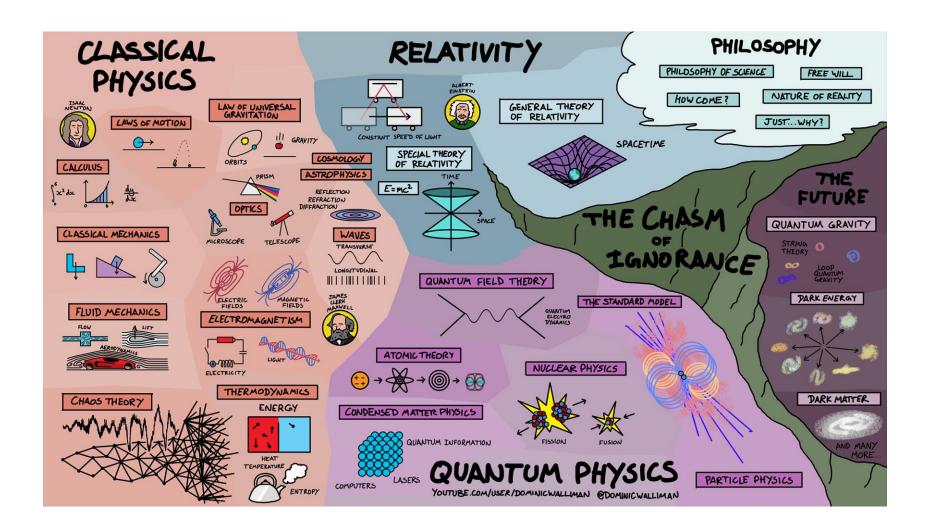


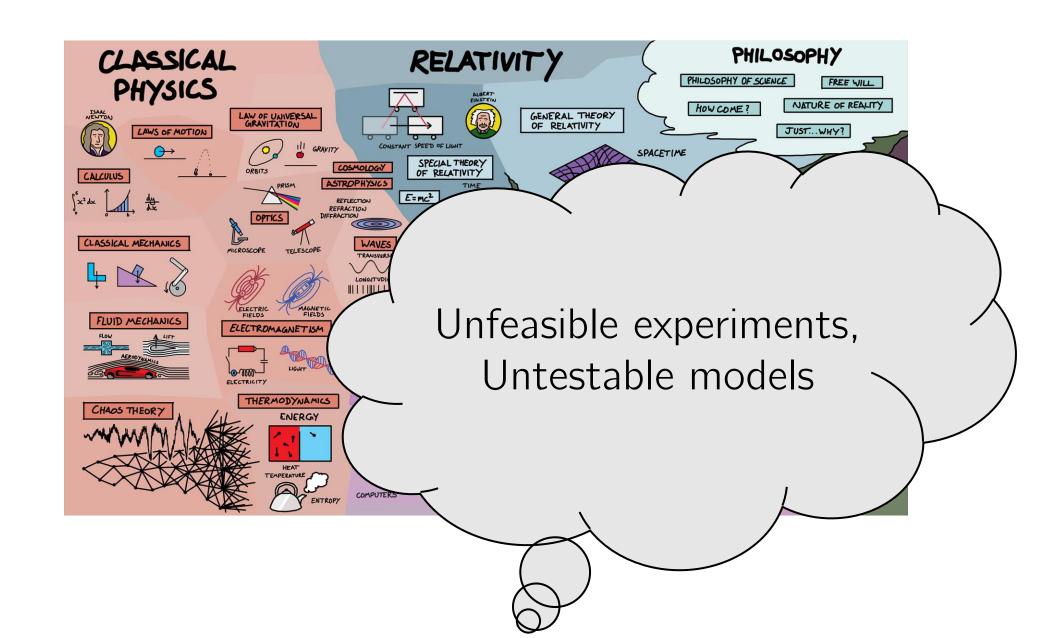
# Introduction to Paraxial Fluids of Light - 1

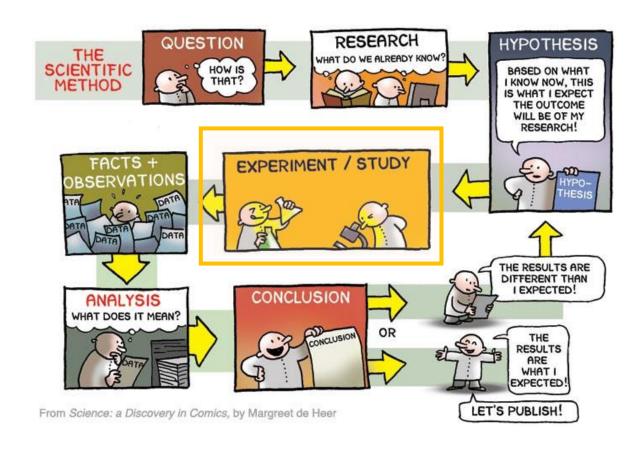
Nuno A. Silva, Tiago D. Ferreira INESC TEC and University of Porto, Portugal

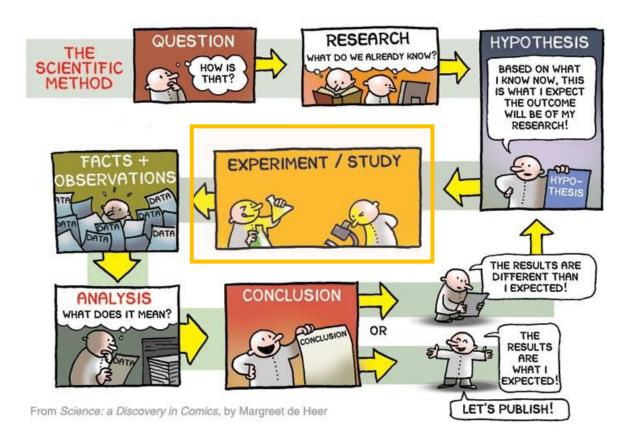
# 1//Analogue and Quantum Simulations

My Usual Approach

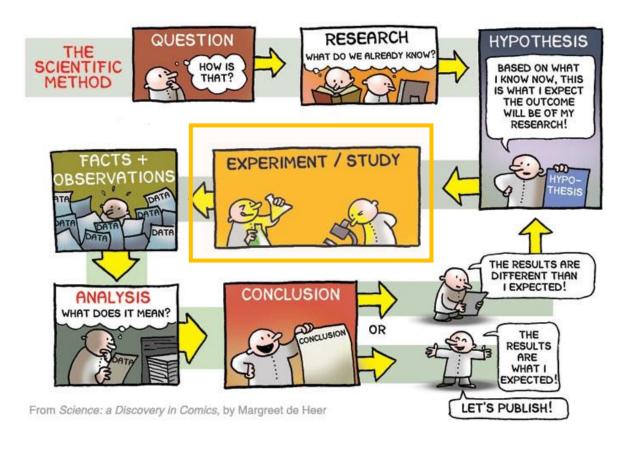






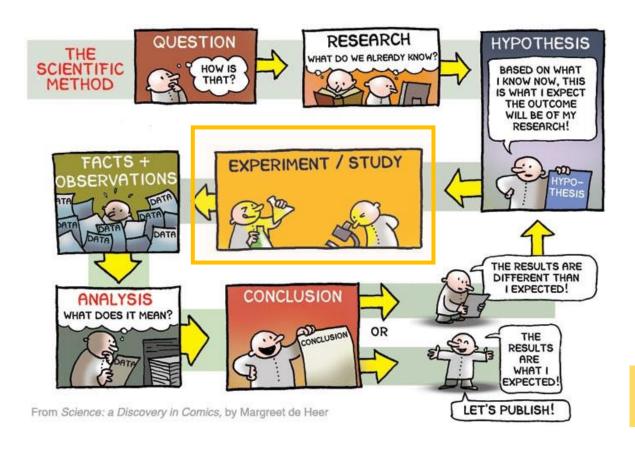


Incompleteness of our research without experiment



Incompleteness of our research without experiment

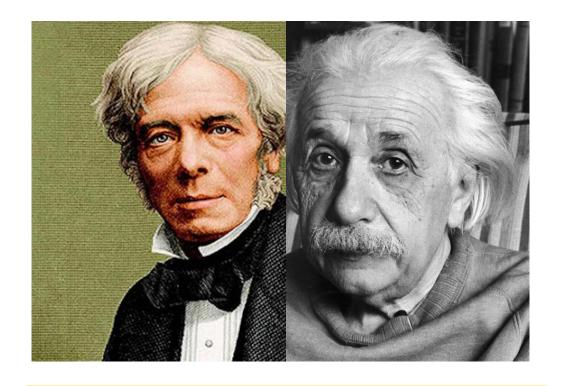
Recreate the physics with analogue systems allow to bypass this challenge



