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K. Balasubramanian and Dingguo Dai

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# Potential energy surfaces for $\text{Ir} + \text{H}_2$ and $\text{Ir}^+ + \text{H}_2$ reactions

K. Balasubramanian<sup>a)</sup> and Dingguo Dai

Department of Chemistry, Arizona State University, Tempe, Arizona 85287-1604

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Potential energy surfaces of 10 electronic states of  $\text{IrH}_2$  and 12 electronic states of  $\text{IrH}_2^+$  are computed. A complete active space multi-configuration self-consistent field (CAS-MCSCF) followed by multireference configuration interaction (MRCI) calculations which included up to 270 000 configurations are employed. In addition spin-orbit effects are studied using the relativistic configuration interaction (RCI) method. It is found that the  $\text{Ir}(^2F)$  state inserts spontaneously into  $\text{H}_2$  to form a stable  $\text{IrH}_2$  molecule in a  $^2A_1$  ground state which is 28 kcal/mol more stable than  $\text{Ir}(^4F) + \text{H}_2$  in the absence of spin-orbit effects. The spin-orbit coupling of the quartet and doublet states provides nonzero transition probability for the insertion of  $\text{Ir}(^4F)$  state into  $\text{H}_2$ . The  $^3P$  state of  $\text{Ir}^+$  was found to insert spontaneously into  $\text{H}_2$  to form the  $^3A_2$  ground state of  $\text{IrH}_2^+$  which is 30 kcal/mol more stable than  $\text{Ir}^+(^5D) + \text{H}_2$ . An excited singlet state of  $\text{Ir}^+$  also was found to insert into  $\text{H}_2$  spontaneously. The spin-orbit couplings of quintet and triplet states of  $\text{IrH}_2^+$  at the crossing of these curves provide nonzero transition probability for the insertion of  $\text{Ir}^+(^5D)$  into  $\text{H}_2$ . The bent  $E$  ground state of  $\text{IrH}_2$  in the  $C_{2v}$  group was found to be a 63%  $^2A_1$ , 16%  $^2B_1$  and 17%  $^2A_2$  mixture. This strong mixing induces a large H–Ir–H bond angle change of  $9.5^\circ$  in the  $E(\text{III})$  state of  $\text{IrH}_2$ . The  $^3A_2$  ( $A_1$ ) ground state of  $\text{IrH}_2^+$  was found to be a 63%  $^3A_2$ , 15%  $^3B_2$ , 12%  $^3B_1$  and 7%  $^1A_1$  mixture. This strong spin-orbit mixing induces a  $\theta_e$  change of almost  $9^\circ$  in the ground state of  $\text{IrH}_2^+$ . The adiabatic ionization potential including spin-orbit effects for  $\text{IrH}_2$  and  $\text{Ir}$  are calculated as 8.2 and 8.6 eV, respectively. The ground state of  $\text{IrH}_2$  was found to be ionic ( $\mu_e = 2.2$  D) with  $\text{Ir}^+ \text{H}^-$  polarity exhibiting  $\text{Ir}(6s^{0.75}6p^{0.12}5d^{8.06})$  hybridization. The  $\text{IrH}_2^+$  ion in its  $^3A_2$  state exhibits  $\text{Ir}^+(6s^{0.62}6p^{0.12}5d^{7.51})$  hybridization character.

## I. INTRODUCTION

The reactivity pattern of transition metal clusters and their ions with small molecules such as  $\text{H}_2$  are the topics of several experimental and theoretical studies in recent years.<sup>1–16</sup> In particular, gas phase reactions of transition metal atoms and ions with  $\text{H}_2$  have been studied experimentally quite extensively.<sup>2,3,5,6</sup> The reactivities of transition metal clusters in dissociative chemisorption with  $\text{H}_2$  are found to vary dramatically with cluster size.<sup>1,7–9</sup> For example, Fe, Nb and Co clusters exhibit strong size dependency while Ni clusters do not exhibit such a variation.

Beauchamp, Armentrout and their co-workers<sup>2,5,6</sup> have studied the energetics and dynamics of the reactions of single transition metal atoms and ions with  $\text{H}_2$  as well as  $\text{CH}_4$ . The M–H, M–H<sup>+</sup>, M–CH<sub>3</sub> as well as M–CH<sub>3</sub><sup>+</sup> bond energies for several transition metal atoms have been measured using these experiments. Knight, Weltner and co-workers and other authors,<sup>4,10</sup> have used the rare gas matrix isolation electron spin resonance (ESR) techniques and gas phase-phase techniques to investigate the reactivities of species such as  $\text{Pd}^+ + \text{H}_2$  as well as transition metal hydrides such as  $\text{PdH}$ ,  $\text{MnH}$ , etc. All of these studies have resulted in a wealth of information.

Theoretical studies of  $\text{MH}_2$  and  $\text{MH}_2^+$  species for the transition metal atoms are now restricted to the first two row transition metal atoms with the exception of  $\text{PtH}_2$  which has been studied extensively both in our lab and by other investigators.<sup>15</sup> Similarly,  $\text{AuH}_2$  has been studied.<sup>16</sup> Such studies

have revealed considerable electronic state dependencies on the reactivities. In general, low spin electronic states of the metal atoms insert into  $\text{H}_2$  forming stable bent  $\text{MH}_2$  species while the high spin state of the metal atoms often have to surmount large barriers for insertion into  $\text{H}_2$ . For example, theoretical calculations on  $\text{TcH}_2$  and  $\text{RuH}_2$ <sup>25</sup> have revealed that  $\text{Tc}(^6S)$  atom has to surmount a large barrier for insertion into  $\text{H}_2$  while an excited doublet state inserts into  $\text{H}_2$  spontaneously. Similarly, the  $^5F$  state of  $\text{Ru}$  atom has to surmount a large barrier for insertion into  $\text{H}_2$  while the excited  $^3F$  and  $^1G$  states insert into  $\text{H}_2$ .<sup>25</sup> A recent study on  $\text{TcH}_2^+$  and  $\text{RuH}_2^+$ <sup>17</sup> revealed that  $\text{Tc}^+(^7S)$  state has to surmount a large barrier while a low-lying excited  $\text{Tc}^+(^5D)$  state inserts into  $\text{H}_2$  spontaneously. To the contrary  $\text{Ru}^+$  in its  $^4F$  ground state readily inserts into  $\text{H}_2$ . Such theoretical studies also have provided considerable insight into the nature of M–H bonds, the role of  $d$  orbitals in these bonds and in the dissociative chemisorptive process, the stability of  $\text{MH}_2$ , etc.

The  $\text{IrH}_2$  and  $\text{RhH}_2$  moieties<sup>18–22</sup> are found in hydroborate complexes or by chelation of diphosphine complexes. Eisenberg and Fisher<sup>18</sup> mention that no set of complexes in catalytic activation of substrates have been more vigorously studied in the last two decades than those of Rh and Ir. The complexes of both Rh and Ir in  $d^8$  ionic states possess rich electron density and thus oxidation chemistry. Both Ir and Rh are important materials in catalysis. Experimentally it is known that electron richness of Ir yields considerably more stable complexes. Crabtree *et al.*<sup>19</sup> have also studied several Ir dihydride and trihydride complexes.

<sup>a)</sup> Camille and Henry Dreyfus Teacher–Scholar.

There are no accurate *ab initio* calculations including relativistic and electron correlation effects for IrH<sub>2</sub> and IrH<sub>2</sub><sup>+</sup> species at present. The objective of this study is to investigate the potential energy surfaces of IrH<sub>2</sub> and IrH<sub>2</sub><sup>+</sup> arising from both the ground state and low-lying excited states of Ir and Ir<sup>+</sup> with the intent of juxtaposing the reactivity of different states of Ir and Ir<sup>+</sup> with H<sub>2</sub>. We employ complete active space MCSCF (CAS/MCSCF followed by multireference CI (MRCI) calculations which included up to 270 000 configurations. In addition spin-orbit effects are studied using the relativistic CI (RCI) method. We employ a large (5s5p4d) semicore and valence Gaussian basis for Ir.

## II. METHOD OF COMPUTATION

We employed a complete active space multiconfiguration self-consistent field (CAS-MCSCF) method to generate the orbitals for higher order MRCI calculations. Both CAS-MCSCF and MRCI calculations were made using relativistic effective core potentials (RECPs) which retained the outer 5s<sup>2</sup>5p<sup>6</sup>5d<sup>7</sup>6s<sup>2</sup> shells of the Ir atom in the valence space. Ross *et al.*<sup>23</sup> have generated such RECPs including the spin-orbit coupling term for the Ir atom. We start with their (5s5p3d) basis set. The first two large exponents *s* and *p* functions are for describing the 5s and 5p orbitals of Ir atom and to yield nodal structures for the 6s and 6p orbitals. The diffuse exponents in the 5p basis used here also generate the 6p orbital of Ir which hybridizes with 5d and 6s near the bending region. Our starting 5d basis set is triplet-zeta in quality. To this are added another set of diffuse six component *d* functions with  $\alpha_d = 0.0552$ . In earlier studies on several second row MH<sub>2</sub> species the effect of 10 component 4f functions were tested. It was established that these functions change bond lengths by only 0.001–0.003 Å. Consequently, in the present study the 4f functions were not included since it is evident that the effect of these functions is rather small. Our final basis thus consisted of a (5s5p4d) Gaussian basis set.

The IrH<sub>2</sub> and IrH<sub>2</sub><sup>+</sup> species were oriented on the YZ plane with the z axis bisecting the H–Ir–H bond angle. In the CAS-MCSCF calculations we included all orbitals of Ir and H which correlated into Ir (6s), Ir (5d) and H (1s) orbitals at infinite separation. In the C<sub>2v</sub> symmetry this active space comprised four a<sub>1</sub> orbitals, two b<sub>2</sub> orbitals and one each of b<sub>1</sub> and a<sub>2</sub> symmetries. Then the outer 11 electrons of IrH<sub>2</sub> and 10 electrons of IrH<sub>2</sub><sup>+</sup> were distributed in all possible ways among the eight active orbitals described above. Thus CAS-MCSCF calculations were made in a full CI space of valence electrons and valence orbitals. The 5s and 5p semicore orbitals were allowed to relax at the CAS-MCSCF but no excitations from these orbitals were allowed.

We used a multireference singles + doubles CI (MRCI) method to include the effect of higher-order electron correlation effects not included in the zeroth-order CAS/MCSCF method. The MRCI calculations included single + double excitations from all configurations in the CAS-MCSCF wave functions with coefficients >0.07. Separate CAS-MCSCF calculations were made for each state of different spatial symmetry in C<sub>2v</sub> or spin multiplicity. Our CAS-MCSCF calculations of IrH<sub>2</sub> included up to 264 CSFs

while MRCI calculations included up to 270 000 configurations. Similarly, the CAS-MCSCF of IrH<sub>2</sub><sup>+</sup> included up to 384 CSFs while MRCI calculations include up to 200 000 configurations.

The spin-orbit splitting of the Ir atom is known to be large. We included the effect of spin-orbit term using a relativistic configuration interaction method described before by one of the authors.<sup>26</sup> The RCI calculations were made near the potential wells of three low-lying doublet states of IrH<sub>2</sub> and at the crossings of bending potential energy surfaces of quartet and doublet states. These calculations were carried out in the C<sub>2v</sub><sup>2</sup> double group. For example, the <sup>2</sup>A<sub>1</sub> state near its  $\theta_e$  with a leading configuration open shell spin of  $\alpha$  included in the RCI, reference configurations for a <sup>2</sup>B<sub>1</sub> state with open shell spin  $\beta$ , <sup>2</sup>A<sub>2</sub> state with spin  $\alpha$  and a <sup>2</sup>B<sub>2</sub> state in the spin  $\beta$ . In addition a few quartet states with appropriate spin function combinations in the C<sub>2v</sub><sup>2</sup> double group were included. Subsequently, single and double excitations were included in the RCI. The same set of calculations was repeated without the spin-orbit integrals. The differences in the energies and geometries were then applied as corrections to the CAS/MRCI results.

The RCI calculations made at the quartet–doublet curve-crossing included all possible quartet states with appropriate spin function combinations and the <sup>2</sup>A<sub>1</sub> state with the objective of studying the transition probability for the quartet to doublet nonadiabatic transition at curve crossings.

The RCI calculations of IrH<sub>2</sub><sup>+</sup> were made on the spin-orbit states arising from the <sup>3</sup>A<sub>2</sub>, <sup>3</sup>B<sub>1</sub> and <sup>3</sup>B<sub>2</sub> states. The <sup>3</sup>A<sub>2</sub> state correlates into A<sub>1</sub>, B<sub>2</sub> and B<sub>1</sub> states in the C<sub>2v</sub><sup>2</sup> double group. The <sup>3</sup>B<sub>1</sub> state correlates into A<sub>1</sub>, B<sub>2</sub>, and A<sub>2</sub> symmetries while the <sup>3</sup>B<sub>2</sub> state gives rise to A<sub>1</sub>, B<sub>1</sub> and A<sub>2</sub> symmetries. Pitzer and Winter<sup>27</sup> have worked out the appropriate symmetry-adapted spin functions in the C<sub>2v</sub><sup>2</sup> double group. All nine spin-orbit states of IrH<sub>2</sub><sup>+</sup> were studied here. The RCI calculations of the A<sub>1</sub> state included all low-lying reference configurations in the C<sub>2v</sub><sup>2</sup> double group which have the A<sub>1</sub> symmetry. The RCI calculations include up to 26 reference configurations and single + double excitations from these configurations.

The RCI calculations at the crossings of triplet and quintet states included all possible quintet states with appropriate spin function combination for a given triplet state correlating into a spin-orbit state in C<sub>2v</sub><sup>2</sup>.

All CAS-MCSCF/MRCI calculations were made using one of the author's<sup>29</sup> modified versions of ALCHEMY II codes<sup>28</sup> to include RECPs. The RCI calculations were done using the general RCI method for polyatomics described in Ref. 26. The spin-orbit integrals for the RCI were obtained using Pitzer's ARGOS integral codes.<sup>30</sup> It should be mentioned that Hay and Wadt<sup>31</sup> have also generated RECPs for Ir which work satisfactorily except for the spin-orbit effects which are not included in the Hay–Wadt ECPs.

