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#### The team:

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Diego Ortiz, UNIR (Math).
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Wladimir Banda, Hamburg University (Physics).

Me, UTE (Applied Math).

#### What we study:

Role of galactic winds and outflows in galaxy evolution.

Remove gas and metals from the disk and nuclear regions of star-forming galaxies and deposit them in the circumgalactic medium.

#### What we want to understand:

The presence of cold gas (clouds) in such outflows.

#### The Wind/Shock - Cloud simulations:

Transport via momentum transfer from hot gas?

- · In purely hydrodynamic regimes: Too many instabilities, cloud gets destroyed rapidly.
- Recent simulations show that magnetic stresses can aid cloud acceleration and survival.

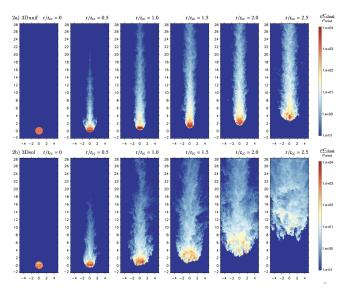
#### In this talk:

 Tools for a systematic statistical study of the effect of magnetic fields.

Need: Probaility distribution over magnetic fileds.

(Herr W.: For magnetic fields with compact support, sometimes symmetric and div-free! And, depending on the day, turbelent too.)

Need: Probaility distribution over magnetic fileds.



Need: Probaility distribution over  $f : \mathbb{R}^m \to \mathbb{R}^n$ . Gaussian Process: A proba. distribution over a function space.

$$f(x) \sim \mathbf{GP}\left(0, \ k(x, x')\right)$$

For any  $\mathbf{x} := [x_1, ..., x_n]^T$ ,

$$\mathbf{f}(\mathbf{x}) \sim \mathbf{N} \left( \begin{bmatrix} 0 \\ \vdots \\ 0 \end{bmatrix}, \begin{bmatrix} k(x_1, x_1) \cdots k(x_1, x_n) \\ & \ddots \\ k(x_n, x_1) \cdots k(x_n, x_n) \end{bmatrix} \right)$$

where  $\mathbf{f}(\mathbf{x}) := [f(x_1), ..., f(x_n)]^T$ .

#### What does that mean? To simulate:

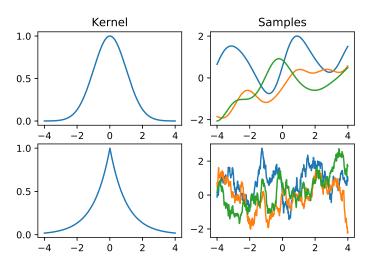
- 1. Choose a vector space a covariance function k.
- 2. Fix a set of input points  $\mathbf{x} := [x_1, ..., x_n]^T$ .
- 3. To simulate f(x), build the covariance matrix K(x,x), draw from N(0, K(x,x)).

where 
$$\mathbf{K}(\mathbf{x}, \mathbf{x})[i, j] = k(x_i, x_j)$$
.

Different covariance functions, different function spaces.

**Which function space?** You get choose by chossing/building the covariance function k.

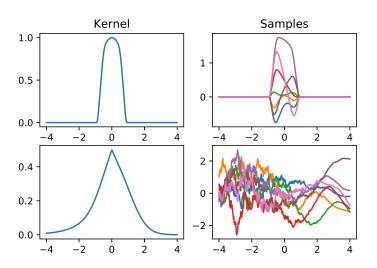
Which function space? Regularity depends on k.



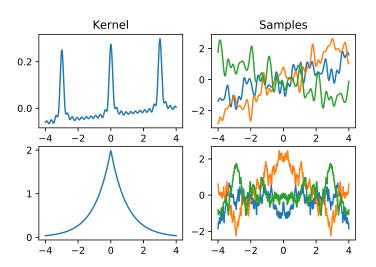
**Algebra of covariance functions** Can combine covariance functions to encode other characteristics:

- $\cdot \alpha k$  is a covariance function.
- $\cdot k(\phi(x), \phi(x'))$  is a covariance function.
- $k_1 \times k_2$  is a covariance function like an AND covariance function, high vals if both are.
- $k_1 + k_2$  is a covariance function like an OR covariance function, high vals if one is.

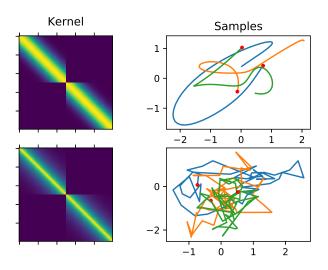
**Algebra of covariance functions** Can combine covariance functions to encode other characteristics:  $k_1 \times k_2$ 



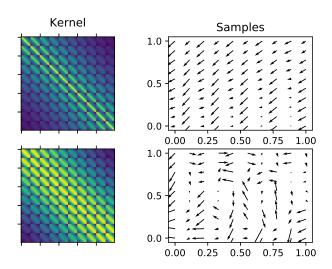
**Algebra of covariance functions** Can combine covariance functions to encode other characteristics:  $k_1 + k_2$ 



**A process**  $R \to R^2$  Is a distribution over doodles.



**A process**  $R^2 \rightarrow R^2$  Is a distribution over vector fields.



In fact you can encode more:

#### Div-Free:

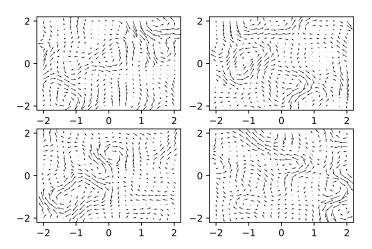
$$\nabla f = 0$$

i.e. Processes that satisfy a linear contraint.

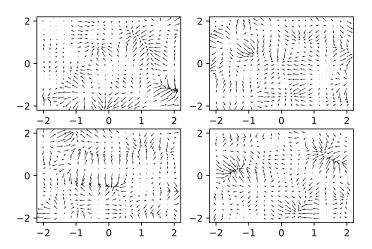
$$\mathcal{L}f = 0 \tag{1}$$

(Idea: Find an operator  $\mathcal G$  such that  $\mathcal L\mathcal G=0$ . Then,  $\mathcal Gf$  satisfies (1).)

**Div-Free**: Samples from a div-free **GP**.



Curl-Free: Samples from a curl-free GP.



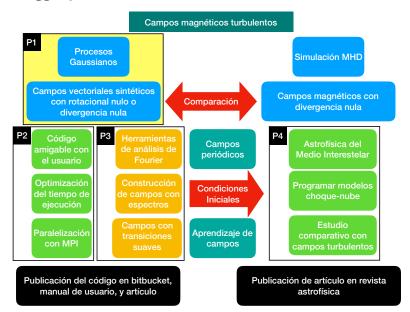
#### So far:

- · Distributions over function spaces.
- Mathematically Expressive: Can encode regularity, symmetries, changes, support.

#### In progress:

- · Computationally NOT Expressive: Autodiff code in progress.
- Turbulence NOT clear: Studying the relationship between kernels and energy spectral decay.

#### A bigger picture:



#### A bigger picture: (In alphabetical order)

- · Computer Science, Math, Physics.
- · Australia, Ecuador, Germany.

#### Some refferences:

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