Incompleteness of our research without experiment

Recreate the physics with analogue systems allow to bypass this challenge

**Simplistic and Naive** 

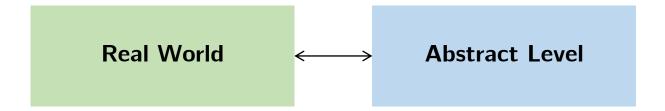


False Dychotomy

experimental vs theoretical physics

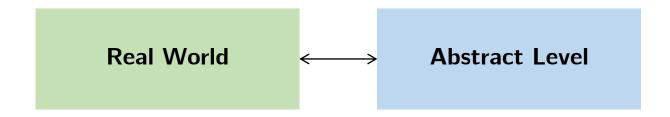
# A better perspective

# The Role of a Physicist



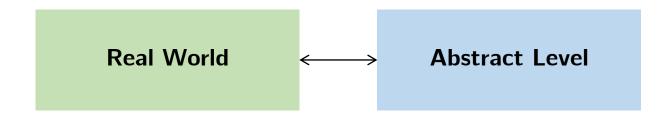
# The Role of a Physicist

Explore how nature works, translating dynamics and phenomena into **mathematical models** 

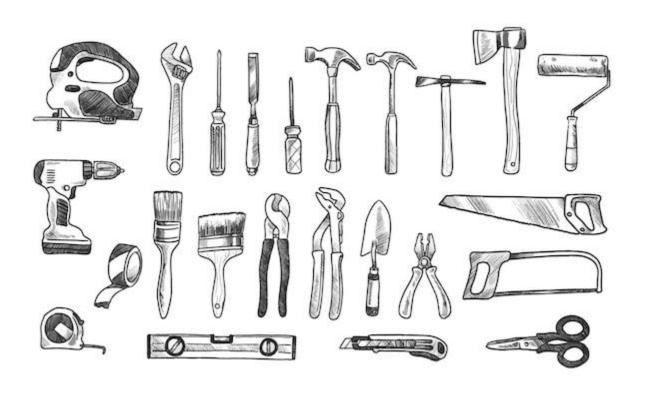


# The Role of a Physicist

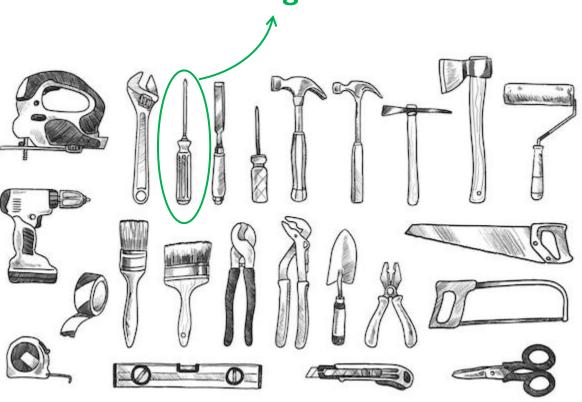
Explore how nature works, translating dynamics and phenomena into **mathematical models** 



How do we approach a complex problem?



#### **Analogies**



# The Role of Analogies

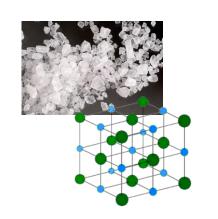
A mental construction to simplify understanding of complex situations by bridging two well-known concepts

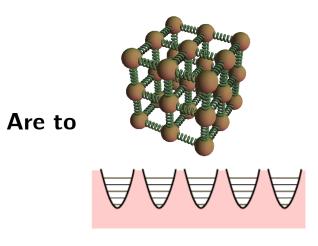


# The Role of Analogies

Provide simpler approaches to complex problems

ex. Harmonic Oscillator





# **Building Models**

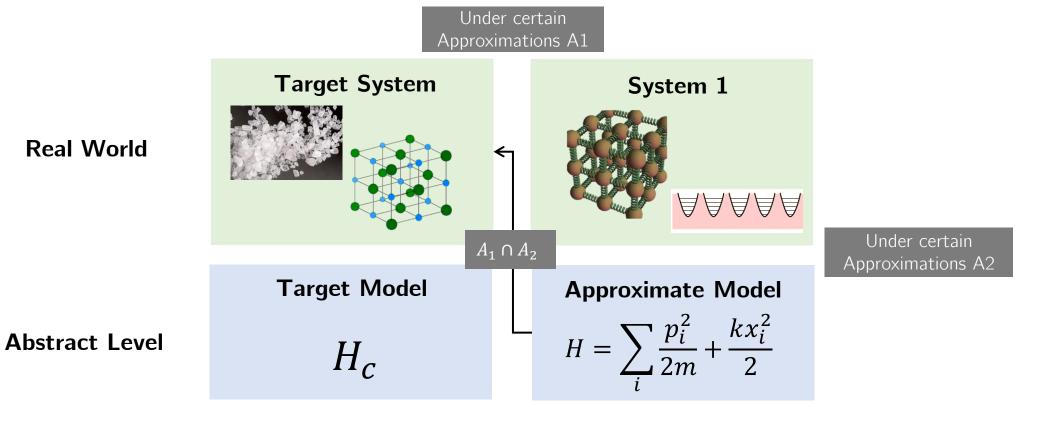
Analogies play an important role in building models, bridging abstract to realworld

**Target System Target Model Abstract Level**  $H_{c}$ 

**Real World** 

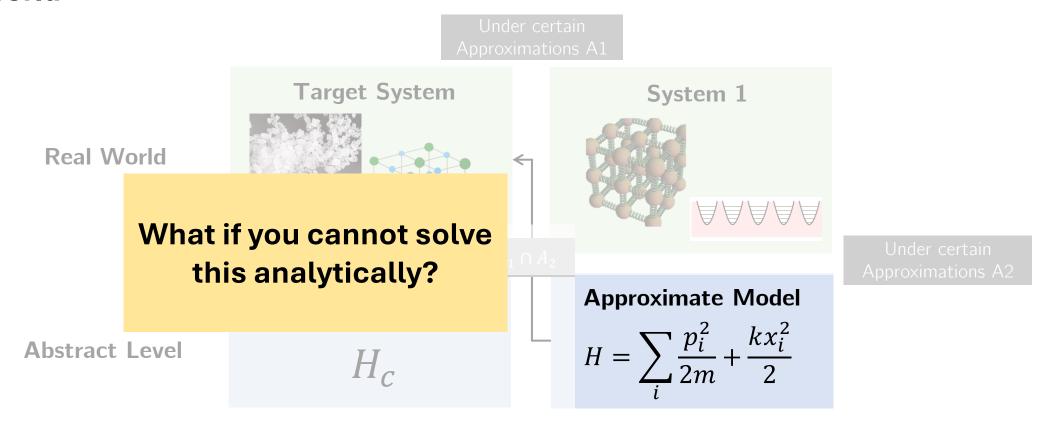
# **Building Models**

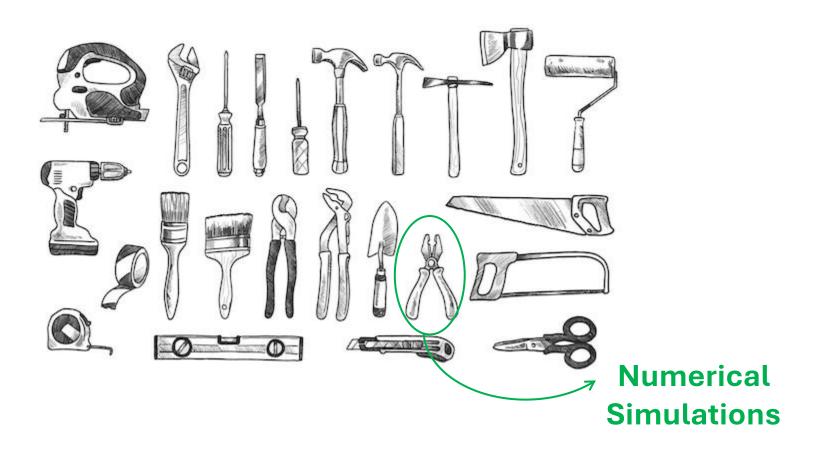
Analogies play an important role in building models, bridging abstract to real-world

