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Citation: *The Journal of Chemical Physics* **15**, 631 (1947); doi: 10.1063/1.1746620

View online: <http://dx.doi.org/10.1063/1.1746620>

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## Contribution to the Study of Electronic Spectra of Bent Triatomic Molecules

JULES DUCHESNE\* AND B. ROSEN

*Department of Chemical Physics and Institute of Astrophysics, University of Liège, Liège, Belgium*

(Received June 9, 1947)

The electronic absorption spectra of the  $\text{SO}_2$ ,  $\text{SeO}_2$ , and  $\text{TeO}_2$  molecules have been studied under widely varied experimental conditions. The vibrational analysis of the ultraviolet absorption region of  $\text{SO}_2$  with center of gravity near 2000Å reveals that it is formed by at least two electronic systems with the following vibrational constants:

System  $\alpha_1 \cdots \nu_0 = 42,170 \text{ cm}^{-1}$ ,  $\nu_1' = 963 \text{ cm}^{-1}$ ,  $\nu_2' = 379 \text{ cm}^{-1}$ ,  $x'\nu_2' = 2 \text{ cm}^{-1}$ .

System  $\alpha_2 \cdots \nu_0 = 44,236 \text{ cm}^{-1}$ ,  $\nu_1' = 775 \text{ cm}^{-1}$ ,  $\nu_2' = 375 \text{ cm}^{-1}$ .

The presence in the same region of a third electronic system, with  $\nu_2' = 360 \text{ cm}^{-1}$  and  $\nu_1' = 845 \text{ cm}^{-1}$ , seems highly probable.

The character of the vibrational structure of the band systems and a preliminary analysis of the rotational structure of several bands have enabled us to obtain some information about the molecular structure of the excited states.

The vibrational analysis of the  $c$ -system of  $\text{SeO}_2$  in the visible, and the  $\text{TeO}_2$  bands in the visible and ultraviolet, have enabled us to correlate the electronic states of the three similar molecules.

## INTRODUCTION

A CONSIDERABLE amount of work has been done in the last few years, chiefly by Mulliken and his collaborators, in the field of electronic spectra of triatomic molecules. During the war we were engaged on similar problems and started investigations on the electronic spectra of the  $\text{SO}_2$ ,  $\text{SeO}_2$ , and  $\text{TeO}_2$  molecules. The wartime conditions prevented us from publishing our results, with the exception of one paper on the  $b$ -system of  $\text{SeO}_2$ .<sup>1</sup> The present paper deals with the analysis of the  $a$ -system of  $\text{SO}_2$ , in the far ultraviolet (center of gravity near 2000Å), the  $c$ -system of  $\text{SeO}_2$  in the visible, and with  $\text{TeO}_2$  bands in the visible and near ultraviolet.

## HISTORICAL CONSIDERATIONS

The  $a$ -system of  $\text{SO}_2$  was first studied in absorption by V. Henri and by Wieland.<sup>2</sup> They correlated the observed bands with three fundamental frequencies of the ground level and assigned the frequencies 960 and  $379.5 \text{ cm}^{-1}$  to the upper level. In 1933, Chow<sup>3</sup> performed a new

analysis and suggested that the bands belong to three electronic systems, each of them being characterized by a frequency of about  $750\text{--}770 \text{ cm}^{-1}$  in the upper level.

Chow's analysis was criticized in 1938 by Price and Simpson<sup>4</sup> who considered that all the most prominent bands in the region  $2300\text{--}2000\text{Å}$  form a single progression with  $\nu_2' = 380 \text{ cm}^{-1}$ .

The  $a$ -system of  $\text{SO}_2$  was observed by Chow<sup>3</sup> in emission as well as in absorption. Later, Lotmar<sup>5</sup> studied the monochromatically-excited resonance spectra in  $\text{SO}_2$  vapor and Kornfeld<sup>6</sup> extended the emission spectra, using electrodeless excitation.

While Chow and Lotmar interpreted the system using an asymmetrical vibration,  $\nu_3''$ , Kornfeld succeeded in arranging the observed bands using only symmetrical frequencies in both states. In view of the application of the selection rules and the discussion of molecular configurations, these divergencies must be cleared up. On the other hand, the values for the band heads given by different authors are in some cases widely divergent, so that new measures of this system seemed to be of interest. For the present study, we were able to use some spectra taken by V. Henri in 1922 and 1930. The dispersion

\* Associate fellow of the Belgian National Foundation of Scientific Researches.

<sup>1</sup> J. Duchesne and B. Rosen, *Physica* **8**, 540 (1941).

<sup>2</sup> V. Henri, *The Structure of Molecules*, edited by P. Debye (Blackie and Son Limited, London, 1932), p. 121; K. Wieland, *Nature* **130**, 847 (1932).

<sup>3</sup> Tung-Ching Chow, *Phys. Rev.* **44**, 638 (1933).

<sup>4</sup> W. C. Price and D. M. Simpson, *Proc. Roy. Soc. A* **165**, 272 (1938).

<sup>5</sup> W. Lotmar, *Zeits. f. Physik* **83**, 765 (1933).

<sup>6</sup> G. Kornfeld, *Trans. Faraday Soc.* **32**, 1487 (1936).

was either 2A/mm or 5A/mm at 2200A, the pressure varied from 0.5 to 100 mm, and the temperature from  $-80^{\circ}$  to  $500^{\circ}$ . The absorption lengths varied from 5 cm to one meter.

### VIBRATIONAL ANALYSIS

In the first column of Table I we give the mean value for the band heads obtained on a series of plates taken at  $-80^{\circ}\text{C}$  and corresponding to widely varied conditions of pressure and

absorption paths. The dispersion was 2A/mm at 2200A, and the difference between individual measures did not exceed  $3\text{ cm}^{-1}$ . Some additional band heads measured only on low dispersion plates, and marked with an asterisk, are also included. For these bands the precision is not so good. This is also the case for the last bands situated towards the ultraviolet, where accurate measures of our high dispersion plates are no longer possible.

TABLE I.

