# Geometries and Stabilities of the Carbon Clusters with the Rhodium Impurity: A Computational Investigation

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A density functional study of the  $RhC_n(n = 1-6)$  clusters with different spin states has been carried out systematically by using the B3LYP/Lan2DZ method. The equilibrium geometries associated with total energies and natural populations of  $RhC_n$  (n = 1-6) clusters are calculated and presented. Stabilities and electronic properties are discussed in detail. The relative stabilities in term of the calculated fragmentation energies show that the lowest-energy  $RhC_n$  clusters with rhodium atom being located at terminal of carbon chain are the linear geometries and the ground states of the  $RhC_n$  clusters alternate between doublet (for n-odd members) and quartet (for n-even members) states. Furthermore, the calculated fragmentation energies of the  $RhC_n$  show strong even—odd alternations: the  $RhC_n$  clusters with an odd number of carbon atoms are more stable than those with an even number ones. In addition, we comment on the charge transfer and chemical bonding properties within the clusters.

#### 1. Introduction

To understand a variety of chemical properties of carbon clusters (e.g., fullerene and nanotube carbon clusters), many experimental and theoretical investigations have been studied in the past decades.  $^{1-3}$  Small carbon clusters were studied with great interest because they are intermediates in chemical vapor deposition of carbon clusters and can be exactly identified in interstellar and circumstellar media by radiowave or infrared spectroscopy in astrophysics. In the interstellar medium, the reactivity of small carbon clusters shows their light reactivity because of the quasicollisionless conditions, and highly linear chains carbon clusters are observed. Furthermore, the discovery and successful preparation of  $C_{60}$  and other fullerenes have spawned a new branch of chemistry.

Heteroatom-doped carbon clusters result in significant changes in the geometrical and electronic structures, which may eventually be used to elaborate new types of materials with interesting applications and novel properties. In this respect, it is very important to have some knowledge about the behavior of transition metal-doped carbon clusters. This will allow the identification of possible systematic trends that could help us understand the structure of these materials and get information that could be useful to make extrapolations for some properties. Much interest has been focused on nonmetallic and metallic carbide clusters, such as SC<sub>n</sub><sup>5</sup>, SiC<sub>n</sub>,<sup>6</sup> PC<sub>n</sub>,<sup>7</sup> ClC<sub>n</sub><sup>8</sup>, AlC<sub>n</sub>,<sup>9,10</sup>  $MgC_{n}$ , 11,12 and  $CaC_{n}$  and  $CaC_{n}$  clusters. Recently, these kinds of experimental studies have been extended to heavier transition metal carbon clusters combined with mass spectrometry, the ion drift tube technique, high-resolution optical spectroscopy, photoelectron spectroscopy, ESR spectroscopy for YC, NbC, PdC, and RhC etc.15-24

Surprisingly, there are few theoretical studies on the structures and stabilities of the transition metal doped carbon clusters. A theoretical study of the linear PbC<sub>n</sub> clusters and the transition metal carbides of CrC<sub>n</sub>, YC<sub>n</sub>, LaC<sub>n</sub>, NbC<sub>n</sub>, and ScC<sub>n</sub><sup>25–30</sup> clusters have recently been investigated systemically. A number of carbides containing rhodium atoms, i.e., RhC,<sup>31</sup> RhN,<sup>32</sup> RhCH<sub>3</sub>,<sup>31</sup> RhHCO,<sup>34</sup> and Rh(CO<sub>2</sub>),<sup>34</sup> have also been studied theoretically.

Although the diatomic RhC has been studied both experimentally and theoretically, to the best of our knowledge, no experimental or theoretical investigation on the RhC<sub>n</sub> (n=1-6) clusters has been reported. To reveal the unusual properties of the rhodium-doped carbon clusters, the aim of the studies reported here is to provide a detailed investigation of equilibrium geometries, charge-transfer properties, relative stabilities, fragmentation energies [D(n, n-1)], atomic averaged binding energies [ $E_b(n)$ ], and HOMO–LUMO gaps of these RhC<sub>n</sub> clusters. The analyzed results help us understand chemical and physical properties of the Rhodium doped carbon clusters.