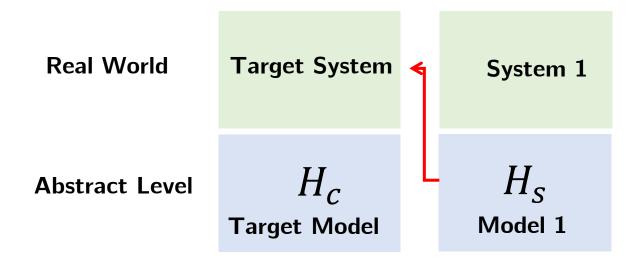


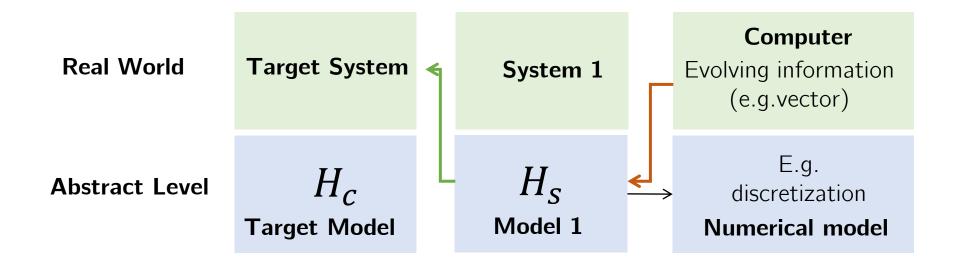
# **Building Models**

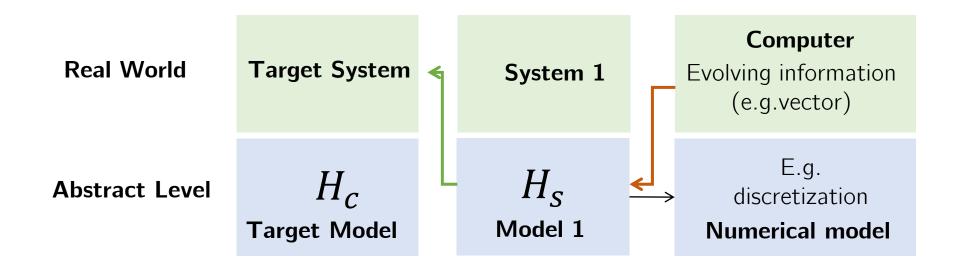
Analogies play an important role in building models, bridging **abstract to real-world** 





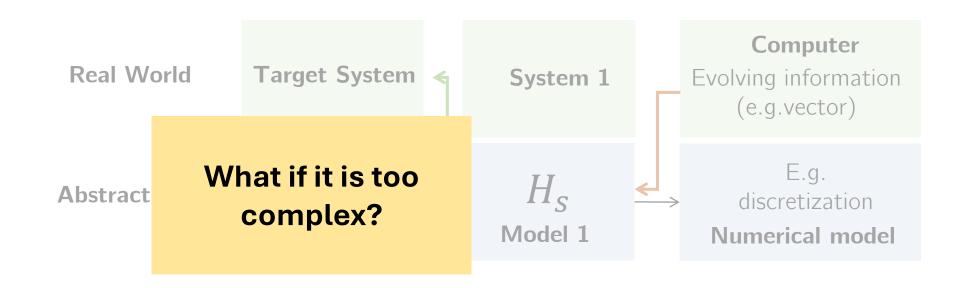






#### **Numerical Simulations**

**Simulation:** utilize a system to provide solutions for a model – from **real-world to abstract level** 



#### **Numerical Simulations**

**Simulation:** utilize a system to provide solutions for a model – from **real-world to abstract level** 



Your PC ran into a problem and needs to restart. We're just collecting some error info, and then we'll restart for you.

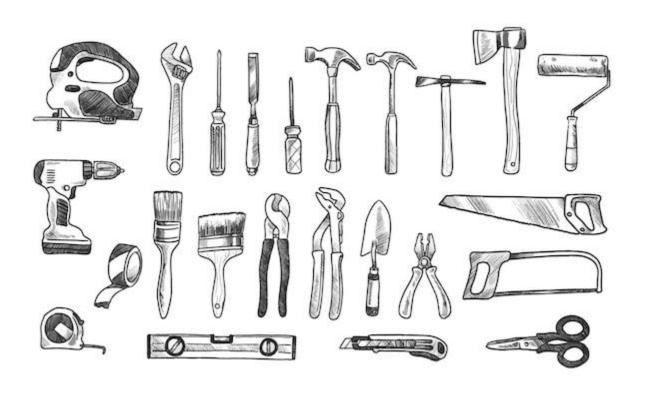
#### 20% complete

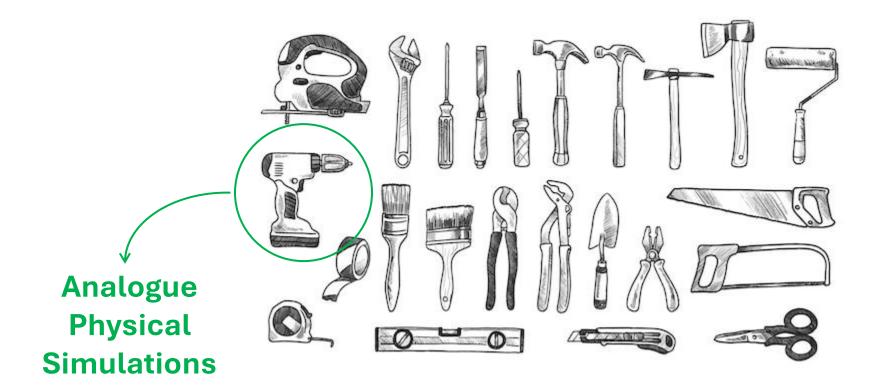


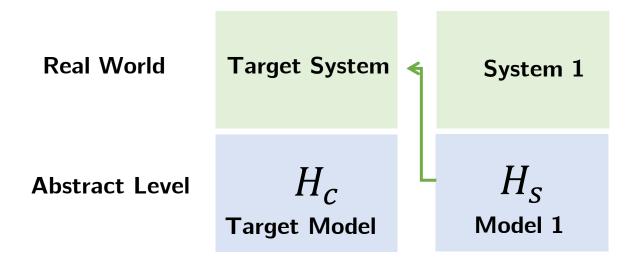
For more information about this issue and possible fixes, visit https://www.windows.com/stopcode

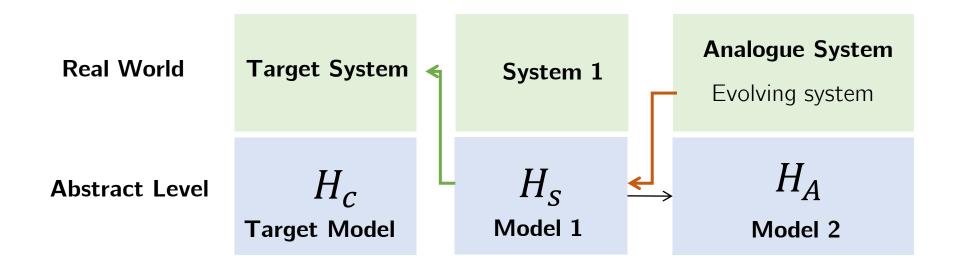
If you call a support person, give them this info:

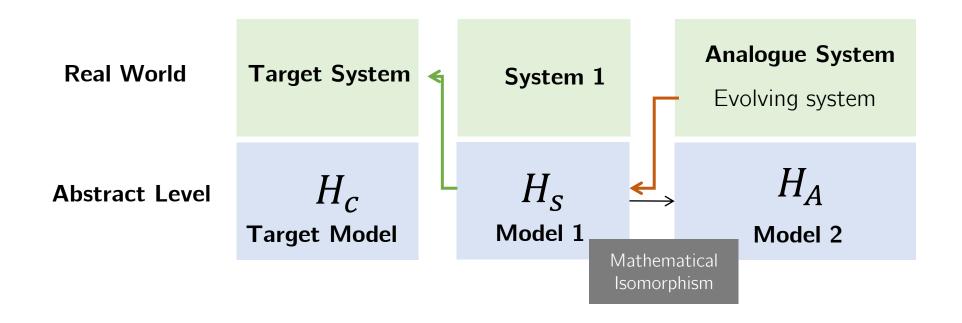
Stop code: CRITICAL PROCESS\_DIED

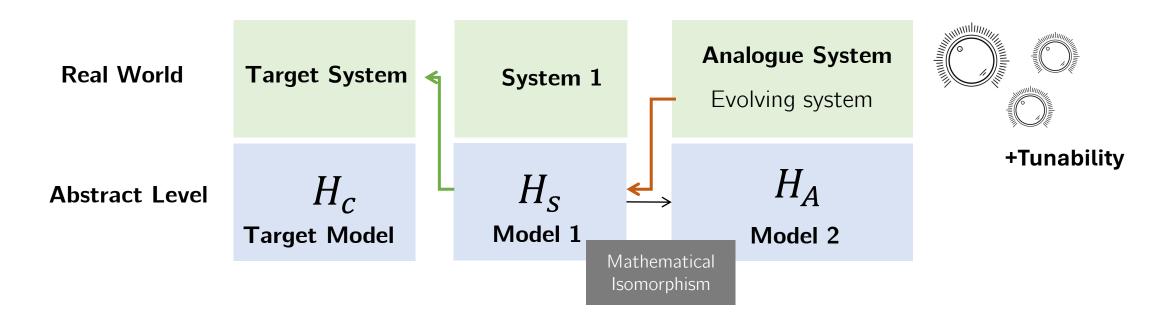




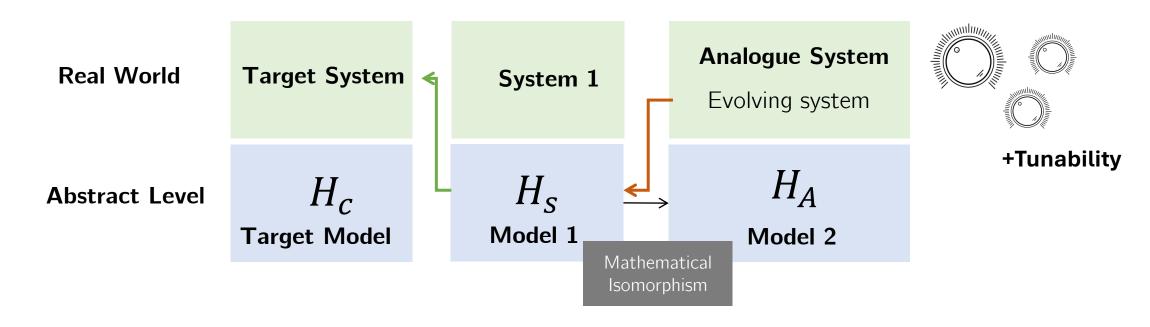






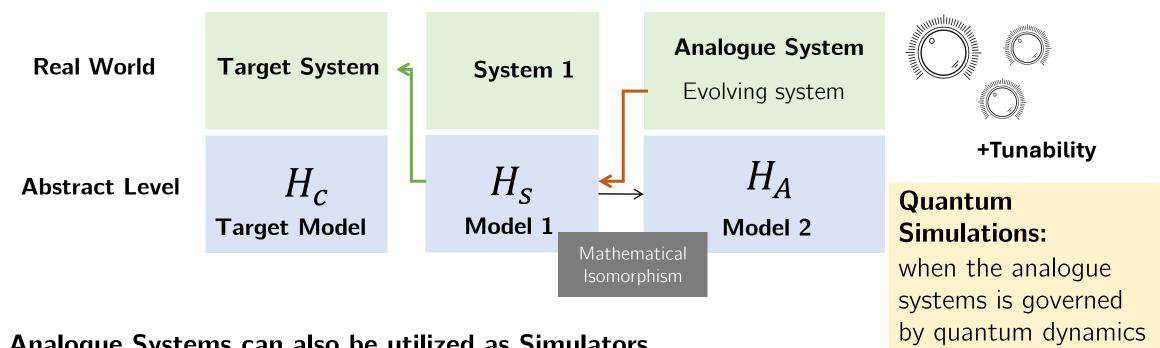


Analogue Systems can also be utilized as Simulators



#### **Analogue Systems can also be utilized as Simulators**

- +May Circumvent issues such as discretization, computer hardware limitation
- +May allow faster simulation times for problems of higher complexity



#### **Analogue Systems can also be utilized as Simulators**

- +May Circumvent issues such as discretization, computer hardware limitation
- +May allow faster simulation times for problems of higher complexity

# **Key takeaways**

1

**Tunable analogue quantum simulators** may speed research of complex systems

#### Macroscopic Quantum Phenomena

**Quantum fluid:** a generic term for a system that displays quantum dynamics at the macroscopic level

#### Macroscopic Quantum Phenomena

**Quantum fluid:** a generic term for a system that displays quantum dynamics at the macroscopic level

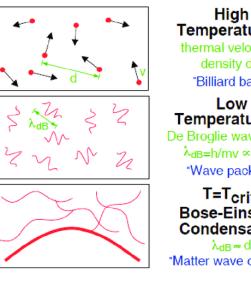
E.g. a Bose-Einstein Condensate (BEC)







Nobel Prize in 2001 (Cornell, Ketterle, Wieman)



**Temperature T:** thermal velocity v density d<sup>-3</sup> "Billiard balls" Low Temperature T: De Broglie wavelength  $\lambda_{dB}=h/mv \propto T^{-1/2}$ "Wave packets" T=T<sub>crit</sub>:

**Bose-Einstein** Condensation  $\lambda_{dB} \approx d$ "Matter wave overlap"

# Bose -Einstein Condensation

 $\hat{\psi}(\mathbf{r},t)$ 

$$\hat{\psi}(\mathbf{r},t)$$

$$\frac{\partial}{\partial t}\hat{\psi}(\mathbf{r},t) = -\frac{i}{\hbar}\left[\hat{\psi}(\mathbf{r},t),\hat{H}\right]$$
 Heisenberg Equation

$$\hat{\psi}(\mathbf{r},t)$$

$$\frac{\partial}{\partial t}\hat{\psi}(\mathbf{r},t) = -\frac{i}{\hbar}\left[\hat{\psi}(\mathbf{r},t),\hat{H}\right]$$
 Heisenberg Equation

$$\hat{H} = \int \hat{\psi}^{\dagger}(\mathbf{r}, t) \left[ \frac{\mathbf{p}^2}{2M} + V_{\text{ext}}(\mathbf{r}, t) \right] \hat{\psi}(\mathbf{r}, t) d\mathbf{r} + \hat{H}_{\text{int}}.$$
 Hamiltonian

$$\hat{\psi}(\mathbf{r},t)$$

$$\frac{\partial}{\partial t}\hat{\psi}(\mathbf{r},t) = -\frac{i}{\hbar}\left[\hat{\psi}(\mathbf{r},t),\hat{H}\right]$$
 Heisenberg Equation

$$\hat{H} = \int \hat{\psi}^{\dagger}(\mathbf{r}, t) \left[ \frac{\mathbf{p}^2}{2M} + V_{\text{ext}}(\mathbf{r}, t) \right] \hat{\psi}(\mathbf{r}, t) d\mathbf{r} + \hat{H}_{\text{int}}.$$
 Hamiltonian

$$\hat{H}_{\text{int}} = \frac{1}{2} \int d\mathbf{r} \int d\mathbf{r}' \hat{\psi}^{\dagger}(\mathbf{r}, t) \hat{\psi}^{\dagger}(\mathbf{r}', t) V(\mathbf{r} - \mathbf{r}') \hat{\psi}(\mathbf{r}, t) \hat{\psi}(\mathbf{r}', t),$$
 Interaction Hamiltonian

#### Mean Field Theory

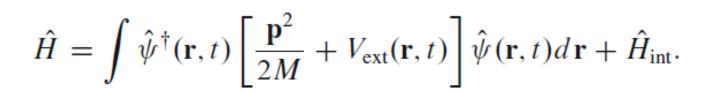
$$\hat{\psi}(\mathbf{r},t) = \Phi(\mathbf{r},t) + \delta \hat{\psi}(\mathbf{r},t), \quad \Phi(\mathbf{r},t) \equiv \langle \hat{\psi}(\mathbf{r},t) \rangle,$$

$$n_0(\mathbf{r},t) = |\Phi(\mathbf{r},t)|^2$$

#### Mean Field Theory

$$\hat{\psi}(\mathbf{r},t) = \Phi(\mathbf{r},t) + \delta \hat{\psi}(\mathbf{r},t), \quad \Phi(\mathbf{r},t) \equiv \langle \hat{\psi}(\mathbf{r},t) \rangle,$$

$$n_0(\mathbf{r},t) = |\Phi(\mathbf{r},t)|^2$$



$$\hat{H}_{\text{int}} = \frac{1}{2} \int d\mathbf{r} \int d\mathbf{r}' \hat{\psi}^{\dagger}(\mathbf{r}, t) \hat{\psi}^{\dagger}(\mathbf{r}', t) V(\mathbf{r} - \mathbf{r}') \hat{\psi}(\mathbf{r}, t) \hat{\psi}(\mathbf{r}', t),$$

#### Mean Field Theory

$$\hat{\psi}(\mathbf{r},t) = \Phi(\mathbf{r},t) + \delta \hat{\psi}(\mathbf{r},t), \quad \Phi(\mathbf{r},t) \equiv \langle \hat{\psi}(\mathbf{r},t) \rangle,$$

$$n_0(\mathbf{r},t) = |\Phi(\mathbf{r},t)|^2$$

$$\frac{\partial}{\partial t}\Phi(\mathbf{r},t) = -\frac{i}{\hbar} \left[ \frac{\mathbf{p}^2}{2M} + V_{\text{ext}}(\mathbf{r},t) + \int \Phi^{\dagger}(\mathbf{r}',t)V(\mathbf{r} - \mathbf{r}')\Phi(\mathbf{r},t)d\mathbf{r}' \right] \Phi(\mathbf{r},t).$$