D. and R.	Chow	D. and R.	Chow	D. and R.	Chow	D. and R.	Chow
{42,170(?) 42,183(?) $\alpha_1$		{44,663 44,677 $\alpha_1$	44,663 44,674	{46,218 46,227 $\alpha_3$	46,217		47,707 47,742
		44,885 $\alpha_1$		46,353 $\alpha_3$	46,348		47,889
{42,576 42,589 $\alpha_1$	42,575 42,589	{44,966* 44,980 $\alpha_2$	44,978 44,987		46,444 46,464		47,922 47,965
	42,759						48,000
{42,955 42,968 $\alpha_1$	42,954 42,966	{45,032 45,046 $\alpha_1$	45,030	{46,471* 46,484 $\alpha_2$	47,472 46,482	48,015* $\alpha_2$	48,005 48,083
	43,018	45,160(?)		46,574 $\alpha_3$	46,570		48,121
{43,138 43,148 $\alpha_1$	43,137 43,145	45,247* $\alpha_1$			46,697		48,355
		{45,337 45,354 $\alpha_2$	45,333 45,351	46,704* $\alpha_3$	46,704	48,390* $\alpha_2$	48,375
{43,332 43,343 $\alpha_1$	43,329 43,340	{45,399 45,413 $\alpha_1$	45,393	46,820* $\alpha_2$	46,818	$\alpha_2$	48,724
	43,407				46,829	$\alpha_2$	48,781
{43,539 43,552 $\alpha_1$	43,538 43,550	{45,499 45,601* $\alpha_1$	45,572	{46,866 46,878 $\alpha_2$	46,865 46,874	$\alpha_2$	49,133 49,477
				46,925* $\alpha_3$		$\alpha_2$	49,550
{43,701 43,713 $\alpha_1$			45,672 45,686	47,057* $\alpha_3$	47,052	$\alpha_2$	49,910
	43,714				47,074	$\alpha_2$	50,248
{43,919 43,930 $\alpha_1$	43,918 43,929	{45,710 45,726 $\alpha_2$	45,706 45,725		47,198 47,223	$\alpha_2$	50,302
44,086 $\alpha_1$	44,085 44,100	45,759 $\alpha_1$	45,757		47,232		
		{45,858 45,869 $\alpha_3$	45,859	47,257* $\alpha_2$	47,246		
44,236 $\alpha_2$	44,236				47,391		
{44,293 44,306 $\alpha_1$	44,293 44,305		45,923 46,077		47,548 47,576		
44,491(?) $\alpha_1$		46,085* $\alpha_2$	46,084		47,623		
44,543(?) $\alpha_2$	44,570 44,603	{46,106 46,122 $\alpha_2$	46,104 46,118	47,627* $\alpha_2$	47,636		

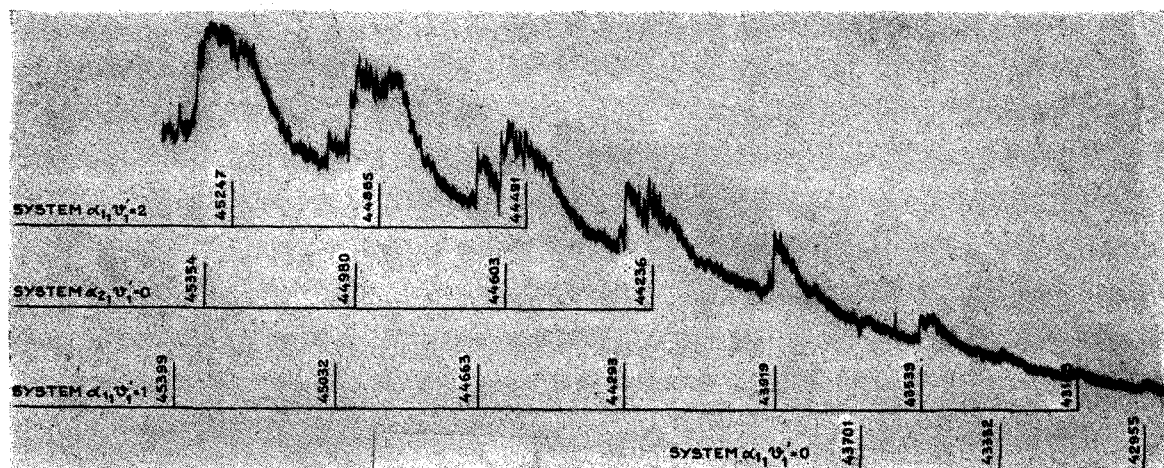


FIG. 1.

The third column of Table I contains the band heads observed by Chow<sup>3</sup> at room temperature. The two sets of values are in general in quite satisfactory agreement.

In 1931, V. Henri<sup>2</sup> pointed out that he had observed 120 band heads between 2445 and 2100Å. A systematic study of our plates reveals, however, that many of these band heads are, in fact, sub-heads in the fine structure. This point will be discussed in detail below. In any case, the bands given in the first column of Table I are by far the most characteristic of the whole system. The analysis of the spectrum shows that the occurrence of the  $380\text{ cm}^{-1}$   $\nu_2'$  frequency is beyond any doubt, but we cannot agree with the statement of Price and Simpson<sup>4</sup> that all the prominent bands are members of a single  $\nu_2'$  progression.

The microphotometer record of Fig. 1, corresponding to the long wave-length end of the spectrum, shows clearly that there are at least three superposed series with characteristic intensity distribution around a single maximum. Moreover, intensity-temperature variations indicate definitely that the observed series cannot arise from excited vibrational levels in the normal state. Thus, the distance between the corresponding members of two series must be correlated either to a frequency (fundamental or combination) of the upper level or to the distance between two neighboring electronic states.

The essential results of our new interpretation are summarized in Tables II to V. In Table II the columns represent  $\nu_2'$  series with  $\nu_2' \sim 380\text{ cm}^{-1}$ . Taking into account the constant difference of about  $963\text{ cm}^{-1}$  between bands of the first

TABLE II. System  $\alpha_1$ .

$\nu_2' \backslash \nu_1'$	0	1	2
0	{42,170 ( ? ) 968 42,183 ( ? ) 965 406	{43,138 (3) 43,148 (2) 402	
1	{42,576 (2) 963 42,589 (1) 963 379	{43,539 (4) 952 43,552 (3) 379	44,491 394
2	{42,955 (3) 964 42,968 (2) 962 376	{43,919 (5) 966 43,930 (4) 375	44,885 362
3	{43,332 (3) 961 43,343 (2) 963 370	{44,293 (5) 954 44,306 (4) 370	45,247 354
4	{43,701 962 43,713 964 373	{44,663 (5) 938 44,677 (4) 369	45,601
5	{ — 44,086 (1) 960	{45,032 (3) 45,046 (2) 367	
6		{45,399 (2) 45,413 (1) 360	
7		{45,759 (1) —	

TABLE III. System  $\alpha_2$ .