## III. RESULTS AND DISCUSSION

### A. Energy separations of Ir and Ir<sup>+</sup>

Table I shows our computed energy separations of pertinent quartet and doublet states of Ir and <sup>5</sup>D, <sup>3</sup>F and <sup>1</sup>G states

TABLE I. Atomic energy separations of Ir and Ir<sup>+</sup> obtained from asymptotic molecular energy separations at the dissociation limit.

| Configuration | Atomic state              | $E(\text{cm}^{-1})$ |                    |
|---------------|---------------------------|---------------------|--------------------|
|               |                           | Theory              | Expt. <sup>a</sup> |
| $5d^7 6s^2$   | Ir ( $^4F$ )              | 0                   | 0                  |
| $5d^8 6s^1$   | Ir ( $^2F$ )              | 9195                | 9086               |
| $5d^7 6s^1$   | Ir <sup>+</sup> ( $^5D$ ) | 0                   | ...                |
|               | Ir <sup>+</sup> ( $^3F$ ) | 4442                | ...                |
|               | Ir <sup>+</sup> ( $^1G$ ) | 11171               | ...                |

<sup>a</sup>  $J$ -weighted average energy separation from Ref. 32. Our theoretical IP of Ir is 8.6 eV compared to an experimental value of 9 eV in Ref. 32.

of Ir<sup>+</sup>. As seen from Table I, the ground state of Ir is a  $^4F$  state arising from the  $5d^7 6s^2$  configuration in agreement with established experimental atomic spectra.<sup>32</sup> The  $^2F$  excited state arising from  $5d^8 6s^1$  state is computed at  $\sim 9195 \text{ cm}^{-1}$  while the corresponding  $J$ -weight atomic separation from Moore's Table is  $\sim 9086 \text{ cm}^{-1}$ . Note that there is a low-lying Ir ( $^2P$ ) state arising from the same configuration<sup>32</sup> but we believe that the state to which doublet states of IrH<sub>2</sub> dissociate is Ir ( $^2F$ ) since four PES go to the same limit and a  $^2\Phi_u$  state goes to this limit.

The ionization potential of Ir atom is calculated as 7.1 eV at the CAS-MCSCF level and 8.3 eV at the MRCI level. The spin-orbit correction to the more accurate MRCI result yields an IP of 8.6 eV compared to an experimental IP of 9 eV.<sup>32</sup> Therefore the energy separations of IrH<sub>2</sub> relative to IrH<sub>2</sub><sup>+</sup> are not as well computed as the energy separations of neutral states or the energy separations of different states of IrH<sub>2</sub><sup>+</sup> relative to the ground state of IrH<sub>2</sub><sup>+</sup>.

There are no atomic spectral data listed in Moore's Tables for Ir<sup>+</sup>.<sup>32</sup> Our theoretical calculations predict the quintet states of IrH<sub>2</sub><sup>+</sup> dissociate to Ir<sup>+</sup> in the ( $5d^7 6s^1$ ) configuration. For some states ( $5d^7 6s^1$ ) and ( $5d^6 6s^2$ ) mixing was found. The ground state of the isoelectronic atom is ( $5d^6 6s^2$ )  $^5D$ . Since  $5d^6 6s^2$ – $5d^7 6s^1$  mixing was found we assign the Ir<sup>+</sup> state as  $^5D$ . Ir<sup>+</sup>  $^5D$ – $^3F$  separation for our asymptotic separation at the dissociation limit is computed as  $\sim 4400 \text{ cm}^{-1}$  while the Ir<sup>+</sup>  $^5D$ – $^1G$  separation is calculated as  $11\,200 \text{ cm}^{-1}$ .

## B. IrH<sub>2</sub>

### 1. Potential energy surfaces of IrH<sub>2</sub>

Figure 1 shows the bending PES of ten electronic states of doublet, quartet and sextet spin multiplicities. The quartet states of IrH<sub>2</sub> arise from the Ir ( $^4F$ ) ground state with the electronic configuration  $5d^7 6s^2$  (see Table I). The doublet states correlate into Ir ( $^2F$ ) arising from the  $5d^8 6s^1$  configuration. Note that the assignment to Ir ( $^2F$ ) is based on the fact that a linear doublet state of IrH<sub>2</sub> correlates into  $^2\Phi_u$  and thus the angular momentum of the atomic state must be at least three. This is the most compatible atomic state with  $5d^8 6s^1$  configuration in this region. The sextet states are most interesting in that they correlate into Ir

( $^6F$ ) + H( $^2S$ ) + H( $^2S$ ). Consequently, the energy separations of quartet and sextet states at the dissociation limit correspond to the  $D_e$  of H<sub>2</sub>. The sextet states are bound with respect to Ir ( $^4F$ ) + H + H as anticipated.

Table II shows the equilibrium geometries of the bent electronic states of IrH<sub>2</sub> at both the CAS-MCSCF and MRCI levels of theory. As seen from both Table II and Fig. 1, the ground state of IrH<sub>2</sub> is a  $^2A_1$  state with  $r_e = 1.54 \text{ \AA}$  and  $\theta_e = 92^\circ$ . We expect the bond length to increase by  $0.005 \text{ \AA}$  if  $4f$  functions are included. There are two low-lying nearly degenerate excited  $^2B_1$  and  $^2A_2$  states for IrH<sub>2</sub>. Three quartet electronic states are clustered near 1–1.1 eV region while the  $^2B_2$  state is 1.8 eV above the ground state. The sextet states are noticeably higher but below the  $D_e$  of H<sub>2</sub>.

In general higher-order correlation effects in the MRCI tend to shrink the Ir–H bonds although the changes in bond angles of the bent states are rather small ( $1^\circ$ – $6^\circ$ ). The ground state  $r_e$  changes at the MRCI level by  $-0.04 \text{ \AA}$  compared to the CAS-MCSCF level.

As seen from Table II, energy separations of excited electronic states are more sensitive to higher-order correlation effects. Among the doublet states, the  $^2A_1$  ground state is more stabilized and thus the energy separations of the  $^2B_1$  and  $^2A_2$  states are much higher. The CASSCF and MRCI separations of the quartet states relative to the bent  $^2A_1$  state are however, very comparable. The sextet states are more stabilized relative to the  $^2A_1$  ground state at the MRCI level.

The bond lengths of the high spin states (quartet, sextet) are noticeably longer than the low spin doublet states. This is anticipated since the low spin states can form more

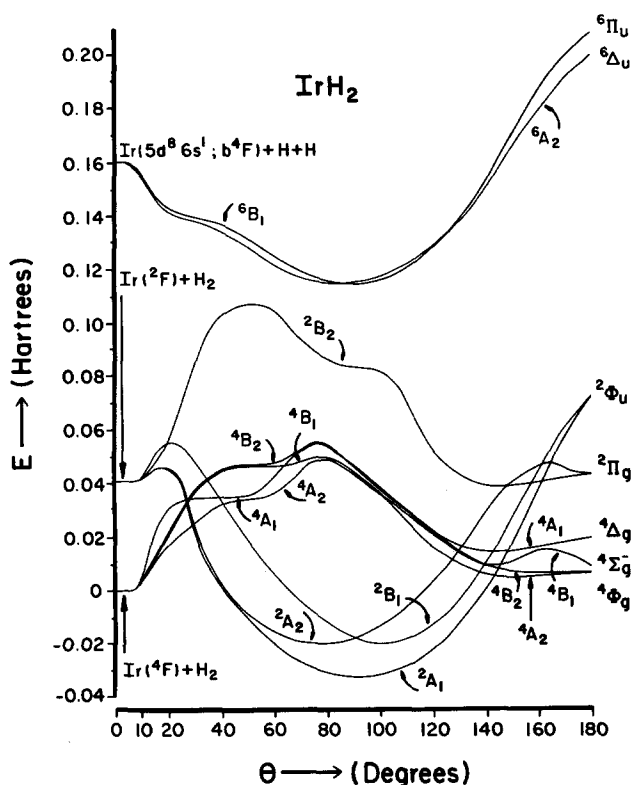
FIG. 1. Bending potential energy surfaces of 10 electronic states of IrH<sub>2</sub>.

TABLE II. Properties of the bent electronic states of IrH<sub>2</sub> (without spin-orbit effect).

| State                       | CASSCF           |           |                       | MRSDCI           |           |                       |
|-----------------------------|------------------|-----------|-----------------------|------------------|-----------|-----------------------|
|                             | $\theta_e$ , deg | $R_e$ , Å | $E$ , eV <sup>a</sup> | $\theta_e$ , deg | $R_e$ , Å | $E$ , eV <sup>b</sup> |
| <sup>2</sup> A <sub>1</sub> | 91.8             | 1.575     | 0.00                  | 91.9             | 1.540     | 0.00                  |
| <sup>2</sup> B <sub>1</sub> | 102.2            | 1.578     | 0.34                  | 105.6            | 1.548     | 0.65                  |
| <sup>2</sup> A <sub>2</sub> | 75.9             | 1.568     | 0.34                  | 74.5             | 1.541     | 0.68                  |
| <sup>4</sup> A <sub>2</sub> | 148.0            | 1.684     | 1.04                  | 142.5            | 1.640     | 1.02                  |
| <sup>4</sup> B <sub>1</sub> | 143.5            | 1.689     | 1.25                  | 137.7            | 1.658     | 1.12                  |
| <sup>4</sup> A <sub>1</sub> | 144.4            | 1.694     | 1.26                  | 139.7            | 1.656     | 1.13                  |
| <sup>2</sup> B <sub>2</sub> | 145.9            | 1.673     | 1.92                  | 141.7            | 1.631     | 1.80                  |
| <sup>6</sup> A <sub>2</sub> | 84.4             | 1.737     | 3.99                  | 81.8             | 1.728     | 3.87                  |
| <sup>6</sup> B <sub>1</sub> | 88.3             | 1.728     | 3.99                  | 88.2             | 1.722     | 3.95                  |

<sup>a</sup>The zero CASSCF energy is for the <sup>2</sup>A<sub>1</sub> bent minimum, and is -105.056 63 hartree.

<sup>b</sup>The zero MRSDCI energy is for the <sup>2</sup>A<sub>1</sub> bent minimum, and is -105.175 03 hartree.

stable bond pairs with the H atoms compared to the high spin states which are relatively less bound.

As seen from Fig. 1, the Ir (<sup>4</sup>F) ground state does not insert into H<sub>2</sub> spontaneously while the excited <sup>2</sup>F state inserts into H<sub>2</sub> spontaneously to form three stable doublet states namely, <sup>2</sup>A<sub>1</sub>, <sup>2</sup>A<sub>2</sub> and <sup>2</sup>B<sub>1</sub>. The barriers in the quartet states are about 29 kcal/mol for insertion into H<sub>2</sub>. However, the saddle points occur after the crossing of the quartet and doublet states. The lowest (<sup>4</sup>A<sub>2</sub>) surface crosses the <sup>2</sup>A<sub>1</sub> state at ~11 kcal/mol above the Ir (<sup>4</sup>F) + H<sub>2</sub> limit. Although the bond lengths at the crossing point are different, because of the large spin-orbit coupling of Ir which we discuss later, the quartet-doublet spin-orbit coupling matrix elements near the crossing are substantial enough to induce transition from the quartet surface to the doublet surface at sufficiently high relative velocities of collisions. Consequently, we estimate that the Ir (<sup>4</sup>F) ground state atom has to surmount a much smaller barrier of 10–15 kcal/mol for insertion into H<sub>2</sub> since the nonadiabatic transition probability for the quartet-doublet transition is nonnegligible. This is indeed a much smaller barrier for a high spin ground state compared to any first and second row transition metal atoms.

All quartet states of IrH<sub>2</sub> correlate into two *E* states in the C<sub>2v</sub> double group. The <sup>2</sup>A<sub>1</sub> state gives rise to an *E* state. For example, the <sup>2</sup>A<sub>1</sub> state with  $\alpha$  spin can mix with <sup>4</sup>A<sub>2</sub> with appropriate spin combinations. Similarly, the <sup>2</sup>A<sub>2</sub> state with  $\alpha$  spin can mix with <sup>4</sup>A<sub>1</sub> state with the spin combinations  $\alpha\beta\alpha$ ,  $\alpha\alpha\beta$ ,  $\beta\alpha\alpha$  and  $\beta\beta\beta$ . Note that the <sup>4</sup>A<sub>1</sub>-<sup>2</sup>A<sub>2</sub> crossing is very close to the <sup>4</sup>A<sub>1</sub>-<sup>2</sup>A<sub>1</sub> crossing. In the next section, we describe the spin-orbit matrix elements coupling these states. Hence Ir + H<sub>2</sub> PES are very interesting from the standpoint of spin-orbit induced enhancement of reactivity with H<sub>2</sub>.

There are two sets of sextet surfaces for IrH<sub>2</sub>. The electronic states which correlate into Ir (<sup>6</sup>D) + H<sub>2</sub> are only roughly ~5000 cm<sup>-1</sup> above the sextet surfaces studied here at the dissociation limit. This <sup>6</sup>D state arises from the

TABLE III. Properties of the linear electronic states of IrH<sub>2</sub> (without spin-orbit effect).<sup>a</sup>

| Leading configuration  | State                                    | CASSCF    |          | MRSDCI    |          |
|--|--|-----------|----------|-----------|----------|
|  |  | $R_e$ , Å | $E$ , eV | $R_e$ , Å | $E$ , eV |
| 1 $\sigma_g^2$ 2 $\sigma_g^1$ $\sigma_u^2$ 6 $\pi_g^3$ $\pi_g^3$             | <sup>4</sup> Φ <sub>g</sub>              | 1.719     | 1.07     | 1.694     | 1.29     |
| 1 $\sigma_g^2$ 2 $\sigma_g^1$ $\sigma_u^2$ 6 $\pi_g^2$ $\pi_g^4$             | <sup>4</sup> Σ <sub>g</sub> <sup>-</sup> | 1.711     | 1.12     | 1.685     | 1.28     |
| 1 $\sigma_g^2$ 2 $\sigma_g^2$ $\sigma_u^2$ 6 $\pi_g^3$ $\pi_g^2$             | <sup>4</sup> Δ <sub>g</sub>              | 1.737     | 1.43     | 1.711     | 1.43     |
| $\sigma_g^2$ $\sigma_u^2$ 6 $\pi_g^3$ $\pi_g^3$                              | <sup>2</sup> Π <sub>g</sub>              | 1.711     | 2.06     | 1.684     | 2.17     |
| 1 $\sigma_g^2$ 2 $\sigma_g^2$ $\sigma_u^2$ 6 $\pi_g^3$ $\pi_g^3$             | <sup>2</sup> Φ <sub>u</sub>              | 1.770     | 2.84     | 1.713     | 2.91     |
| 1 $\sigma_g^1$ 2 $\sigma_g^1$ $\sigma_u^2$ 6 $\pi_g^3$ $\pi_g^3$             | <sup>6</sup> Δ <sub>u</sub>              | 1.678     | 6.35     | 1.639     | 7.40     |
| 1 $\sigma_g^1$ 2 $\sigma_g^1$ $\sigma_u^2$ 6 $\pi_g^2$ $\pi_g^4$ 2 $\pi_u^1$ | <sup>6</sup> Π <sub>u</sub>              | 1.694     | 6.58     | 1.686     | 7.56     |

<sup>a</sup>All energies relative to the bent <sup>2</sup>A<sub>1</sub> minimum.

