

PHYS-GA-2000: Computational Physics Modeling the Dynamics of Pluto's Upper Atmosphere to Calculate Escape

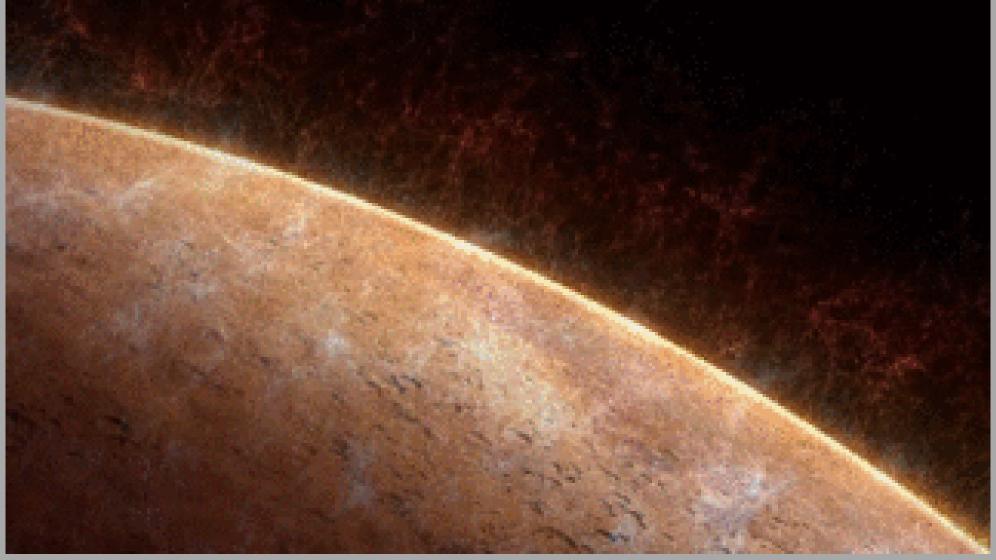
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http://las.colorado.edu/home/maven/2016/10/19/maven-observes-ups-and-downs-of-water-escape-from-mars/

MYU

<u>Motivation</u> → Background → Results



https://phys.org/news/2015-04-hubble-space-telescope-ground.html



https://www.spaceanswers.com/news/hottest-and-most-massive-double-star-found-by-ground-based-telescopes1/

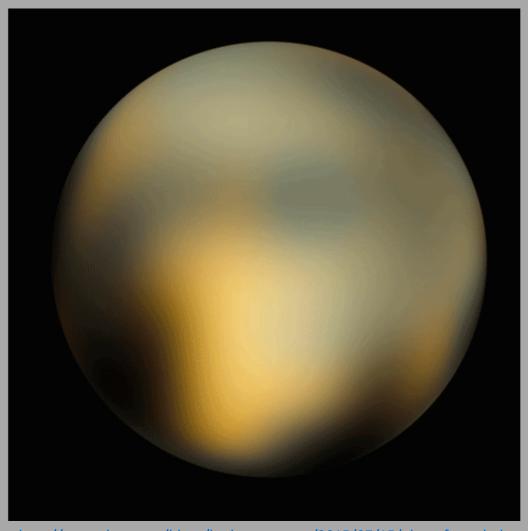


https://nasa.tumblr.com/post/166016245839/chasing-the-shadow-of-neptunes-moon-triton





http://for-all-mankind.tumblr.com/post/159273055401/cassini-prepares-for-final-orbital-grand-finale



http://www.slate.com/blogs/bad_astronomy/2015/07/15/pluto_from_hub_ble_to_new_horizons.html







Factors that can drive escape:

Solar/Magnetospheric Heating

Impacts

Gravity

Surrounding Planets/Moons

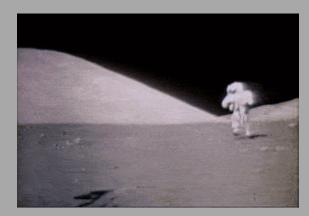


Factors that can drive escape:

