\phi(y) : +[\phi^{+}(y), \phi^{-}(x)]$$

Thus altogether:

$$T\phi(x)\phi(y) =: \phi(x)\phi(y) : +D_f(x-y)$$
 (4.3.3)

as
$$\Theta(x^0 - y^0)[\phi^+(x), \phi^-(y)] + \Theta(y^0 - x^0)[\phi^+(y), \phi^-(x)] = D_F(x - y)$$
.

Worth noting that $D_F(x - y)$ is still a c-number, not operator (yet). Thus it can be pulled out of any matrix element or expectation value.

We now define "contraction":

$$\phi(x_1)\phi(x_2) = D_F(x_1 - x_2)$$
(4.3.4)

Thus we can remove the fields from the product leaving only the propagators:

$$T\phi(x)\phi(y) =: \phi(x)\phi(y) : +\phi(x)\phi(y) \tag{4.3.5}$$

General form of Wick's theorem for arbitary number of fields

$$T\phi(x_1)\dots\phi(x_n) =: \phi(x_1)\dots\phi(x_n): +: \text{(sum over all possible contractions)}:$$
 (4.3.6)

Example with four fields:

$$T(\phi_{1}\phi_{2}\phi_{3}\phi_{4}) =: \phi_{1}\phi_{2}\phi_{3}\phi_{4} :$$

$$+ \phi_{1}\phi_{2} : \phi_{3}\phi_{4} : + \phi_{1}\phi_{3} : \phi_{2}\phi_{4} : + \phi_{1}\phi_{4} : \phi_{2}\phi_{3} : + \phi_{2}\phi_{3} : \phi_{1}\phi_{4} : + \phi_{2}\phi_{4} : \phi_{1}\phi_{3} : + \phi_{3}\phi_{4} : \phi_{1}\phi_{2} :$$

$$+ \phi_{1}\phi_{2}\phi_{3}\phi_{4} + \phi_{1}\phi_{3}\phi_{2}\phi_{4} + \phi_{1}\phi_{4}\phi_{2}\phi_{3}$$

Thus

$$\langle 0|T(\phi_1\phi_2\phi_3\phi_4)|0\rangle = D_F(x_1-x_2)D_F(x_3-x_4) + D_F(x_1-x_3)D_F(x_2-x_4) + D_F(x_1-x_4)D_F(x_2-x_3)$$

which can be visually represented as

Proof of the general theorem by *induction* in the number of fields (see exercise). The idea is to suppose it is true for $\phi_2 \dots \phi_m$, $x_1^0 > x_{k>1}^0$. Then

$$T\phi_1\phi_2\dots\phi_m = (\phi_1^+ + \phi_1^-)T\phi_2\dots\phi_m$$
$$= (\phi_1^+ + \phi_1^-)[:\phi_2\dots\phi_m: + : \text{contractions}:]$$

 ϕ_1^- can stay as it is part of $(:\phi_1\phi_2\dots\phi_m:)$. But ϕ_1^+ needs to be comuted past all ϕ_1^- operators, giving rise to additional contractions $\phi_1\phi_2$.

Consequences

• $n = 2k + 1, k \in \mathbb{N}$

$$\langle 0|T\phi_1\dots\phi_m|0\rangle=0$$

• $n = 2k, k \in \mathbb{N}$

$$\langle 0|T\phi_1\dots\phi_m|0\rangle = \sum_{\text{pairing of fields}} D_F(x_{i_1}-x_{i_2})\dots D_F(x_{i_{m-1}}-x_{i_m})$$

4.3.1 Wick's theorem and the S-Matrix

Apply Wick's theorem to correlation functions $\langle 0|T\{\phi_1\dots\phi_m\}S|0\rangle$ n-th term in the perturbative expansion of S with $\phi(x_1) := \phi_1$.

$$\frac{1}{n!} \left(\frac{-i\lambda}{4!} \right)^n \int d^4 y_1 \dots d^4 y_n \langle 0 | T\{\phi_1 \dots \phi_m \phi^4(y_1) \dots \phi^4(y_n)\} | 0 \rangle$$

Example with m = 4, n = 1

$$-\frac{i\lambda}{4!} \int \mathrm{d}^4x \langle 0|T\phi_1\phi_2\phi_3\phi_4\phi^4(x)|0\rangle$$

$$= -\frac{i\lambda}{4!} \int \mathrm{d}^4x \langle 0|\phi_1\phi_2\phi_3\phi_4\phi(x)\phi(x)\phi(x)\phi(x)|0\rangle + 23 \text{ permutations}$$

$$-\frac{i\lambda}{4!} \int \mathrm{d}^4x \langle 0|\phi_1\phi_2\phi_3\phi_4\phi(x)\phi(x)\phi(x)\phi(x)|0\rangle + 11 \text{ permutations} + 5 \text{ similar}$$

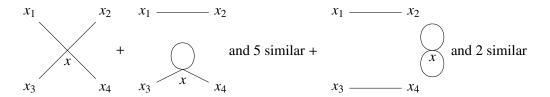
$$-\frac{i\lambda}{4!} \int \mathrm{d}^4x \langle 0|\phi_1\phi_2\phi_3\phi_4\phi(x)\phi(x)\phi(x)\phi(x)|0\rangle + 2 \text{ permutations} + 2 \text{ similar}$$

$$= -i\lambda \int \mathrm{d}^4x D_F(x_1 - x)D_F(x_2 - x)D_F(x_3 - x)D_F(x_4 - x)$$

$$-\frac{i\lambda}{2}D_F(x_1 - x_2) \int \mathrm{d}^4x D_F(x_3 - x)D_F(x_4 - x)D_F(x - x) + 5 \text{ similar}$$

$$-\frac{i\lambda}{8}D_F(x_1 - x_2)D_F(x_3 - x_4) \int \mathrm{d}^4x D_F(x - x) + 2 \text{ similar}$$

Permutation means permutation of $\phi(x)$ and similar means exchanging ϕ_i , $i \in 1, 2, 3, 4$ without changing the diagram. Represented in Feynman diagrams:



In fact $D_F(x - x) = D_F(0)$ diverges!

Example with m = 0, n = 1 vacuum diagram

$$-\frac{i\lambda}{4!} \int d^4x \langle 0|T\phi^4(x)|0\rangle$$
$$= -\frac{i\lambda}{8} [D_F(0)]^2 \int d^4x$$
$$= \underbrace{x}$$

Example: 2nd order S-matrix term

$$\frac{1}{2!} \left(\frac{-i\lambda}{4!} \right)^4 \int d^4x d^4y \langle 0| T\phi_1 \phi_2 \phi_3 \phi_4 \phi^4(x) \phi^4(y) |0\rangle$$

It has many contractions and some of the fully connected ones are of the type there are

 (4×3) [choose $\phi(x)$] \times (4×3) [choose $\phi(y)$] \times 2[x-y-cont.] \times 2(x-y-symm.) + 2 similar, exchanging external points

$$= \frac{(-i\lambda)^2}{2} \int d^4x d^4y D_F(x_1 - x) D_F(x_2 - x) D_F(x_3 - y) D_F(x_4 - y) [D_F(x - y)]^2 + 2 \text{ similar}$$

$$= x_1 \qquad x_2 \qquad x_1 \qquad x_4 \qquad x_5 \qquad x_4 \qquad x_5 \qquad x_4 \qquad x_5 \qquad x_5 \qquad x_5 \qquad x_6 \qquad x_6$$

Symmetry factors A lot of the contractions eliminate the factors $\frac{1}{n!} \left(\frac{1}{4!}\right)^4$ in the denominators; the $\frac{1}{4!}$

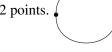
was chosen as to yield $\sim -i\lambda$

See examples above. Sometimes, factors are not completely cancelled and thus procedure gets "reversed". Divide diagrams by $symmetry\ factor\ \stackrel{\wedge}{=}\ the$ "missing factors".

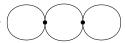
Where does it come from?

• factor 2 from the line that starts and ends at the same point.

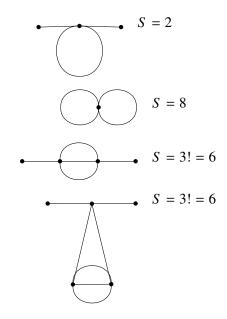
• two (or more) lines linking the same 2 points.



• 2 vertices can be equivalent.



When in doubt, can always go back to Wick's theorem and count the contractions explicitely. Examples:



Summary of Feynman rules

$$\langle 0|T\phi_1\dots\phi_m\exp\left(-\frac{i\lambda}{4!}\int d^4x\phi^4(x)\right)|0\rangle$$

= sum of all diagrams with m external points;

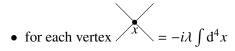
usually organised by number of internal points (i.e. power of λ).

Each diagram built cut of

- propagators
- vertices (n)
- external points (m)

Feynman rules in position space Analytic expression obtained by combining

• for each propagator $\overset{x}{\bullet} = D_F(x-y)$



- for each external point $\stackrel{x}{\bullet}$ = 1
- divide diagram by its symmetry factor S

Since the propagator $D_F(x-y)=\int \frac{\mathrm{d}^4p}{(2\pi)^4}\frac{i}{p^2-m^2+i\epsilon}e^{-ip(x-y)}$. It is actually simpler to express these in momentum space instead.

The way to do it is to sssign a momentum p to each propagator. (direction arbitary)



- assign e^{ipy} to y-vertex (arrow out)
- assign e^{-ipx} to x-vertex (arrow in)
- $\frac{i}{p^2 m^2 + i\epsilon}$ to the line and the integration $\int \frac{d^4p}{(2\pi)^4}$

At vertex *x*:

$$p_{1} \qquad p_{2}$$

$$= -i\lambda \int d^{4}x e^{-i(p_{1}+p_{2}+p_{3})x+ip_{4}x}$$

$$= -i\lambda (2\pi)^{4} \delta^{(4)}(p_{1}+p_{2}+p_{3}-p_{4})$$

This imposes momentum conservation at vetex. $\delta^{(4)}$ -functions make some of the momentum integrals trivial, always with $(2\pi)^4$ cancelled appropriately.

Momentum space Feynman rules

• propagator
$$\xrightarrow{p} = \frac{i}{p^2 - m^2 + i\epsilon}$$

• vertex (position integrated out)
$$= -i\lambda$$

- external points $\begin{cases} e^{-ipx} & \text{incoming} \\ e^{+ipx} & \text{outgoing} \end{cases}$
- impose momentum conservation at each vertex
- integrate over each undetermined momentum $\int \frac{d^4p}{(2\pi)^4}$
- divide by symmetry factor

e.g.:

$$x \stackrel{p}{\longleftarrow} y = (-i\lambda) \int \frac{\mathrm{d}^4 p}{(2\pi)^4} \frac{\mathrm{d}^4 1}{(2\pi)^4} \left(\frac{i}{p^2 - m^2 + i\epsilon}\right)^2 \frac{i}{q^2 - m^2 + i\epsilon} e^{-ip(x-y)}$$

Vacuum diagrams

Disconnected pieces in Feynman diagrams are pretty bad. Not only $D_F(0) = \int \frac{d^4p}{(2\pi)^4} \frac{i}{p^2 - m^2 + i\epsilon}$ is divergent (that will be taken care of later), it also contains an integral $\int d^4x$ const. thus divergent once more!

Typical diagram contributing to 2-point function. one piece connected to x and y, plus disconnected pieces.

Call disconnected pieces
$$V_i \in \left\{ \begin{array}{c} \\ \\ \end{array} \right\}$$
. Points are connected inter-

nally, but not to external points.