5d<sup>7</sup>6s6p configuration of Ir.<sup>32</sup> However, we expect the sextet surfaces from Ir (<sup>6</sup>D) + H<sub>2</sub> to have large barriers for insertion into H<sub>2</sub>. The sextet states correlating into Ir (<sup>6</sup>D) + H(2s) + H(2s) exhibit bent minima in the PES of <sup>6</sup>B<sub>1</sub> and <sup>6</sup>A<sub>2</sub> states in Fig. 1. Therefore the sextet minima of IrH<sub>2</sub> with  $\theta_e$  close to the  $\theta_e$  of the <sup>2</sup>A<sub>1</sub> ground state would provide channels for photoinduced dissociation of H<sub>2</sub> into H + H if excitation to the sextet states can be accomplished.

Table III shows the geometries and energy separations of linear states of IrH<sub>2</sub>. In general linear stationary points form longer Ir-H bonds since the Ir-H overlaps are considerably smaller for linear geometries. A primary reason for this is that in the linear states, the Ir (*d*<sub>xy</sub>) orbital cannot participate in bonding. The primary overlap in the  $\sigma_g$  orbitals of linear states is through the Ir (5d<sub>σ</sub>) + H(1s) and Ir (6s) + H(1s) interactions.

The linear electronic states are in general above the corresponding bent states with the exception of the <sup>2</sup>B<sub>2</sub> state for which the <sup>2</sup>B<sub>2</sub>-<sup>2</sup>Π<sub>g</sub> energy separation is small as this state forms an obtuse angle minimum primarily through the Ir *d*<sup>7</sup>s<sup>2</sup> + *d*<sup>7</sup>sp mixing.

The <sup>2</sup>B<sub>2</sub> bending PES has a shoulder after the saddle point. This shoulder arises from an avoided crossing. The saddle points in some of the quartet states arise from several avoided crossings discussed in Sec. IV A 1.

## 2. Spin-orbit effects for IrH<sub>2</sub>

Table IV shows the effect of spin-orbit (SO) coupling on the low-lying doublet state of IrH<sub>2</sub> near their minima. As

TABLE IV. Properties of the bent states of IrH<sub>2</sub> including spin-orbit effect.

| State                                  | $\theta_e$ , deg |       | $R_e$ , Å |       | $E$ , eV <sup>a</sup> |       |
|--|------------------|-------|-----------|-------|-----------------------|-------|
|  | SO               | No SO | SO        | No SO | SO                    | No SO |
| E(I) ( <sup>2</sup> A <sub>1</sub> )   | 91.5             | 91.9  | 1.545     | 1.540 | 0.00                  | 0.00  |
| E(II) ( <sup>2</sup> A <sub>2</sub> )  | 78.1             | 74.5  | 1.541     | 1.541 | 1.26                  | 0.68  |
| E(III) ( <sup>2</sup> B <sub>1</sub> ) | 96.1             | 105.6 | 1.547     | 1.548 | 1.33                  | 0.65  |

<sup>a</sup>The zero MRSDCI energy including spin-orbit effect is for the E(I) (<sup>2</sup>A<sub>1</sub>) bent minimum.

seen from Table IV, the geometry of the  $E(I)$  state which is predominantly  $^2A_1$ , changes a bit due to spin-orbit coupling although it is stabilized to a significant extent. The most dramatic change in the H-Ir-H bond angle is seen in the  $E(III)$  state which is predominantly  $^2B_1$ . The  $\theta_e$  of this state is significantly reduced ( $\sim 9.5^\circ$ ).

As seen from Fig. 1, the  $^2A_2$  and  $^2B_1$  states cross. In the  $C_{2v}^2$  double group both states give rise to  $E$  states. That is, in the presence of SO coupling, the  $^2A_2$  state with the open shell spin  $\alpha$  undergoes significant spin-orbit mixing with the  $^2B_2$  state with open shell spin  $\beta$ , since

$$\langle a_1 \alpha | H^{so} | b_1 \beta \rangle \neq 0,$$

$$\langle a_1 \alpha | H^{so} | b_2 \beta \rangle \neq 0,$$

$$\langle a_1 \alpha | H^{so} | a_2 \alpha \rangle \neq 0.$$

Near the crossing regions these two states therefore undergo avoided crossing when SO coupling is included. Since the  $^2B_1$  state has a large bond angle while the  $^2A_2$  state has a much smaller bond angle, the  $\theta_e$  of the  $E(II)$  state is increased due to this spin-orbit contamination while the  $\theta_e$  of the  $E(III)$  state is decreased significantly ( $\sim 9.5^\circ$ ) due to SO coupling. Consequently, these states of IrH<sub>2</sub> present very interesting examples for which SO coupling makes a dramatic impact on bond angles. The only other case documented up to now is the  $^3B_1$  ( $A_1$ ) state of PbH<sub>2</sub><sup>26</sup> for which  $\theta_e$  changes significantly due to the  $^3B_1$  ( $A_1$ )– $^1A_1$  ( $A_1$ ) spin-orbit mixing. The  $^2B_2$  state does not make much contribution to any of the low-lying doublet states since this state is significantly higher in energy.

As seen from Table IV, when spin-orbit effects are included energy separations are almost doubled compared to the  $^2A_1$ – $^2A_2$  separation in the absence of spin-orbit effects. This is primarily a result of spin-orbit stabilization of the  $^2A_1$  state due to contamination of  $^2A_1$  with  $^2A_2$  and  $^2B_1$  states.

The crossings of the quartet surfaces of IrH<sub>2</sub> arising from Ir ( $^4F$ ) + H<sub>2</sub> with the doublet surfaces, especially  $^2A_1$ , are very interesting from the standpoint of SO coupling (see Fig. 1). If the SO coupling between the quartet and doublet states is sufficiently large, indeed the transition probability for a transition from quartet to doublet surfaces will be suffi-

ciently large to enhance the reactivity of the Ir ( $^4F$ ) state with a much smaller barrier.

Table V shows the magnitude of the spin-orbit coupling matrix elements of the lowest  $^2A_1$  state with various quartet states near the curve crossing (theta) obtained from our relativistic CI matrix. As seen from Table V, the spin-orbit coupling of the  $^2A_1$  state with especially the  $^4B_1$  state is very large as a typical matrix element between two primitive Slater determinants is of the order of 0.010 88 hartrees. While some of this coupling will contribute to the  $^2B_1$  state which is obtained only with a sign change of the spin functions, the final RCI wave function reveals that the  $^2A_1$ – $^4B_1$  coupling is nonnegligible. The primary state contributing to the facilitation of the Ir ( $^4F$ ) insertion into H<sub>2</sub> is the  $^4B_1$  state as seen from Table V. The  $^4A_2$  state couples only to a small but nonnegligible extent. The transition probability can be significant enough that a measurable number of quartet→doublet surface hoppings occur to enhance the reactivity of Ir ( $^4F$ ). Consequently, due to the nonnegligible spin-orbit coupling between  $^2A_1$  and  $^4B_1$ , we conclude that Ir ( $^4F$ ) will insert into H<sub>2</sub> through transitions facilitated by the spin-orbit coupling.

The analysis of the relativistic CI calculations which included both doublet and quartet states at the curve crossing revealed that the contribution to the  $^2A_1$  state near crossing by  $^4B_1$  is 1.5%, while  $^4B_2$  and  $^4A_2$  states made only 0.3% and 0.05% contribution, respectively, near the curve crossings. This is fully consistent with the spin-orbit matrix elements coupling these states.

### 3. Dissociation energies

Table VI shows the dissociation energies of the low-lying states of IrH<sub>2</sub> with respect to Ir ( $^4F$ ) + H<sub>2</sub> in the absence of spin-orbit effects. In general MRCI calculations stabilize the ground state much more compared to CASSCF calculations. The  $^2A_1$  ground state of IrH<sub>2</sub> is 28 kcal/mol more stable than Ir ( $^4F$ ) + H<sub>2</sub> in the absence of spin-orbit effects. However, as noted above spin-orbit effects are significant for IrH<sub>2</sub>. The Ir ( $^4F$ ) atomic state is stabilized by 10.4 kcal/mol due to spin-orbit coupling. The  $E(I)$  state is stabilized

TABLE V. The magnitudes of the spin-orbit matrix elements for IrH<sub>2</sub> and IrH<sub>2</sub><sup>+</sup> at the doublet-quartet and triplet-quintet crossings, respectively.

|                           | IrH <sub>2</sub>                      |                                     |                                      |                                       |                                     |
|---------------------------|---------------------------------------|-------------------------------------|--------------------------------------|---------------------------------------|-------------------------------------|
|                           | $^4A_2$ ( $\alpha\alpha\beta$ )       | $^4A_2$ ( $\alpha\beta\alpha$ )     | $^4B_1$ ( $\beta\beta\alpha$ )       | $^4B_1$ ( $\alpha\beta\beta$ )        |                                     |
| $^2A_1 \alpha$            | 0.000 47                              | 0.000 47                            | 0.010 88                             | 0.010 88                              |                                     |
|                           | IrH <sub>2</sub> <sup>+</sup>         |                                     |                                      |                                       |                                     |
|                           | $^3B_2$ ( $\alpha\alpha\alpha\beta$ ) | $^3B_2$ ( $\alpha\beta\beta\beta$ ) | $^5A_2$ ( $\alpha\alpha\beta\beta$ ) | $^5B_1$ ( $\alpha\alpha\alpha\beta$ ) | $^5B_1$ ( $\alpha\beta\beta\beta$ ) |
| $^3A_2$ ( $\alpha\beta$ ) | 0.003 497                             | 0.003 497                           | 0.000 049                            | 0.003 56                              | 0.003 56                            |
|                           | $^5B_2$ ( $\beta\alpha\alpha\alpha$ ) | $^5B_2$ ( $\beta\beta\beta\alpha$ ) | $^5A_2$ ( $\beta\beta\alpha\alpha$ ) | $^5B_1$ ( $\beta\alpha\alpha\alpha$ ) | $^5B_1$ ( $\beta\beta\beta\alpha$ ) |
| $^3A_2$ ( $\beta\alpha$ ) | 0.003 497                             | 0.003 497                           | 0.000 49                             | 0.003 56                              | 0.003 56                            |

<sup>a</sup> Units are in hartrees. The spin combination shows the spin functions in the order of open shell orbitals in that Slater determinant. Only nonzero matrix elements are shown.

TABLE VI. Dissociation energies (Ir–H<sub>2</sub>) of IrH<sub>2</sub> with respect to Ir (<sup>4</sup>F) + H<sub>2</sub> (without spin–orbit effect).

| State                       | CASSCF (kcal/mol) | MRSDCI (kcal/mol) |
|-----------------------------|-------------------|-------------------|
| <sup>2</sup> A <sub>1</sub> | 21                | 28                |
| <sup>2</sup> B <sub>1</sub> | 13                | 13                |
| <sup>2</sup> A <sub>2</sub> | 13                | 12                |
| <sup>4</sup> A <sub>2</sub> | –4                | 5                 |
| <sup>4</sup> B <sub>1</sub> | –8                | 2                 |
| <sup>4</sup> A <sub>1</sub> | –9                | 2                 |
| <sup>2</sup> B <sub>2</sub> | –24               | –13               |

by 9.8 kcal/mol. Therefore, the effective stability of the  $E(I)$  state of IrH<sub>2</sub> with respect to the lowest Ir (<sup>4</sup>F<sub>9/2</sub>) + H<sub>2</sub> is 27.4 kcal/mol. The stability of this state with respect to the Ir (<sup>2</sup>F<sub>5/2</sub>) state is considerably higher. Based on the atomic <sup>4</sup>F<sub>9/2</sub>–<sup>2</sup>F<sub>5/2</sub> separation of 10 579 cm<sup>–1</sup>,<sup>32</sup> we estimate that the dissociation energy of the  $E(I)$  state of IrH<sub>2</sub> with respect to Ir (<sup>2</sup>F<sub>5/2</sub>) + H<sub>2</sub> is 57.6 kcal/mol. Consequently, both with respect to Ir (<sup>4</sup>F<sub>9/2</sub>) and Ir (<sup>2</sup>F<sub>5/2</sub>), the  $E(I)$  ground state of IrH<sub>2</sub> is very stable. For a product to have long enough life times to be seen in the supersonic beam it should be at least 19 kcal/mol more stable than the dissociated M + H<sub>2</sub>. Since IrH<sub>2</sub> is ~27 kcal/mol more stable than Ir (<sup>4</sup>F<sub>9/2</sub>) + H<sub>2</sub>, it should definitely be a very stable molecule and detectable in a supersonic beam. Up to now the isolated IrH<sub>2</sub> molecule has not been experimentally characterized, although as noted in the introduction IrH<sub>2</sub> is characterized as a subunit in larger organometallic complexes. We hope that this study will stimulate such experimental investigations of IrH<sub>2</sub>.