#### Gross-Pitaevskii equation

$$\frac{\partial}{\partial t}\Phi(\mathbf{r},t) = -\frac{i}{\hbar} \left[ \frac{\mathbf{p}^2}{2M} + V_{\text{ext}}(\mathbf{r},t) + \int \Phi^{\dagger}(\mathbf{r}',t)V(\mathbf{r} - \mathbf{r}')\Phi(\mathbf{r},t)d\mathbf{r}' \right] \Phi(\mathbf{r},t).$$

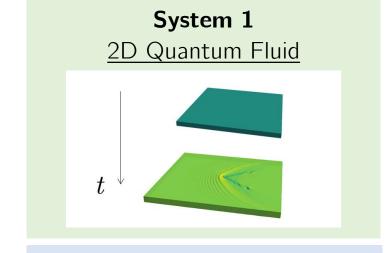
Binary Collisions only  $V(\mathbf{r} - \mathbf{r}') = g \, \delta(\mathbf{r} - \mathbf{r}')$ ,  $g = 4\pi \hbar^2 \frac{a}{M}$ .

$$i\hbar \frac{\partial}{\partial t} \Phi(\mathbf{r}, t) = H\Phi(\mathbf{r}, t) , \quad H \equiv \left[ -\frac{\hbar^2 \nabla^2}{2M} + V_{\text{ext}}(\mathbf{r}, t) + g |\Phi(\mathbf{r}, t)|^2 \right].$$

Gross-Pitaevskii equation (GPE)
Nonlinear Schrödinger Equation (NLSE)

# **Analogue Simulation**

**Real World** 



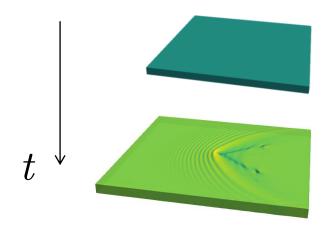
**Abstract Level** 

$$i\hbar\partial_t\Psi+rac{\hbar^2}{2m}
abla^2\Psi+g\left|\Psi
ight|^2\Psi+V_{ext}\Psi=0$$
2D Gross-Pittaevskii



# An analogue quantum fluid

#### Bose Einstein Condensate



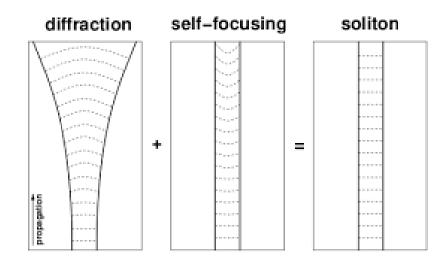
$$i\hbar\partial_t\Psi + \frac{\hbar^2}{2m}\nabla^2\Psi + g|\Psi|^2\Psi + V_{ext}\Psi = 0$$



Nonlinear Schrödinger Equation (NLSE) is an interesting nonlinear playground

# Nonlinear Schrödinger Equation (NLSE) is an interesting nonlinear playground

e.g. Solitons – self-localized wave solutions



Nonlinear Schrödinger Equation (NLSE) is an interesting nonlinear playground and universal model in nonlinear physics appearing in:

Quantum Physics;

Fluid Dynamics;

Plasma Physics;

And so on...

Nonlinear Schrödinger Equation (NLSE) is an interesting nonlinear playground and universal model in nonlinear physics appearing in:

Quantum Physics; Fluid Dynamics; Plasma Physics; And so on...

#### And nonlinear optics

It can be argued that NLO have been around since 1931 (two-photon absorption)

It can be argued that NLO have been around since 1931 (two-photon absorption)

But only after the deployment of the first laser (Maiman 1960) it caught the interest

It can be argued that NLO have been around since 1931 (two-photon absorption)

But only after the deployment of the first laser (Maiman 1960) it caught the interest

Second-harmonic generation (Franken 1961)

It can be argued that NLO have been around since 1931 (two-photon absorption)

But only after the deployment of the first laser (Maiman 1960) it caught the interest

Second-harmonic generation (Franken 1961)

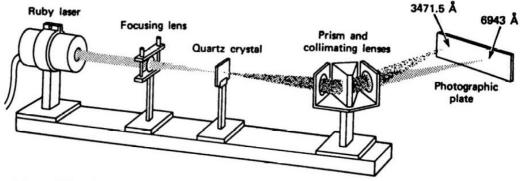


Figure 12.1. Arrangement used in the first experimental demonstration of second-harmonic generation [1]. A ruby-laser beam at  $\lambda = 0.694~\mu m$  is focused on a quartz crystal, causing the generation of a (weak) beam at  $\frac{1}{2}\lambda = 0.347~\mu m$ . The two beams are then separated by a prism and detected on a photographic plate.

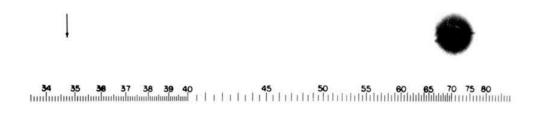


FIG. 1. A direct reproduction of the first plate in which there was an indication of second harmonic. The wavelength scale is in units of 100 A. The arrow at 3472 A indicates the small but dense image produced by the second harmonic. The image of the primary beam at 6943 A is very large due to halation.

### **Maxwell Equations**

$$abla imes oldsymbol{E} egin{array}{lll}
abla imes oldsymbol{E} &=& -rac{\partial oldsymbol{B}}{\partial t} \\
abla imes oldsymbol{D} &=& \sigma_{\mathrm{ext}} \\
abla imes oldsymbol{H} &=& oldsymbol{J}_{\mathrm{ext}} + rac{\partial oldsymbol{D}}{\partial t} \\
abla imes oldsymbol{B} &=& 0,
onumber \end{array}$$

# **Maxwell Equations**

$$abla imes oldsymbol{E} egin{array}{lll}
abla imes oldsymbol{E} &=& -rac{\partial oldsymbol{B}}{\partial t} \\
abla imes oldsymbol{D} &=& \sigma_{\mathrm{ext}} \\
abla imes oldsymbol{H} &=& oldsymbol{J}_{\mathrm{ext}} + rac{\partial oldsymbol{D}}{\partial t} \\
abla imes oldsymbol{B} &=& 0,
onumber \end{array}$$

(absence of free charges or sources)

$$\nabla^2 \mathbf{E} - \frac{1}{c^2} \frac{\partial^2 \mathbf{E}}{\partial t^2} = \frac{1}{\epsilon_0 c^2} \frac{\partial^2 \mathbf{P}}{\partial t^2}$$

# **Maxwell Equations**

$$abla imes oldsymbol{E} egin{array}{lll}
abla imes oldsymbol{E} &=& -rac{\partial oldsymbol{B}}{\partial t} \\
abla imes oldsymbol{D} &=& \sigma_{\mathrm{ext}} \\
abla imes oldsymbol{H} &=& oldsymbol{J}_{\mathrm{ext}} + rac{\partial oldsymbol{D}}{\partial t} \\
abla imes oldsymbol{B} &=& 0,
onumber \end{array}$$

$$\nabla^2 \mathbf{E} - \frac{1}{c^2} \frac{\partial^2 \mathbf{E}}{\partial t^2} = \frac{1}{\epsilon_0 c^2} \frac{\partial^2 \mathbf{P}}{\partial t^2}$$

Wave equation

# Wave equation

$$\nabla^2 \mathbf{E} - \frac{1}{c^2} \frac{\partial^2 \mathbf{E}}{\partial t^2} = \frac{1}{\epsilon_0 c^2} \frac{\partial^2 \mathbf{P}}{\partial t^2}$$

#### **Polarization**

### Wave equation

$$\nabla^2 \mathbf{E} - \frac{1}{c^2} \frac{\partial^2 \mathbf{E}}{\partial t^2} = \frac{1}{\epsilon_0 c^2} \frac{\partial^2 \mathbf{P}}{\partial t^2}$$

#### **Polarization**

$$P(\boldsymbol{r},t) = \epsilon_0 \, \boldsymbol{\chi} \left[ \boldsymbol{E}(\boldsymbol{r},t) \right] \boldsymbol{E}(\boldsymbol{r},t)$$

$$= \underbrace{\epsilon_0 \, \boldsymbol{\chi}^{(1)} \cdot \boldsymbol{E}(\boldsymbol{r},t) + \epsilon_0 \, \boldsymbol{\chi}^{(2)} : \boldsymbol{E}(\boldsymbol{r},t)^2}_{\boldsymbol{P}^{(1)}(\boldsymbol{r},t)} + \underbrace{\epsilon_0 \, \boldsymbol{\chi}^{(3)} : \boldsymbol{E}(\boldsymbol{r},t)^3}_{\boldsymbol{P}^{(3)}(\boldsymbol{r},t)} + \dots$$