 V_i can occur n_i -times, then

[diagram] = [connected pieces]
$$\times \prod_{i} \frac{1}{n!} (V_i)^{n_i}$$

The factorial is the symmetry factor of n_i disconnected copies of V_i .

Then

$$\langle 0|T\phi_1 \dots \phi_n S|0\rangle$$

$$= \sum_{\text{connected}} \sum_{\text{all}\{n_i\}} [\text{connected}] \times \prod_i \frac{1}{n_i!} (V_i)^{n_i}$$

$$= \left(\sum_{\text{[connected]}}\right) \times \sum_{\text{all}\{n_i\}} \left(\prod_i (i) \frac{1}{n_i!} (V_i)^{n_i}\right)$$

$$= \left(\sum_{\text{[connected]}}\right) \times \prod_i \left(\sum_{n_i} \frac{1}{n_i!} (V_i)^{n_i}\right)$$

$$= \left(\sum_{\text{[connected]}}\right) \times \exp\left(\sum_i V_i\right)$$

Thus

$$\times$$
 exp(sum of all DISCONNECTED diagrams) (4.3.8)

Obvious from the above:

$$\langle 0|S|0\rangle = \langle 0|T\{\exp\left(-\frac{i\lambda}{4!}\int d^4x\phi^4(x)\right)\}|0\rangle = \exp(\text{sum of all vacuum bubbles})$$

Conclusion from the (unproven) formula for n-point correlation functions in the true, interacting vacuum:

$$\langle \Omega | T \phi_1 \dots \phi_m | \Omega \rangle = \frac{\langle 0 | T \phi_1 \dots \phi_m S | 0 \rangle}{\langle 0 | S | 0 \rangle}$$

$$= \sum \text{(connected diagrams with m external points)}$$
(4.3.9)

Here: "connected" means connected to any external point. External points do not have to linked to each other.

4.4 S-matrix elements and Feynman diagrams

What is the correlation function in interacting vacuum $\langle \Omega | T\phi_1 \dots \phi_m | \Omega \rangle$ good for? For scattering, shouldn't we rather look at $\langle p_1 \dots p_m | S | p_A p_B \rangle$ with the perturbative expansion of S as before?

Decompose the S-matrix

$$S_{fi} = \delta_{fi} + iT_{fi} = \delta_{fi} + i(2\pi)^4 \delta^{(4)}(p_f - p_i) M_{fi}$$
(4.4.1)

 M_{fi} is the invariant matrix element, used to calculate cross section etc..

Zeroth term in the expansion of S

$$\langle p_1 p_2 | p_A p_B \rangle = \sqrt{2E_1 2E_2 2E_A 2E_B} \langle 0 | a_1 a_2 a_A^+ a_B^+ | 0 \rangle$$

$$= 2E_A 2E_B (2\pi)^6 \left\{ \delta^{(3)} (\boldsymbol{p}_A - \boldsymbol{p}_1) \delta^{(3)} (\boldsymbol{p}_B - \boldsymbol{p}_2) + \delta^{(3)} (\boldsymbol{p}_A - \boldsymbol{p}_2) \delta^{(3)} (\boldsymbol{p}_B - \boldsymbol{p}_1) \right\}$$

This actually is "no scattering", part of the 1 in the S-matrix.

First term is

$$\langle p_1 p_2 | T \left(-\frac{i\lambda}{4!} \int d^4 x \phi^4(x) \right) | p_A p_B \rangle$$

$$\stackrel{\text{wick}}{=} \langle p_1 p_2 | : \left(-\frac{i\lambda}{4!} \int d^4 x \phi^4(x) + \text{contractions} \right) : | p_A p_B \rangle$$

However now the expectation value of a normal-ordered expression doesn't vanish!

$$\phi^{+}(x)|\mathbf{p}\rangle = \int \frac{\mathrm{d}^{3}k}{(2\pi)^{3}\sqrt{2E_{k}}} a_{\mathbf{k}}e^{-ikx}\sqrt{2E_{p}}a_{\mathbf{p}}^{+}|0\rangle$$

$$= \int \frac{\mathrm{d}^{3}k}{(2\pi)^{3}\sqrt{2E_{k}}} e^{-ikx}\sqrt{2E_{p}}\delta^{(3)}(\mathbf{k} - \mathbf{p})|0\rangle$$

$$= e^{-ipx}|0\rangle$$

So in general, need two field operators to annihilate the in-state and m fields operators to create the outstates

New type of Feynman diagram to deal with external states. Define contractions of field operators with external states according to

$$\begin{aligned}
\phi(x) | \mathbf{p} \rangle &= e^{-ipx} | 0 \rangle \\
\langle \mathbf{p} | \phi(x) &= e^{+ipx} | 0 \rangle
\end{aligned}$$

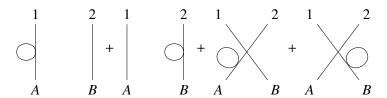
How does this work for $p_A p_B \to p_1 p_2$ in ϕ^4 at $O(\lambda)$? The above contains 3 types of terms: $\phi \phi \phi \phi$, $\phi \phi \phi \phi \phi$ and $\phi \phi \phi \phi \phi$.

1. $\phi\phi\phi\phi$ allows full contractions with all external states. There is 4! possibilities

$$A \qquad B \qquad = 4! \frac{-i\lambda}{4!} \int d^4x e^{-i(p_A + p_B - p_1 - p_2)x} = -i\lambda \underbrace{(2\pi)^4 \delta^{(4)}(p_A + p_B - p_1 - p_2)}_{\text{Prefactor in definition of } i\mathcal{M}}$$

 $i\mathcal{M}$ receives a contributuon $-i\lambda!$

2. $\phi\phi\phi\phi$ leaves 2 operators to connect to external particles. Momentum conservation at each vertex. Still trivial!



Only fully connected Feynman diagrams contribute to iT/iM!

3.

4.4.1 Feynman rules (with external lines)

Position space calculate *iT* by summing overall fully connected diagrams with

• propagator
$$\overset{x}{\bullet} = D_F(x-y)$$

• vertex
$$= -i\lambda \int d^4x$$

• external lines "in"
$$\xrightarrow{x} \stackrel{p}{\longleftarrow} = e^{-ip \cdot x}; \qquad \xrightarrow{x} \stackrel{p}{\longrightarrow} = e^{ip \cdot x}$$

• divide diagram by its symmetry factor $\frac{1}{S}$

Momentum space We have seen it before. Now (with external lines) all positions are integrated over. Result is a function of external momenta only. Integrating out all momentum-conserving δ -distribution yields <u>overall</u> momentum conservation: $(2\pi)^4 \delta^{(4)}(P_f - P_i)$

Momentum space Feynman rules for calculating *iM*:

• internal propagator
$$\overset{x}{\bullet} = \frac{i}{p^2 - M^2 + i\epsilon}$$

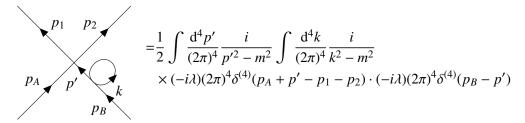
• vertex
$$x = -i\lambda$$

• external lines ("in" or "out")
$$x \leftarrow p$$
 = 1

• impose 4-momentum conservation at each vertex

- integrate over all <u>undetermined</u> momenta $\int \frac{d^4p}{(2\pi)^4}$
- divide diagram by its symmetry factor $\frac{1}{S}$

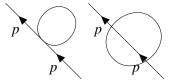
There is still trouble in there. Consider the next-to-leading contribution to the scattering amplitude



This contains the internal propagator $\frac{i}{P_R^2 - m^2 + i\epsilon}$, but all the external particle are on their mass-shell, i.e.

$$P_A^2 = P_B^2 = P_1^2 = P_2^2 = m^2 \implies \frac{i}{P_B^2 - m^2} = \frac{i}{0}$$

In Addition to having <u>fully connected</u> diagrams, also need to confine ourselves to <u>amputated</u> diagrams: disregard all these diagrams with loops attached to external legs.



These diagrams represent the transition from the <u>free</u> to the interacting asymptotic states.

Lehmann-Symanzik-Zimmermann (LSZ) reduction formula

Proof on relation between correlation functions and S-matrix elements will be provided later.

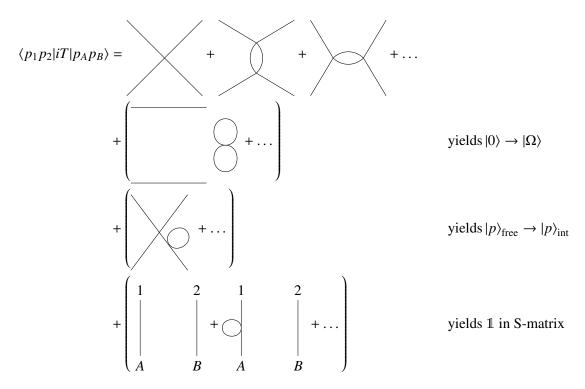
$$\prod_{i=1}^{n} \int d^{4}x_{i}e^{ip_{i}\cdot x_{i}} \prod_{j=1}^{m} \int d^{4}y_{j}e^{-ik_{j}\cdot y_{i}} \langle \Omega|T\phi(x_{1})\dots\phi(x_{n})\phi(y_{1})\dots\phi(y_{m})|\Omega\rangle$$

$$\stackrel{LSZ}{=} (disconnected stuff) + \underbrace{\prod_{i=1}^{n} \frac{\sqrt{z}i}{p_{i}^{2} - m^{2} + i\epsilon} \prod_{j=1}^{m} \frac{\sqrt{z}i}{k_{j}^{2} - m^{2} + i\epsilon} \langle p_{1}\dots p_{n}|S|k_{i}\dots k_{m}\rangle$$
remove poles from external legs
$$(4.4.2)$$

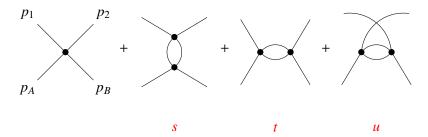
z is the wave-function renormalization factor.

Then amend feynman rules above

consider only fully connected, amputated diagrams



All allowed scattering diagrams $2 \to 2$ in ϕ^4 up to $O(\lambda^2)$:



Define the Lorentz-invariant quantities, Mandelstam variables:

$$s = (p_A + p_B)^2$$
, $t = (p_A - p_1)^2$, $u = (p_A - p_2)^2$ (4.4.3)

$$p_{A} + p_{B} + k$$

$$p_{A} = \frac{1}{2}(-i\lambda)^{2} \int \frac{d^{4}k}{(2\pi)^{4}} \frac{i}{k^{2} - m^{2} + i\epsilon} \frac{i}{(p_{A} + p_{B} + k)^{2} - m^{2} + i\epsilon} =: \frac{1}{2}(-i\lambda)^{2}iJ(s)$$

Then the complete invariant amplitude is

$$M = -\lambda - \frac{\lambda^2}{2} (J(s) + J(t) + J(u))$$
 (4.4.4)

4.5 Scattering cross section

This section is based on Itzykson & Zuber, Chapter 5.1.

The aim is to relate (differential) cross section to reduced/invariant matrix element M_{fi} . First we describe the initial states not as momentum eigenstates $|p_A p_B\rangle$, but as wave packets.

$$|i\rangle = \int \frac{\mathrm{d}^3 k_A}{(2\pi)^3 2 k_A^0} \frac{\mathrm{d}^3 k_B}{(2\pi)^3 2 k_B^0} f(k_A) g(k_B) \, |k_A k_B\rangle$$

with $f(k_A)$, $g(k_B)$ strongly peaked at $k_A \approx p_A$, $k_B \approx p_B$.