TABLE VII. Properties of bent electronic states of IrH<sub>2</sub><sup>+</sup> (without spin–orbit effect).

| State                       | CASSCF           |           |                       | MRSDCI           |           |                       |
|-----------------------------|------------------|-----------|-----------------------|------------------|-----------|-----------------------|
|                             | $\theta_e$ , deg | $R_e$ , Å | $E$ , eV <sup>a</sup> | $\theta_e$ , deg | $R_e$ , Å | $E$ , eV <sup>b</sup> |
| <sup>3</sup> A <sub>2</sub> | 69.8             | 1.573     | 0.00                  | 70.2             | 1.551     | 0.00                  |
| <sup>3</sup> B <sub>1</sub> | 96.9             | 1.582     | 0.14                  | 92.4             | 1.552     | 0.04                  |
| <sup>3</sup> A <sub>1</sub> | 57.3             | 1.582     | 0.29                  | 53.6             | 1.554     | 0.19                  |
| <sup>3</sup> B <sub>2</sub> | 82.3             | 1.604     | 0.57                  | 82.0             | 1.566     | 0.55                  |
| <sup>1</sup> A <sub>1</sub> | 60.4             | 1.570     | 0.73                  | 59.0             | 1.538     | 0.66                  |
| <sup>1</sup> A <sub>2</sub> | 74.8             | 1.572     | 1.52                  | 73.5             | 1.540     | 1.34                  |
| <sup>1</sup> B <sub>1</sub> | 92.2             | 1.579     | 1.62                  | 92.2             | 1.544     | 1.45                  |
| <sup>1</sup> B <sub>2</sub> | 80.9             | 1.598     | 1.98                  | 81.2             | 1.565     | 1.78                  |
| <sup>5</sup> B <sub>1</sub> | 142.9            | 1.684     | 2.24                  | 141.2            | 1.651     | 2.00                  |
| <sup>5</sup> A <sub>2</sub> | 148.1            | 1.678     | 2.24                  | 145.0            | 1.635     | 2.00                  |
| <sup>5</sup> A <sub>1</sub> | 144.4            | 1.685     | 2.27                  | 142.5            | 1.652     | 2.03                  |

<sup>a</sup> The zero CASSCF energy is for the <sup>3</sup>A<sub>2</sub> bent minimum and is –104.793 41 hartree.

<sup>b</sup> The zero MRSDCI energy is for the <sup>3</sup>A<sub>2</sub> bent minimum and is –104.877 69 hartree.

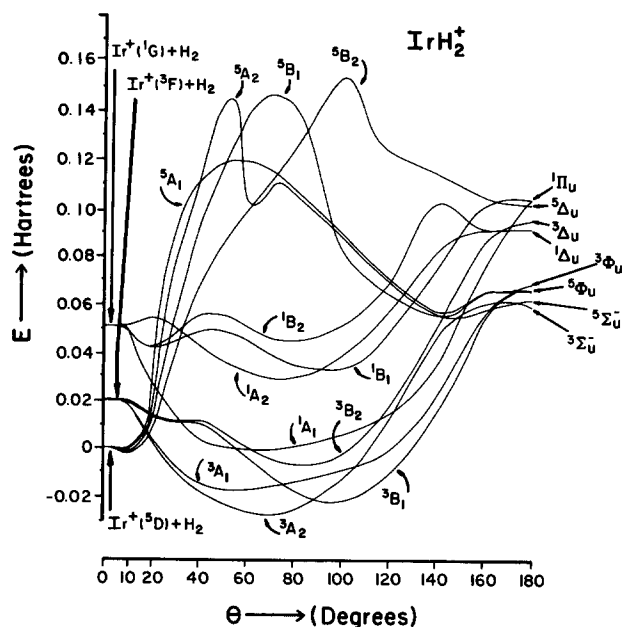
## C. IrH<sub>2</sub><sup>+</sup>

### 1. Potential energy surfaces of IrH<sub>2</sub><sup>+</sup>

Figure 2 shows the bending potential energy surfaces of the IrH<sub>2</sub><sup>+</sup> positive ion. Experimental information on Ir<sup>+</sup> is not available as Moore's Table<sup>32</sup> contains no information on the ground state or excited states of Ir<sup>+</sup>. However, based on the well characterized, energy separations of the isoelectronic Os atom,<sup>32</sup> we believe that our predicted ground state of Ir<sup>+</sup> as well as the energy separations of the excited states are reasonable. Since we did not study the heptet states of IrH<sub>2</sub><sup>+</sup>, the possibility of a heptet ground state of Ir<sup>+</sup> cannot be ruled out, but it is evident that the heptet states will not form low-lying minima. Nor will the heptet state insert into H<sub>2</sub> spontaneously.

As seen from Fig. 2, the low-lying Ir<sup>+</sup> (<sup>3</sup>F) state inserts into H<sub>2</sub> spontaneously while the Ir<sup>+</sup> (<sup>5</sup>D) state has to surmount a large barrier by itself to insert into H<sub>2</sub>. However, as we discuss later the spin–orbit coupling plays a key role in changing the course of reactivity of Ir<sup>+</sup> (<sup>5</sup>D) state through spin–orbit coupling of the triplet and quintet states of IrH<sub>2</sub><sup>+</sup>.

The insertion of Ir<sup>+</sup> (<sup>3</sup>F) into H<sub>2</sub> leads to at least four bent minima of <sup>3</sup>A<sub>2</sub>, <sup>3</sup>B<sub>1</sub>, <sup>3</sup>A<sub>1</sub> and <sup>3</sup>A<sub>2</sub> symmetries. Table VII shows the geometries and energy separations of the bent electronic states of IrH<sub>2</sub><sup>+</sup>. As seen from Table VII, the <sup>3</sup>A<sub>2</sub> and <sup>3</sup>B<sub>1</sub> electronic states of IrH<sub>2</sub><sup>+</sup> are nearly degenerate. At the MRCI level <sup>3</sup>A<sub>2</sub>–<sup>3</sup>B<sub>1</sub> energy separation is only 0.9 kcal/mol. As seen from Fig. 2, at the lower CAS-MCSCF level this separation is a bit larger. The <sup>3</sup>A<sub>2</sub> forms a minimum at  $\theta_e = 70.2^\circ$  and  $r_e = 1.551$  Å. The <sup>3</sup>B<sub>1</sub> minimum has a much larger bond angle of  $92^\circ$ . The bond lengths of all triplet minima are quite comparable. With the exception of the <sup>3</sup>B<sub>2</sub> state of IrH<sub>2</sub><sup>+</sup>, all other triplet states become closer to the <sup>3</sup>A<sub>2</sub> ground state at the higher MRCI level. This suggests that the <sup>3</sup>A<sub>2</sub> state of IrH<sub>2</sub><sup>+</sup> is slightly destabilized relative to <sup>3</sup>B<sub>1</sub> and <sup>3</sup>A<sub>1</sub> states due to higher-order correlation effects.

FIG. 2. Bending PES of 12 electronic states of IrH<sub>2</sub><sup>+</sup>.

The singlet states of IrH<sub>2</sub><sup>+</sup> arise from the insertion of Ir<sup>+</sup> (<sup>1</sup>G) into H<sub>2</sub>. Note that our assignment of the singlet dissociation into Ir<sup>+</sup> (<sup>1</sup>G) is tentative but based on the known experimental data on Os.<sup>32</sup> As seen from Fig. 2, the insertion of Ir<sup>+</sup> (<sup>1</sup>G) into H<sub>2</sub> is spontaneous. Among various singlet states thus formed, the <sup>1</sup>A<sub>1</sub> state is especially more stable. At the final MRCI level, the <sup>1</sup>A<sub>1</sub> state is 0.66 eV above the <sup>3</sup>A<sub>2</sub> state. The <sup>1</sup>A<sub>1</sub> state forms a minimum near 60° but the bending surface is somewhat shallow for this state. All other singlet states form bent minima with  $\theta_e = 73^\circ$ – $92^\circ$ . The energy separations of these states are also significantly higher.

In comparing the <sup>2</sup>A<sub>1</sub> ground state of IrH<sub>2</sub> with the two lowest states of IrH<sub>2</sub><sup>+</sup>, we find that  $\Delta\theta_e$  for the <sup>3</sup>B<sub>1</sub> state is almost zero. Similarly  $\Delta r_e$  is only +0.012 Å. This means that the Franck–Condon factor for the IrH<sub>2</sub> (<sup>2</sup>A<sub>1</sub>) → IrH<sub>2</sub><sup>+</sup> (<sup>3</sup>B<sub>1</sub>) transition will be large and thus the photoelectron spectrum corresponding to this transition will show a very narrow and sharp peak without much vibrational structure.

The IrH<sub>2</sub> (<sup>2</sup>A<sub>1</sub>) → IrH<sub>2</sub><sup>+</sup> (<sup>3</sup>A<sub>2</sub>) transition involves a large bond angle change ( $\Delta\theta_e = -22^\circ$ ) and somewhat smaller  $r_e$  changes. Consequently, this transition would have a smaller Franck–Condon factor and the spectral peak will be broad for this transition.

As seen from Fig. 2, although the quintet surfaces have large barriers, they cross the triplet surfaces much closer to the dissociation limit. The sharp barriers in all quintet surfaces arise from avoided crossings which we discuss in Sec. IV B 1.

The <sup>5</sup>B<sub>2</sub> surface crosses the <sup>3</sup>A<sub>2</sub> and <sup>3</sup>A<sub>1</sub> surfaces. Similarly, all of the quintet surfaces cross the <sup>3</sup>A<sub>2</sub> and <sup>3</sup>A<sub>1</sub> surfaces in this region. The energy barrier at the crossing is only 2.5 kcal/mol. The <sup>5</sup>B<sub>2</sub> state correlates into B<sub>1</sub>, B<sub>2</sub>, 2A<sub>1</sub>, and A<sub>2</sub> states in the C<sub>2v</sub> double group. The <sup>3</sup>A<sub>2</sub> state correlates into A<sub>1</sub>, B<sub>1</sub> and B<sub>2</sub> states while the <sup>3</sup>A<sub>1</sub> state correlates into A<sub>2</sub>, B<sub>1</sub> and B<sub>2</sub>. Consequently, the interaction of the spin-orbit state of a given symmetry arising from the quintet states with the corresponding spin-orbit state arising from <sup>3</sup>A<sub>1</sub> and <sup>3</sup>A<sub>2</sub> could be significant. The spin-orbit coupling near the crossing regions are discussed in the next subsection. The nonadiabatic transition probability for the quintet surfaces to hop to triplet states through significant spin-orbit contamination in the region of curve crossings is found to be nonnegligible. Since the energy barrier at the crossing region is only 2.5 kcal/mol, the probability of Ir<sup>+</sup> in its ground state inserting into H<sub>2</sub> is nonzero. Spin-orbit effect therefore seems to be the most important effect in governing the reactivity of Ir and Ir<sup>+</sup> towards H<sub>2</sub>.

The crossing of <sup>3</sup>A<sub>2</sub> and <sup>3</sup>B<sub>1</sub> surfaces, <sup>3</sup>A<sub>1</sub> and <sup>3</sup>B<sub>1</sub> surfaces and <sup>1</sup>A<sub>1</sub> and <sup>3</sup>B<sub>1</sub> surfaces, etc., seen in Fig. 2 induce significant spin-orbit contaminations among states of the same symmetry in the presence of spin-orbit coupling. Crossing of <sup>1</sup>B<sub>1</sub> with <sup>1</sup>A<sub>2</sub> will, however, not induce any spin-orbit coupling among these states since <sup>1</sup>B<sub>1</sub> correlates into B<sub>1</sub> while <sup>1</sup>A<sub>2</sub> correlates into the A<sub>2</sub> state when spin-orbit effects are included.

Table VIII shows the energy separations of linear electronic states of IrH<sub>2</sub><sup>+</sup>. A critical comparison of the Ir–H

TABLE VIII. Properties of the linear electronic states of IrH<sub>2</sub><sup>+</sup> (without spin-orbit effect).<sup>a</sup>

| Leading configuration                                   | State                                    | CASSCF    |          | MRSDCI    |          |
|---|--|-----------|----------|-----------|----------|
|   |  | $R_e$ , Å | $E$ , eV | $R_e$ , Å | $E$ , eV |
| $1\sigma_g^2 2\sigma_g^1 \sigma_u^1 \delta_g^2 \pi_g^4$ | <sup>3</sup> Σ <sub>u</sub> <sup>−</sup> | 1.748     | 2.35     | 1.712     | 2.57     |
| $1\sigma_g^2 2\sigma_g^1 \sigma_u^1 \delta_g^2 \pi_g^4$ | <sup>5</sup> Σ <sub>u</sub> <sup>−</sup> | 1.741     | 2.41     | 1.692     | 2.50     |
| $1\sigma_g^2 2\sigma_g^1 \sigma_u^1 \delta_g^3 \pi_g^3$ | <sup>5</sup> Φ <sub>u</sub>              | 1.763     | 2.55     | 1.718     | 2.67     |
| $1\sigma_g^2 2\sigma_g^1 \sigma_u^1 \delta_g^3 \pi_g^3$ | <sup>3</sup> Φ <sub>u</sub>              | 1.813     | 2.60     | 1.755     | 2.92     |
| $\sigma_g^2 \sigma_u^1 \delta_g^3 \pi_g^4$              | <sup>1</sup> Δ <sub>u</sub>              | 1.740     | 3.23     | 1.687     | 3.25     |
| $\sigma_g^2 \sigma_u^1 \delta_g^3 \pi_g^4$              | <sup>3</sup> Δ <sub>u</sub>              | 1.736     | 3.34     | 1.691     | 3.37     |
| $1\sigma_g^2 2\sigma_g^2 \sigma_u^1 \delta_g^3 \pi_g^2$ | <sup>5</sup> Δ <sub>u</sub>              | 1.839     | 3.50     | 1.800     | 3.55     |
| $\sigma_g^2 \sigma_u^1 \delta_g^4 \pi_g^3$              | <sup>1</sup> Π <sub>u</sub>              | 1.797     | 3.52     | 1.742     | 3.67     |

<sup>a</sup> All energies relative to the <sup>3</sup>A<sub>2</sub> bent minimum.

bond lengths of the linear and bent states reveals that in general, linear states form significantly longer Ir–H bonds compared to the bent states. This is consistent with the corresponding bond lengths of linear states of IrH<sub>2</sub>. In general, linear states are destabilized relative to the bent states at the higher MRCI level of theory. It is quite interesting that high spin states form more stable linear minima while for the low spin triplet and singlet states the linear stationary points are saddle points as evidenced from Fig. 2.

## 2. Spin-orbit effects for IrH<sub>2</sub><sup>+</sup>

Table IX shows the effect of spin-orbit coupling on the low-lying triplet states of IrH<sub>2</sub><sup>+</sup> near their minima. As seen from Table IX the spin-orbit effects are very significant for IrH<sub>2</sub><sup>+</sup>. All the spin-orbit components of the <sup>3</sup>A<sub>2</sub> state are stabilized relative to the <sup>3</sup>A<sub>2</sub> state.