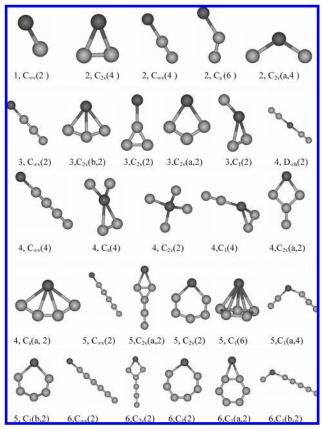
## 2. Computational Details

All calculations were performed with the hybrid Becke (B3) exchange and the Lee, Yang and Parr (LYP) correlation (B3LYP) functionals in combination with the Los Alamos ECP D95V and double- $\zeta$  basis sets (LanL2DZ) as implemented in the Gaussian 98 code.35-40 For each stationary point of the clusters, the stability is examined by calculating the harmonic vibrational frequencies. If an imaginary frequency is found, a relaxation along the coordinates of the imaginary vibrational mode is carried out until a true local minimum is reached. Therefore, the reported geometries and total energies for each stable cluster and its stable isomers correspond to local minima. However, as the number of isomers increases rapidly with the cluster size, it is very difficult to determine the global minimum simply according to the calculated energies of the isomers. To find the stable  $RhC_n$  (n = 1-6) geometries, we considered first some previous theoretical or experimental geometries that are

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**Figure 1.** Equilibrium geometries of the  $RhC_n$  (n=1-6) clusters. For example, the  $C_1(2)$  above represents the doublet state with  $C_1$  symmetry.

available, we varied the structure starting from high-symmetry structure and reducing the symmetry. The spin of the RhC<sub>n</sub> clusters were restricted to S = 1/2, 3/2, and 5/2. The spin contamination is zero and can be neglected.

### 3. Results and Discussion

The calculated results of bond lengths and bond angles, together with total energies, are listed in Table 1; Milliken atomic net populations are listed in Table 2. Natural populations and natural electron configurations of the RhC<sub>n</sub> (n=1-6) clusters are listed in Table 3, whereas the equilibrium geometries of RhC<sub>n</sub> (n=1-6) are displayed in Figure 1.

**3.1. Geometry and Stability.** To check the reliability of our calculations, we calculated the bond lengths of C<sub>2</sub> and Rh<sub>2</sub> dimers at the UB3LYP/LanL2DZ level. The calculated results show that the C-C bond length of C<sub>2</sub> dimer is 1.281 Å, which is in good agreement with those of B3LYP/ccp-VDZ(1.262 Å);<sup>41</sup> and that the Rh-Rh bond length (2.321 Å) corresponds to the previous results of the UHF/SHIM-SU method<sup>42</sup> with relativistic core potential (2.283 Å). On the basis of the tests above, the B3LYP/LANL2DZ method is believed to yield an adequate description of the bonding properties of C and Rh atoms in small molecular units.

3.1.1. RhC. The equilibrium geometry of the RhC cluster with  $C_{\infty \nu}$  symmetry is optimized at the (U)B3LYP/LanL2DZ level (Figure 1). The Rh—C bond length depends on spin state and is elongated as spin increases monotonically. The total energies of RhC cluster are increased apparently as spin increases (Table 1). So the doublet RhC cluster is obviously the lowest-energy isomer and the ground state, and the corresponding electronic state is  $^2\Sigma$ . This feature is in good agreement with an available calculated result for RhC<sup>42</sup> and is similar to those for TaC, <sup>43</sup>

VC,<sup>24</sup> and CoC;<sup>44</sup> however, it differs from results for MC (M = Sc, Nb, Al, and Na) clusters.<sup>45–48,9</sup>

The calculated Rh–C distance for the ground state RhC (1.669 Å) is also in good agreement with the previous theoretical results of Skaarup et al. (1.65 Å)<sup>49</sup> and the experimental value given by Norman et al.<sup>50</sup> (1.655 Å); however, it is slightly shorter than the previous calculated datum at the UHF/RECP level.<sup>34</sup> The frequency of the ground state is calculated to be 1057.8 cm<sup>-1</sup>, comparson with an available experimental result (1000  $\pm$  60 cm<sup>-1</sup>) is made.<sup>22</sup>