We can write the transition amplitude to the final state $|f\rangle \propto |p_1p_2\rangle$ (note: normalisation not the same)

$$A_{fi} = \int \frac{d^3k_A}{(2\pi)^3 2k_A^0} \frac{d^3k_B}{(2\pi)^3 2k_B^0} f(k_A) g(k_B) \langle f|iT|k_A k_B \rangle$$

$$= \int \frac{d^3k_A}{(2\pi)^3 2k_A^0} \frac{d^3k_B}{(2\pi)^3 2k_B^0} f(k_A) g(k_B) (2\pi)^4 \delta^{(4)} (\underbrace{p_f}_{=p_1+p_2} -k_A - k_B) iM(f, k_A, k_B)$$

Thus the transition probablity:

$$\omega_{fi} = (2\pi)^{8} \int \frac{\mathrm{d}^{3}k_{A}}{(2\pi)^{3}2k_{A}^{0}} \frac{\mathrm{d}^{3}k_{B}}{(2\pi)^{3}2k_{B}^{0}} \frac{\mathrm{d}^{3}q_{A}}{(2\pi)^{3}2q_{A}^{0}} \frac{\mathrm{d}^{3}q_{B}}{(2\pi)^{3}2q_{B}^{0}} f(k_{A})g(k_{B})f(q_{A})^{*}g(q_{B})^{*}$$

$$\times \underbrace{\delta^{(4)}(p_{f} - k_{A} - k_{B})\delta^{(4)}(p_{f} - q_{A} - q_{B})}_{=\delta^{(4)}(q_{A} + q_{B} - k_{A} - k_{B})\delta^{(4)}(p_{f} - p_{A} - p_{B})} \underbrace{M(f, k_{A}, k_{B})M^{*}(f, q_{A}, q_{B})}_{\approx |M(f, p_{A}, p_{B})|^{2}}$$

Using the fourier representation of delta function $\delta^{(4)}(q_A + q_B - k_A - k_B) = (2\pi)^{-4} \int d^4x e^{i(k_A + k_B - q_A - q_B) \cdot x}$

$$= \int d^{4}x \underbrace{\int \frac{d^{3}k_{A}}{(2\pi)^{3}2k_{A}^{0}} \frac{d^{3}q_{A}}{(2\pi)^{3}2q_{A}^{0}} e^{i(k_{A}-q_{A})\cdot x} f(k_{A}) f^{*}(q_{A})}_{:=|\tilde{f}(x)|^{2}} \times \underbrace{\int \frac{d^{3}k_{B}}{(2\pi)^{3}2k_{B}^{0}} \frac{d^{3}q_{B}}{(2\pi)^{3}2q_{B}^{0}} e^{i(k_{B}-q_{B})\cdot x} g(k_{B}) g^{*}(q_{B}) (2\pi)^{4} \delta^{(4)}(p_{f}-p_{A}-p_{B}) \cdot |M(f,p_{A},p_{B})|^{2}}_{:=|\tilde{g}(x)|^{2}}$$

Using Fourier transformation $\tilde{g}(x) := \int \frac{d^3q}{(2\pi)^3 2q^0} e^{iq \cdot x} g(q)$

$$= \int d^4x |\tilde{f}(x)|^2 |\tilde{g}(x)|^2 (2\pi)^4 \delta^{(4)}(p_f - p_A - p_B) \cdot |M(f, p_A, p_B)|^2$$

note that $M(f, p_A, p_B)$ and $M(p_1, p_2, p_A, p_B)$ have different normalisation.

We now consider transition probabilty per unit volume per unit time:

$$\frac{\mathrm{d}\omega_{fi}}{\mathrm{d}V\mathrm{d}t} = (\mathrm{incident\ flux}) \cdot (\mathrm{target\ density}) \cdot \mathrm{d}\sigma$$

with $d\sigma$ the infinitismal cross section for scattering into final state $\langle f|$.

Product (incident flux) · (target density) denotes overlap of wave function. Necessary condition!

Covariant renormalization of states $\langle \boldsymbol{p} | \boldsymbol{q} \rangle \sim 2 p^0 \delta^3(\boldsymbol{p} - \boldsymbol{q})$ means the number of particles per unit volume is $2 p_A^0 |\tilde{f}(x)|^2$ and $2 p_B^0 |\tilde{g}(x)|^2$, respectively.

Assume

in target rest frame. Then $2p_B^0 = 2m_B$ and target density $= 2m_B|\tilde{g}(x)|^2$ Incident flux $= |\mathbf{v}_A| \cdot 2p_A^0 |\tilde{f}(x)|^2 = 2|\mathbf{p}_A||\tilde{f}(x)|^2$ since $|\mathbf{v}_A| = |\mathbf{p}_A|/p_A^0$. Then

$$d\sigma = (2\pi)^4 \delta^{(4)} (p_f - p_A - p_B) \frac{1}{4m_B |\mathbf{p_A}|} |M(f, p_A, p_B)|^2$$

for $A + B \rightarrow 1 + 2$ processes

$$= \int_{\Delta} \frac{\mathrm{d}^3 p_1}{(2\pi)^3 2p_1^0} \frac{\mathrm{d}^3 p_2}{(2\pi)^3 2p_2^0} (2\pi)^4 \delta^{(4)}(p_1 + p_2 - p_A - p_B) \frac{1}{4m_B |\boldsymbol{p_A}|} |M(p_1, p_2, p_A, p_B)|^2$$

with Δ energy-momentum resolution of 4-momentum of final state $|f\rangle$.

Covariant form of

$$m_B \cdot |\mathbf{p}_A| = m_B \sqrt{(p_A^0)^2 - m_A^2} = \sqrt{(p_A \cdot p_B)^2 - m_A^2 m_B^2} =: F$$
 (4.5.1)

This is scattering into arbitary final state subject to 4-momentum conservation: $p_A + p_B = p_1 + p_2$. Consider now <u>differential</u> cross section for scattering into a particular infinitismal solid angle $d\Omega$, hence specific momentum dp_1 , dp_2 variations:

$$d\sigma = \frac{1}{4F} \prod_{f} \frac{d^{3}p_{f}}{(2\pi)^{3}2p_{f}^{0}} (2\pi)^{4} \delta^{(4)}(p_{A} + p_{B} - \sum_{f} p_{f})|M|^{2}$$

$$f = 1, 2 \frac{1}{4F} \frac{d^{3}p_{1}}{(2\pi)^{3}2p_{1}^{0}} \frac{d^{3}p_{2}}{(2\pi)^{3}2p_{2}^{0}} (2\pi)^{4} \delta^{(4)}(p_{i} - p_{f})|M|^{2}$$

$$= \frac{1}{64\pi^{2}F} \frac{d^{3}p_{1}}{E_{1}} \frac{d^{3}p_{2}}{E_{2}} \delta^{(4)}(p_{1} + p_{2} - p_{i})|M|^{2}$$

$$\begin{bmatrix} \int \frac{d^{3}p_{1}}{E_{1}} \frac{d^{3}p_{2}}{E_{2}} \delta^{(4)}(p_{1} + p_{2} - p_{i}) \\ CMS \int d|\mathbf{p}_{1}|d\Omega_{1} \frac{|\mathbf{p}_{1}|^{2}}{E_{1}E_{2}} \delta(E_{1} + E_{2} - E_{i}) \\ = \int d(E_{1} + E_{2}) \frac{d|\mathbf{p}_{1}|}{d(E_{1} + E_{2})} d\Omega_{1} \frac{|\mathbf{p}_{1}|^{2}}{E_{1}E_{2}} \delta(E_{1} + E_{2} - E_{i}) \\ = \frac{|\mathbf{p}_{1}|^{2}}{E_{1}E_{2}} \left(\frac{|\mathbf{p}_{1}|}{E_{1}} + \frac{|\mathbf{p}_{1}|}{E_{2}}\right)^{-1} d\Omega_{1} \\ = \frac{|\mathbf{p}_{1}|d\Omega_{1}}{E_{1} + E_{2}} = \frac{|\mathbf{p}_{1}|d\Omega_{1}}{E_{i}} \\ \frac{d\sigma}{d\Omega} = \frac{1}{64\pi^{2}} \frac{|\mathbf{p}_{1}|}{F \cdot E_{i}} |M|^{2}$$

$$(4.5.2)$$

Rewrite all kinematical factors in terms of $s = (p_A + p_B)^2 = (p_1 + p_2)^2$. Define the function

$$\lambda(x, y, z) := x^2 + y^2 + z^2 - 2(xy + xz + yz) \tag{4.5.3}$$

then

$$F = \sqrt{(p_A \cdot p_B)^2 - m_A^2 m_B^2} = \frac{1}{2} \lambda^{\frac{1}{2}} (s, m_A^2, m_B^2) = \sqrt{s} |\boldsymbol{p}_i|$$

$$\lambda(s, m_A^2, m_B^2) = s^2 - 2s(m_A^2 + m_B^2) - (m_A^2 - m_B^2)^2 = (s - (m_A + m_B)^2)(s - (m_A - m_B)^2)$$

$$= (2p_A \cdot p_B - 2m_A \cdot m_B) \cdot (2p_A \cdot p_B + 2m_A \cdot m_B) = 4 \left[(p_A p_B)^2 - m_A^2 m_B^2 \right]$$

$$p_A = (c \sqrt{s}, \boldsymbol{p}_i), c \in [0, 1] \rightarrow m_A^2 = c^2 s - |\boldsymbol{p}_i|^2$$

$$p_B = ((1 - c) \sqrt{s}, -\boldsymbol{p}_i) \rightarrow m_B^2 = (1 - c)^2 s - |\boldsymbol{p}_i|^2$$

$$= 4 \left[\left((c(1 - c)s + p_i^2)^2 + (c^2 s - p_i^2)((1 - c)^2 s - p_i^2) \right] = 4s|\boldsymbol{p}_i|^2$$

$$|\mathbf{p}_f| = \sqrt{E_{1,2}^2 - m_{1,2}^2} = \frac{1}{2\sqrt{s}} \lambda^{\frac{1}{2}}(s, m_1^2, m_2^2)$$

 $E_i = \sqrt{s}$

$$\frac{d\sigma}{d\Omega_{CMS}} = \frac{1}{64\pi^2 s} \frac{|\mathbf{p}_f|}{|\mathbf{p}_i|} |M|^2 = \frac{1}{64\pi^2 s} \sqrt{\frac{\lambda(s, m_1^2, m_2^2)}{\lambda(s, m_2^2, m_2^2)}} |M|^2$$
(4.5.4)

Decay rate instead of cross section means no "incident flux" to divide by, only "target density"

$$d\Gamma = \frac{1}{2m_A} \prod_f \frac{d^3 p_f}{(2\pi)^3 2p_f^0} (2\pi)^4 \delta^{(4)}(p_A - \sum_f p_f) |M|^2$$
 (4.5.5)

Particles with spin (unpolarized): sum over outgoing or average over initial spins

$$|M|^2 \to \frac{1}{(2s_A + 1)(2s_B + 1)} \sum_{s_i, s_f} |M_{fi}|^2$$
 (4.5.6)

Symmetry factor $|M|^2 \to \frac{1}{s}|M|^2$ with $s = \prod_i k_i!$ if there are k_i identical particles of species i in the final states.

If 1 and 2 are identical, then facotr $\frac{1}{s} = \frac{1}{2}$ on the right hand side.