The change in the H–Ir–H bond angle of the <sup>3</sup>A<sub>2</sub> (A<sub>1</sub>) component is dramatic (~9.3°) mainly due to the strong mixing of <sup>3</sup>A<sub>2</sub> (A<sub>1</sub>) with <sup>3</sup>B<sub>2</sub> (A<sub>1</sub>) as well as <sup>3</sup>B<sub>1</sub> (A<sub>1</sub>) and <sup>1</sup>A<sub>1</sub> (A<sub>1</sub>) states. Since the nearly degenerate <sup>3</sup>B<sub>1</sub> state has a much larger  $\theta_e$  of 92.4°, the strong mixing of <sup>3</sup>B<sub>1</sub> (A<sub>1</sub>) with

TABLE IX. Properties of the bent states of IrH<sub>2</sub><sup>+</sup> including spin-orbit effect.

| State   | $\theta_e$ , deg | $R_e$ , Å | $E$ , eV |
|---|------------------|-----------|----------|
| <sup>3</sup> A <sub>2</sub> (A <sub>1</sub> ) | 79.5             | 1.557     | 0.00     |
| <sup>3</sup> A <sub>2</sub> (B <sub>1</sub> ) | 72.8             | 1.556     | 0.19     |
| <sup>3</sup> A <sub>2</sub> (B <sub>2</sub> ) | 71.4             | 1.552     | 0.27     |
| <sup>3</sup> A <sub>2</sub>                   | 70.2             | 1.551     | 0.38     |
| <sup>3</sup> B <sub>1</sub> (A <sub>1</sub> ) | 83.6             | 1.548     | 0.34     |
| <sup>3</sup> B <sub>1</sub> (B <sub>2</sub> ) | 85.6             | 1.550     | 0.39     |
| <sup>3</sup> B <sub>1</sub> (A <sub>2</sub> ) | 89.6             | 1.545     | 0.39     |
| <sup>3</sup> B <sub>1</sub>                   | 92.4             | 1.552     | 0.42     |
| <sup>3</sup> B <sub>2</sub> (A <sub>2</sub> ) | 83.1             | 1.566     | 0.88     |
| <sup>3</sup> B <sub>2</sub> (B <sub>1</sub> ) | 82.1             | 1.560     | 1.01     |
| <sup>3</sup> B <sub>2</sub> (A <sub>1</sub> ) | 81.1             | 1.560     | 1.09     |
| <sup>3</sup> B <sub>2</sub>                   | 82.0             | 1.566     | 0.93     |



$^3A_2$  ( $A_1$ ) changes the  $\theta_e$ s of the corresponding components significantly. Note that the  $\theta_e$  of the  $^3B_1$  ( $A_1$ ) decreases by 9° while the  $\theta_e$  of the  $^3A_2$  ( $A_1$ ) states increases by almost the same amount. The other spin-orbit components are less influenced.

The energy separations are sensitive to the inclusion of spin-orbit coupling as expected for IrH<sub>2</sub><sup>+</sup>. As seen from Table IX the  $^3A_2$  ( $A_1$ )– $^3A_2$  ( $B_2$ ) energy separation is 0.27 eV while these states are degenerate in the absence of SO coupling.

Table V shows the spin-orbit matrix elements between the two primitive determinants of the  $^3A_2$  states of IrH<sub>2</sub><sup>+</sup> with various quintet states near the region of quintet-triplet curve crossings at small bond angles (Fig. 2). As discussed before for IrH<sub>2</sub>, the transition probability for Ir<sup>+</sup> + H<sub>2</sub> collision to hop from a quintet to triplet surface will depend on the spin-orbit coupling matrix element since the energy separation at the curve crossing is related to the spin-orbit matrix elements.

As seen from Table V, the SO coupling between the  $^3A_2$  state with determinants that give rise to  $^5B_2$  and  $^5B_1$  are significant (total of eight matrix elements each = 0.0035 hartrees). Consequently, the quintet-triplet SO coupling for IrH<sub>2</sub><sup>+</sup> is quite comparable to the quartet-doublet SO coupling for IrH<sub>2</sub> at similar curve crossings.

The analysis of the final RCI wave function at curve crossing region reveals that the  $^3A_2$  ( $A_1$ ) state is composed of 69%  $^3A_2$ , 15%  $^3B_2$ , 11%  $^3B_1$ , 1.5%  $^5B_2$  and 1.3%  $^5B_1$ . Hence the total quintet contamination to the  $^3A_2$  ( $A_1$ ) state is 2.8% through primary configurations and up to 1% through secondary excited quintet configurations. Similar composition was found for other spin-orbit states arising from the triplet states.

The above analyses of RCI wave functions and SO coupling matrix in the curve-crossing region reveal that the non-adiabatic transition probability for the quintet surfaces to hop to the triplet states is nonnegligible. This contrasts with the higher second and first row MH<sub>2</sub> species for which the spin-orbit contamination of different spin multiplets are smaller. Therefore we conclude that a nonzero number of nonadiabatic quintet-triplet transitions will occur in the curve-crossing regions thereby facilitating spin-orbit enhanced reactivity of the quintet state of Ir<sup>+</sup>.

### 3. Adiabatic ionization potential of IrH<sub>2</sub>

The adiabatic IP of IrH<sub>2</sub> was obtained both with and without SO effects. At the lower CAS-MCSCF level of theory the IP is only 7.2 eV while at the superior MRCI level of theory the adiabatic IP of IrH<sub>2</sub><sup>+</sup> is 8.1 eV. Since the spin-orbit effects of IrH<sub>2</sub> and IrH<sub>2</sub><sup>+</sup> are comparable, the net result of this effect is to increase the IP by 0.1 eV. The final IP of IrH<sub>2</sub> including SO coupling is calculated as 8.2 eV. The CAS-MCSCF level of theory underestimates the IP by almost 0.9 eV. This is not surprising in view of the large electron correlation effects for IrH<sub>2</sub>. Since the adiabatic IP of IrH<sub>2</sub> has been determined, the photoelectron spectra of IrH<sub>2</sub> can be predicted using the energy separations of electronic states of IrH<sub>2</sub> and IrH<sub>2</sub><sup>+</sup>. We note that our final SO corrected MRCI IP of 8.2 eV is likely to be smaller than the

TABLE X. Dissociation energies of IrH<sub>2</sub><sup>+</sup> without spin-orbit effect.<sup>a</sup>

| State   | CASSCF (kcal/mol) | MRSDCI (kcal/mol) |
|---------|-------------------|-------------------|
| $^3A_2$ | 17                | 30                |
| $^3B_1$ | 14                | 29                |
| $^3A_1$ | 11                | 25                |
| $^3B_2$ | 4                 | 17                |
| $^1A_1$ | 0                 | 14                |
| $^1A_2$ | −18               | 1                 |
| $^1B_1$ | −20               | −4                |
| $^1B_2$ | −28               | −11               |
| $^5A_2$ | −34               | −16               |
| $^5B_1$ | −34               | −16               |
| $^5A_1$ | −35               | −17               |

<sup>a</sup>With respect to Ir<sup>+</sup> ( $^5D$ ) + H<sub>2</sub>.

true value since our calculations underestimate the IP of the Ir atom by 0.4 eV compared to the experimental value.

### 4. Dissociation energies of IrH<sub>2</sub><sup>+</sup>

Table X shows the dissociation energies of IrH<sub>2</sub><sup>+</sup> in the absence of spin-orbit effects with respect to Ir<sup>+</sup> ( $^5D$ ) + H<sub>2</sub> dissociation limit. As seen from Table X, the IrH<sub>2</sub><sup>+</sup>  $^3A_2$  ground state is 30 kcal/mol more stable than Ir ( $^5D$ ) + H<sub>2</sub>. When compared to the dissociation energy of IrH<sub>2</sub> we find that IrH<sub>2</sub><sup>+</sup> is 2 kcal/mol more stable compared to IrH<sub>2</sub> for dissociation in the absence of SO coupling. As evidenced from Table X, the MRCI calculations increase the  $D_e$  of IrH<sub>2</sub><sup>+</sup> by almost 13 kcal/mol indicating that electron correlation effects have significant effect especially on binding energies.

Although the higher singlet states of IrH<sub>2</sub><sup>+</sup> are not bound with respect to Ir<sup>+</sup> ( $^5D$ ) + H<sub>2</sub>, they are bound with respect to their own singlet dissociation limits, viz., Ir<sup>+</sup> ( $^1G$ ) + H<sub>2</sub>. For such a heavy ion however, the crossing of quintet surfaces with these singlet surfaces in Fig. 2, would lead to predissociation into Ir<sup>+</sup> ( $^5D$ ) + H<sub>2</sub>. Consequently, all  $D_e$ 's are listed with respect to Ir<sup>+</sup> ( $^5D$ ) + H<sub>2</sub>.

Next we consider the effect of spin-orbit coupling on the  $D_e$ . The  $^3A_2$  ( $A_1$ ) state of IrH<sub>2</sub><sup>+</sup> in the  $C_{2v}^2$  double group is stabilized by 0.38 eV due to spin-orbit coupling. The Ir<sup>+</sup> ( $^5D_4$ ) state is stabilized by ~0.3 eV due to spin-orbit coupling. But the  $A_1$  spin-orbit state dissociates into Ir<sup>+</sup> ( $^5D_0$ ) + H<sub>2</sub> which is estimated at 0.7 eV above the  $^5D_4$  state based on the electronic states of Os. Consequently, the  $D_e$  of the  $A_1$  spin-orbit state of IrH<sub>2</sub><sup>+</sup> with respect to Ir<sup>+</sup> ( $^5D_0$ ) + H<sub>2</sub> is estimated as to be 18 kcal/mol more than its value in the absence of SO coupling or 48 kcal/mol.

## IV. THE NATURE OF ELECTRONIC STATES AND PES OF IrH<sub>2</sub> AND IrH<sub>2</sub><sup>+</sup>

### A. IrH<sub>2</sub>

#### 1. The CI compositions, the shapes of PES, and the origins of saddle points

Table XI shows the contributions of leading configurations to the MRSDCI wave function of IrH<sub>2</sub>. While the

TABLE XI. Leading configurations in the MRSDCI of the bent states of IrH<sub>2</sub>.

| State                       | Coefficient | Configuration   |                 |                 |                 |                 |                 |                 |                 |
|-----------------------------|-------------|-----------------|-----------------|-----------------|-----------------|-----------------|-----------------|-----------------|-----------------|
|                             |             | 1a <sub>1</sub> | 2a <sub>1</sub> | 3a <sub>1</sub> | 4a <sub>1</sub> | 1b <sub>2</sub> | 2b <sub>2</sub> | 1b <sub>1</sub> | 1a <sub>2</sub> |
| <sup>2</sup> A <sub>1</sub> | 0.951       | 2               | 2               | 1               | 0               | 2               | 0               | 2               | 2               |
| <sup>2</sup> B <sub>1</sub> | −0.951      | 2               | 2               | 2               | 0               | 2               | 0               | 1               | 2               |
| <sup>2</sup> A <sub>2</sub> | −0.953      | 2               | 2               | 2               | 0               | 2               | 0               | 2               | 1               |
| <sup>4</sup> A <sub>2</sub> | −0.961      | 2               | 2               | 1               | 0               | 2               | 1               | 1               | 2               |
| <sup>4</sup> B <sub>1</sub> | −0.959      | 2               | 2               | 1               | 0               | 2               | 1               | 2               | 1               |
| <sup>4</sup> A <sub>1</sub> | 0.959       | 2               | 2               | 2               | 0               | 2               | 1               | 1               | 1               |
| <sup>2</sup> B <sub>2</sub> | 0.579       | 2               | 2               | 0               | 0               | 2               | 1               | 2               | 2               |
|                             | 0.514       | 2               | 1               | 1               | 0               | 2               | 1               | 2               | 2               |
|                             | −0.400      | 2               | 0               | 2               | 0               | 2               | 1               | 2               | 2               |
|                             | −0.361      | 2               | 2               | 2               | 0               | 2               | 1               | 0               | 2               |
| <sup>6</sup> A <sub>2</sub> | 0.958       | 2               | 1               | 1               | 1               | 2               | 1               | 1               | 2               |
| <sup>6</sup> B <sub>1</sub> | 0.959       | 2               | 1               | 1               | 1               | 2               | 1               | 2               | 1               |

lowest <sup>2</sup>A<sub>1</sub>, <sup>2</sup>B<sub>1</sub>, <sup>2</sup>A<sub>2</sub> and three quartet states are simple in composition this is not so for the <sup>2</sup>B<sub>2</sub> state. This explains the shape of the PES of the <sup>2</sup>B<sub>2</sub> state in Fig. 1. The obtuse  $\theta_e$  in the <sup>2</sup>B<sub>2</sub> state arises from avoided crossings of  $1a_1^2 2a_1^2 1b_2^2 2b_2 1b_1^2 1a_2^2$  with  $1a_1^2 2a_1 3a_1 1b_2^2 2b_2 1b_1^2 1a_2^2$ ,  $1a_1^2 3a_1^2 1b_2^2 2b_2 1b_1^2 1a_2^2$ , and  $1a_1^2 2a_1^2 3a_1^2 1b_2^2 2b_2 1a_2^2$  configurations. The shoulder in the <sup>2</sup>B<sub>2</sub> PES also is due to avoided crossings. All other bent electronic states of IrH<sub>2</sub> are well represented by a single configuration (Table XI).

