In[75]:= Clear["Global`*"]

Defining the variables

G: graviational constant

 M_{gc} : GC mass

M_{BH}: BH mass

 μ : M_{BH} / M_{gc}

r: GC - centric radius

a: Plummer scale length

x: r/a

In[76]:= \$Assumptions = $\{G > 0, M_{gc} > 0, \mu > 0, r > 0, a > 0, x > 0\}$

Out[76]= $\left\{ G > 0, M_{gc} > 0, \mu > 0, r > 0, a > 0, x > 0 \right\}$

GC (Plummer)

In[77]:= PhiPlummer[x_] :=
$$-\frac{G * M_{gc}}{a \sqrt{1 + x^2}}$$
;

PhiPlummer[x]

Out[78]=
$$-\frac{G M_{gc}}{a \sqrt{1 + x^2}}$$

In[79]:= PlummerDensity[x_] :=
$$\frac{3 \text{ M}_{gc}}{4 \text{ Pi a}^3} (1 + x^2)^{-5/2}$$
;

PlummerDensity[x]

Out[80]=
$$\frac{3 M_{gc}}{4 a^3 \pi (1 + x^2)^{5/2}}$$

BH (Kepler)

The black hole is assumed to be at the center of the GC.

Therefore, its potential is given by:

In[81]:= PhiKepler[x_] :=
$$-\frac{G * \mu * M_{gc}}{a * x}$$

PhiKepler[x]

Out[82]=
$$-\frac{G \mu M_{gc}}{a x}$$

Velocity Dispersion

Create the potential

$$\text{Out[84]= } - \frac{\text{G M}_{gc}}{\text{a } \sqrt{1 + x^2}} - \frac{\text{G } \mu \text{ M}_{gc}}{\text{a } x}$$

$$\text{Out[86]=} \ \ \frac{ \ \, G \ x \ M_{gc} }{ \ \, a \ \left(1 + x^2 \right)^{3/2} } + \frac{ \ \, G \ \mu \ M_{gc} }{ \ \, a \ x^2 }$$

The number density is given only be the Plummer potential (except directly at 0)

$$In[87] = v[x_] := PlummerDensity[x] / M_{gc}$$

Out[88]=
$$\frac{3}{4 \ a^3 \ \pi \ \left(1 + x^2\right)^{5/2}}$$

Actually do it

$$ln[89]:= sig2 = \frac{1}{v[x]} Integrate[v[xp] * dPhi[xp], \{xp, x, Infinity\}]$$

$$\begin{array}{l} \text{Out[89]=} \;\; -\frac{1}{6\;a\;x\;\left(1+x^2\right)} G\;\left(\!\!\!-6\;\mu-36\;x^2\;\mu-70\;x^4\;\mu-56\;x^6\;\mu-16\;x^8\;\mu+16\;x^3\;\sqrt{1+x^2}\;\mu+48\;x^5\;\sqrt{1+x^2}\;\mu+16\;x^7\;\sqrt{1+x^2}\;\mu+x\;\sqrt{1+x^2}\;\left(-1+16\;\mu\right)\;\right) \; \text{M}_{gc} \end{array}$$