4.6 Feynman rules for fermions

Consider the simplest interacting theory with fermions, Yukawa-theory. We will treat QED later.

$$\mathcal{L} = \frac{1}{2} \partial_{\mu} \phi \partial^{\mu} \phi - \frac{M^2}{2} \phi^2 + \bar{\psi} (i \partial \!\!\!/ - m) \psi - g \bar{\psi} \psi \phi \tag{4.6.1}$$

Feynman rules will involve:

• scalar
$$\overset{x}{\bullet}$$
 = $D_F(x-y) = \int \frac{\mathrm{d}^4}{(2\pi)^2} \frac{i}{p^2 - M^2 + i\epsilon} e^{-ip(x-y)}$

• fermions
$$x, \alpha$$
 y, β $= S_F(x-y)_{\alpha\beta} = \int \frac{\mathrm{d}^4}{(2\pi)^4} \frac{i(p+m)}{p^2-m^2+i\epsilon} e^{-ip(x-y)}$

• vertices
$$---- = -ig \int d^4x$$

What previous steps need reconsideration due to the <u>anticommutating</u> fermion operators? Interaction Hamiltonina $\sim \bar{\psi}\psi\phi$ and in general compose of <u>even</u> number of fermion fields (spin conservation and fermion number conservation). Thus there is no problem with time-ordered exponential in definition of S-matrix. (Time ordering always takes two or even number of fields.)

Remember the relation

$$T(\psi_{\alpha}(x)\bar{\psi}_{\beta}(x)) = -\bar{\psi}_{\beta}(x)\psi_{\alpha}(x) \text{ when } y^{0} > x^{0}$$

$$(4.6.2)$$

Similarly in normal product:

$$: \psi^+ \psi^- = -\psi^- \psi^+ : \tag{4.6.3}$$

Then Wick's theorem is formally the same as before

$$T(\psi_{\alpha}(x)\bar{\psi}_{\beta}(x)) =: \psi_{\alpha}(x)\bar{\psi}_{\beta}(x) : + \psi_{\alpha}(x)\bar{\psi}_{\beta}(x)$$

note by definition $\psi \psi = \bar{\psi} \bar{\psi} = 0$

Thus contractions inside normal-ordered products would be

$$: \psi_1 \psi_2 \bar{\psi}_3 \bar{\psi}_4 := -\psi_1 \bar{\psi}_3 : \psi_2 \bar{\psi}_4 := -S_F(x_1 - x_3) : \psi_2 \bar{\psi}_4 :$$

because of the additional operator exchange.

We will want to consider fermion-(anti-)fermion scattering. Leading contribution at $O(g^2)$:

$$\frac{1}{2!}(-ig)^2 \int d^4x d^4y \langle p', k'| T\bar{\phi}(x)\phi(x)\phi(x)\bar{\phi}(y)\phi(y)|p,k\rangle$$

Contractions with initial-/final-state fermions?

$$\phi^{+}(x)|p,s\rangle = \int \frac{\mathrm{d}^{3}k}{(2\pi)^{3}\sqrt{2E_{k}}} \sum_{r} a_{k}^{r} u_{r}(k) e^{-ik\cdot x} \sqrt{2E_{p}} a_{p}^{s\dagger} |0\rangle$$
$$= e^{-ip\cdot x} u_{s}(p) |0\rangle$$

So define

$$\overline{\psi(x)|p,s}\rangle = e^{-ip\cdot x}u_s(p)$$

$$\langle p,s|\psi(x) = e^{ip\cdot x}\overline{u}_s(p)$$
(4.6.4)

note, though, for antifermion states $|p', s'\rangle$:

$$\overline{\psi}(x)|p,s\rangle = e^{-ip'\cdot x}\overline{v}_{s'}(p')$$

$$\langle p',s'|\psi(x) = e^{ip'\cdot x}v_{s'}(p')$$
(4.6.5)

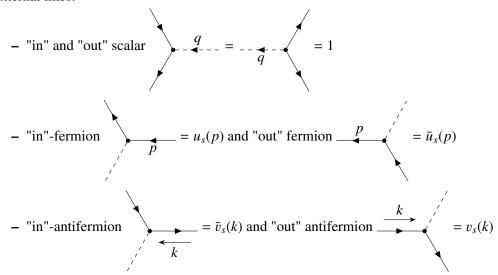
In short ψ | \rangle contractes with a fermion, $\langle \neg \psi \rangle$ with an antifermion; vice verse for $\bar{\psi}$.

Momentum space feynman rule for iM

• internal propagators --- $\stackrel{q}{\longrightarrow} --- = \frac{i}{q^2 - M^2 + i\epsilon};$ $\stackrel{\beta}{\longrightarrow} \frac{q}{\longrightarrow} \alpha = \frac{i}{q^2 - M^2 + i\epsilon} = \frac{i(p + m)_{\alpha\beta}}{p^2 - m^2 + i\epsilon}$

• vertex
$$\beta = -ig \int d^4x = ig \delta_{\beta \alpha}$$

• external lines:



- impose energy-momentum conservation at each vertex
- integrate over undetermined (loop) momenta
- include an overall sign for the diagram

Note

- <u>Arrowas</u> on the fermion lines by convention denote <u>fermion</u> (or charge) <u>flow</u>. They must flow consistently through the diagram. (≡ fermion number conservation) (Only potential confusion: external antifermion lines)
- No symmetry factors (except vacuum bubbles $\frac{1}{s} = \frac{1}{2}$). $\bar{\psi}\psi\phi$ allows for unambiguous contractions.
- Dirac indices are summed over at each vertex

$$\mathcal{L}_{\text{int}} \approx \bar{\psi}_{\alpha}(x)\psi_{\alpha}(x)\phi(x)$$

(p + m) terms in propagator are matrix-multiplied contracted with external spinors, e.g.

$$\frac{p_3}{p_2} \qquad p_1 \qquad p_0 \qquad \sim \bar{u}_{\alpha}(p_3) \frac{i(\not p+m)_{\alpha\beta}}{p_2^2 - m^2 + i\epsilon} \frac{i(\not p+m)_{\beta\gamma}}{p_1^2 - m^2 + i\epsilon} u_{\gamma}(p_0)$$

• closed fermion loop

It always (also with more propagators/couplings) involves an overal (-1) and a trace Tr(...).

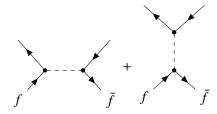
Examples

• fermion-fermion scattering to lowest order $O(g^2)$

$$iM = p - \frac{k'}{k} + \frac{p'}{k}$$

$$= (-ig)^2 \left\{ \bar{u}(p')u(p) \frac{i}{\underbrace{(p'-p)^2 - M^2 + i\epsilon}} \bar{u}(k')u(k) - \bar{u}(p')u(k) \frac{i}{\underbrace{(p'-k)^2 - M^2 + i\epsilon}} \bar{u}(k')u(p) \right\}$$

• fermion-antifermion scattering



These are tree diagrams. Thus there is no undetermined momenta to integrate.

5 Quantum Electrodynamics (QED)

5.1 Classical Electrodynamics and Maxwell's equations

We have the gauge potential $A^{\mu}=(A^0, \mathbf{A})=(\phi, \mathbf{A})$ & $A_{\mu}=(A^0, -\mathbf{A})=(\phi, -\mathbf{A})$ and the field strength tensor $F_{\mu\nu}=\partial_{\mu}A_{\nu}-\partial_{\nu}A_{\mu}$.

Then

- electric field $E_i = F_{0i} = \partial_0 A_i \partial_i A_0 \rightarrow \mathbf{E} = -\dot{\mathbf{A}} \nabla \phi$
- magnetic field $B^i = -\frac{1}{2} \epsilon^{ijk} F_{jk} \to \boldsymbol{B} = \nabla \times \boldsymbol{A}$

Lagrangian density $\mathcal{L}_{EM} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} = -\frac{1}{2}(\boldsymbol{E}\cdot\boldsymbol{E}-\boldsymbol{B}\cdot\boldsymbol{B})$. The field equation $\partial_{\mu}\left(\frac{\partial\mathcal{L}}{\partial(\partial_{\mu}A_{\nu})}\right) - \frac{\partial\mathcal{L}}{\partial A_{\nu}} = 0$ leads to

$$\partial_{u}F^{\mu\nu} = 0 \tag{5.1.1}$$

it is half of Maxwell's equations (in vacuum).

The other half are Bianchi identities following from the definition of $F_{\mu\nu}$:

$$\begin{split} \partial_{\lambda}F_{\mu\nu} + \partial_{\mu}F_{\nu\lambda} + \partial_{\nu}F_{\lambda\mu} &= 0 \Leftrightarrow \epsilon^{\sigma\lambda\mu\nu}\partial_{\lambda}F_{\mu\nu} = 0 \\ \text{or } \partial_{\lambda}\tilde{F}^{\sigma\lambda} &= 0, \ \tilde{F}^{\sigma\lambda} &= \frac{1}{2}\epsilon^{\sigma\lambda\mu\nu}F_{\mu\nu} \end{split}$$

In terms of E and B:

$$\nabla \cdot \mathbf{E} = 0$$
, $\dot{\mathbf{E}} = \nabla \times \mathbf{B}$ dynamical equations $\nabla \cdot \mathbf{B} = 0$, $\dot{\mathbf{B}} = -\nabla \times \mathbf{E}$ Bianchi identities

Remarks

• Lagrangian density does not depend on \dot{A}_0 , since A_0 is not really dynamical.

$$\nabla \cdot \boldsymbol{E} = 0 \rightarrow \nabla^2 A_0 + \nabla \cdot \dot{\boldsymbol{A}} = 0$$

Solve this <u>Poisson</u> equation for $A_0(\mathbf{x}, t) = \frac{1}{4\pi} \int \mathrm{d}^3 y \frac{\nabla \cdot \dot{\mathbf{A}}(\mathbf{y}, t)}{|\mathbf{y} - \mathbf{x}|}$. Thus A_0 is given in terms of the other components of A.

gauge invariance: field strength tensor invariant under the transformation A_μ → A_μ − ∂_μX due to commuting derivatives. This leads to gauge invariance of Maxwell equations.
 Choose X to satisfy ∂_μ∂^μX = ∂²X = ∂_μA^μ allows us to demand the condition (Lorenz condition)

$$\partial_{\mu}A^{\mu} = 0 \tag{5.1.2}$$

such that A_{μ} belongs to the "Lorenz gauge" and reduces the degrees of freedom from 4 to 3.

- Further freedom is eliminated by adding any X with $\partial^2 X = 0$, e.g. $\frac{\partial}{\partial t} X = A_0$. Then we get the Coulomb or radiation gauge

$$A_0 = 0, \ \nabla \cdot \mathbf{A} = 0 \tag{5.1.3}$$

Note: vice versa imposing $\nabla \cdot \mathbf{A} = 0$ first, yields $A_0 = 0$ (using Lorenz condition?). In Coulomb gauge:

$$E = -\dot{\mathbf{A}}. \ B = \nabla \times \mathbf{A}, \ \nabla \times \mathbf{A} = 0$$
$$- \ddot{\mathbf{A}} = \dot{\mathbf{E}} \stackrel{\text{Maxwell}}{=} \nabla \times \mathbf{B} = \nabla \times (\nabla \times \mathbf{A}) = \nabla (\underbrace{\nabla \cdot \mathbf{A}}_{=0}) - \nabla^2 \mathbf{A}$$
$$\Rightarrow \partial^2 \mathbf{A} = 0$$

This wave equation is massless KG equation for each spatial component.

Then the solutions are obvious: $\mathbf{A} = \boldsymbol{\epsilon} e^{-ik \cdot x}$ with $k^2 = 0$ and $\boldsymbol{\epsilon} \cdot \boldsymbol{k} = 0$. The polarization vector $\boldsymbol{\epsilon}$ is transverse to \boldsymbol{k} .