At  $\theta = 20^\circ$  the <sup>2</sup>B<sub>2</sub> state is 27%  $1a_1^2 2a_1^2 3a_1 4a_1 1b_2 1b_1^2 1a_2^2$ , 25%  $1a_1^2 2a_1^2 3a_1^2 4a_1 1b_2^2 1b_1 1a_2$ , 19%  $1a_1^2 2a_1^2 3a_1^2 1b_2 1b_1^2 1a_2^2$ , etc. At  $\theta = 40^\circ$  this state becomes 94%  $1a_1^2 2a_1^2 3a_1^2 1b_2 1b_1^2 1a_2^2$ . Consequently, the saddle point in this region of the <sup>2</sup>B<sub>2</sub> state arises from this avoided crossing. At  $\theta = 60^\circ$ , this state becomes 89%  $1a_1^2 2a_1^2 3a_1 1b_2^2 2b_2 1b_1^2 1a_2^2$ . This avoided crossing explains the shoulder in the <sup>2</sup>B<sub>2</sub> PES beyond the saddle point. At  $\theta = 80^\circ$ , this state is 63%  $1a_1^2 2a_1 3a_1 1b_2^2 2b_2 1b_1^2 1a_2^2$  and 14%  $1a_1^2 2a_1^2 1b_2^2 2b_2 1b_1^2 1a_2^2$  and 11%

$1a_1^2 3a_1^2 1b_2^2 2b_2 1b_1^2 1a_2^2$ . This explains the second shoulder. The composition of the <sup>2</sup>B<sub>2</sub> state is quite complex near its obtuse  $\theta_e$  as seen from Table XI.

The <sup>4</sup>A<sub>2</sub> state at  $\theta = 20^\circ$  is composed of 98%  $1a_1^2 2a_1^2 3a_1 4a_1 1b_2^2 1b_1^2 1a_2$ . At  $\theta = 80^\circ$ , it is still dominated by this configuration. Consequently, the broad saddle points in the <sup>4</sup>A<sub>2</sub> and other quartet states are natural barriers and arise from changes in the antibonding orbital interactions to bonding interactions while the saddle point in the <sup>2</sup>B<sub>2</sub> state at very acute  $\theta$  occurs due to avoided crossings discussed above.

Among the linear electronic states of IrH<sub>2</sub>, the <sup>4</sup>Δ<sub>g</sub>, <sup>4</sup>Φ<sub>g</sub>, <sup>6</sup>Δ<sub>u</sub>, <sup>6</sup>Π<sub>u</sub>, <sup>2</sup>Φ<sub>u</sub> states are well represented by one or two configurations. But the <sup>4</sup>Σ<sub>g</sub><sup>−</sup> and <sup>2</sup>Π<sub>g</sub> states need a multiconfigurational treatment. The mixing of the  $\sigma_g \delta_g^2$  and  $\sigma_g \pi_g^2$  configurations in the <sup>4</sup>Σ<sub>g</sub><sup>−</sup> state are very noticeable (Table XII).

## 2. Dipole moments

Table XIII shows the dipole moments ( $\mu_e$ ) of the low-lying states of IrH<sub>2</sub>. All  $\mu_e$  in Table XIII have Ir<sup>+</sup>H<sup>−</sup> polarity. The large  $\mu_e$ s of IrH<sub>2</sub> indicate the strong ionicities of the IrH bonds. The  $\mu_e$ s of <sup>6</sup>A<sub>2</sub> and <sup>6</sup>B<sub>1</sub> states are significantly smaller as expected since these are high spin states. For the high spin states, the electron transfer from Ir to H is expected to be significantly smaller and thus the dipole moment is small for <sup>6</sup>A<sub>2</sub> and <sup>6</sup>B<sub>1</sub>.

The <sup>2</sup>A<sub>1</sub> and <sup>2</sup>A<sub>2</sub> states have significantly larger  $\mu_e$ s compared to <sup>2</sup>B<sub>1</sub>. As seen from Table II, the  $\theta_e$ s of the <sup>2</sup>A<sub>2</sub> and <sup>2</sup>A<sub>1</sub> states are much smaller than the  $\theta_e$  of the <sup>2</sup>B<sub>1</sub> state. In general, as the bond angle is reduced,  $\mu_e$  should become larger. For this reason the  $\mu_e$  of the <sup>2</sup>B<sub>2</sub> state is only 0.985 D since its  $\theta_e \sim 142^\circ$ . The large  $\mu_e$ s of the <sup>2</sup>A<sub>1</sub> and <sup>2</sup>B<sub>2</sub> states of IrH<sub>2</sub> are consistent with the electron rich character of Ir. The  $\mu_e/r_e$  values for the <sup>2</sup>A<sub>1</sub> and <sup>2</sup>A<sub>2</sub> states are 1.42 and 1.44 D/Å, respectively.

The smaller dipole moments of the quartet states are consistent with both their high spin characters and the obtuse  $\theta_e$ s of these states as seen from Table II.

TABLE XII. Important configurations in the MRSDCI of the linear states of IrH<sub>2</sub>.<sup>a</sup>

| State                                    | Coefficient(s) | Configuration   |                 |                |                |                |                 |                 |  |
|--|----------------|-----------------|-----------------|----------------|----------------|----------------|-----------------|-----------------|--|
|  |                | 1σ <sub>g</sub> | 2σ <sub>g</sub> | σ <sub>u</sub> | δ <sub>g</sub> | π <sub>g</sub> | 1π <sub>u</sub> | 2π <sub>u</sub> |  |
| <sup>4</sup> Φ <sub>g</sub>              | 0.684, −0.684  | 2               | 1               | 2              | 3              | 3              | 0               | 0               |  |
| <sup>4</sup> Σ <sub>g</sub> <sup>−</sup> | 0.849          | 2               | 1               | 2              | 2              | 4              | 0               | 0               |  |
|  | 0.461          | 2               | 1               | 2              | 4              | 2              | 0               | 0               |  |
| <sup>4</sup> Δ <sub>g</sub>              | −0.960         | 2               | 2               | 2              | 3              | 2              | 0               | 0               |  |
| <sup>2</sup> Φ <sub>u</sub>              | 0.596, −0.554, | 2               | 2               | 1              | 3              | 3              | 0               | 0               |  |
|  | −0.367, 0.297  |                 |                 |                |                |                |                 |                 |  |
| <sup>2</sup> Π <sub>g</sub>              | −0.575         | 2               | 0               | 2              | 4              | 3              | 0               | 0               |  |
|  | −0.418, 0.400  | 2               | 1               | 2              | 3              | 3              | 0               | 0               |  |
|  | 0.306, 0.297,  | 2               | 2               | 2              | 2              | 3              | 0               | 0               |  |
|  | −0.256         |                 |                 |                |                |                |                 |                 |  |
| <sup>6</sup> Δ <sub>u</sub>              | 0.981          | 1               | 1               | 2              | 1              | 3              | 3               | 0               |  |
| <sup>6</sup> Π <sub>u</sub>              | 0.991          | 1               | 1               | 2              | 0              | 2              | 4               | 1               |  |

<sup>a</sup> Multiple CSFs for a given configuration have multiple coefficients in a line.

TABLE XIII. Dipole moment for the bent minimum of IrH<sub>2</sub>.

| State                       | $\mu$ , D <sup>a</sup> |
|-----------------------------|------------------------|
| <sup>2</sup> A <sub>1</sub> | 2.18                   |
| <sup>2</sup> B <sub>1</sub> | 1.79                   |
| <sup>2</sup> A <sub>2</sub> | 2.22                   |
| <sup>4</sup> A <sub>2</sub> | 1.09                   |
| <sup>4</sup> B <sub>1</sub> | 1.25                   |
| <sup>4</sup> A <sub>1</sub> | 1.20                   |
| <sup>2</sup> B <sub>2</sub> | 0.99                   |
| <sup>6</sup> A <sub>2</sub> | 0.61                   |
| <sup>6</sup> B <sub>1</sub> | 0.47                   |

<sup>a</sup>Polarity is Ir<sup>+</sup>H<sup>-</sup>.

### 3. Mulliken populations

Table XIV shows the Ir *s*, *p*, and *d* Mulliken populations of the electronic states of IrH<sub>2</sub>. Due to the complexity of the Ir hybridization, Mulliken populations are by far one of the most useful quantities to gain insight into the nature of bonding in transition metal compounds although Mulliken populations are basis-set-dependent. As seen from Table XIV, in the ground state of IrH<sub>2</sub>, Ir atom exhibits 5d<sup>8.1</sup>6s<sup>0.75</sup>6p<sup>0.12</sup> hybridization. The remaining electron density from Ir is transferred to H atoms. This agrees reasonably with the fact that the Ir (<sup>2</sup>A<sub>1</sub>) state arises from the 5d<sup>8</sup>6s<sup>1</sup> (<sup>2</sup>F) configuration of the Ir atom. The 6*p* participation in the <sup>2</sup>A<sub>1</sub> state is noticeable. In general, all doublet electronic states have considerably larger *d* populations compared to the quartet and sextet states.

The <sup>2</sup>B<sub>1</sub> and <sup>2</sup>A<sub>2</sub> electronic states of IrH<sub>2</sub> exhibit a more complex 5d<sup>8</sup>6s<sup>1</sup>, 5d<sup>7</sup>6s<sup>2</sup> hybridization since their 6*s* Mulliken populations are larger while the *d* populations are smaller.

The 6*p* population in the quartet and sextet states are especially large. We are not surprised by the large 6*p* popula-

TABLE XIV. Mulliken population analyses for the electronic states of IrH<sub>2</sub>.

| State                                    | Gross           |                 |                 | Ir-H Overlaps |
|--|-----------------|-----------------|-----------------|---------------|
|  | Ir ( <i>s</i> ) | Ir ( <i>p</i> ) | Ir ( <i>d</i> ) |               |
| <sup>2</sup> A <sub>1</sub>              | 0.753           | 0.123           | 8.057           | 1.117         |
| <sup>2</sup> B <sub>1</sub>              | 1.350           | 0.118           | 7.453           | 1.132         |
| <sup>2</sup> A <sub>2</sub>              | 1.318           | 0.104           | 7.546           | 1.116         |
| <sup>4</sup> A <sub>2</sub>              | 0.752           | 0.309           | 7.541           | 1.043         |
| <sup>4</sup> B <sub>1</sub>              | 1.346           | 0.320           | 6.944           | 1.079         |
| <sup>4</sup> A <sub>1</sub>              | 1.372           | 0.324           | 6.911           | 1.086         |
| <sup>2</sup> B <sub>2</sub>              | 0.823           | 0.264           | 7.551           | 0.967         |
| <sup>6</sup> A <sub>2</sub>              | 1.006           | 0.706           | 6.834           | 0.272         |
| <sup>6</sup> B <sub>1</sub>              | 1.013           | 0.722           | 6.803           | 0.233         |
| <sup>4</sup> Φ <sub>g</sub>              | 0.857           | 0.384           | 7.289           | 1.096         |
| <sup>4</sup> Σ <sub>g</sub> <sup>-</sup> | 0.892           | 0.392           | 7.263           | 1.088         |
| <sup>4</sup> Δ <sub>g</sub>              | 1.466           | 0.419           | 6.663           | 1.129         |
| <sup>2</sup> Π <sub>g</sub>              | 0.861           | 0.378           | 7.295           | 1.079         |
| <sup>2</sup> Φ <sub>u</sub>              | 1.503           | 0.127           | 7.472           | 0.724         |

tion in the <sup>6</sup>A<sub>2</sub> and <sup>6</sup>B<sub>1</sub> state since minima in these PES arise although correlating into Ir (<sup>4</sup>F) + H (<sup>2</sup>S) + H (<sup>2</sup>S) there is a nearby Ir (<sup>6</sup>D) state arising from the 5d<sup>7</sup>6s6*p* configuration.

In general, linear electronic states of a given spin multiplicity have enhanced 6*p* character compared to the bent states of the same spin multiplicity. This is in accord with the strong participation of the 6*p* orbital in the 1σ<sub>u</sub> orbital in overlapping with H (1*s*) orbitals.

The Ir-H overlaps are significantly larger for the bent <sup>2</sup>A<sub>1</sub>, <sup>2</sup>B<sub>1</sub> and <sup>2</sup>A<sub>2</sub> states compared to the <sup>6</sup>A<sub>2</sub> and <sup>6</sup>B<sub>1</sub> (0.23–0.28) states. The Ir-H overlaps of these doublet states are also larger than the overlaps for the quartet states. This is consistent with the bond lengths of these states. The <sup>6</sup>A<sub>2</sub> and <sup>6</sup>B<sub>1</sub> states form a shallow minima. In general, high spin states have smaller overlaps and therefore longer bond lengths compared to the low-spin doublet states of IrH<sub>2</sub>.

### 4. Relativistic CI compositions and spin-orbit mixings

Table XV shows the composition of RCI wave functions including spin-orbit coupling near the minima of three lowest states of IrH<sub>2</sub>. We have already discussed the spin-orbit coupling matrix elements and the composition of RCI wave functions near quartet-doublet crossings. As seen from Table XV, the <sup>2</sup>A<sub>1</sub> state in the absence of spin-orbit coupling becomes 63% <sup>2</sup>A<sub>1</sub>, 16% <sup>2</sup>B<sub>1</sub> and 17% <sup>2</sup>A<sub>2</sub>. As seen from Fig. 1, near the well of the <sup>2</sup>A<sub>1</sub> state of IrH<sub>2</sub>, the <sup>2</sup>B<sub>1</sub> and <sup>2</sup>A<sub>2</sub> states cross and consequently, they make almost the same contribution. Hence the spin-orbit contamination is very large for IrH<sub>2</sub>. Therefore we designate the lowest spin-orbit state as the *E*(I) state due to large spin-orbit contamination. The *E*(II) state which is predominantly <sup>2</sup>A<sub>2</sub> is 65% <sup>2</sup>A<sub>2</sub>, 29% <sup>2</sup>A<sub>1</sub> and only 2% <sup>2</sup>B<sub>1</sub> near its minimum. The *E*(III) state is 65% <sup>2</sup>B<sub>1</sub> and 29% <sup>2</sup>A<sub>2</sub>.

The large spin-orbit mixings discussed above not only change the energy separations of electronic states of IrH<sub>2</sub> significantly but also induce substantial changes in the bond angles (up to 9°) in the *E*(III) state. Note that the contribution of the <sup>2</sup>B<sub>2</sub> state is negligible since this state is much higher in energy.