### Wave equation

$$\nabla^2 \mathbf{E} - \frac{1}{c^2} \frac{\partial^2 \mathbf{E}}{\partial t^2} = \frac{1}{\epsilon_0 c^2} \frac{\partial^2 \mathbf{P}}{\partial t^2}$$

$$P(\boldsymbol{r},t) = \epsilon_0 \chi \left[ \boldsymbol{E}(\boldsymbol{r},t) \right] \boldsymbol{E}(\boldsymbol{r},t)$$

$$= \underbrace{\epsilon_0 \chi^{(1)} \cdot \boldsymbol{E}(\boldsymbol{r},t) + \epsilon_0 \chi^{(2)} : \boldsymbol{E}(\boldsymbol{r},t)^2}_{\boldsymbol{P}^{(1)}(\boldsymbol{r},t)} + \underbrace{\epsilon_0 \chi^{(3)} : \boldsymbol{E}(\boldsymbol{r},t)^3 + \dots}_{\boldsymbol{P}^{(3)}(\boldsymbol{r},t)} + \underbrace{\epsilon_0 \chi^{(3)} : \boldsymbol{E}(\boldsymbol{r},t)^3 + \dots}_{\boldsymbol{P}^{(3)}(\boldsymbol{r},t)}$$

 $\chi^{(n)}$  Susceptibilities

$$P_j^{(n)}(\boldsymbol{r},t) = \epsilon_0 \sum_{i_1...i_n} \int_{-\infty}^{\infty} \chi_{j i_1...i_n}^{(n)}(\boldsymbol{r}-\boldsymbol{r}_1,\ldots,\boldsymbol{r}-\boldsymbol{r}_n;t-t_1,\ldots,t-t_n) \times$$

$$E_{i_1}(\boldsymbol{r}_1,t_1)\ldots E_{i_n}(\boldsymbol{r}_m,t_n) \,\mathrm{d}\boldsymbol{r}_1\ldots\mathrm{d}\boldsymbol{r}_1 \,\mathrm{d}t_1\ldots\mathrm{d}t_n$$

$$P_{j}^{(n)}(\boldsymbol{r},t) = \epsilon_{0} \sum_{i_{1}...i_{n}} \int_{-\infty}^{\infty} \chi_{j i_{1}...i_{n}}^{(n)}(\boldsymbol{r}-\boldsymbol{r}_{1},\ldots,\boldsymbol{r}-\boldsymbol{r}_{n};t-t_{1},\ldots,t-t_{n}) \times E_{i_{1}}(\boldsymbol{r}_{1},t_{1})\ldots E_{i_{n}}(\boldsymbol{r}_{m},t_{n}) d\mathbf{r}_{1}\ldots d\mathbf{r}_{1} dt_{1}\ldots dt_{n}$$

Tensor that must respect the properties of the media

$$P_{j}^{(n)}(\boldsymbol{r},t) = \epsilon_{0} \sum_{i_{1}...i_{n}} \int_{-\infty}^{\infty} \chi_{j i_{1}...i_{n}}^{(n)}(\boldsymbol{r}-\boldsymbol{r}_{1},\ldots,\boldsymbol{r}-\boldsymbol{r}_{n};t-t_{1},\ldots,t-t_{n}) \times E_{i_{1}}(\boldsymbol{r}_{1},t_{1})\ldots E_{i_{n}}(\boldsymbol{r}_{m},t_{n}) d\mathbf{r}_{1}\ldots d\mathbf{r}_{1} dt_{1}\ldots dt_{n}$$

Tensor that must respect the properties of the media

What does it mean for isotropic media?

### Neumann's principle

coordinate transformations (inversion, mirror image and rotation) **T** of field and polarization vectors **E** and **P** 

$$\begin{pmatrix} E_x' \\ E_y' \\ E_z' \end{pmatrix} = \begin{pmatrix} T_{x'x} & T_{x'y} & T_{x'z} \\ T_{y'x} & T_{y'y} & T_{y'z} \\ T_{z'x} & T_{z'y} & T_{z'z} \end{pmatrix} \begin{pmatrix} E_x \\ E_y \\ E_z \end{pmatrix}$$
(2.47)

$$\mathbf{E}' = \mathbf{T} \cdot \mathbf{E} \tag{2.48}$$

$$\mathbf{E} = \mathbf{T}^{-1} \cdot \mathbf{E}' \tag{2.49}$$

### Neumann's principle

#### employing Einstein's summation convention

$$E_{i} = (T_{ii'})^{T} E_{i'} = T_{i'i} E_{i'}$$
(2.51)

$$P_i = T_{i'i}P_{i'} \tag{2.52}$$

#### relations in the two coordinate systems

$$P_i^{(n)} = \varepsilon_0 \chi_{ij\dots s}^{(n)} E_j \cdots E_s, \qquad (2.53)$$

$$P_{i'}^{(n)} = \varepsilon_0 \chi_{i'j'\ldots s'}^{(n)} E_{j'} \cdots E_{s'}, \qquad (2.54)$$

#### Then

$$T_{i'i}P_i^{(n)} = P_{i'}^{(n)} = \varepsilon_0 T_{i'i} \chi_{ij...s}^{(n)} T_{j'j} E_{j'} T_{k'k} E_{k'} \cdots T_{s's} E_{s'}$$
 (2.55)

$$\chi_{i'j'...s'}^{(n)} = T_{i'i}T_{j'j}\cdots T_{s's}\chi_{ij...s}^{(n)}$$
(2.56)

# Neumann's principle

Example: let's consider inversion  $T_{i'i} = (-1)\delta_{i'i}$ 

susceptibility tensor of the inverted medium

$$\chi_{i'j'...s'}^{(n)} = (-1)^{n+1} \chi_{ij...s}^{(n)}$$
(2.57)

If the medium is invariant under inversion, it follows for *n*=even

$$\chi_{ij\dots s}^{(n)} = (-1)^{n+1} \chi_{ij\dots s}^{(n)} = 0 \tag{2.58}$$

i.e., in an inversion symmetric medium, the susceptibility tensors of even orders vanish

$$P(\boldsymbol{r},t) = \epsilon_0 \boldsymbol{\chi} \left[ \boldsymbol{E}(\boldsymbol{r},t) \right] \boldsymbol{E}(\boldsymbol{r},t)$$

$$= \underbrace{\epsilon_0 \boldsymbol{\chi}^{(1)} \cdot \boldsymbol{E}(\boldsymbol{r},t)}_{\boldsymbol{P}^{(1)}(\boldsymbol{r},t)} + \underbrace{\epsilon_0 \boldsymbol{\chi}^{(2)} : \boldsymbol{E}(\boldsymbol{r},t)^2}_{\boldsymbol{P}^{(2)}(\boldsymbol{r},t)} + \underbrace{\epsilon_0 \boldsymbol{\chi}^{(3)} : \boldsymbol{E}(\boldsymbol{r},t)^3}_{\boldsymbol{P}^{(3)}(\boldsymbol{r},t)} + \dots$$

#### **Polarization**

$$P(\boldsymbol{r},t) = \epsilon_0 \chi \left[ \boldsymbol{E}(\boldsymbol{r},t) \right] \boldsymbol{E}(\boldsymbol{r},t)$$

$$= \underbrace{\epsilon_0 \chi^{(1)} \cdot \boldsymbol{E}(\boldsymbol{r},t)}_{P^{(1)}(\boldsymbol{r},t)} + \underbrace{\epsilon_0 \chi^{(2)} : \boldsymbol{E}(\boldsymbol{r},t)^2}_{P^{(2)}(\boldsymbol{r},t)} + \underbrace{\epsilon_0 \chi^{(3)} : \boldsymbol{E}(\boldsymbol{r},t)^3}_{P^{(3)}(\boldsymbol{r},t)} + \dots$$

**Isotropic media** – Polarization aligned with electric field

No longer a tensor, a number is sufficient

#### **Polarization**

$$P(\boldsymbol{r},t) = \epsilon_0 \, \boldsymbol{\chi} \left[ \boldsymbol{E}(\boldsymbol{r},t) \right] \boldsymbol{E}(\boldsymbol{r},t)$$

$$= \underbrace{\epsilon_0 \, \boldsymbol{\chi}^{(1)} \cdot \boldsymbol{E}(\boldsymbol{r},t)}_{\boldsymbol{P}^{(1)}(\boldsymbol{r},t)} + \underbrace{\epsilon_0 \, \boldsymbol{\chi}^{(2)} : \boldsymbol{E}(\boldsymbol{r},t)^2}_{\boldsymbol{P}^{(2)}(\boldsymbol{r},t)} + \underbrace{\epsilon_0 \, \boldsymbol{\chi}^{(3)} : \boldsymbol{E}(\boldsymbol{r},t)^3}_{\boldsymbol{P}^{(3)}(\boldsymbol{r},t)} + \dots$$