Can write the lagrangian in Coulmb gauge

$$\mathcal{L}_{\rm EM} = \frac{1}{2}\dot{\boldsymbol{A}}\dot{\boldsymbol{A}} - \frac{1}{2}\boldsymbol{B}\cdot\boldsymbol{B}$$

Then the conjugate momentum to \mathbf{A} is $\mathbf{\Pi} = \frac{\partial \mathcal{L}}{\partial \dot{A}} = \dot{\mathbf{A}} = -\mathbf{E}$. It has only 3 components, there is no conjugate momentum to A_0 !. Because of Coulomb gauge $\mathbf{\Pi}$ is subject to the constraint $\nabla \cdot \mathbf{\Pi} = 0$

Hamiltonian

$$H_{\rm EM} = \int d^3x \left(\frac{1}{2} \mathbf{\Pi} \cdot \mathbf{\Pi} + \frac{1}{2} \mathbf{B} \cdot \mathbf{B} \right)$$

5.2 Quantizing the Maxwell field

We would like to impose canonical commutation relations, à la

$$[A_i(\mathbf{x}), A_j(\mathbf{y})] = [\Pi_i(\mathbf{x}), \Pi_j(\mathbf{y})] = 0$$
$$[A_i(\mathbf{x}), \Pi_j(\mathbf{y})] = i\delta_{ij}\delta^{(3)}(\mathbf{x} - \mathbf{y})$$

However this cannot be true. Take either derivative of the last equation and it needs to vanish due to $\nabla \cdot \mathbf{A} = \nabla \cdot \mathbf{\Pi} = 0$. But

$$[\partial^i A_i(\boldsymbol{x}), \Pi_k(\boldsymbol{y})] = i\delta_{ij}\partial^i \delta^{(3)}(\boldsymbol{x} - \boldsymbol{y})$$

here the derivative is takev with respect to \mathbf{x} , i.e. $\partial^i = \frac{\partial}{\partial x_i}$. Replace δ_{ij} by Δ_{ij}

$$[\partial^{i} A_{i}(\mathbf{x}), \Pi_{j}(\mathbf{y})] = i\Delta_{ij}\partial^{i} \frac{1}{(2\pi)^{3}} \int d^{3}k e^{i\mathbf{k}\cdot(\mathbf{x}-\mathbf{y})}$$
$$= -\frac{1}{(2\pi)^{3}} \int d^{3}k (k^{i}\Delta_{ij}) e^{i\mathbf{k}\cdot(\mathbf{x}-\mathbf{y})} \stackrel{!}{=} 0$$

it works for $\Delta_{ij} = \delta_{ij} - \frac{k_i k_j}{k^2}$ in momentum space or $\Delta_{ij} = \delta_{ij} - \nabla^{-2} \partial_i \partial_j$ in position space.

$$[A_i(\mathbf{x}), \Pi_i(\mathbf{y})] = i \left(\delta_{ij} - \nabla^{-2} \partial_i \partial_j \right) \delta^{(3)}(\mathbf{x} - \mathbf{y})$$
(5.2.1)

As before we have the mode expansion

$$\begin{aligned} \boldsymbol{A}(\boldsymbol{x}) &= \int \frac{\mathrm{d}^3 k}{(2\pi)^3 \sqrt{2|\boldsymbol{k}|}} \left(\boldsymbol{a}_{\boldsymbol{k}} e^{i\boldsymbol{k}\cdot\boldsymbol{x}} + \boldsymbol{a}_{\boldsymbol{k}}^{\dagger} e^{-i\boldsymbol{k}\cdot\boldsymbol{x}} \right) \\ \boldsymbol{\Pi}(\boldsymbol{x}) &= \int \frac{\mathrm{d}^3 k}{(2\pi)^3} (-i) \sqrt{\frac{|\boldsymbol{k}|}{2}} \left(\boldsymbol{a}_{\boldsymbol{k}} e^{i\boldsymbol{k}\cdot\boldsymbol{x}} - \boldsymbol{a}_{\boldsymbol{k}}^{\dagger} e^{-i\boldsymbol{k}\cdot\boldsymbol{x}} \right) \end{aligned}$$

with $\mathbf{k} \cdot \mathbf{a_k} = \mathbf{k} \cdot \mathbf{a_k^{\dagger}} = 0$.

Introduce 2 orthogonal polarization vectors $\boldsymbol{\epsilon}^{(1)}(\boldsymbol{k})$ and $\boldsymbol{\epsilon}^{(2)}(\boldsymbol{k})$ for each \boldsymbol{k} .

$$\mathbf{a}_{k} = a_{k}^{(1)} \boldsymbol{\epsilon}^{(1)} + a_{k}^{(2)} \boldsymbol{\epsilon}^{(2)} = \sum_{\lambda=1}^{2} a_{k}^{(\lambda)} \boldsymbol{\epsilon}^{(\lambda)}(\mathbf{k})$$
with $\mathbf{k} \cdot \boldsymbol{\epsilon}^{(1)}(\mathbf{k}) = \mathbf{k} \cdot \boldsymbol{\epsilon}^{(2)}(\mathbf{k}) = 0$, $\boldsymbol{\epsilon}^{(\lambda)} \cdot \boldsymbol{\epsilon}^{(\lambda;)} = \delta_{\lambda \lambda'}$

Creation and annihilation operator have the standard commutation relations

$$[a_{\mathbf{k}}^{(\lambda)}, a_{\mathbf{k}'}^{(\lambda')\dagger}] = (2\pi)^3 \delta_{\lambda\lambda'} \delta^{(3)}(\mathbf{k} - \mathbf{k}')$$

$$(5.2.2)$$

and all other commutators vanish. Geometrically, still possible to write including the unphysical longitudinal components:

$$[\boldsymbol{a_k}, \boldsymbol{a_l}] = [\boldsymbol{a_k^{\dagger}}, \boldsymbol{a_l^{\dagger}}] = 0$$
$$[\boldsymbol{a_k^i}, \boldsymbol{a_l^{j\dagger}}] = (2\pi)^3 \left(\delta^{ij} - \frac{k^i k^j}{k^2}\right) \delta^{(3)}(\boldsymbol{k} - \boldsymbol{l})$$

 $a_{\pmb{k}}^{(\lambda)}$ and $a_{\pmb{k}}^{(\lambda)\dagger}$ create and destroy photons of momentum \pmb{k} , energy $|\pmb{k}|$ and (electric) polarization along $\pmb{\epsilon}^{(\lambda)}(\pmb{k})$.

Next steps are analogout to KG theory.

Hamiltonian

$$H = \frac{1}{2} \int d^3x \left(\mathbf{E}^2 + \mathbf{B}^2 \right) = \frac{1}{2} \int d^3x \left(\dot{\mathbf{A}}^2 + (\nabla \times \mathbf{A}) \cdot (\nabla \times \mathbf{A}) \right)$$

using identity $\mathbf{A} \cdot (\mathbf{B} \times \mathbf{C}) = \mathbf{B} \cdot (\mathbf{C} \times \mathbf{A})$

$$= \frac{1}{2} \int d^3x \left(\dot{\boldsymbol{A}}^2 + \boldsymbol{A} \cdot \boldsymbol{\nabla} \times (\boldsymbol{\nabla} \times \boldsymbol{A}) \right)$$

using the identity $\nabla \times (\nabla \times \mathbf{A}) = \nabla (\nabla \cdot \mathbf{A} - \nabla^2 \mathbf{A})$

$$= \frac{1}{2} \int d^3x \left(\dot{\boldsymbol{A}}^2 - \boldsymbol{A} \cdot \nabla^2 \boldsymbol{A} + \boldsymbol{A} \cdot \nabla (\nabla \cdot \boldsymbol{A}) \right)$$

using coulomb gauge condition

$$= \frac{1}{2} \int d^3x \left(\dot{\boldsymbol{A}}^2 - \boldsymbol{A} \cdot \nabla^2 \boldsymbol{A} \right)$$

the first term vanishes and use normal ordering

$$= \int \frac{\mathrm{d}^3 k}{(2\pi)^3} |\mathbf{k}| \mathbf{a}_{\mathbf{k}}^{\dagger} \cdot \mathbf{a}_{\mathbf{k}} = \sum_{k=1}^2 \int \frac{\mathrm{d}^3 k}{(2\pi)^3} |\mathbf{k}| a_{\mathbf{k}}^{(\lambda \dagger)} a_{\mathbf{k}}^{\lambda}$$

Heisenberg field

$$\boldsymbol{A}(\boldsymbol{x},t) = \int \frac{\mathrm{d}^3 k}{(2\pi)^3} \frac{1}{\sqrt{2|\boldsymbol{k}|}} \left(\boldsymbol{a}_{\boldsymbol{k}} e^{-ik \cdot x} + \boldsymbol{a}_{\boldsymbol{k}}^{\dagger} e^{ik \cdot x} \right)$$

Photon propagator

$$\langle 0|TA_{i}(x)A_{j}(y)|0\rangle =: D_{ij}^{tr}(x-y) = \int \frac{\mathrm{d}^{4}k}{(2\pi)^{4}} \frac{i}{k^{2}+i\epsilon} \left(\delta_{ij} - \frac{k_{i}k_{j}}{|\mathbf{k}|^{2}}\right) e^{-ik\cdot(x-y)}$$
(5.2.3)

tr stands for transverse: photon polarization perpendicular to its momentum. This is **NOT** the final version of the photon propagator!

5.3 Inclusion of matter - QED

$$\mathcal{L}_{\text{QED}} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \bar{\psi} (i\not\!\!D - m) \psi \tag{5.3.1}$$

where $D_{\mu} = \partial_{\mu} + ieA_{\mu}$ is the (gauge) covariante derivative

$$= \mathcal{L}_{EM} + \mathcal{L}_D - e \underbrace{\bar{\psi} \gamma^{\mu} \psi A_{\mu}}_{i^{\mu}}$$
 (5.3.2)

Field equations would be

$$\partial_{\mu}F^{\mu\nu}=ej^{\nu}\qquad (iD\!\!\!/-m)\psi=0$$

where ej^{ν} is the electromagnetic 4-current.

Gauge invariance under the transformation

$$\begin{cases} \psi(x) \longmapsto \psi'(x) = e^{ie\chi(x)} \psi \\ A_{\mu}(x) \longmapsto A'_{\mu}(x) = A_{\mu}(x) - \partial_{\mu}\chi(x) \end{cases}$$

To check the consistence: cavariant derivative transforms like $D_{\mu} \longmapsto D'_{\mu} \psi'(x) = e^{ie\chi(x)} D_{\mu} \psi(x)$. Since the adjoint spinor transforms like $\bar{\psi}(x) \longmapsto \bar{\psi}'(x) = \bar{\psi}(x)e^{-ie\chi(x)}$, the Lagrangian and field equations are gauge invariant.