The large spin-orbit contamination of the three low-lying states will also have significant impact on the dipole moments of these states. Since the *E*(I) state is 63% <sup>2</sup>A<sub>1</sub>,

TABLE XV. The contributions of the various states in the RCI of IrH<sub>2</sub>.

| Relativistic state                             | Weights in percent   |
|--|--|
| <i>E</i> (I) ( <sup>2</sup> A <sub>1</sub> )   | 1a <sub>1</sub> <sup>2</sup> 2a <sub>1</sub> <sup>2</sup> 3a <sub>1</sub> <sup>1</sup> 1b <sub>2</sub> <sup>2</sup> 1b <sub>2</sub> <sup>2</sup> 1a <sub>2</sub> <sup>2</sup> ( <sup>2</sup> A <sub>1</sub> , 63%) |
|  | 1a <sub>1</sub> <sup>2</sup> 2a <sub>1</sub> <sup>2</sup> 3a <sub>1</sub> <sup>1</sup> 1b <sub>2</sub> <sup>2</sup> 1b <sub>1</sub> <sup>1</sup> 1a <sub>2</sub> <sup>2</sup> ( <sup>2</sup> B <sub>1</sub> , 16%) |
|  | 1a <sub>1</sub> <sup>2</sup> 2a <sub>1</sub> <sup>2</sup> 3a <sub>1</sub> <sup>1</sup> 1b <sub>2</sub> <sup>2</sup> 1b <sub>2</sub> <sup>2</sup> 1a <sub>2</sub> <sup>1</sup> ( <sup>2</sup> A <sub>2</sub> , 17%) |
| <i>E</i> (II) ( <sup>2</sup> A <sub>2</sub> )  | 1a <sub>1</sub> <sup>2</sup> 2a <sub>1</sub> <sup>2</sup> 3a <sub>1</sub> <sup>1</sup> 1b <sub>2</sub> <sup>2</sup> 1b <sub>2</sub> <sup>2</sup> 1a <sub>2</sub> <sup>2</sup> ( <sup>2</sup> A <sub>2</sub> , 65%) |
|  | 1a <sub>1</sub> <sup>2</sup> 2a <sub>1</sub> <sup>2</sup> 3a <sub>1</sub> <sup>1</sup> 1b <sub>2</sub> <sup>2</sup> 1b <sub>1</sub> <sup>2</sup> 1a <sub>2</sub> <sup>2</sup> ( <sup>2</sup> A <sub>1</sub> , 29%) |
|  | 1a <sub>1</sub> <sup>2</sup> 2a <sub>1</sub> <sup>2</sup> 3a <sub>1</sub> <sup>1</sup> 1b <sub>2</sub> <sup>2</sup> 1b <sub>1</sub> <sup>1</sup> 1a <sub>2</sub> <sup>2</sup> ( <sup>2</sup> B <sub>1</sub> , 2%)  |
| <i>E</i> (III) ( <sup>2</sup> B <sub>1</sub> ) | 1a <sub>1</sub> <sup>2</sup> 2a <sub>1</sub> <sup>2</sup> 3a <sub>1</sub> <sup>1</sup> 1b <sub>2</sub> <sup>2</sup> 1b <sub>1</sub> <sup>1</sup> 1a <sub>2</sub> <sup>2</sup> ( <sup>2</sup> B <sub>1</sub> , 65%) |
|  | 1a <sub>1</sub> <sup>2</sup> 2a <sub>1</sub> <sup>2</sup> 3a <sub>1</sub> <sup>1</sup> 1b <sub>2</sub> <sup>2</sup> 1b <sub>2</sub> <sup>1</sup> 1a <sub>2</sub> <sup>2</sup> ( <sup>2</sup> A <sub>2</sub> , 29%) |
|  | 1a <sub>1</sub> <sup>2</sup> 2a <sub>1</sub> <sup>2</sup> 3a <sub>1</sub> <sup>1</sup> 1b <sub>2</sub> <sup>2</sup> 1b <sub>1</sub> <sup>2</sup> 1a <sub>2</sub> <sup>2</sup> ( <sup>2</sup> A <sub>1</sub> , 1%)  |

16%  $^2B_1$ , and 17%  $^2A_2$  and since the  $\mu_e$  of  $^2B_1$  is significantly smaller, we estimate the  $\mu_e$  of the ground state of IrH<sub>2</sub><sup>+</sup> to be reduced to 2.0 D based on a simple weighted average of the  $\mu_e$ s of the three doublet states as determined by the RCI composition. The dipole moment of the  $E(III)$  state to the contrary will increase.

## B. IrH<sub>2</sub><sup>+</sup>

### 1. The CI compositions and the nature of PES and the origins of saddle points in PES

Table XVI shows the leading configurations in the MRSDCI wave functions of bent electronic states of IrH<sub>2</sub><sup>+</sup>. All four triplet states of IrH<sub>2</sub><sup>+</sup> are well described by a single leading configuration. With the exception of  $^1A_1$  all other singlet states are well represented by their leading configurations. The  $^1A_1$  state is, however, a mixture of  $1a_1^2 2a_1^2 1b_2^2 1b_1^2 1a_2^2$ ,  $1a_1^2 2a_1^2 3a_1^2 1b_2^2 1b_1^2$ , and  $1a_1^2 3a_1^2 1b_2^2 1b_1^2 1a_2^2$  configurations.

The saddle points in the  $^5A_2$  state and  $^5B_2$  state are somewhat narrow while the  $^5A_1$  and  $^5B_1$  states have relatively broad barriers. This contrast needs an explanation. Also the  $^5A_2$  state has a sharp barrier followed by a minimum and another saddle point prior to the formation of a linear minimum (Fig. 2). These features in the  $^5A_2$  PES are attributed to several avoided crossings.

The composition of the  $^5A_2$  state at  $\theta = 20^\circ$  is 98%  $1a_1^2 2a_1^2 3a_1 4a_1 1b_2 1b_1 1a_2^2$ . At  $\theta = 40^\circ$  this state becomes 93%  $1a_1^2 2a_1^2 3a_1 4a_1 1b_2 1b_1 1a_2^2$  and 5%  $1a_1^2 2a_1 3a_1 1b_2 2b_2 1b_1 1a_2^2$ . At  $\theta = 60^\circ$ , the  $^5A_2$  state becomes 93%  $1a_1^2 2a_1 3a_1 4a_1 1b_2^2 1b_1^2 1a_2$  and 2%  $1a_1^2 2a_1^2 3a_1 4a_1 1b_2 1b_1 1a_2^2$ . Note that at this angle this state completely changes its composition. The avoided crossing between the  $1a_1^2 2a_1 3a_1 4a_1 1b_2^2 1b_1^2 1a_2$  and  $1a_1^2 2a_1^2 3a_1 4a_1 1b_2 1b_1 1a_2^2$  configurations of IrH<sub>2</sub><sup>+</sup> in the  $^5A_2$  state lead to the narrow barrier followed by a minimum. At  $\theta = 80^\circ$ , this state undergoes yet another avoided crossing since its composition changes to 93%  $1a_1^2 2a_1 3a_1 1b_2^2 2b_2 1b_1 1a_2^2$ . This configuration continues to dominate for  $\theta > 100^\circ$  also. Consequently, the double hump

in the  $^5A_2$  PES of IrH<sub>2</sub><sup>+</sup> is clearly a manifestation of these avoided crossings.

The  $^5B_1$  state is predominantly  $1a_1^2 2a_1^2 3a_1 4a_1 1b_2 1b_1^2 1a_2$  at  $\theta = 20^\circ$  and  $40^\circ$ . At  $\theta = 60^\circ$ , however, it becomes 65%  $1a_1^2 2a_1^2 3a_1 4a_1 1b_2 1b_1^2 1a_2$  and 17%  $1a_1^2 2a_1 3a_1 1b_2^2 2b_2 1b_1^2 1a_2$ . At  $\theta = 80^\circ$ , this state changes to 55%  $1a_1^2 2a_1 3a_1 1b_2^2 2b_2 1b_1^2 1a_2$  and 31%  $1a_1^2 2a_1^2 3a_1 4a_1 1b_2 1b_1^2 1a_2$ . The saddle point in the  $^5B_1$  state therefore arises from this avoided crossing. At  $\theta = 100^\circ$  and larger angles the  $1a_1^2 2a_1 3a_1 1b_2^2 2b_2 1b_1^2 1a_2$  configuration dominates in the  $^5B_1$  state.

The  $^5B_2$  state is predominantly  $1a_1^2 2a_1^2 3a_1 4a_1 1b_2^2 1b_1 1a_2$  for  $10^\circ < \theta < 100^\circ$ . At  $\theta = 120^\circ$  it becomes 84%  $1a_1^2 2a_1^2 3a_1^2 1b_2 2b_2 1b_1 1a_2$  and only 8%  $1a_1^2 2a_1^2 3a_1 4a_1 1b_2^2 1b_1 1a_2$ . Therefore the saddle point and the shape of the PES of the  $^5B_2$  state is due to this avoided crossing since this occurs between  $\theta = 100^\circ$  and  $120^\circ$ . Note that the saddle point in this PES occurs in this region. At larger bond angles the  $1a_1^2 2a_1^2 3a_1^2 1b_2 2b_2 1b_1 1a_2$  configuration dominates.

The  $^5A_1$  state is dominated by the  $1a_1^2 2a_1^2 3a_1^2 4a_1 1b_2 1b_1 1a_2$  configuration for  $10^\circ < \theta < 40^\circ$ . At  $\theta = 60^\circ$  this state becomes 89%  $1a_1^2 2a_1^2 3a_1 1b_2^2 2b_2 1b_1 1a_2$  configuration. Consequently, the saddle point in this region is due to this avoided crossing. For larger bond angles the  $1a_1^2 2a_1^2 3a_1 1b_2^2 2b_2 1b_1 1a_2$  configuration dominates.

The above analysis of the composition of the CI wave functions reveals that all saddle points in the quintet surfaces arise from avoided crossings. The general trend is that at small angles only the  $1b_2$  orbital which is predominantly Ir ( $d$ ) is occupied. At larger bond angles the occupancy of the  $2b_2$  orbital becomes more significant as the overlap between Ir ( $5d\pi$ ), Ir ( $6p_y$ ) (smaller amount) and H<sub>2</sub> ( $1s$ )–H<sub>2</sub> ( $1s$ ) orbital becomes larger as  $\theta$  increases. This minimizes the antibonding interactions through favorable bonding interaction between Ir ( $5d\pi$ ), Ir ( $6p_y$ ), and H<sub>2</sub> ( $1\sigma_u$ ) orbital. In those states for which this interaction occurs more readily, the dissociation of H<sub>2</sub> is facilitated. In general, low spin states readily rearrange to form strong Ir ( $5d\pi$ ) + Ir( $6p_y$ ) + H<sub>2</sub> ( $1\sigma_u$ ) bonding interaction leading to the dissociation of H<sub>2</sub>.

TABLE XVI. Leading configurations in the MRSDCI of the bent states of IrH<sub>2</sub><sup>+</sup>.

| State   | Coefficient | Configuration   |                 |                 |                 |                 |                 |                 |                 |
|---------|-------------|-----------------|-----------------|-----------------|-----------------|-----------------|-----------------|-----------------|-----------------|
|         |             | 1a <sub>1</sub> | 2a <sub>1</sub> | 3a <sub>1</sub> | 4a <sub>1</sub> | 1b <sub>2</sub> | 2b <sub>2</sub> | 1b <sub>1</sub> | 1a <sub>2</sub> |
| $^3A_2$ | 0.958       | 2               | 2               | 1               | 0               | 2               | 0               | 2               | 1               |
| $^3B_1$ | 0.955       | 2               | 2               | 1               | 0               | 2               | 0               | 1               | 2               |
| $^3A_1$ | 0.959       | 2               | 1               | 1               | 0               | 2               | 0               | 2               | 2               |
| $^3B_2$ | −0.956      | 2               | 2               | 2               | 0               | 2               | 0               | 1               | 1               |
| $^1A_1$ | −0.799      | 2               | 2               | 0               | 0               | 2               | 0               | 2               | 2               |
|         | 0.414       | 2               | 2               | 2               | 0               | 2               | 0               | 2               | 0               |
|         | 0.324       | 2               | 0               | 2               | 0               | 2               | 0               | 2               | 2               |
| $^1A_2$ | 0.955       | 2               | 2               | 1               | 0               | 2               | 0               | 2               | 1               |
| $^1B_1$ | −0.955      | 2               | 2               | 1               | 0               | 2               | 0               | 1               | 2               |
| $^1B_2$ | −0.939      | 2               | 2               | 2               | 0               | 2               | 0               | 1               | 1               |
| $^5B_1$ | 0.961       | 2               | 1               | 1               | 0               | 2               | 1               | 2               | 1               |
| $^5A_2$ | −0.955      | 2               | 1               | 1               | 0               | 2               | 1               | 1               | 2               |
| $^5A_1$ | 0.963       | 2               | 2               | 1               | 0               | 2               | 1               | 1               | 1               |

### 2. Mulliken populations

Table XVII shows the Ir ( $6s$ ,  $6p$  and  $5d$ ) Mulliken populations and the total Ir–H overlap populations for various electronic state of IrH<sub>2</sub><sup>+</sup>. As seen from Table XVII, in the  $^3A_2$  ground state the Ir<sup>+</sup> exists in an effective configuration of Ir  $6s^{0.62} 6p^{0.12} 5d^{7.5}$  as compared to the  $^2A_1$  ground state of IrH<sub>2</sub> with an effective population of  $6s^{0.75} 6p^{0.12} 5d^{8.1}$ . Consequently, significant electronic density is lost by the  $5d$  orbital upon ionization of IrH<sub>2</sub>. The  $6p$  population is nearly the same between neutral and the ionic species. The Ir–H overlaps of neutral and ionic species are also similar.

The quintet states have somewhat larger  $6p$  populations compared to the triplet and singlet states. The  $^1A_1$  state has a much smaller  $6s$  population and a slightly larger  $5d$  population. Therefore in the  $^1A_1$  state the Ir<sup>+</sup> ion exhibits  $5d^6 6s^1$  and  $5d^7$  configurational mixing.