3.1.2. RhC<sub>2</sub>. Five RhC<sub>2</sub> isomers with  $C_s$ ,  $D_{\infty h}$ ,  $C_{2v}$ ,  $C_{2v}$ (a), and  $C_{\infty \nu}$  symmetries were considered. The linear  $D_{\infty h}$  geometry is proved to be an unstable structure while the  $C_{2\nu}$  and  $C_{2\nu}(a)$ clusters as well as the linear  $C_{\infty \nu}$  with different spin states are found to be stable geometries. As seen from the Table 1, the bond lengths and the total energies of the Rh-capped  $C_{2v}$  isomer have an evident correlation with the spin; the Rh-C bond length and the total energies increase monotonical as spin S goes from S = 1/2, to 3/2, and to 5/2, resulting in the weakness of the Rh-C bond interaction and the stability of  $C_{2v}$  isomer. Like the  $C_{2v}$  isomer, the Rh-C bond length and the bond angle  $\angle$ CRhC in the  $C_{2\nu}(a)$  isomer are increased monotonously with the spin going from S = 1/2, to 3/2, and to 5/2. The total energy of the quartet  $C_{2\nu}(a)$  structure, however, is lower than those of the doublet and sextet isomers. The Rh-terminated linear  $C_{\infty \nu}$ isomer has similar feature to that of the  $C_{2\nu}(a)$  isomer; The total energy of the quartet isomer is the lowest one with the Rh-C and C-C bond lengths being the smallest ones. The bent  $C_s$ structure is a distortion of the Rh-capped  $C_{2\nu}$  isomer, but the geometry optimization of the doublet and quartet  $C_s$  isomers fail to be converged. The total energy of the sextet  $C_s$  isomer is higher than those of the other isomers.

On the basis of the calculated total energies, it is confirmed that the linear quartet  $C_{\infty\nu}$  geometry with electronic state of  $^4\Sigma$  is the lowest-energy isomer and is consequently selected as the ground state. This finding is different from those for the most stable doublet LaC<sub>2</sub>,  $^{28}$  NaC<sub>2</sub>,  $^{48}$  ScC<sub>2</sub>,  $^{51}$  and YC<sub>2</sub><sup>27</sup> isomers.

3.1.3.  $RhC_3$ . The RhC<sub>3</sub> clusters with  $C_2v$ ,  $C_8$ ,  $C_1$ , and the linear  $C_{\infty v}$  isomers are taken into account. All of them are considered as ground-state candidates. The calculated results reveal that the scooplike  $C_1$  isomer has higher total energy and cannot be competitive with the other isomers in stability. Exchanging the location of the Rh and the apex C atom results in a planar geometry: the three carbon atoms form an equilateral triangle and the Rh atom bonds to one of the carbon atoms, which we call the kitelike  $C_{2v}$  isomer. The calculated results prove that the kite-like  $C_{2v}$  geometry is a stable structure. From the Table 1, it can be seen that the total energy and the C–C bond length of the kitelike quartet  $C_{2v}$  structure are lower than those of the doublet and sextet states, thus the kitelike quartet  $C_{2v}$  geometry with electronic state of  $^4B_2$  is the most stable structure.

Guided by the previous theoretical results on the  ${\rm YC_3}^{27}$  and  ${\rm LaC_3}^{28}$  clusters, two planar  $C_{2\nu}({\rm a})$  and  $C_{2\nu}({\rm b})$  isomers were considered. One is a ringlike  $C_{2\nu}({\rm a})$  structure in which the Rh atom is inserted into the  ${\rm C}_n$  ring, the other is a fanlike  $C_{2\nu}({\rm b})$  structure with the Rh atom bonding to three carbon atoms. For these two isomers, the total energy increases monotonically as spin varies from S=1/2, to 3/2, and to 5/2. The calculated results indicate that the total energy of the doublet  $C_{2\nu}$  (a) isomer with electronic state of  ${}^2{\rm A}_1$  is slightly lower by about 0.005 hartree than that of the doublet  $C_{2\nu}({\rm b})$  isomer with electronic state of  ${}^2{\rm B}_1$ . The linear  $C \sim \nu$  RhC<sub>3</sub> isomer was also optimized. The

TABLE 1: Geometries, Total Energies, and Electronic States of  $RhC_n$  (n=1-6) Clusters