$v_2' \backslash v_1'$	0	1	2	3	4	5	6
0	44,236 367						
1	44,603 377						
2	{44,966 44,980 (779) 373}	— 45,759 363					
3	{45,337 (769) 45,354 (768) 372}	{46,106 46,122 363}					
4	{45,710 (761) 45,726 (758) 375}	{46,471 (780) 46,484 395}	47,257 (758) 370	48,015 (766) 375	48,781* (769) 352	49,550* (752) 360	50,302*
5	46,085 (781) 359	{46,866 (755) 46,878}	47,627 (763)	48,390 (743) 334	49,133* (777) 344	49,910* 338	
6	{(46,444)* (46,464)* 356}			48,724* (753)	49,477* (771)	50,248*	
7	46,820						

\* Measured by Chow only.

two columns of Table II, we obtain the second characteristic frequency mentioned by V. Henri and correlated with  $\nu_1'$ . All the values  $963 \pm 380$   $\text{cm}^{-1}$  are of course alternative values for  $\nu_1'$ . All the bands in the first two columns seem to have the same fine structure. For instance, they are all doublets with the same mean separation of about 13  $\text{cm}^{-1}$ .

The first band in the second column of Table II shows irregular behavior inasmuch as its distance from the next one has an abnormally high value of 402  $\text{cm}^{-1}$ . On the other hand, the doublet separation is definitely smaller than for the other bands (it is only 8  $\text{cm}^{-1}$  according to Chow's measure). Both kinds of irregularities result probably from a perturbation affecting the first level of the second series (the first doublet of the first series is too faint and accurate measures are impossible; thus it is not possible to state definitely whether this level is also

perturbed). The other levels seem regularly spaced, indicating an anharmonicity  $x'\nu_2'$  of about 2  $\text{cm}^{-1}$ .

The distribution of intensity in the first two columns of Table II is very regular around a single maximum occurring for  $\nu_2'=2$  or 3. The intensities indicated (between brackets) are rough visual estimations, valid only when comparing bands in the same column. It must be noted that bands of the first column appear only at pressures about 10 times larger than those necessary to observe the bands of the second one.

A third group of bands can be fitted into the scheme of Table II as a  $\nu_1'=2$  series. They are blended by strong bands of other series (see Fig. 1) and it is rather difficult to estimate their intensity. Nevertheless, it seems that the maximum intensity for the bands of Table II is reached for the  $(1'\nu_2'0)$  series, and this seems normal in view of the assumed ratio  $\nu_1'/\nu_1''$

=963/1151. The following series could not be observed because their position coincides with strong bands belonging to another electronic transition.

The first column of Table III contains a very strong series, with nearly the same characteristic frequency difference of about  $375\text{ cm}^{-1}$ . It also consists of doublets but, contrary to the first two series of Table II, the strongest head is situated towards the ultraviolet, and the doublet separation is larger ( $17\text{ cm}^{-1}$  on the average).

The series given in the second column of Table III is not so regular, and at first sight it seems that it could form a continuation of the first one, the bands 46106–46122 following the bands 45710–45726. This arrangement, which would suppose some perturbations, cannot be definitely ruled out. But several arguments can be advanced against it. First of all, the characters of the bands in the first and second columns are somewhat different, the latter being less regular, for instance in regard to the relative intensity in the doublets. The fine structure seems also to be different, and the structures of the bands in the second column are on the whole more complex, probably as a consequence of superpositions of at least two series. A possible confirmation of this point of view is given by the observations made by Chow (at room temperature) of 4 bands at 46444, 46464, 46472, and 46482  $\text{cm}^{-1}$ , instead of two bands 46471–46484 observed by us. These 4 bands can be interpreted as nearly superposed doublets belonging to two series. It is to be noted that the first doublet cannot be interpreted as arising from excited vibrational levels of the lower state. It is included (between brackets) in the first column of Table III.

Assuming the existence of two separate series, it is possible to interpret two additional bands which otherwise cannot be easily fitted in the scheme (bands 46085 and 46820  $\text{cm}^{-1}$ ).

The constant  $\Delta\nu$  difference of about  $770 \pm 370\text{ cm}^{-1}$  between the first two columns of Table III must correspond to the frequency  $\nu_1'$ . We admit with Chow that the most probable value is  $\nu_1' = 770\text{ cm}^{-1}$ . The next columns of Table III contain further bands measured towards the ultraviolet, as well as the most characteristic bands measured by Chow. In this region the characteristic frequency difference of about 770

TABLE IV. System  $\alpha_3$ .

	0	1	2
0	45,499 (854)	46,353 (845)	47,198*
	359		
1	{45,858 (846) 45,869}	46,704 (844)	47,548*
	359		
2	{46,218 (839) 46,227}	47,057 (832)	47,889*
	356		
3	46,574		
	351		
4	46,925		

\* Measured by Chow.

$\text{cm}^{-1}$  is clearly visible on the records given in Chow's paper. It is predominant further to the ultraviolet, according to Herreng,<sup>7</sup> and to not yet published results obtained in our laboratory by P. Migeotte.\*\*

For the system of Table III, the  $\nu_1'$  series are much more developed than for the system of Table II corresponding to the decreasing  $\nu_1'/\nu_1''$  ratio.

TABLE V. Absorption bands of  $\text{SO}_2$  arising from excited vibrational levels of the ground state.

$\nu$	Arising from:	Corresponding band arising from (0''0''0'')	$\Delta\nu$
{41,620 41,633}	(2''0''0'')	{43,919 43,930}	2,299 2,297
41,807	(1''0''0'')	42,955	1,148
{41,987 41,998}	{(1''0''0'') (1''0''0'')}	{43,138 43,149}	1,151 1,151
{42,059 42,073}	{(0''1''0'') (0''1''0'')}	{42,576 42,589}	517 516
{42,249 42,387}	{(1''1''0'') (1''0''0'')}	43,919 43,539	1,670 1,152
{42,766 42,776}	{(1''0''0'') (1''0''0'')}	{43,919 43,930}	1,153 1,154
43,020	(0''1''0'')	43,539	519

Mean values:  $\nu_1'' = 1,151.5\text{ cm}^{-1}$  from infra-red\*  $\nu_1' = 1,151.38$   
 $\nu_2'' = 517.3$   $\nu_2' = 517.84$

<sup>7</sup> P. Herreng, Rev. d'optique 15, 413 (1936).