**Isotropic media** – Polarization aligned with electric field

No longer a tensor, a number is sufficient

$$P^{(n)}(\boldsymbol{r},t) = \epsilon_0 \int_{-\infty}^{\infty} \chi^{(n)}(t-t_1,\ldots,t-t_n) \times E(\boldsymbol{r},t_1) \ldots E(\boldsymbol{r},t_n) dt_1 \ldots dt_n.$$

#### **Polarization**

$$P^{(n)}(\boldsymbol{r},t) = \epsilon_0 \int_{-\infty}^{\infty} \chi^{(n)}(t-t_1,\ldots,t-t_n) \times E(\boldsymbol{r},t_1) \ldots E(\boldsymbol{r},t_n) dt_1 \ldots dt_n.$$

$$E(\mathbf{r},t) = \frac{1}{2} \left[ \mathcal{E}(\mathbf{r}) e^{i\omega t} + \mathcal{E}^*(\mathbf{r}) e^{-i\omega t} \right].$$

Third-order polarizations – polarization oscillating at  $+\omega$  four-wave mixing

$$\binom{3}{2} = 3$$

#### Wave equation for a cubic nonlinear media

$$\nabla^2 \mathcal{E}(\mathbf{r}) + \frac{\omega^2}{c^2} \left[ 1 + \chi^{(1)}(\omega) \right] \mathcal{E}(\mathbf{r}) = -\frac{3}{4} \frac{\omega^2}{c^2} \chi^{(3)}(\omega) \left| \mathcal{E}(\mathbf{r}) \right|^2 \mathcal{E}(\mathbf{r}).$$

$$\mathcal{E}(\mathbf{r}_{\perp}, z) = \mathcal{E}_0(\mathbf{r}_{\perp}, z) e^{ik(\omega)z}$$

Physical meaning – small angle deviations from the propagation axis z

$$|\nabla_{\perp}^2 \mathcal{E}_0|/k^2 \sim |\partial_z \mathcal{E}_0|/k \ll 1$$

$$\mathcal{E}(\mathbf{r}_{\perp}, z) = \mathcal{E}_0(\mathbf{r}_{\perp}, z) e^{ik(\omega)z}$$

Physical meaning – small angle deviations from the propagation axis z

$$|\nabla_{\perp}^2 \mathcal{E}_0|/k^2 \sim |\partial_z \mathcal{E}_0|/k \ll 1$$

$$\mathcal{E}(\mathbf{r}_{\perp}, z) = \mathcal{E}_0(\mathbf{r}_{\perp}, z) e^{ik(\omega)z}$$

Physical meaning – small angle deviations from the propagation axis z

$$\left| \frac{\partial^2 E}{\partial z^2} \right| \ll \left| 2ik \frac{\partial E}{\partial z} \right|$$

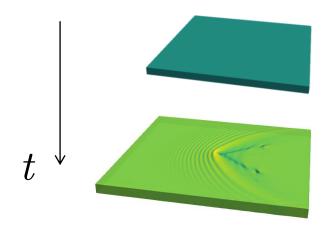
At the end we get (adding an absorption alpha)

$$i \,\partial_z \,\mathcal{E}_0(\mathbf{r}_\perp, z) = \left[ -\frac{1}{2k} \mathbf{\nabla}_\perp^2 - \frac{i\alpha}{2} - \frac{3}{8} \frac{k}{n_0^2} \,\chi^{(3)}(\omega) \left| \mathcal{E}_0\left(\mathbf{r}_\perp, z\right) \right|^2 \right] \mathcal{E}_0(\mathbf{r}_\perp, z),$$

Addition of a linear refractive index (obstacle-like, delta n)

$$i \,\partial_z \,\mathcal{E}_0(\mathbf{r}_\perp, z) = \left[ -\frac{1}{2k} \mathbf{\nabla}_\perp^2 - \frac{i\alpha}{2} - k \,\frac{\delta n(\mathbf{r}_\perp, z)}{n_0} - \frac{3}{8} \frac{k}{n_0^2} \,\chi^{(3)}(\omega) \left| \mathcal{E}_0\left(\mathbf{r}_\perp, z\right) \right|^2 \right] \mathcal{E}_0(\mathbf{r}_\perp, z). \tag{2.10}$$

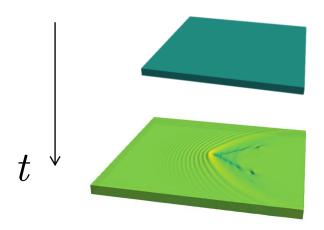
#### Bose Einstein Condensate



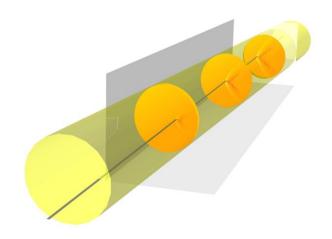
$$i\hbar\partial_t\Psi + \frac{\hbar^2}{2m}\nabla^2\Psi + g|\Psi|^2\Psi + V_{ext}\Psi = 0$$



#### Bose Einstein Condensate

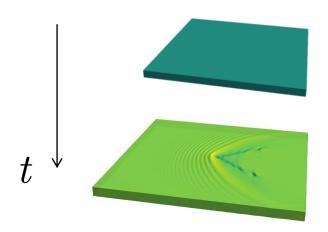


$$i\hbar\partial_t\Psi + \frac{\hbar^2}{2m}\nabla^2\Psi + g|\Psi|^2\Psi + V_{ext}\Psi = 0$$



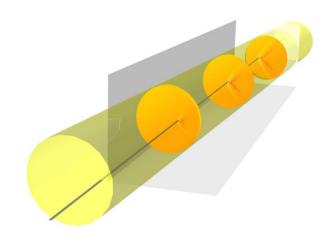
$$i\partial_z \Omega_p + \frac{1}{2} \nabla_\perp^2 \Omega_p + V \Omega_p + G |\Omega_p|^2 \Omega_p = 0$$

#### Bose Einstein Condensate



$$i\hbar\partial_t\Psi + \frac{\hbar^2}{2m}\nabla^2\Psi + g|\Psi|^2\Psi + V_{ext}\Psi = 0$$

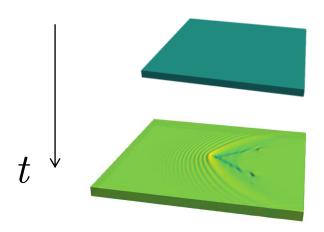
Time $t$
Density $ \Psi ^2$
Interaction via collisions $g$
Atomic mass $m$
Potential (Confinement, Barrier) $V_{ext}$



$$i\partial_z \Omega_p + \frac{1}{2} \nabla_\perp^2 \Omega_p + V \Omega_p + G |\Omega_p|^2 \Omega_p = 0$$

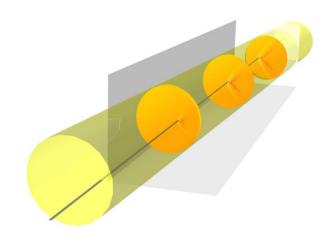
Propagation Distance $z$
Intensity $ \Omega_p ^2$
Interaction via nonlinearity $G$
Effective mass through diffraction $k_p$
Linear refractive index $V$

#### Bose Einstein Condensate



$$i\hbar\partial_t\Psi + \frac{\hbar^2}{2m}\nabla^2\Psi + g|\Psi|^2\Psi + V_{ext}\Psi = 0$$

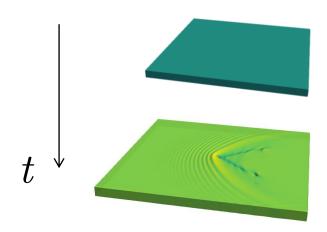
Time $t$
Density $ \Psi ^2$
Interaction via collisions $g$
Atomic mass $m$
Potential (Confinement, Barrier) $V_{ext}$



$$i\partial_z \Omega_p + \frac{1}{2} \nabla_\perp^2 \Omega_p + V \Omega_p + G |\Omega_p|^2 \Omega_p = 0$$

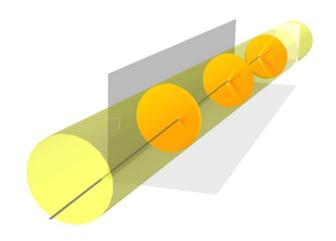
Propagation Distance $z$
Intensity $ \Omega_p ^2$
Interaction via nonlinearity $G$
Effective mass through diffraction $k_p$
Linear refractive index $V$

#### Bose Einstein Condensate



$$i\hbar\partial_t\Psi + \frac{\hbar^2}{2m}\nabla^2\Psi + g|\Psi|^2\Psi + V_{ext}\Psi = 0$$