• Solar/Magnetospheric Heating

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https://giphy.com/gifs/moon-nasa-3cJucqJLw37vq

Surrounding Planets/Moons*



https://thumbs.gfycat.com/FaithfulWealthyBanta mrooster-max-1mb.gif

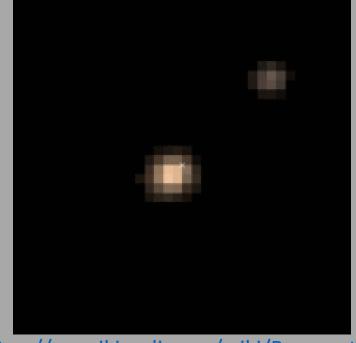


Factors that can drive escape:

Solar/Magnetospheric Heating

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Gravity



https://en.wikipedia.org/wiki/Barycenter

* Hoey, W.A., Yeoh, S.K., Trafton, L.M., Goldstein, D.B. and Varghese, P.L., 2017. Rarefied gas dynamic simulation of transfer and escape in the Pluto–Charon system. *Icarus*, *287*, pp.87-102.

Surrounding Planets/Moons*



Solar Heating:

• Radiation (UV, EUV), Cycle

Composition (N2, CH4)

Atmospheric Thickness

• Distance (~30-50 AU)

Gravity:

Planet Mass (Pluto)

Molecular Mass (N2, CH4)

Radial Temperature



Important Fluid Properties:

Altitude, r [km]

Number Density, n(r, T) [#/m3]

Temperature, T(r, n) [K]

Bulk Flow Velocity, u(r, n) [m/s]

Conservation of Mass

Conservation of Momentum

Conservation of Energy

Equation of State



Conservation of Mass

$$4\pi r^2 n(r)u(r) = \phi$$

Conservation of Momentum

$$nm\frac{d}{dr}\left(\frac{1}{2}u(r)^2\right) + \frac{dp(r)}{dr} = nm\frac{d}{dr}\left(\Phi_G(r)\right)$$

Conservation of Energy
$$\frac{d}{dr} \left(\phi \left(C_P T + \frac{1}{2} m u(r)^2 - \Phi_G(r) \right) - 4 \pi r^2 \kappa \left(T(r) \right) \frac{dT(r)}{dr} \right) = 4 \pi r^2 q(r)$$

Equation of State

$$p = n k_B T$$



Number Density

$$n(r) = n_0 \left(\frac{T_0}{T(r)}\right) \exp\left[-\int \left(\frac{\lambda(r)}{r} + \frac{m\frac{d}{dr}(u(r)^2)}{2k_B T(r)}\right) dr\right]$$

Temperature
$$\frac{dT(r)}{dr} = \left(\frac{1}{4 \pi r^2 \kappa (T(r))}\right) \left[-\left(\phi_E + Q(r)\right) + \phi \left(C_P T(r) + \frac{1}{2} m u(r)^2 - \Phi_G(r)\right)\right]$$



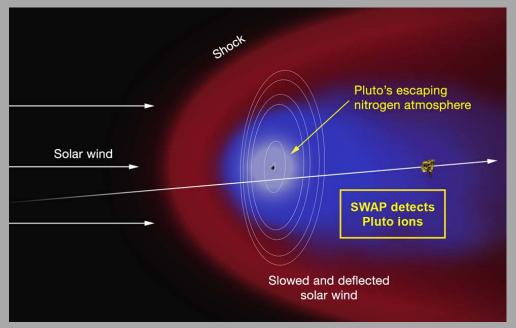
Total Heating Rate

$$Q(r) = 4\pi \int_{r_0} r^2 q(r) dr$$

Local Heating Rate
$$q(r) = \frac{dF(r)}{dr} = \chi \varepsilon \sigma n(r) F(r) \exp \left(-\frac{\tau(r)}{\mu}\right)$$