\*\* According to preliminary results of P. Migeotte, it seems probable that a new electronic transition is responsible for bands in the 1900–1800 Å region.

The two schemes of Tables II and III contain nearly all the strong bands observed by us and by Chow, with the exception of the bands denoted  $\alpha_3$  in the Table I. On account of their great relative intensity, it is difficult to correlate the latter with an asymmetrical frequency. For this reason, the possibility that they belong to a third electronic system must be taken into consideration. For this hypothetical third system, the scheme of Table IV with  $\nu_1' = 845$  and  $\nu_2' = 360 \text{ cm}^{-1}$  seems to be the most probable one.

The third column of Table IV corresponds to bands observed by Chow only. They cannot be explained as arising from excited vibrational levels of the ground state. A study further to the ultraviolet is of course necessary before reaching definite conclusions concerning the reality of this system.

Under the conditions of our experiments, bands corresponding to excited vibrational levels of the ground state are practically lacking at  $-80^\circ$ , but they are easily observed at  $20^\circ$ . An accurate measure of these bands is possible only on the long wave-length side of the  $(0'0'0') - (0''0''0'')$  band of the  $\alpha_1$  system, and the mean values obtained for several bands at temperatures ranging up to  $335^\circ$  are given in Table V. For  $\nu_1''$  and  $\nu_2''$  the mean values obtained are in good agreement with the values given by Barker<sup>8</sup> on the basis of infra-red analysis.  $\Delta\nu$  values for the first two bands of Table V indicate a small but definite anharmonicity for  $\nu_1''$ .

#### ROTATIONAL STRUCTURE

The dispersion used did not allow a complete resolution of the fine structure, but some very characteristic features of this structure were observed. For instance, the coarse  $K$ -structure was analyzed for several bands situated at the long wave-length end of the spectrum, where blendings are less frequent than in other parts of the spectrum.

It can be seen easily that the structure of all bands in the first two columns of Table II is similar, thus reinforcing our previous assumption that these bands belong to the same electronic transition. On the other hand, we have observed two series of sub-bands ( $p$  and  $r$ ) in the  $K$ -

structure of several  $\alpha_1$  bands; this indicates that they belong to the perpendicular type. Using the second differences, we obtain  $C' - C'' = -0.7 \pm 0.10 \text{ cm}^{-1}$ . Taking into account the value  $C'' = 1.8 \text{ cm}^{-1***}$  and the fact that the  $K$ -structure is degraded towards the red, it can be stated that  $C' = 1.1 \pm 0.10 \text{ cm}^{-1}$ .

With the dispersion used, the  $J$ -structure remains completely unresolved, but there are indications that the convergence is slow.

For this reason, it can be reasonably admitted that the change of  $B$  correlated with the electronic transition is small; even the sense of the degradation in the  $J$ -structure is not clear, but it seems to be rather degraded in the same direction as the  $K$ -structure, i.e., toward the red. This statement is enhanced by the lack of marked absorption on the short wave-length side of the  $K$ -heads, even at high pressure.

It should be of great interest to resolve the fine structure of the  $\alpha_2$  bands given in the first two columns of Table III. All that can be said for the time being is that it seems to be very different from the structure of the  $\alpha_1$  bands (Table II). Our previous arguments in favor of separating the observed bands into at least two electronic systems are thus reinforced.†

#### INTERPRETATION OF THE RESULTS

The molecular orbitals for  $C_{2v}$  symmetry are distributed among four classes, following the symmetry of the wave functions with respect to the planes of symmetry of the molecule. These classes have been described by Mulliken as  $a_1$ ,  $a_2$ ,  $b_1$ , and  $b_2$ . On this basis and by the application of the molecular orbital theory, Mulliken<sup>9</sup> has proposed a scheme of the electronic levels of angular molecules and particularly of  $\text{SO}_2$ . He has given a preliminary interpretation of the observed electronic spectra of this molecule and

\*\*\* In our previous letter (Nature **157**, 692 (1946)) we had erroneously given the value  $C'' = 1.6 \text{ cm}^{-1}$ . This error does not affect our results.

† It must be stated, however, that the alternative arrangement, according to which the bands of Table II and Table III form a single electronic system, though rather improbable cannot be definitely ruled out. In fact, from a strictly numerical point of view it would be possible to consider the first two columns of Table II and the columns of Table III as consecutive  $v'$  series with  $\nu_1'$  increasing from  $965 \pm 380$  to  $1062 \pm 375$  and reaching  $1150 \pm 370 \text{ cm}^{-1}$  for  $\nu_1' = 2$ .

<sup>9</sup> R. S. Mulliken, Rev. Mod. Phys. **14**, 204 (1942).

<sup>8</sup> E. F. Barker, Rev. Mod. Phys. **14**, 198 (1942).

has considered the  $a$ -system as resulting from an allowed transition  $3b_2 \rightarrow 5a_1$ .

In the first part of this paper we have explained that the " $a$ -system" must be divided into at least 2 electronic transitions corresponding to Tables II and III and described, respectively, as  $\alpha_1$  and  $\alpha_2$ .

The molecular constants of the upper state of the two systems are the following:

$$\begin{aligned}\alpha_1 \cdots \nu_1' &= 963 \text{ cm}^{-1}, \nu_2' = 379 \text{ cm}^{-1}, \\ &\quad x'\nu_2' = 2 \text{ cm}^{-1}, \nu_0 = 42,170 \text{ cm}^{-1}. \\ \alpha_2 \cdots \nu_1' &= 775 \text{ cm}^{-1}, \nu_2' = 375 \text{ cm}^{-1}, \\ &\quad \nu_0 = 44,236 \text{ cm}^{-1}.\end{aligned}$$

The high transition probability for both systems and the lack of antisymmetrical vibrations indicate that they are presumably due to allowed transitions; it must be stated however that the  $\alpha_2$  bands are much more intense than the  $\alpha_1$  bands. As already remarked, the  $\alpha_1$  bands appear to be of perpendicular type; according to group theoretical rules they must thus correspond to one of the three transitions:  $a_1 - a_1$ ,  $a_1 - b_1$ , and  $a_2 - b_2$ . The first assignment is partially in agreement with the theoretical scheme proposed by Mulliken.<sup>††</sup> In fact this scheme predicts two transitions in the 2000A region, one of them corresponding to  $4a_1 - 5a_1$  ( $\perp$  type), the other to  $3b_2 - 5a_1$  ( $\parallel$  type). It seems thus possible to correlate  $\alpha_1$  with the first and  $\alpha_2$  with the second one. The relative intensities of both systems are also in agreement with Mulliken's prediction. This interpretation, however, cannot be definitely accepted because Mulliken predicts  $\Delta\alpha > 0$  for the  $4a_1 - 5a_1$  transition, while, as we shall see below, we have deduced  $\Delta\alpha < 0$  from our analysis of rotational structure.