Again we choose Coulomb gauge $\nabla \cdot \mathbf{A} = 0$, then equation for A^0 :

$$\partial_{i}F^{i0} = ej^{0}$$

$$\Rightarrow -\nabla^{2}A^{0} = ej^{0} = e\bar{\psi}\gamma^{0}\psi$$

$$= e\bar{\psi}\gamma^{0}\psi = e\psi^{\dagger}\psi$$

$$= e\rho(x)$$

$$A^{0}(\mathbf{x}, t) = e\int d^{3}y \frac{\rho(\mathbf{y}, t)}{4\pi |\mathbf{x} - \mathbf{y}|}$$
(5.3.3)

We want to derive the interaction Hamiltonian. Note

$$\int d^3x \frac{1}{2} \boldsymbol{E}^2 = \int d^3x \frac{1}{2} \left(\dot{\boldsymbol{A}} + \boldsymbol{\nabla} A^0 \right)^2$$

cross terms vanish after integration by parts due to $\nabla \cdot \dot{\bf A} = 0$

$$= \int d^3x \frac{1}{2} \left(\dot{\mathbf{A}}^2 + (\nabla A^0)^2 \right)$$

$$= \int d^3x \frac{1}{2} \left(\dot{\mathbf{A}}^0 - A^0 \nabla^2 A^0 \right)$$

$$= \int d^3x \frac{1}{2} \dot{\mathbf{A}}^2 + \underbrace{\frac{e^2}{2} \int d^3x d^3y \frac{\rho(\mathbf{x})\rho(\mathbf{y})}{4\pi |\mathbf{x} - \mathbf{y}|}}_{=\frac{e^2}{2} j^0 A_0}$$

Combined Hamiltonian

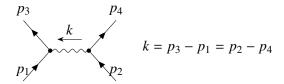
$$H = \int d^3x \left\{ \frac{1}{2} \mathbf{\Pi} \cdot \mathbf{\Pi} + \frac{1}{2} \mathbf{B} \cdot \mathbf{B} + i \bar{\psi} \mathbf{\gamma} \cdot \nabla \psi + m \bar{\psi} \psi \right\}$$
 free photon and fermion

$$+ \frac{e^2}{2} \int d^3x d^3y \frac{\rho(\mathbf{x})\rho(\mathbf{y})}{4\pi |\mathbf{x} - \mathbf{y}|} - e \int d^3x \mathbf{j} \cdot \mathbf{A}$$
 interactions

where $\rho = \psi^{\dagger} \psi = \bar{\psi} \gamma^0 \psi$, $\mathbf{j} = \bar{\psi} \boldsymbol{\gamma} \psi$ for 2 types of interactions.

5.4 Lorentz-invariant propagator

Consider e^-e^- scattering at $O(e^2)$



We expect this to involve

- spinors for external fermions
- $-ie\gamma^{\mu}$
- Photon propagator $D_{\mu\nu}(x-y)$

What we have found in Coulomb gauge is actually

- vertices $ie\gamma^i$, transverse propagator $D_{\mu\nu}^{\text{tr}}(x-y)$
- vertices $\pm ie\gamma^0$, instantaneous Coumlomb interaction $\frac{1}{4\pi|\mathbf{x}-\mathbf{y}|}\delta(x^0-y^0)$

Effectively combine these propagators terms into $D_{\mu\nu}^{\text{Coul}}(x-y)$, where the $D_{00}^{\text{Coul}}(x-y) = \frac{1}{4\pi |\mathbf{x}-\mathbf{y}|} \delta(x^0-y^0)$. This component in momentum space is simply

$$\int \frac{\mathrm{d}^3 k}{(2\pi)^3} \frac{e^{i\boldsymbol{k}\cdot\boldsymbol{r}}}{|\boldsymbol{k}|^2} = \frac{1}{4\pi|\boldsymbol{r}|}$$

Therefore Coumlomb propagator in momentum space:

$$D_{\mu\nu}^{\text{Coul}}(k) = \begin{cases} \frac{i}{k^2 + i\epsilon} \left(\delta_{ij} - \frac{k_i k_j}{|\mathbf{k}|^2} \right) & \mu = i, \nu = j \\ \frac{i}{|\mathbf{k}|^2} & \mu = \nu = 0 \\ 0 & \text{otherwise} \end{cases}$$

Consider contraction to scattering amplitude from vertex at x:

$$\sim e\bar{u}(p_3)\gamma^{\mu}u(p_1)e^{i(p_3-p_1)x}$$

current conservation $\partial_{\mu}j^{\mu}=0$ written in momentum space

$$\underbrace{(p_3-p_1)_\mu}_{k_\mu}\bar{u}(p_3)\gamma^\mu u(p_1)=0$$

so in the complete diagram $D_{\mu\nu}^{\mathrm{Coul}}$ occurs in a form

$$a^{\mu}D_{\mu\nu}^{\text{Coul}}(k)b^{\nu}$$

$$=a^{0}\frac{i}{|\mathbf{k}|^{2}}b^{0}+a^{i}\left[\frac{i}{k^{2}+i\epsilon}\left(\delta_{ij}-\frac{k_{i}k_{j}}{|\mathbf{k}|^{2}}\right)\right]b^{j}$$

where $k^{\mu}a^{\mu} = 0, k_{\mu}a^{\mu} = 0$

$$=i\left[\frac{\boldsymbol{a}\cdot\boldsymbol{b}}{k^{2}}\underbrace{-\frac{k_{0}^{2}a_{0}b_{0}}{k^{2}|\boldsymbol{k}|^{2}}+\frac{a_{0}b_{0}}{|\boldsymbol{k}|^{2}}}_{=\frac{-k_{0}^{2}a_{0}b_{0}+a_{0}b_{0}(k_{0}^{2}-|\boldsymbol{k}|^{2})}{k^{2}|\boldsymbol{k}|}}\right]$$

$$=\frac{i}{k^2}(\boldsymbol{a}\cdot\boldsymbol{b}-a_0b_0)=-\frac{i}{k}a_\mu b^\mu$$

Conclusion in this diagram (and in fact, in general), we may replace the $D_{\mu\nu}^{\text{Coul}}(k)$ by the manifestly Lorentz covariant propagator

$$D_{\mu\nu}(k) = -\frac{ig_{\mu\nu}}{k + i\epsilon} \tag{5.4.1}$$

This can be generalised to

$$D_{\mu\nu}(k) = -\frac{i}{k^2 + i\epsilon} \left(g_{\mu\nu} - (1 - \alpha) \frac{k_{\mu} k_{\nu}}{k^2} \right)$$
 (5.4.2)

as, by current consevation, additional term doesn't contribute.

Feynman gauge $\alpha = 1$; Landau gauge $\alpha = 0$.

Remark one can also try to quantise photons in a manifestly covariant way, imposing Lorentz gauge $\partial_{\mu}A^{\mu}=0$

$$[A_{\mu}(\boldsymbol{x}), \Pi_{\nu}(\boldsymbol{y})] = ig_{\mu\nu}\delta^{(3)}(\boldsymbol{x} - \boldsymbol{y})$$

This is trouble since $\Pi^0 = \frac{\partial \mathcal{L}}{\partial \dot{A}_0} = 0$. This cannot hold!

We thus change the Lagrangian

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{2\alpha}(\partial_{\mu}A^{\mu})^2$$

with "gauge fixing term". The equation of motion from this is

$$\partial^2 A^{\mu} - \left(1 - \frac{1}{\alpha}\right) \partial^{\mu} (\partial_{\lambda} A^{\lambda}) = 0$$

e.g. $\alpha = 1$ is the Feynman gauge.

With this Lagrangian we can the 0th component of conjugate momentum

$$\Pi^0 = -\frac{1}{\alpha} \partial_\mu A^\mu$$

but this seems as bad as before!

We cannot impose Coulomb gauge condition $\partial_{\mu}A^{\mu}=0$ as an operator identity. Instead demand a weaker condition $\langle \text{out}|\partial_{\mu}A^{\mu}|\text{in}\rangle=0$ for all physical states.

This in turn tells us which states are actually physical. The 4 polarisation states consist of physical, timelike(scalar) and longitudinal states. The negative-norm states cancel each ther out (Gopta-Bleuler formalism).

Feynman rules for QED diagrams constucted from electron (positron) — and photon ,, rules for fermions are valid as before.

In addition

• vertex
$$\mu = -ie\gamma^{\mu};$$

- photon propagator $\begin{picture}(1,0) \put(0,0){\line(1,0){100}} \put($
- external photons

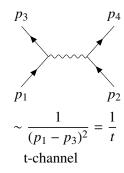
$$\mu \stackrel{K_{\text{in}}}{\longleftarrow} \nu = \epsilon_{\mu}$$

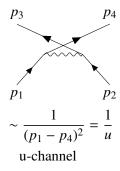
$$\mu \xrightarrow{k_{\text{out}}} \nu = \epsilon_{\nu}^{*}$$

 ϵ_{μ} polarisation vector of in/out photon and ϵ_{μ}^* for out photon required for complex (circular) polarisation.

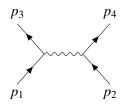
5.5 QED process at tree level

Example $e^-e^- \rightarrow e^-e^-$

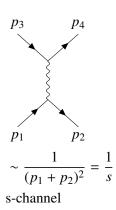




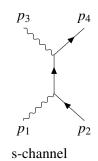
Example $e^-e^+ \rightarrow e^-e^+$

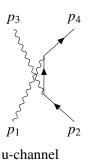


$$\sim \frac{1}{(p_1 - p_3)^2} = \frac{1}{t}$$
t-channel

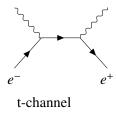


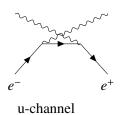
Compton scattering $\gamma e^- \to \gamma e^-$





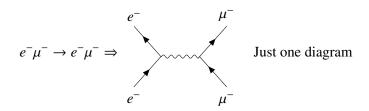
Example $e^+e^- o \gamma\gamma$

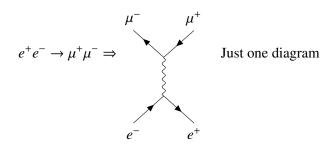




These are important for lifetime of positronium.

All these amplitudes are $O(e^2)$, $\alpha = \frac{e^2}{4\pi} = \frac{1}{137.036}$ the fine structure constant. Muons μ^{\pm} , like electrons, just ca. 200 times heavier.





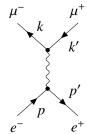
 μ^{\pm} decay into e^{\pm} and neutrinos in weak interactions.

For tree level diagrams the photon propagator does not need to have the $i\epsilon$ in the denominator, since we will never be able to see a singularity/pole.

5.5.1 Some hints and tricks for cross section calculations

See application in exercises!

Example $e^+e^- \rightarrow \mu^+\mu^-$



$$\begin{split} iM &= \bar{v}_e^{s'}(-ie\gamma^\mu)u_e^s(p) \left. \frac{-ig_{\mu\nu}}{s} \right|_{s=q^2} \bar{u}_\mu^r(k)(-ie\gamma^\nu)v_\mu^{r'}(k') \\ &= \frac{ie^2}{s} \left(\bar{v}_e(p')\gamma^\mu u_e(p) \right) \left(\bar{u}_\mu(k)\gamma_\mu v_\mu(k') \right) \end{split}$$

See section ??, $|M|^2$ is needed for cross section. M^* involves things like

$$(\bar{\nu}\gamma^{\mu}u)^* = (\bar{\nu}\gamma^{\mu}u)^{\dagger} = u^{\dagger}\gamma^{\mu\dagger}\gamma_0^{\dagger}\nu$$
$$= u^{\dagger}\gamma_0\gamma^{\mu}\gamma_0\gamma_0\nu = \bar{u}\gamma^{\mu}\nu$$

So

$$|M|^2 = \frac{e^4}{s^2} \left[\bar{v}(p') \gamma^{\mu} u(p) \bar{u}(p) \gamma^{\nu} v(p') \right]_{e^{\pm}} \cdot \left[\bar{u}(k) \gamma_{\mu} v(p) \bar{v}(k') \gamma_{\nu} u(k) \right]_{\mu^{\pm}}$$

Unpolarized scattering= $\frac{1}{4} \sum_{r,s,r',s'} |M|^2$.