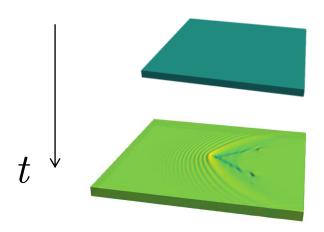
Time $t$
Density $ \Psi ^2$
Interaction via collisions $g$
Atomic mass $m$
Potential (Confinement, Barrier) $V_{ext}$



$$i\partial_z \Omega_p + \frac{1}{2} \nabla_\perp^2 \Omega_p + V \Omega_p + G |\Omega_p|^2 \Omega_p = 0$$

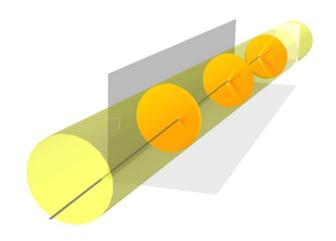
Propagation Distance $z$
Intensity $( \Omega_p ^2)$
Interaction via nonlinearity $G$
Effective mass through diffraction $k_p$
Linear refractive index $V$

#### Bose Einstein Condensate



$$i\hbar\partial_t\Psi + \frac{\hbar^2}{2m}\nabla^2\Psi + g|\Psi|^2\Psi + V_{ext}\Psi = 0$$

Time $t$
Density $ \Psi ^2$
Interaction via collisions $g$
Atomic mass $m$
Potential (Confinement, Barrier) $V_{ext}$



$$i\partial_z \Omega_p + \frac{1}{2} \nabla_\perp^2 \Omega_p + V \Omega_p + G |\Omega_p|^2 \Omega_p = 0$$

Propagation Distance $z$	
Intensity $ \Omega_p ^2$	
Interaction via nonlinearity $G$	
Effective mass through diffraction k	$\overline{\epsilon_p}$
Linear refractive index $V$	

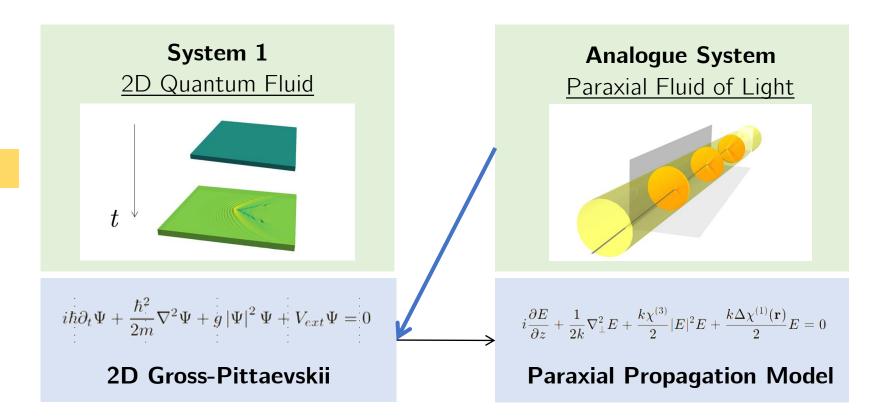
#### **Analogue Quantum Simulation**



**+Tunability** 

**Real World** 

**Abstract Level** 



Light Speed Simulation of a 2D Quantum Fluid under GPE With no discretization

$$i\partial_z E + \frac{1}{2}\nabla_\perp^2 E + V(r_\perp) E + g|E|^2 E = 0$$

Madelung transformation

$$E\left(\mathbf{r}_{\perp},z\right) = \sqrt{\rho\left(\mathbf{r}_{\perp},z\right)}e^{i\phi\left(\mathbf{r}_{\perp},z\right)}$$

Set of Hydrodynamic like equations

$$v = \nabla \phi$$
,

$$\boldsymbol{v} = \nabla \phi, \qquad V_B \equiv \frac{\nabla^2 \sqrt{\rho}}{2\sqrt{\rho}}$$

Set of Hydrodynamic like equations

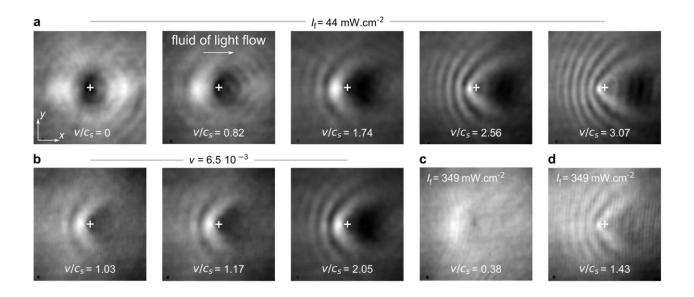
$$v = \nabla \phi$$
,

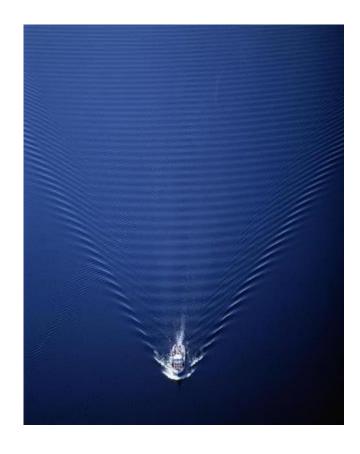
$$oldsymbol{v} = oldsymbol{\nabla} \phi, \qquad V_B \equiv rac{
abla^2 \sqrt{
ho}}{2\sqrt{
ho}}$$

$$\partial_z \rho + \nabla \left( \rho \mathbf{v} \right) = 0$$

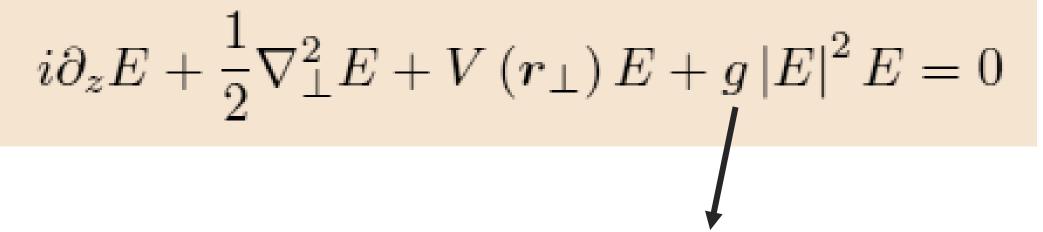
$$\partial_z v + (v \cdot \nabla) v = \nabla (V + V_B + g\rho)$$

Light shall behave like a fluid





# But we said quantum fluid, where is the quantum part?



In this self-interaction term, that emulates BECs

Small excitations spectra

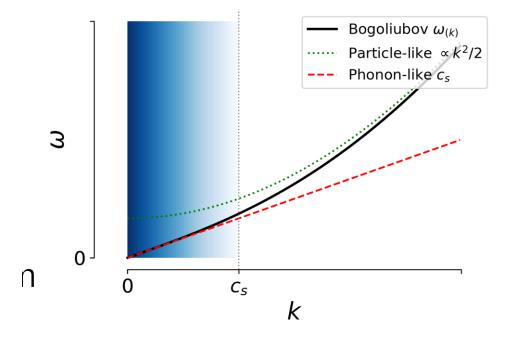
Bogoliubov dispersion relation

$$\omega_{\mathrm{Bog}}(\boldsymbol{k}_{\perp}) = \sqrt{k_{\perp}^2 \left(\frac{k_{\perp}^2}{4} + G|\Omega_p^0|^2\right)}.$$

Small excitations spectra

Bogoliubov dispersion relation

$$\omega_{\text{Bog}}(\mathbf{k}_{\perp}) = \sqrt{k_{\perp}^2 \left(\frac{k_{\perp}^2}{4} + G|\Omega_p^0|^2\right)}.$$

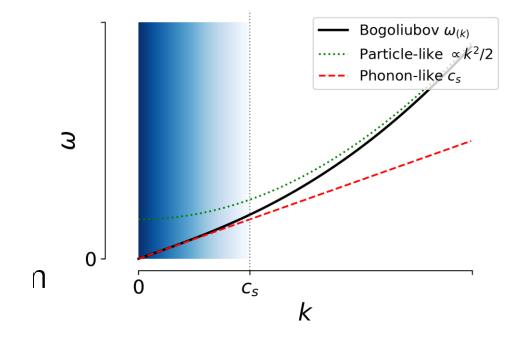


Small excitations spectra

Bogoliubov dispersion relation

$$\omega_{\text{Bog}}(\mathbf{k}_{\perp}) = \sqrt{k_{\perp}^2 \left(\frac{k_{\perp}^2}{4} + G|\Omega_p^0|^2\right)}.$$

Landau Criteria  $\, v < c \,$ 



Small excitations spectra

Bogoliubov dispersion relation

$$\omega_{\text{Bog}}(\mathbf{k}_{\perp}) = \sqrt{k_{\perp}^2 \left(\frac{k_{\perp}^2}{4} + G|\Omega_p^0|^2\right)}.$$

Landau Criteria

$$v < c_s$$

Superfluid properties absence of drag, loss of momentum