| cluster            | symmetry                         | spin                     | electronic state   | $R1^a$ (Å)              | $R2^b$ (Å)              | $\alpha^c$              | $E_{\rm b}{}^d$ (Hartree)           |
|--------------------|----------------------------------|--------------------------|--|-------------------------|-------------------------|-------------------------|-------------------------------------|
| RhC $C_{\infty v}$ |                                  | 1/2                      | $^{2}\Sigma$ $^{4}\Sigma$ $^{6}\Sigma$ $^{2}B_{1}$                                     | 1.669                   |                         |                         | -147.4691                           |
|                    | -                                | 3/2                      | $4\overline{\sum}$   | 1.798                   |                         |                         | -147.4149                           |
|                    |                                  | 5/2                      | $6\sum$  | 2.035                   |                         |                         | -147.3484                           |
| $RhC_2$            | $C_{\sim v}$                     | 1/2                      | $^{2}\mathrm{B}_{1}$   | 1.916                   | 1.350                   | 180.00                  | -185.4438                           |
|                    |                                  | 3/2                      | $^4\mathrm{B}_1$   | 1.812                   | 1.341                   | 180.00                  | -185.4871                           |
|                    |                                  | 5/2                      | ${}^{6}B_{1}$  | 1.846                   | 1.521                   | 180.00                  | -185.3575                           |
|                    | $C_{2v}$                         | 1/2                      | ${}^{2}B_{1}$  | 1.933                   | 1.390                   | 42.34                   | -185.4759                           |
|                    |                                  | 3/2                      | ${}^{4}B_{1}$  | 2.038                   | 1.357                   | 39.44                   | -185.4754                           |
|                    | a                                | 5/2                      | ${}^{6}A_{1}$  | 2.041                   | 1.526                   | 43.91                   | -185.3405                           |
|                    | $C_{2v}(\mathbf{a})$             | 1/2                      | ${}^{2}B_{2}$  | 1.780                   |                         | 103.21                  | -185.4126                           |
|                    |                                  | 3/2                      | ${}^{4}A_{2}$  | 1.789                   |                         | 111.98                  | -185.4133                           |
|                    |                                  | 5/2                      | $^{6}\mathrm{B}_{2}$ $^{6}\mathrm{A''}$ $^{2}\Sigma$ $^{6}\Sigma$ $^{2}\mathrm{B}_{2}$ | 1.891                   | 1.250                   | 166.10                  | -185.3464                           |
| D1 C               | $C_{ m s} \ C_{\infty_{\cal V}}$ | 5/2                      | 2 <b>X</b>   | 1.949                   | 1.359                   | 137.79                  | -185.4009                           |
| $RhC_3$            | $C_{\infty v}$                   | 1/2                      | <sup>2</sup> ∑<br><sub>4</sub> ∑   | 1.781                   | 1.307                   |                         | -223.5922                           |
|                    |                                  | 3/2                      | -Σ<br>6Σ   | 1.882                   | 1.295                   |                         | -223.4702                           |
|                    | C                                | 5/2<br>1/2               | 2 <b>D</b>   | 2.022                   | 1.291                   |                         | -223.4692                           |
|                    | $C_{2v}$                         | 3/2                      | -В <sub>2</sub><br>4 <b>р</b>  | 1.868<br>1.912          | 1.415                   |                         | -223.5165                           |
|                    |                                  | 5/2<br>5/2               | ${}^{4}B_{2}^{2}$ ${}^{6}A_{1}$  | 2.002                   | 1.411                   |                         | -223.5199 $-223.3965$               |
|                    | $C_{-}(a)$                       | 1/2                      | 2 A  | 2.083                   | 1.450                   | 74.92                   | -223.5963 $-223.5553$               |
|                    | $C_{2\nu}(\mathbf{a})$           | 3/2                      | ${}^{2}A_{1}$ ${}^{4}B_{1}$  | 2.212                   | 1.357<br>1.341          | 67.86                   | -223.4919                           |
|                    |                                  | 5/2<br>5/2               | ${}^{6}\mathrm{A}_{1}$   | 1.993                   | 1.431                   | 68.84                   |                                     |
|                    | C. (b)                           | 5/2<br>1/2               | ${}^{2}B_{1}$  | 2.113                   |                         | 75.65                   | -223.4360<br>-223.5503              |
|                    | $C_{2v}(\mathbf{b})$             | 3/2                      | ${}^{4}B_{1}$  | 1.927                   | 1.365<br>1.454          | 75.05<br>76.06          | -223.5503 $-223.4811$               |
|                    |                                  | 5/2<br>5/2               | ${}^{6}\mathrm{B}_{1}$   | 1.927                   | 1.434                   | 68.42                   | -223.4811 $-223.4360$               |
|                    | $C_1$                            | 1/2                      | ${}^{2}A$  | 2.207                   | 1.318                   | 116.17                  | -223.4940                           |