It is also to be noticed that, while the  $\alpha_1$  system is weak relatively to  $\alpha_2$ , it is much stronger than the  $b$ -system at 3000A, corresponding to  $1a_2 - 2b_1$ ; this seems also contrary to Mulliken's predictions.

For this reason, the other possibilities mentioned above cannot be disregarded.

As the transition  $a_1 - b_1$  has already been correlated with the  $c$ -system ( $\lambda_{\text{max}} \sim 3700\text{A}$ ),<sup>9</sup> we have to emphasize only the third one, i.e.,

<sup>††</sup> We are very indebted to Professor Mulliken for his valuable suggestion concerning this possibility, erroneously neglected by us in our previous letter (Nature 157, 692 (1946)).

$1a_2 - 4b_2$ . However this would require, according to Mulliken's scheme, 10 ev, instead of about 5 ev, according to our analysis. But it must be noted that even the energetic position of  $3b_2$  has not been determined with certainty experimentally and that according to Mulliken all the calculations have not much more than a qualitative meaning. If  $\alpha_1$  corresponds really to  $1a_2 - 4b_2$ , it should be necessary to reconsider theoretically the variation of  $4b_2$  with respect to the apex angle  $\alpha$  of the molecule.

On the other hand, we have said that  $B'$  must be nearly equal to  $B''$  while  $C'$  is very different from  $C''$ . If so, we have a case already discussed by Metropolis<sup>10</sup> and related to  $\Delta\alpha < 0$  and  $\Delta r \sim 0$ . Then from the variation of  $C$  and  $B$ , we can conclude that the angle  $\alpha$  has decreased by about  $20^\circ$ , while  $r_e$  has slightly increased by excitation.

These conclusions are in complete agreement with the general considerations of Metropolis<sup>10</sup> and would explain the fact that the  $\alpha_1$  system is not well developed toward high  $\nu_1'$  values.

For the  $\alpha_2$  system, the lack of rotational data prevents a more extensive analysis. A transition such as  $3b_2 \rightarrow 5a_1$ , as proposed by Mulliken for the whole  $a$ -system, seems not to be excluded and would give rise to parallel bands. In any event, the angle  $\alpha$  cannot be determined by the analysis of the vibrational potential function of the excited state as it has been done by Metropolis<sup>11</sup> for the  $b$ -system. The reason lies in the fact that the value of  $\alpha$  is very sensitive to the cross terms for angular molecules and, as it has been proved by one of us,<sup>12</sup> the cross terms cannot be neglected in such molecules. Furthermore, as the parallel character of the bands in the  $b$ -system must be considered with caution, we shall understand that the assigned transition ( $1a_2 \rightarrow 2b_1$ ) cannot be considered as being established with certainty. The full discussion of the scheme proposed by Mulliken must thus be postponed until new experimental data farther to the ultraviolet will be available. On the other hand, it seems that the  $\alpha_2$  system is much more developed in the sense of high  $\nu_1'$  values than the  $\alpha_1$  system, and this indicates that in the

<sup>10</sup> N. Metropolis, Phys. Rev. 60, 283 (1941).

<sup>11</sup> N. Metropolis, Phys. Rev. 60, 295 (1941).

<sup>12</sup> J. Duchesne, Mémoires Soc. Roy. Sci. (Liège) 1, fasc. 2, 429 (1943).



former  $r_e$  must be modified by excitation as well as  $\alpha$ .

Before new steps in the interpretation of the  $\text{SO}_2$  spectrum can be undertaken, it would be important to check the reality of the third  $\alpha_3$  system of Table IV, to get more information concerning the distribution of intensity in the bands, and to have more reliable data concerning bands in the 2000–1800Å region.

P. Migeotte in our laboratory is now engaged on such experiments.

## SeO<sub>2</sub>

### Introduction

The electronic diffraction patterns of  $\text{SeO}_2$  vapor, studied by Palmer and Elliott,<sup>13</sup> indicate an angular symmetrical structure of the  $\text{SeO}_2$  molecule, with a Se–O distance of 1.61Å. The diffraction patterns being relatively insensitive to the apex angle, its value has not yet been determined.

The  $\text{SO}_2$ ,  $\text{SeO}_2$ , and  $\text{TeO}_2$  molecules belong to the same class of 18-electron molecules, and their comparative spectroscopic study presents, therefore, considerable interest. Up to now, three electronic systems of  $\text{SeO}_2$  have been reported between 2000 and 5000Å. Preliminary descriptions of these systems were given by several authors,<sup>14</sup> and in a recent paper<sup>1</sup> we have analyzed in some detail the vibrational structure of the  $b$ -system situated between 2300 and 3300Å. In the present paper we deal with the vibrational study of the  $c$ -system situated in the visible, and we will try to develop some general conclusions regarding the electronic structure of the excited states. The analysis of the  $a$ -system in the farther ultraviolet is now in progress and will be published later.

### Vibrational Analysis

For the present study we have used a two-prism Bourguet spectrograph for the visible, giving a dispersion of about 30Å/mm at 4500Å, and a Jobin quartz spectrograph giving a dispersion of about 15Å/mm at 3000Å. The source

of background radiation was an Osram Nitratungsten-strip filament bulb. The absorption of the  $\text{SeO}_2$  vapor was investigated in 4-meter, 25-cm, and 2-cm absorption cells at temperatures varying between 300 and 500°C.

The  $\text{SeO}_2$  crystals were carefully purified by repeated distillations in vacuum and introduced in a side tube sealed to the main absorption cell. This side tube could be heated independently from the furnace heating the main cell, and thus the vapor pressure could be varied in the usual manner independently of the temperature. We used pressures between 50 and 450 mm.