Now $\bar{v}\gamma^{\mu}u$, $\bar{u}\gamma^{\nu}v$ etc. are scalars in Dirac/spinor space:

$$\sum_{s,s'} \bar{v}_{s'} p' \gamma^{\mu} u_{s}(p) \bar{u}_{s}(p) \gamma^{\nu} v_{s'}(p')$$
(taking trace of scalar)
$$= \sum_{s,s'} \operatorname{Tr} \left(\bar{v}_{s'} p' \gamma^{\mu} u_{s}(p) \bar{u}_{s}(p) \gamma^{\nu} v_{s'}(p') \right)$$

$$= \sum_{s,s'} \operatorname{Tr} \left(v_{s'}(p') \bar{v}_{s'}(p') \gamma^{\mu} u_{s}(p) \bar{u}_{s}(p) \gamma^{\nu} \right)$$

using spin sums

$$= \operatorname{Tr} \left((p' - m) \gamma^{\mu} (p + m) \gamma^{\nu} \right)$$

Trace technology

- remember $Tr\gamma_{\mu} = 0$
- $\operatorname{Tr}(\gamma^{\mu}\gamma^{\nu}) = 4g_{\mu\nu}$
- Tr(odd number of γ) = 0
- $\operatorname{Tr}(\gamma_{\mu}\gamma_{\nu}\gamma_{\alpha}\gamma_{\beta}) = 4\left(g_{\mu\nu}g_{\alpha\beta} + g_{\mu\beta}g_{\nu\alpha} g_{\mu\alpha}g_{\nu\beta}\right)$
- more rules involving γ_5 (weak interactions!)

So

$$Tr((p'-m)\gamma^{\mu}(p+m)\gamma^{\nu}) = 4(p'^{\mu}p^{\nu} + p'^{\nu}p^{\mu} - g^{\mu\nu}(p \cdot p' + m^2))$$

Mandelstam variables with 4 equal masses, center-of-mass system (CMS):

$$p = (E, \mathbf{p}), \quad p' = (E, -\mathbf{p}), \quad k = (E, \mathbf{k}), \quad \theta = \measuredangle(\mathbf{p}, \mathbf{k})$$

$$s \stackrel{\text{CMS}}{=} (p + p')^2 = 4E^2$$
 (5.5.1)

$$t = (p - k)^{2} = -(\mathbf{p} - \mathbf{k})^{2} = -2|\mathbf{p}|^{2}(1 - \cos\theta)$$
 (5.5.2)

$$u = (p' - k)^2 = -2|\mathbf{p}|^2(1 + \cos\theta)$$
 (5.5.3)

$$|\mathbf{p}|^2 = E^2 - m^2 = \frac{s}{4} - m^2 \tag{5.5.4}$$

Only 2 Mandelstam variables are independent.

$$s + t + u = (p_1 + p_2)^2 + (p_1 - p_3)^2 + (p_1 - p_4)^2$$

$$= p_1^2 + p_2^2 + p_3^2 + p_4^2 + 2p_1 \underbrace{(p_1 + p_2 - p_3 - p_4)}_{=0}$$

$$= \sum_i m_i^2 = \text{const}$$

photon polarisation sums Analogy to fermion spin sums before, Feynman rules for external photons involve $\epsilon^{(*)}_{\mu}$; e.g. Compton amplitude of the form

$$M \sim \epsilon_{\mu}^*(p_3)\epsilon_{\nu}(p_1)T^{\mu\nu}$$

Thus

$$\sum_{\text{spin, pol.}} |M|^2 = \sum_{\text{spin, pol.}} \epsilon_{\mu}^*(p_3) \epsilon_{\alpha}(p_3) \epsilon_{\beta}^*(p_1) \epsilon_{\nu}(p_1) T^{\mu\nu} T^{\alpha\beta^*}$$

How can we simplify $\sum_{pol} \epsilon_{\mu}^*(k) \epsilon_{\nu}(k)$? Again we have only 2 physical polarisation states, but want to do it in a covariant form.

Assume a simpler process (than Compton) with a single external photon, $\epsilon_{\mu}^{*}(k)M^{\mu}$. Choose

$$k^{\mu} = (k, 0, 0, k), \quad \epsilon^{\mu}_{(1)} = (0, 1, 0, 0), \quad \epsilon^{\mu}_{(2)} = (0, 0, 1, 0)$$

so
$$\sum_{\text{pol}} |\epsilon_{\mu}^*(k) M^{\mu}|^2 = |M_1|^2 + |M_2|^2$$

Remember that photon coupled source j^{μ} , current sonvervation $\partial_{\mu}j^{\mu}=0$. We will see (next term) this holds in general as Ward identity

$$k_{\mu}M^{\mu} = 0 \tag{5.5.5}$$

In exercises, show $p_{3\mu}T^{\mu\nu} = 0 = p_{1\nu}T^{\mu\nu}$ for Compton Here $kM^0 - kM^3 = 0 \Rightarrow M^0 = M^3$ and we can rewrite

$$\sum_{pol} \epsilon_{\mu}^{*} \epsilon_{\nu} M^{\mu} M^{*\nu} = |M_{1}|^{2} + |M_{2}|^{2} + \underbrace{|M_{3}|^{2} - |M_{0}|^{2}}_{=0} = -g_{\mu\nu} M^{\mu} M^{*\nu}$$

so effectively

$$\sum_{pol} \epsilon_{\mu}^*(k)\epsilon_{\nu}(k) = -g_{\mu\nu} \tag{5.5.6}$$

side remark

- KG propagator $\frac{i}{p^2 M^2 + i\epsilon}$
- Dirac propagator $\frac{i(\not p+m)}{p^2-m^2+i\epsilon} = \frac{i\sum_s u_s(p)\bar{u}_s(p)}{p^2-m^2+i\epsilon}$
- Photon propagator $\frac{-ig_{\mu\nu}}{p^2+i\epsilon} = \frac{i\sum_{pol}\epsilon_{\mu}^*(p)\epsilon_{\nu}(p)}{p^2+i\epsilon}$

6 Radiative corrections

6.1 Optical theorem

We have seen in Advanced Quantum Theory that tree diagrams are in general <u>real</u>. So there is no imaginary parts. Need to restore perturbatively in higher-order corrections. Then the optical theorem is valid again.

S-matrix is unitary: $S^{\dagger}S = 1$ with S = 1 + iT. Thus

$$-i(T-T^{\dagger}) = T^{\dagger}T$$

We take matrix element for $k_1k_2 \rightarrow p_1p_2$ scattering. On RHS, insert a complete set of states,

$$\langle p_1 p_2 | T^{\dagger} T | k_1 k_2 \rangle = \sum_n \prod_{i=1}^n \int \frac{\mathrm{d}^3 q_i}{(2\pi)^3 2E_i} \langle p_1 p_2 | T^{\dagger} | q_1 \dots q_n \rangle \langle q_1 \dots q_n | T | k_1 k_2 \rangle$$

Reduce $T_{fi} = (2\pi)^4 \delta^{(4)}(p_f - p_i) M_{fi}$ and omitting overal $(2\pi)^4 \delta^{(4)}(p_f - p_i)$

$$-i\left[\mathcal{M}(k_1k_2 \to p_1p_2) - \mathcal{M}^*(p_1p_2 \to k_1k_2)\right]$$

$$= \underbrace{\sum \prod_{i=1}^n \int \frac{\mathrm{d}^3q_i}{(2\pi)^3 2E_i}}_{\text{invariant phase-space volume element}} \mathcal{M}^*(p_1p_2 \to q_1 \dots q_n) \mathcal{M}(k_1k_2 \to q_1 \dots q_n) (2\pi)^4 \delta^{(4)}(k_1 + k_2 - \sum_i q_i)$$

So optical theorem, for forward scattering $(p_1 = k_1, p_2 = k_2)$ reads (see ??)

Im
$$\mathcal{M}(k_1k_2 \to k_1k_2) = 2F\sigma_{\text{tot}}(k_1k_2 \to \text{anything})$$

$$2\sqrt{s}|f_i^{\text{CMS}}| = \lambda^{\frac{1}{2}}(s, m_1^2, m_2^2)$$

Optical theorem for Feynman diagrams Consider a specific diagram contributing to the imaginary part, e.g. in ϕ^4 -theory.

$$i\mathcal{M}(s) = \frac{\lambda^2}{2} \int \frac{d^4q}{(2\pi)^4} \frac{1}{[(p_s/2 - q)^2 - M^2 + i\epsilon][(p_s/2 + q)^2 - M^2 + i\epsilon]}$$
(6.1.1)

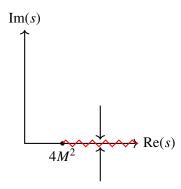
From optical theorem: Im $\mathcal{M}(s < 4M^2) = 0$, so $\mathcal{M}(s < 4M^2) \in \mathbb{R}$, (Since it is physical case, the cross section must vanish) when regarding $\mathcal{M}(s)$ as an analytic function of s beyond what physical S-matrix element allow.

Schwarz reflection principle If (in some region) analytic function $\mathcal{M}(s)$ is <u>real</u> at least for a finite, nonvanishing interval $\in \mathbb{R}$, then

$$\mathcal{M}(s^*) = \mathcal{M}^*(s) \tag{6.1.2}$$

Hence

$$\mathcal{M}(s+i\epsilon) - \mathcal{M}(s-i\epsilon) \equiv \operatorname{disc}\mathcal{M}(s) = \mathcal{M}(s+i\epsilon) - \mathcal{M}^*(s+i\epsilon) = 2i\operatorname{Im}\mathcal{M}(s+i\epsilon)$$



Onset of imaginary part for $s \le 4M^2$ necessarily leads to a "branch cut", a nontrivial discontinuity in the comlex energy plane. The branch cut is equivalent to $\sqrt{4M^2 - s}$. Function has discontinuity, a cut, on real axis.

How can we calculate the discontinuity (= imaginary part) of the above diagram?

Use centre-of-mass system $p_s = (\sqrt{s}, \mathbf{0})$. Poles from propagators

$$\frac{s}{4} \mp \sqrt{sq^0 + q^2 - M^2 + i\epsilon} = 0$$

$$\Leftrightarrow (q^0)^2 \pm \sqrt{sq^0 + \frac{s}{4} - |\mathbf{q}|^2 - M^2 + i\epsilon} = 0$$

first propagator
$$q^0 = +\frac{\sqrt{s}}{2} \pm (\sqrt{M^2 + |\boldsymbol{q}|^2} - i\epsilon) = +\frac{\sqrt{s}}{2} \pm (E_q - i\epsilon)$$
 second propagator
$$q^0 = -\frac{\sqrt{s}}{2} \pm (E_q - i\epsilon)$$

$$Im(q^{0})$$

$$-\frac{\sqrt{s}}{2} - E_{q} + i\epsilon \qquad + \frac{\sqrt{s}}{2} - E_{q} + i\epsilon \qquad \rightarrow \operatorname{Re}(q^{0})$$

$$-\frac{\sqrt{s}}{2} + E_{q} + i\epsilon \qquad + \frac{\sqrt{s}}{2} + E_{q} + i\epsilon$$

If we close the contour of the q_0 integration in the <u>lower</u> half plane, we only pick up the 2 residues at $\mp \frac{\sqrt{s}}{2} + E_q - i\epsilon$. As E_q is positive, only $-\frac{\sqrt{s}}{2} + E_q - i\epsilon$ from second propagator contirbutes to discontinuity.