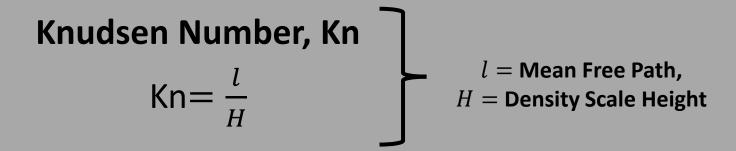
Vertical Optical Depth

$$\tau(r) = \int_{r_{\text{max}}} \chi \, \sigma \, n(r) \, dr$$



https://www.nasa.gov/nh/pluto-wags-its-tail





Exobase, r(Kn~1)

Above Kn~1, Fluid Equations FAIL;
Can still receive valuable information at Exobase

Jeans Parameter
$$\lambda(r) = \frac{GMm}{rk_BT(r)}$$

 $\lambda_x \leq 2$: Hydrodynamic Escape, $\lambda_x > 10$: Jean କୁଲିକ୍ଲେମ୍ଫ୍ରେ ape Rates

$$\phi \to \phi_J = \pi r_x^2 n_x v_{th,x} (1 + \lambda_x) \exp(-\lambda_x)$$

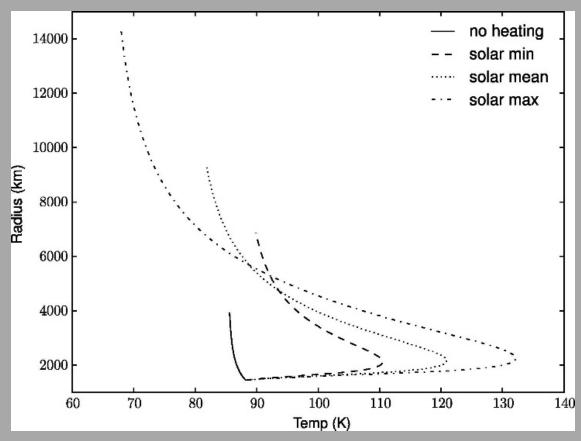
Energy Outflow*

$$\phi_E \to \phi_{E,J} = (k_B T_x) \left(2 + \frac{1}{1 + \lambda_x} \right) \phi_J$$

Jeans, J.H., 1925.
The Dynamical
Theory of Gases, 444.

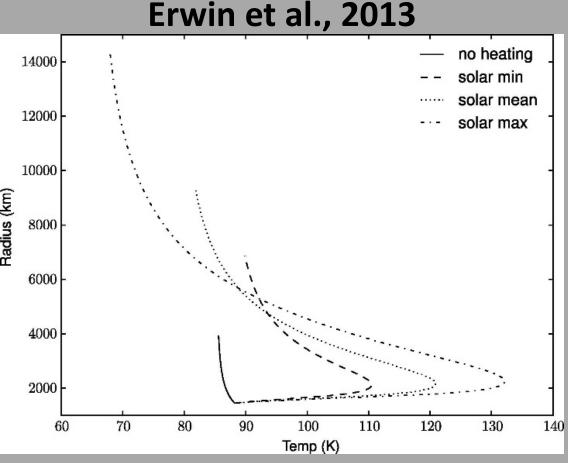
* ()_{$$\chi$$} = Exobase

Erwin et al., 2013



Erwin, J., Tucker, O.J. and Johnson, R.E., 2013. Hybrid fluid/kinetic modeling of Pluto's escaping atmosphere. *Icarus*, 226(1), pp.375-384.





My Results No Heating 14000 Solar Minimum 12000 <u>E</u> 100000 Radial Distance 8000 6000 4000 2000 60 70 90 100 110 120 80 130 140

Temperature [K]

Erwin, J., Tucker, O.J. and Johnson, R.E., 2013. Hybrid fluid/kinetic modeling of Pluto's escaping atmosphere. *Icarus*, 226(1), pp.375-384.