By varying the conditions of pressure and temperature, the  $\text{SeO}_2$  absorption spectrum was investigated between 5000 and 3400Å. In this region about 100 bands were measured, and the measurements are in general in agreement with those of Evans and of Choong-Shin Piaw.<sup>14</sup> These authors have reported (especially in the ultraviolet part of the spectrum) more bands than we. Some of them must probably be correlated with rotational structure, as we have already noted in the case of the  $b$ -system.<sup>1</sup>

The absorption in the bands of the  $c$ -system is considerably weaker than in the bands of the  $a$ - and  $b$ -systems. The number of molecules needed to reveal absorption in the visible is at least 10 times greater than the number needed to provoke absorption in the ultraviolet.

A characteristic feature of the  $c$ -system is the diffuseness of the bands in the central region and the increased sharpness toward both ends. On both extremes, the bands are relatively well defined and the majority of the heads can be measured with good precision, although errors up to 20  $\text{cm}^{-1}$  cannot be excluded for fainter bands. In the central region, heads cannot be measured at all, and we have determined the position of the maxima of intensity on microphotometric records. Another striking feature of the  $c$ -system is provided by long series of bands extending through all the spectrum with the characteristic periodicity of about 200  $\text{cm}^{-1}$ . These series are given in the columns of Table VI. The 200- $\text{cm}^{-1}$  progression must surely be correlated with the bending frequency  $\nu_2'$  of the excited state. The presence of long  $\nu_2'$  progressions indicates a considerable change of the apex angle of the molecule. The apparent lack of

<sup>13</sup> K. J. Palmer and N. Elliott, J. Am. Chem. Soc. **60**, 1389 (1938).

<sup>14</sup> S. F. Evans, Nature **125**, 528 (1930); Choong-Shin Piaw, C. R. Acad. Sci. Paris **202**, 127 (1936); *ibid.* **203**, 239 (1936); Ann. de physique **10**, 191 (1938).

TABLE VI.

$0_1'v_2'0_3'$	$0_1''0_2''0_3''$	0	1	2	3	4
0	—		20,940 178			
1	—		21,118 206	917	20,201	
2	—		21,324 210			
3	—		<u>21,534*</u> 210	919	20,615 205	
4	—		21,743 192	923	20,820 215	
5	22,855 203	920	21,935 208	900	21,035 203	
6	23,058 201	915	22,143 198	905	21,238 210	
7	23,259 203	918	22,341 199	893	21,448 207	
8	23,462 192	922	22,540 193	885	21,655 188	
9	23,650 184	917	22,733 197	890	21,843 207	
10	23,834 190	904	22,930 206	880	<u>22,050*</u> 190	
11	24,030 200	894	23,136 209	896	22,240 174	
12	24,230 192	885	23,345 165	931	22,414 216	
13	<u>24,422*</u> 200	912	23,510 194	880	22,630 163	21,680 192
14	24,622 198	918	23,704 185	911	22,793 198	921 200
15	24,820 182	931	23,889 201	898	22,991 202	919 211
16	25,002 203	912	24,090 204	897	23,193 197	910 199
17	25,205 185	911	<u>24,294*</u> 186	904	23,390 182	908 214
18	25,390 245	910	24,480 210	908	23,572 190	876 198
19	25,635 181	945	24,690 212	928	23,672 197	868 208
20	25,816	914	24,902	947	23,955 208	853 208
21					24,163 205	855 23,310
22					24,368 196	
23					<u>24,564*</u> 211	
24					24,775 188	
25					24,963	

\* The underlined bands correspond to intensity maxima in the series.

TABLE VII.

$\Delta G_{v_1}''$	<i>b</i> -system mean values	<i>c</i> -system weighted values
0-1	894	911
1-2	893	898
2-3	907	889
3-4	902	903
0-4	3,596	3,601

stretching frequencies means that the valency bond remains practically unchanged. It seems, however, not excluded that some fainter bands in the ultraviolet part of the spectrum are due to a weak excitation of the valence frequency.

The first column of Table VI contains the most prominent bands of the whole system, and must, therefore, undoubtedly be correlated to the zero level of the normal state. The intensity distribution in this series is regular, and the maximum occurs at about  $24,500\text{ cm}^{-1}$ . The assignment of the other series to definite vibrational levels cannot be done without some ambiguity. This is due to lack of regularity in the intensity distribution and to frequent irregularities in the distances between neighboring bands. From a strictly numerical point of view, several arrangements are possible; however, the scheme given in Table VI seems to us to be by far the most reliable; in particular, it gives for the ground level a frequency of about  $900\text{ cm}^{-1}$ , already assigned to the stretching frequency of normal  $\text{SeO}_2$  on the basis of our analysis of the *b*-system.

It is important to check this last conclusion by comparing the mean values of  $\Delta G''$  resulting from the study of the *b*- and the *c*-systems. The agreement of the mean  $\Delta G''$  values, corresponding to both systems and given in Table VII, can be considered as satisfactory and is even improved if we consider only those bands which can be measured with good accuracy. In doing so we obtain for the symmetrical stretching frequency in the ground level of  $\text{SeO}_2$  the value of  $\nu_1'' = 910\text{ cm}^{-1}$ .

In adopting the vibrational scheme of Table

TABLE VIII. Isotopic constitution of selenium.

Atomic weights	74	76	77	78	80	82
Abundance	0.9	9.5	8.3	24	48	9.3

VI for the *c*-system, we must admit that the irregularities in the progressions are due to some kind of perturbation. On the other hand, the occurrence of two intensity maxima observed in the bands of the second and third columns of Table VI (underlined bands) must be correlated either with the special form of potential energy surfaces or with the superposition of at least two series.

We have already noted that the band edges become increasingly sharper toward both ends of the spectrum; whereas on the long wave-length side the bands are shaded toward the red, they are shaded to the opposite sense at the other extremity of the spectrum. In the central part, the sense of the shading is not well defined; however the change of shading can be approximately located at  $22,000\text{ cm}^{-1}$ .