So pinching up the residue equivalent to replacement under q^0 integration

$$\frac{1}{(p_s/2+q)^2-M^2+i\epsilon} \longmapsto \underbrace{-2\pi i}_{\text{orientation of contour}} \delta((p_s/2+q)^2-M^2)$$

Determine the residue of the rest at the pole at $-\frac{\sqrt{s}}{2} + E_q - i\epsilon$

$$M(s) \longmapsto -\frac{\lambda^2}{2} \int \frac{\mathrm{d}^3 q}{(2\pi)^3} \frac{1}{2E_a \sqrt{s}(\sqrt{s} - 2E_a)}$$

With no angular dependence and using substitution (note the limits of integral also change) $d^3q \rightarrow 4\pi |\mathbf{q}|^2 d|\mathbf{q}| = 4\pi |\mathbf{q}| E_q dE_q$

$$= -\frac{\lambda^2}{8\pi^2} \int_M^{\infty} \frac{dE_q \sqrt{E_q^2 - M^2}}{\sqrt{s}(\sqrt{s} - 2E_q)}$$
 (6.1.3)

It has pole at $E_q = \frac{\sqrt{s}}{2}$. The second pole in $\ref{eq:model}$?? at $\frac{\sqrt{s}}{2} + E_q - i\epsilon$ would produce a pole in $\ref{eq:model}$?? for $E_q = -\frac{\sqrt{s}}{2}$, outside the integration range $M \le E_q < \infty$.

- for $\sqrt{s} < 2M$, ?? is manifestly real.
- for $\sqrt{s} > 2M$, the pole at $E_q = \frac{\sqrt{s}}{2}$ in ?? contributes differently depending on $\sqrt{s} \pm i\epsilon$; difference yields discontinuity.

Use

$$\frac{1}{\sqrt{s} - 2E_q \pm i\epsilon} = \underbrace{\frac{P}{\sqrt{s} - 2E_q}}_{\text{real}} \underbrace{\mp i\pi\delta(\sqrt{s} - 2E_q)}_{\text{yields discontinuity}}$$

So for calculation of the discontinuity, have replacement

$$\frac{1}{(p_s/2-q)^2-M^2+i\epsilon}\longmapsto -2\pi i\delta((p_s/2-q)^2-M^2)$$

for other propagator too!

Cuthosky rules (1960) replace cut propagator according to

$$\frac{1}{p^2 - M^2 + i\epsilon} \longmapsto -2\pi i \delta(p^2 - M^2) \tag{6.1.4}$$

to calculate discontinuity across the cut!

Calculateion completed:

disc
$$= i\frac{\lambda^2}{2} \int \frac{d^4q}{(2\pi)^4} 2\pi \delta(q^2 - M^2) 2\pi \delta((p_s - q)^2 - M^2)$$

using
$$d^4q = dq^0 dq |q|^2 d\Omega_q$$
 and $(p_s - q)^2 - M^2 = s - 2\sqrt{s}q^0$

$$= \frac{\lambda^2}{2} \frac{i}{4\pi^2} \int \frac{|q|^2 d|q| d\Omega_q}{2q^0} \delta(s - 2\sqrt{s}q^0)$$

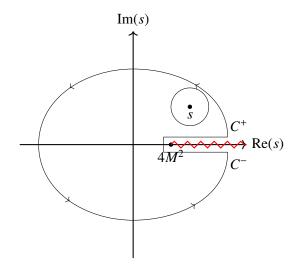
$$= \frac{\lambda^2}{2} \frac{i}{8\pi^2} \int \sqrt{(q^0)^0 - M^2} dq^0 d\Omega_q \delta(s - 2\sqrt{s}q^0)$$

$$= \frac{\lambda^2}{2} \frac{i}{8\pi^2} \frac{\sqrt{s/4 - M^2}}{2\sqrt{s}} \int d\Omega_q$$

$$= \frac{\lambda^2}{2} \frac{i}{8\pi} \sqrt{1 - \frac{4M^2}{s}}$$

$$Im \mathcal{M} = \frac{\lambda^2}{4} \frac{1}{8\pi} \sqrt{1 - \frac{4M^2}{s}}$$

Note $\sigma = \frac{\lambda^2}{32\pi}$ and $2F = s\sqrt{1 - \frac{4M^2}{s}}$. Thus optical theorem is still valid. We can do more. Construct the complete $\mathcal{M}(s)$ from Im $\mathcal{M}(s)$ through a dispersion relation!



Use Cauchy's theorem:

$$\mathcal{M}(s) = \frac{1}{2\pi i} \oint \frac{\mathcal{M}(z)dz}{z - s}$$
 (6.1.5)

dropping the large circle

$$\longmapsto \frac{1}{2\pi i} \int_{C_{+}+C_{-}} \frac{\mathcal{M}(z)dz}{z-s}$$

$$= \frac{1}{2\pi i} \left[\int_{4M^{2}}^{\infty} \frac{M(z+i\epsilon)dz}{z-s} - \int_{4M^{2}}^{\infty} \frac{M(z-i\epsilon)dz}{z-s} \right]$$

$$= \frac{1}{2\pi i} \int_{4M^{2}}^{\infty} \frac{\operatorname{disc}\mathcal{M}(z)dz}{z-s}$$

$$= \frac{1}{\pi} \int_{4M^{2}}^{\infty} \frac{\operatorname{Im}\mathcal{M}(z)dz}{z-s}$$
(6.1.6)

Repeat the exercise for $\frac{\mathcal{M}(s)-\mathcal{M}(0)}{s}$ (no pole introduced!).

$$\operatorname{Im}\left(\frac{\mathcal{M}(s) - \mathcal{M}(0)}{s}\right) = \frac{\operatorname{Im} \mathcal{M}(s)}{s}$$

$$\mathcal{M}(s) - \mathcal{M}(0) = \frac{s}{\pi} \int_{4M^2}^{\infty} \frac{\operatorname{Im} \mathcal{M}(z) dz}{z(z - s)}$$

$$= \frac{\lambda^2}{2} \frac{s}{(4\pi)^2} \int_{4M^2}^{\infty} \frac{dz}{z(z - s)} \sqrt{1 - \frac{4M^2}{z}}$$

$$= \frac{\lambda^2}{2} \frac{1}{8\pi^2} \int_0^1 \frac{\zeta^2}{\zeta^2 - \sigma^2} d\zeta$$

$$= \frac{\lambda^2}{2} \frac{1}{8\pi^2} \left(1 - \frac{\sigma}{2} \log \frac{\sigma + 1}{\sigma - 1}\right) \qquad s < 0 \Leftrightarrow \sigma > 1$$

$$= \frac{\lambda^2}{2} \frac{1}{8\pi^2} \left(1 - \sqrt{-\sigma^2} \arctan \frac{1}{\sqrt{-\sigma^2}}\right) \qquad 0 < s < 4M^2, \sigma^2 < 0$$

$$\frac{1}{8\pi^2} \left(1 - \frac{\sigma}{2} \log \frac{1 + \sigma}{1 - \sigma} + \frac{i\sigma}{16\pi}\right) \qquad s > M^2, 0 < \sigma < 1$$

Note: we are going to calculate this diagram again, noticing that $\int \frac{d^4q}{(q^2...)(q^2...)}$ is logarithmically divergent!. The above representation demonstrates that this divergence resides in M(0)!

6.2 Field-strength renomrlization

What is structure of the propagator $\langle \Omega | T \phi(x) \phi(y) | \Omega \rangle$ at higher orders? At lower order

$$\frac{p}{p} = \frac{i}{p^2 - M^2 + i\epsilon}$$

Beyond this it is not a simple pole. In ϕ^3 -theory _____ branch at $p^2 \le 4M^2$. In ϕ^4 -theory

branch at $p^2 \le 9M^2$. To induce cuts in the analytic structure.

Insert complete set of intermediate states $(x^0 > y^0)$

$$\langle \Omega | T \phi(x) \phi(y) | \Omega \rangle = \sum_{\lambda} \int \frac{\mathrm{d}^3 p}{(2\pi)^3 2 E_p(\lambda)} \langle \Omega | \phi(x) | \lambda_f \rangle \langle \lambda_f | \phi(y) | \Omega \rangle$$

with

 λ multiparticle state

 λ_0 "rest frame", i.e. $\hat{\boldsymbol{P}} | \lambda_0 \rangle = 0$

 λ_{p} boosted to momentum p

Call energy of $\lambda_0 = m_{\lambda}$. From single particle to multi particle $E_{\mathbf{p}}(\lambda) = \sqrt{m_{\lambda}^2 + |\mathbf{p}|^2}$.

$$\begin{split} \langle \Omega | \phi(x) | \lambda_{\pmb{p}} \rangle &= \langle \Omega | e^{i\hat{P}x} \phi(0) e^{-i\hat{P}x} | \lambda_{\pmb{p}} \rangle \\ &= \langle \Omega | \phi(0) | \lambda_{\pmb{p}} \rangle e^{-ipx} \Big|_{p^0 = E_{\pmb{p}}} \end{split}$$

 Ω and $\phi(0)$ are invariant under momentum boost

$$= \left. \langle \Omega | \phi(0) | \lambda_0 \rangle \, e^{-ipx} \right|_{p^0 = E_{\pmb{p}}}$$

$$\langle \Omega | T \phi(x) \phi(y) | \Omega \rangle = \sum_{\lambda} \int \frac{\mathrm{d}^3 p}{(2\pi)^3 2 E_p(\lambda)} e^{-ip(x-y)} |\langle \Omega | \phi(0) | \lambda_0 \rangle|^2$$
(6.2.1)

$$= \sum_{\lambda} \underbrace{\int \frac{\mathrm{d}^4 p}{(2\pi)^4} \frac{i}{p^2 - m_{\lambda}^2 + i\epsilon} e^{-ip(x-y)} |\langle \Omega | \phi(0) | \lambda_0 \rangle|^2}_{D_F(x-y; m_{\lambda}^2) \text{ when combined with } y^0 > x^0}$$
(6.2.2)

(6.2.3)

Formally write this as

$$\langle \Omega | T \phi(x) \phi(y) | \Omega \rangle = \int_0^\infty \frac{\mathrm{d}s}{2\pi} \rho(s) D_F(x - y; s)$$
 (6.2.4)

with $\rho(s)$ the spectral density function.

$$\rho(s) := \sum_{\lambda} (2\pi)\delta(s - m_{\lambda}^2) |\langle \Omega | \phi(0) | \Omega \rangle|^2$$
(6.2.5)

A typical spectral function looks like

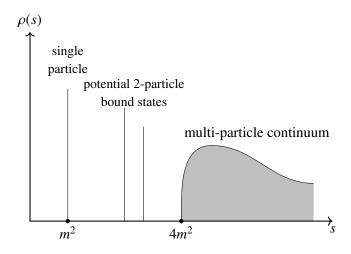


Figure 6.1: typical spectral function

Single particle contribution

$$\rho(s) = 2\pi\delta(s - m^2)z + (\text{contributions} \ge 4m^2)$$
 (6.2.6)

with $z = |\langle \Omega | \phi(0) | \boldsymbol{p} \rangle|^2$ the field-strength renomrlization factor. ($|\boldsymbol{p}\rangle$ is single particle state) Fourier transforming

$$\int d^4x e^{ipx} \langle \Omega | T\phi(x)\phi(0) | \Omega \rangle$$

$$= \int_0^\infty \frac{ds}{2\pi} \rho(s) \frac{i}{p^2 - s + i\epsilon}$$

$$= \frac{iz}{p^2 - m_i^2 \epsilon} + \int_{\sim 4m^2}^\infty \frac{ds}{2\pi} \rho(s) \frac{i}{p^2 - s + i\epsilon}$$

Comparing to free theory: $\langle 0|\phi(0)|\boldsymbol{p}\rangle=1$ hence z=1.