

Theoretical Astroparticle Physik

Homework 6

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1 Graviton Decoupling

- a) Since gravitational interaction is characterised by M_{pl} , through simple dimensional analysis, we estimate the interaction rate

$$\Gamma = \frac{T^5}{M_{pl}^4} \quad (1.1)$$

The Hubble time in radiation-dominated era is $H = T^2/M_{pl}$. The particle of interest decouples when $\Gamma \sim H$, so

$$\frac{\Gamma}{H} = \left(\frac{T}{M_{pl}} \right)^3 \sim 1$$

Thus we have

$$T_{G,dec} \sim M_{pl} = 1.2 \times 10^{19} \text{ GeV} \quad (1.2)$$

It happened rather near the beginning of the Universe. It also justifies the assumption that it happened in the radiation-dominated era.

- b) With this high temperature, all particles in SM were relativistic. They were all in thermal equilibrium, thus $(T_i/T) = 1$. Since this is before recombination and QCD phase transition, we only consider elementary particles.

Ths bosons are the Higg boson , gluons, W^\pm , Z^0 and γ . The fermion part includes quarks and leptons. Note that in SM there is no right-

handed neutrinos.

$$g_{*s}(T_{G,\text{dec}}) = \sum_{\text{bosons}} g_i + \frac{7}{8} \sum_{\text{fermions}} g_i \quad (1.3)$$

$$= (1 \cdot 1 + 1 \cdot 8 \cdot 2 + 2 \cdot 3 + 1 \cdot 3 + 1 \cdot 2) + \frac{7}{8} \cdot 2 \cdot (3 \cdot 2 \cdot 6 + 3 \cdot 2 + 3 \cdot 1) \quad (1.4)$$

$$= 106.75 \quad (1.5)$$

Somehow considering the particles/fields before EWSB doesn't give the correct value. The total degrees of freedom are not the same, since the Higgs loses its one degree of freedom to make W^\pm and Z^0 massive and thus give them in total 3 degrees of freedoms.

- c) In the present Universe, there are only photons and neutrinos still relativistic. Although neutrinos are relativistic, they have already decoupled and not in thermal equilibrium with photons anymore. From the lecture, $T_{\gamma,0}/T_{\nu,0} \simeq (11/4)^{1/3}$

$$g_{*s}(T_0) = 1 \cdot 2 + \frac{7}{8} \cdot 6 \cdot \frac{4}{11} = 3.91 \quad (1.6)$$

Instead of considering only electron-photon component in determining neutrino temperature, we consider all the particles

$$g_{*s}(T)a^3T^3 = \text{const}$$

Using this to compare the time of graviton decoupling and present Universe.

$$\begin{aligned} g_{*s}(T_{G,\text{dec}})a_{G,\text{dec}}^3T_{G,\text{dec}}^3 &= g_{*s}(T_0)a_0^3T_0^3 \\ \Rightarrow T_{G,0}^3 &= \left(T_{G,\text{dec}}\frac{a}{a_0}\right)^3 = T_0^3 \frac{g_{*s}(T_0)}{g_{*s}(T_{G,\text{dec}})} \\ \Rightarrow T_{G,0} &= T_0 \left(\frac{g_{*s}(T_0)}{g_{*s}(T_{G,\text{dec}})}\right)^{1/3} = 0.897 \text{ K} \end{aligned} \quad (1.7)$$

Since graviton is massless, it should be relativistic

$$n_{G,0} = 1 \cdot \frac{\zeta(3)}{\pi^2} T_{G,0}^3 = 7.25 \text{ cm}^{-3} \quad (1.8)$$

As expected, we have far less gravitons floating around than neutrinos, since graviton decoupling happened before neutrino decoupling.

Shouldn't graviton always be massless?