A similar anomaly in the shading was observed by us in the  $\text{TeO}_2$  bands, as will be reported in detail in the last part of this paper. On the other hand, the bands of all known  $\text{SO}_2$  band systems are uniformly degraded toward the red (if we neglect the *J*-structure which has not been observed in the spectra of  $\text{SeO}_2$  and  $\text{TeO}_2$ ). It seems to us that the above-mentioned behavior of the  $\text{SeO}_2$  and  $\text{TeO}_2$  spectra must be explained as due to an isotopic shift and can be correlated with the fact that Se and Te have several abundant isotopes, whereas S has practically only one. We give in Table VIII the isotopic constitution of selenium:

Table VIII shows that the center of gravity of the isotopic group lies on the side of greater atomic weights. Thus the groups of isotopic bands are apparently shaded toward the red on the long wave-length side with respect to the origin, and toward the violet at the opposite extremity. In conformity with this point of view the change in the sense of degradation occurs at about  $22,000\text{ cm}^{-1}$ , and coincides with the origin of the system according to the proposed vibration scheme.

### Interpretation of Results

The present data do not permit stating whether the *c*-bands of  $\text{SeO}_2$  correspond to an allowed or to a forbidden transition. The intensity distribution in the vibrational series indicates strongly that the electronic excitation does

not change appreciably the Se—O distance, whereas the apex angle is considerably modified.

We have already noted that most of the bands are rather diffuse, and this could presumably be due to the fact that the *J*- and *K*-structures are shaded in opposite directions. Not much more can be said for the moment about the structure of the molecule in the upper level.

It is worth while pointing out that no analogy can be established between this system and the *c*-system of the SO<sub>2</sub> molecule, studied by Metropolis and Beutler.<sup>15</sup> For the former,  $\Delta r \cong 0$  and  $\Delta \alpha \neq 0$ , while for the latter  $\Delta r \neq 0$  and  $\Delta \alpha \cong 0$ . On the contrary, the *c*-system of SO<sub>2</sub> is much more similar in regard to these properties to the *b*-system of SeO<sub>2</sub> for which we have previously proved<sup>1</sup> that  $r_e$  is the only parameter to change appreciably.

As Metropolis remarked,<sup>10</sup> the ratio of the moments of inertia  $I_c/I_b$  is practically constant ( $\sim 1.17$ ) in the normal state of a series of angular molecules such as SO<sub>2</sub>, NO<sub>2</sub>, and ClO<sub>2</sub>. If we extend the validity of this rule to SeO<sub>2</sub>, it provides an apex angle of about 130°. On the other hand, we observe that the ratio of the antisymmetrical and symmetrical frequencies of the same molecules also follows such a sort of regularity and equals approximately 1.17. Though the identity of the constants seems purely fortuitous, the extrapolation of this rule to SeO<sub>2</sub> gives us the possibility to locate  $\nu_3''$  at about 1060 cm<sup>-1</sup>. The bending frequency  $\nu_2''$  can also be approximately estimated; in fact the apparent disappearance of this vibration in the spectrum may be explained by assuming that the fundamental frequency  $\nu_2''$  is a multiple of  $\nu_2'$ , the most probable value being  $\nu_2'' = 400$  cm<sup>-1</sup>. If this estimation is correct, the columns of Table VI could result from a superposition of several types of progressions, and the apparent lack of the 400-cm<sup>-1</sup> progression, as well as the anomalies of intensity referred to above, could be satisfactorily explained.

With the three fundamental frequencies thus obtained (910, 400, and 1060 cm<sup>-1</sup>) and by using a valence-deformation force field, we compute an apex angle of about 160°. The reason for the divergency with respect to 130°, which seems

much more plausible, could be explained by the neglect of the cross terms in the potential function. The introduction of a small stretch-stretch cross term ( $f_{12}$ ) of about  $-0.4 \times 10^5$  dyne/cm is sufficient to remove the divergency. On the other hand, by taking into account a stretch-deformation cross term ( $g_e$ ) only, the best value equals about  $+0.4 \times 10^5$  dyne/cm, and it is impossible to obtain quite satisfactory results. However, in order to obtain such results, it is sufficient to add in the potential function a slightly negative value of  $f_{12}$ . Thus, as we have no reason to believe that  $f_{12}$  has a considerable value, while  $g_e = 0$  (or the reverse), it seems established that  $f_{12}$  and  $g_e$  have opposite signs, the former being negative and the latter positive, with numerical values smaller than  $0.4 \times 10^5$  dyne/cm. Furthermore, the presence of the cross terms proves that the apex angle of molecules like SeO<sub>2</sub> and SO<sub>2</sub> cannot be determined by assuming a valency force field, contrary to the assumption by Metropolis for the case of the *c*-system of SO<sub>2</sub>.

### TeO<sub>2</sub>

The TeO<sub>2</sub> absorption bands were first studied by Choong-Shin Piaw.<sup>16</sup> He arranged a part of the observed bands in a vibrational scheme according to the formula:

$$\nu = 25636 + 545\nu_1' - \nu_1'^2 + 206\nu_2' - 3\nu_2'^2 + 675\nu_3' - 2.7\nu_3'^2.$$

Several objections can be raised against this analysis:

1—It seems quite abnormal that at the high temperature used, the frequencies of the ground level do not appear in the spectrum.

2—only about 60 percent of the observed bands fit into the scheme.

3—the selection rules are not respected. It thus seemed necessary to re-investigate the TeO<sub>2</sub> spectrum before any attempt to compare it with the SO<sub>2</sub> and SeO<sub>2</sub> band systems could be undertaken.

### Experimental Procedure

The TeO<sub>2</sub> vapor reacts with quartz at high temperatures necessary to obtain sufficient vapor

<sup>15</sup> N. Metropolis and H. Beutler, Phys. Rev. **57**, 1078 (1940).

<sup>16</sup> Choong-Shin Piaw, C. R. Paris **201**, 1181 (1935); Thesis, Paris (1937).

pressure. The absorption was therefore studied in thin open porcelain tubes filled with carefully purified  $\text{TeO}_2$  crystals. The tube with  $\text{TeO}_2$  was introduced into the electric furnace after the desired temperature was attained. We have thus operated in conditions far from equilibrium and contrary to the experiments concerning the  $\text{SO}_2$  and  $\text{SeO}_2$  absorption we were not able to vary the pressure independently of the temperature. In some cases, it was necessary to avoid excess oxygen and then the absorption tube was provided with water-cooled quartz windows and

filled with nitrogen to prevent rapid distillation of  $\text{TeO}_2$  outside the high temperature zone.