TABLE XVII. Mulliken population analyses for the electronic states of IrH<sub>2</sub><sup>+</sup>.

| State                       | Gross  |        |        | Ir-H<br>Overlaps |
|-----------------------------|--------|--------|--------|------------------|
|                             | Ir (s) | Ir (p) | Ir (d) |                  |
| <sup>3</sup> A <sub>2</sub> | 0.622  | 0.121  | 7.509  | 1.122            |
| <sup>3</sup> B <sub>1</sub> | 0.700  | 0.126  | 7.405  | 1.154            |
| <sup>3</sup> A <sub>1</sub> | 0.555  | 0.109  | 7.572  | 1.018            |
| <sup>3</sup> B <sub>2</sub> | 1.081  | 0.130  | 7.012  | 1.169            |
| <sup>1</sup> A <sub>1</sub> | 0.454  | 0.108  | 7.692  | 1.047            |
| <sup>1</sup> A <sub>2</sub> | 0.614  | 0.130  | 7.508  | 1.130            |
| <sup>1</sup> B <sub>1</sub> | 0.616  | 0.127  | 7.487  | 1.140            |
| <sup>1</sup> B <sub>2</sub> | 1.076  | 0.138  | 7.027  | 1.161            |
| <sup>5</sup> B <sub>1</sub> | 0.719  | 0.323  | 6.987  | 1.071            |
| <sup>5</sup> A <sub>2</sub> | 0.770  | 0.302  | 7.001  | 1.043            |
| <sup>5</sup> A <sub>1</sub> | 0.786  | 0.325  | 6.919  | 1.076            |

The sum of two Ir-H overlap populations of all states are > 1.0 indicating significant Ir-H bonding. This corroborates with shorter Ir-H bond lengths for the triplet and singlet states of IrH<sub>2</sub><sup>+</sup> but somewhat longer bond lengths for the quintet states. The quintet states have longer bond lengths since Ir<sup>+</sup> high spin state does not form strong bond pairs with H (1s) orbitals.

### 3. Composition of the relativistic CI wave functions and the effect of spin-orbit coupling

Table XVIII shows the weights of various configurations in the RCI wave functions of the electronic states of IrH<sub>2</sub><sup>+</sup>. As seen from Table XVIII, various spin-orbit components of IrH<sub>2</sub><sup>+</sup> are heavily contaminated approaching an intermediate coupling. The <sup>3</sup>A<sub>2</sub> (A<sub>1</sub>) ground state of IrH<sub>2</sub><sup>+</sup> is 63% <sup>3</sup>A<sub>2</sub>, 15% <sup>3</sup>B<sub>2</sub>, 12% <sup>3</sup>B<sub>1</sub> and 7% <sup>1</sup>A<sub>1</sub>. This large spin-orbit contamination increases the H-Ir-H bond angle of the ground state of IrH<sub>2</sub><sup>+</sup> by 9.3°. The increase in bond angle is consistent with the larger bond angle of the <sup>3</sup>B<sub>1</sub> and <sup>3</sup>B<sub>2</sub> states which are primary contaminants.

The <sup>3</sup>A<sub>2</sub> (B<sub>1</sub>) state cannot mix with <sup>3</sup>B<sub>1</sub> since there is no spin-orbit state of the B<sub>1</sub> symmetry. Therefore the bond angle of the B<sub>1</sub> component changes by only 3°. The <sup>3</sup>A<sub>2</sub> (B<sub>2</sub>) component is contaminated by only 3% <sup>3</sup>B<sub>1</sub> (B<sub>2</sub>). The <sup>3</sup>B<sub>1</sub>

(B<sub>2</sub>) is further high. Consequently, the bond angle of <sup>3</sup>A<sub>2</sub> (B<sub>2</sub>) changes by only 1°.

The <sup>3</sup>B<sub>1</sub> (A<sub>1</sub>) spin-orbit state is heavily mixed akin to the <sup>3</sup>A<sub>2</sub> (A<sub>1</sub>) state. Therefore the bond angle of <sup>3</sup>B<sub>1</sub> (A<sub>1</sub>) state reduces by ~9° relative to <sup>3</sup>B<sub>1</sub> (NOSO) due to contamination with <sup>3</sup>A<sub>2</sub>. Therefore we conclude that spin-orbit contaminations of different states of IrH<sub>2</sub><sup>+</sup> are very significant.

### V. COMPARISON OF IrH<sub>2</sub> WITH RhH<sub>2</sub>

In a previous study one of the authors and Liao<sup>34</sup> investigated the PES of RhH<sub>2</sub> arising from Rh (<sup>4</sup>F) + H<sub>2</sub> and Rh (<sup>2</sup>F) + H<sub>2</sub>. They found that the Rh (<sup>2</sup>F) atom inserts into H<sub>2</sub> spontaneously while the Rh (<sup>4</sup>F) state has to surmount a large barrier of 56 kcal/mol. The ground state of RhH<sub>2</sub> was found to be a <sup>2</sup>A<sub>1</sub> state with *r*<sub>e</sub> = 1.51 Å and θ<sub>e</sub> = 83°. The <sup>2</sup>A<sub>1</sub> state was found to be 10 kcal/mol more stable than Rh (<sup>4</sup>F) + H<sub>2</sub>. The spin-orbit contaminations of different states of RhH<sub>2</sub> were found to be quite small.

A few dramatic differences are noticed in comparing RhH<sub>2</sub> with IrH<sub>2</sub>. First, the barrier for the insertion of Rh (<sup>4</sup>F) into H<sub>2</sub> (56 kcal/mol) is substantially larger than the barrier for the insertion of Ir (<sup>4</sup>F) into H<sub>2</sub> (~28 kcal/mol). The spin-orbit coupling of the quartet and doublet states near the crossing region provide nonzero transition probability for the quartet states to hop to the doublet states and vice versa thereby reducing the barrier further to less than 4 kcal/mol while such spin-orbit coupling was found to be very small for RhH<sub>2</sub>. Consequently, the Rh atom is found to be much less reactive to H<sub>2</sub> compared to the Ir atom.

The stability of IrH<sub>2</sub> (<sup>2</sup>A<sub>1</sub>) relative to Ir (<sup>4</sup>F) + H<sub>2</sub> is 28 kcal/mol in the absence of spin-orbit coupling compared to only 10 kcal/mol for RhH<sub>2</sub> although the latter value is at the CAS level. Even if improvements from higher-order correlations are taken into consideration for RhH<sub>2</sub> up to 7 kcal/mol, evidently, IrH<sub>2</sub> is considerably more stable than RhH<sub>2</sub>. Note that the IrH<sub>2</sub> E state is 27 kcal/mol relative to Ir (<sup>4</sup>F<sub>9/2</sub>) + H<sub>2</sub> including spin-orbit coupling. The enhanced stability of IrH<sub>2</sub> relative to Ir + H<sub>2</sub> would make it a more attractive candidate for experimental studies since it will have long enough lifetimes to be seen in a supersonic beam.

The spin-orbit effects are substantially larger for IrH<sub>2</sub> compared to RhH<sub>2</sub>. The spin-orbit contamination of different states of RhH<sub>2</sub> was found to be at most 1% while the spin-orbit contamination for IrH<sub>2</sub> was found to add up to 25%–30%. Hence, spin-orbit effects are dramatically larger for IrH<sub>2</sub>. As noted before, a major contribution of spin-orbit coupling is to the H-Ir-H bond angle of the ground state which changes by 9°.

The dipole moment of the ground state of RhH<sub>2</sub> is 2.67 D with Rh<sup>+</sup>H<sup>-</sup> polarity compared to 2.2 D for IrH<sub>2</sub> with the same polarity. The dipole moments of the two species are comparable since the M-H bond lengths for IrH<sub>2</sub> and RhH<sub>2</sub> are 1.545 and 1.54 Å, respectively. The <sup>2</sup>A<sub>1</sub> state of RhH<sub>2</sub> has a smaller bond angle (84°) compared to IrH<sub>2</sub> (92°). But this difference seems to compensate for the more ionicity of the Ir-H bonds so that the dipole moments become comparable. However, the spin-orbit effects have larger impact on the dipole moment of IrH<sub>2</sub> compared to RhH<sub>2</sub>.

TABLE XVIII. The contributions of the various states in the RCI of IrH<sub>2</sub><sup>+</sup>.

| Relativistic state                            | Weight in percent  |
|---|--|
| <sup>3</sup> A <sub>2</sub> (A <sub>1</sub> ) | 63% <sup>3</sup> A <sub>2</sub> , 15% <sup>3</sup> B <sub>2</sub> , 12% <sup>3</sup> B <sub>1</sub> , 7% <sup>1</sup> A <sub>1</sub>     |
| <sup>3</sup> A <sub>2</sub> (B <sub>1</sub> ) | 85% <sup>3</sup> A <sub>2</sub> , 14% <sup>3</sup> B <sub>2</sub>  |
| <sup>3</sup> A <sub>2</sub> (B <sub>2</sub> ) | 94% <sup>3</sup> A <sub>2</sub> , 2.5% <sup>3</sup> B <sub>1</sub> , 2.2% <sup>1</sup> B <sub>2</sub>                                    |
| <sup>3</sup> B <sub>1</sub> (A <sub>1</sub> ) | 60% <sup>3</sup> B <sub>1</sub> , 26% <sup>3</sup> A <sub>2</sub> , 11% <sup>1</sup> A <sub>1</sub>                                      |
| <sup>3</sup> B <sub>1</sub> (B <sub>2</sub> ) | 90% <sup>3</sup> B <sub>1</sub> , 5.3% <sup>1</sup> B <sub>2</sub> , 3.4% <sup>3</sup> A <sub>2</sub>                                    |
| <sup>3</sup> B <sub>1</sub> (A <sub>2</sub> ) | 96% <sup>3</sup> B <sub>1</sub> , 3% <sup>1</sup> A <sub>2</sub>   |
| <sup>3</sup> B <sub>2</sub> (A <sub>2</sub> ) | 96% <sup>3</sup> B <sub>2</sub> , 0.01% <sup>1</sup> A <sub>2</sub> , 0.01% <sup>3</sup> B <sub>1</sub>                                  |
| <sup>3</sup> B <sub>2</sub> (B <sub>1</sub> ) | 75% <sup>3</sup> B <sub>2</sub> , 14% <sup>3</sup> A <sub>2</sub> , 8.6% <sup>1</sup> B <sub>1</sub>                                     |
| <sup>3</sup> B <sub>2</sub> (A <sub>1</sub> ) | 79% <sup>3</sup> B <sub>2</sub> , 8.5% <sup>1</sup> A <sub>1</sub> , 7.9% <sup>3</sup> A <sub>2</sub> , 2.3% <sup>3</sup> B <sub>1</sub> |

We conclude that the main difference between RhH<sub>2</sub> and IrH<sub>2</sub> is in the reactivity and stability. The Ir atom is more reactive and forms more stable Ir–H bonds in the ground state of IrH<sub>2</sub>. The other important contrasts are in the dramatically larger spin–orbit effects which change the bond angles and enhance the nonadiabatic transition probability for the insertion of Ir (<sup>4</sup>F) state into H<sub>2</sub>.

## VI. CONCLUSION

The PES of IrH<sub>2</sub> and IrH<sub>2</sub><sup>+</sup> arising from Ir (<sup>4</sup>F) + H<sub>2</sub>, Ir (<sup>2</sup>F) + H<sub>2</sub>, Ir (<sup>4</sup>F) + H + H, Ir<sup>+</sup> (<sup>5</sup>D) + H<sub>2</sub>, Ir<sup>+</sup> (<sup>3</sup>F) + H<sub>2</sub>, and Ir<sup>+</sup> (<sup>1</sup>G) + H<sub>2</sub> dissociation limits were studied using a complete active space MCSCF (CAS-MCSCF) followed by MRCI and RCI calculations. The ground state of IrH<sub>2</sub> including spin–orbit effects is found to be a *E*(I) (<sup>2</sup>A<sub>1</sub>) state with *r*<sub>e</sub> = 1.54 Å, *θ*<sub>e</sub> = 92° and *D*<sub>e</sub> = 27 kcal/mol relative to Ir (<sup>4</sup>F<sub>9/2</sub>) + H<sub>2</sub>. The Ir (<sup>4</sup>F) atom has to surmount a barrier for insertion into H<sub>2</sub> while the Ir (<sup>2</sup>F) atom spontaneously inserts into H<sub>2</sub> in the absence of spin–orbit coupling. However, spin–orbit coupling of the quartet and doublet surfaces near the curve crossing region reduced the barrier for insertion of Ir (<sup>4</sup>F) into H<sub>2</sub> substantially to 5 kcal/mol through nonvanishing nonadiabatic transition probability facilitated by the spin–orbit coupling. Therefore, we find Ir (<sup>4</sup>F) to be more reactive through spin–orbit coupling.

The spin–orbit effects were found to be significant for both IrH<sub>2</sub> and IrH<sub>2</sub><sup>+</sup>. The *E*(I) (<sup>2</sup>A<sub>1</sub>) ground state of IrH<sub>2</sub> was found to be 63% <sup>2</sup>A<sub>1</sub>, 16% <sup>2</sup>B<sub>1</sub>, and 17% <sup>2</sup>A<sub>2</sub> while the <sup>3</sup>A<sub>2</sub> (*A*<sub>1</sub>) ground state of IrH<sub>2</sub><sup>+</sup> was found to be 63% <sup>3</sup>A<sub>2</sub>, 15% <sup>3</sup>B<sub>2</sub>, 12% <sup>3</sup>B<sub>1</sub>, and 7% <sup>1</sup>A<sub>1</sub>. This large spin–orbit mixing induces dramatic changes in the ground state *θ*<sub>e</sub>'s of the two species (9.5°).

The Ir<sup>+</sup> (<sup>5</sup>D) state has to surmount a barrier for insertion into H<sub>2</sub> in the absence of spin–orbit coupling while both Ir<sup>+</sup> (<sup>3</sup>F) and Ir<sup>+</sup> (<sup>1</sup>G) states insert into H<sub>2</sub> spontaneously. The crossing of quintet surfaces with triplet surfaces reduce the barrier for Ir<sup>+</sup> (<sup>5</sup>D) into H<sub>2</sub> through spin–orbit coupling of the two surfaces significantly to ~5 kcal/mol, thereby providing nonzero nonadiabatic transition probability for the ground state of Ir<sup>+</sup> to insert into H<sub>2</sub>. The IrH<sub>2</sub><sup>+</sup> ion in its <sup>3</sup>A<sub>2</sub> (*A*<sub>1</sub>) ground state was found to be 30 kcal/mol more stable than Ir<sup>+</sup> + H<sub>2</sub>.

The saddle points in some of the quartet states of IrH<sub>2</sub> and all quintet states of IrH<sub>2</sub><sup>+</sup> were found to arise from several avoided crossings. Both IrH<sub>2</sub> and IrH<sub>2</sub><sup>+</sup> states exhibit complex 6s6p5d hybridization as evidenced from their Mulliken populations.

It is concluded that the Ir atom reacts more readily with H<sub>2</sub> compared to Rh. The stability of IrH<sub>2</sub> is also considerably more than RhH<sub>2</sub>.

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