The dispersion used was the same as in the case of  $\text{SeO}_2$ . For the bands situated near 4000Å, it was possible to use a multiple-prism Jobin spectrograph in Littrow mounting, either with 5 quartz prisms giving a dispersion of about 9Å/mm, or with 5 glass prisms with a dispersion of about 2Å/mm at 4000Å.

### Vibrational Analysis

The most characteristic band system of  $\text{TeO}_2$  is situated in the blue and near ultraviolet, between 3000 and 4550Å, with a maximum near 3550Å. In the conditions of our experiments (absorbing layer of a few centimeters, rapid distillation from the middle of the furnace towards the end) it appears at about 900°. At about the same temperature, we have observed very strong absorption bands in the region of 2000Å, but owing to the presence of  $\text{O}_2$ , it has been up to now impossible to investigate these in detail. It is not even excluded that this absorption is due to impurities. Choong Shin-Piaw observed in the ultraviolet above 3000Å only continuous absorption. At somewhat higher temperatures (about 1000°C) very weak and diffuse absorption bands appear in the region between 2450–2650Å†††. It may belong also to some impurities. At a temperature of about 1100°, some weak and very diffuse bands appear between 4650 and 5100Å. We shall refer to them at the end of this paper. At still higher temperature, the  $\text{TeO}_2$  bands gradually disappear, and they are replaced by  $\text{TeO}$  absorption bands which will be discussed in detail elsewhere. In this paper we will deal chiefly with the main system between 3000–4550Å, which will be designed as the *b*-system.

In the *b*-system, the bands situated in the blue are sharp and shaded towards the red. Towards the ultraviolet they gradually become more diffuse, and even the sense of degradation is reversed in the extreme ultraviolet end of the spectrum. This change of shading is probably due to an isotopic effect in the same way as

TABLE IX. Intensity maxima in the bands of the *b*-system of  $\text{TeO}_2$ .

$\nu_{\text{max}}$	$\nu_{\text{max}}$	$\nu_{\text{max}}$
21,950	24,680	{29,460
22,130	24,810	29,490
22,240	25,000	{29,515
{22,380	{25,150	{29,650
22,390	25,170	29,675
{22,400	{25,190	{29,700
{22,560	25,340	29,870
22,570	25,470	29,950
{22,580	25,525	30,080
		30,165 a
{22,650	{25,620	30,320
22,665	25,640	30,400 a
{22,680	{25,660	30,515
	25,825	30,610 a
	26,000	30,750
{22,750	26,080	30,790
22,765	26,160	30,830 a
{22,780	26,280	30,960 b
	26,460	31,050 a
	26,640	31,160 b
{22,920	26,820	31,250 a
22,930	26,920	31,385 b
{22,940	{27,070	31,450 a
	27,095	31,580 b
23,060	{27,120	31,660 a
	27,260	31,800 b
23,200	27,460	31,865 a
	27,550	32,000 b
23,385	27,730	32,090 a
	{27,880	32,205 b
{23,550	27,910	32,275 a
23,575	{27,940	32,410 b
{23,600	28,125	32,500 a
	28,210	32,620 b
{23,710	{28,570	32,700 a
23,730	28,385	32,830 b
{23,750	{28,400	33,040 b
	28,540	
23,870	28,860	
{24,000	{29,020	
24,020	29,030	
{24,040	{29,045	
24,200	29,190	
{24,360	29,345	
24,375		
{24,400		
24,540		

††† The wave numbers of the band heads, which are shaded toward ultraviolet, are the following: 40740, 40260, 39760, 39400, 39260, 39060, 38920, 38740, 38580, 38420, 38240, 38080  $\text{cm}^{-1}$ .

It is worth while to note that the value  $\nu_1''=810\text{ cm}^{-1}$  for the valency frequency of the ground state of  $\text{TeO}_2$  is nearly equal to the vibrational frequency  $\omega''=796$  of  $\text{TeO}$  given by

TABLE X. *b*-system of TeO<sub>2</sub>.[illegible]

Choong-Shin-Piaw.<sup>16</sup> ‡ The same relation is correspondingly valid for the valency frequencies of  $\text{SeO}_2$  and  $\text{SO}_2$  (910 and 1152) as compared with those of  $\text{SeO}$  and  $\text{SO}$  (909 and 1124). Although at present, no theoretical considerations supply a basis for such a relation, it seems nevertheless that the parallelism in the behavior of the six molecules constitutes a further argument in favor of our analysis. The scheme of Table X contains nearly all bands observed between 22,000 and 30,000  $\text{cm}^{-1}$ . In the ultra-violet end of the spectrum, however, there are two rather long series of faint bands which appear only when the absorption in the central part is complete and which cannot be fitted in the scheme. These bands are marked *a* and *b* in Table IX and form two progressions with a characteristic frequency of about 200  $\text{cm}^{-1}$ . On the other hand, some of the weaker bands farther to the visible, which cannot be fitted in the scheme of Table X, are situated at about 200  $\text{cm}^{-1}$  on the violet side of strong bands belonging to the main series. It seems reasonable to admit that 200  $\text{cm}^{-1}$  represents the deformation frequency of the upper state and that each

$v_1'$  level is combined with a series of  $v_2'$  levels. This statement cannot, however, be considered as definitive, as the possibility that the 200- $\text{cm}^{-1}$  progression belongs to a separate electronic transition cannot be ruled out. If all the bands belong to a single system, it would mean that the corresponding transition is accompanied not only by considerable change of the apex angle  $\alpha$  but also of the  $\text{Te}-\text{O}$  distance.

After having established the analogy existing between the *b*-systems of  $\text{SeO}_2$  and  $\text{TeO}_2$ , we have searched for a system of  $\text{TeO}_2$  situated farther toward the red in analogy with the *c*-system of  $\text{SeO}_2$ . At the highest pressure obtainable, corresponding to temperature above 1100°C, some weak and diffuse bands, without marked heads, appear between 4650 and 5100 Å on microphotometer records; we have measured the following maxima: 21,520, 21,030, 20,560, 20,150, 19,750  $\text{cm}^{-1}$ . These bands bear no analogy with the *c*-system of  $\text{SeO}_2$ .

Our best thanks are due to Professor R. S. Mulliken for valuable criticism. Some of the spectra used in this investigation were taken and measured by R. Migeotte and A. Robert; we wish to express to them our thanks for their effective help.

‡ According to the new analysis of  $\text{TeO}$  bands (not yet published) by one of us,  $\omega'' = 790 \text{ cm}^{-1}$ .