|                    | Cl                               | 3/2                      | 4A   | 1.919                   | 1.312                   | 108.32                  | -223.4623                           |
|                    |                                  | 5/2                      | Α<br>6 <b>Λ</b>  | 1.813                   | 1.353                   | 107.04                  | -223.4049                           |
| RhC <sub>4</sub>   | $C_{\infty v}$                   | 1/2                      | $2\sum_{i=1}^{N}$  | 1.915                   | 1.302                   | 107.04                  | -261.6061                           |
| CIIC4              | $C \infty v$                     | 3/2                      | $4\overline{\Sigma}$   | 1.825                   | 1.298                   |                         | -261.6369                           |
|                    |                                  | 5/2                      | $^{6}{ m A}$ $^{2}{ m \sum}$ $^{4}{ m \Sigma}$ $^{6}{ m \Sigma}$ $^{2}{ m A}_{1}$      | 1.906                   | 1.279                   |                         | -261.5366                           |
|                    | $C_{2v}$                         | 1/2                      | $^{2}\overline{\mathbf{A}}_{1}$  | 1.869                   | 1.27)                   | 110.49                  | -261.2672                           |
|                    | $C_{2v}$                         | 3/2                      | ${}^{4}\mathrm{B}_{2}$   | 1.860                   |                         | 112.35                  | -261.2574                           |
|                    | $C_{2v}(\mathbf{a})$             | 1/2                      | ${}^{2}A_{1}$  | 1.879                   | 1.335                   | 99.85                   | -261.5292                           |
|                    | 20(33)                           | 3/2                      | ${}^{4}A_{2}$  | 1.994                   | 1.360                   | 127.32                  | -261.5233                           |
|                    |                                  | 5/2                      | <sup>6</sup> B <sub>1</sub>  | 1.930                   | 1.410                   | 136.17                  | -261.4551                           |
|                    | $C_s$                            | 3/2                      | <sup>6</sup> B₁<br><sup>4</sup> A″   | 1.908                   | 1.317                   | 93.44                   | -261.4015                           |
|                    | ,                                | 5/2                      | <sup>6</sup> A″  | 1.824                   | 1.390                   | 119.92                  | -261.3694                           |
|                    | $C_s(\mathbf{a})$                | 1/2                      | <sup>2</sup> A"  | 2.153                   | 1.327                   | 106.72                  | -261.5960                           |
|                    |                                  | 3/2                      | <sup>4</sup> A"  | 2.228                   | 1.311                   | 101.31                  | -261.5937                           |
|                    |                                  | 5/2                      | <sup>6</sup> A"  | 1.949                   | 1.302                   | 119.92                  | -261.5615                           |
|                    | $C_1$                            | 1/2                      |  | 1.808                   | 1.343                   | 110.35                  | -261.4993                           |
|                    |                                  | 3/2                      |  | 1.814                   | 1.337                   | 106.93                  | -261.5042                           |
|                    |                                  | 5/2                      |  | 1.908                   | 1.320                   | 154.44                  | -261.4751                           |
|                    | $D_{^\infty h}$                  | 1/2<br>5/2               | $2\sum$ $6\sum$ $2\sum$ $4\sum$ $6\sum$  | 1.977                   | 1.283                   |                         | -261.4086                           |
|                    |                                  | 5/2                      | $^6\Sigma$   | 2.069                   | 1.339                   |                         | -261.3578                           |
| $RhC_5$            | $C_{\sim v}$                     | 1/2                      | $^2\Sigma$   | 1.793                   | 1.284                   |                         | -299.7326                           |
|                    |                                  | 3/2                      | $^4\Sigma$   | 1.970                   | 1.271                   |                         | -299.6831                           |
|                    |                                  | 5/2                      |  | 2.096                   | 1.265                   |                         | -299.6030                           |
|                    | $C_{2v}$                         | 1/2                      | ${}^{2}\overline{\mathrm{B}}_{2}$  | 1.912                   | 1.327                   | 86.25                   | -299.6744                           |
|                    | =                                | 5/2                      | ${}^{6}A_{1}$  | 2.158                   | 1.378                   | 120.38                  | -299.6061                           |
|                    | $C_{2v}(\mathbf{a})$             | 1/2                      | ${}^{2}A_{1}$  | 1.994                   | 1.298                   |                         | -299.6278                           |
|                    | a                                | 5/2                      | $^6\mathrm{B}_2$   | 2.041                   | 1.357                   | co 55                   | -299.4945                           |
|                    | $C_1$                            | 1/2                      |  | 2.163                   | 1.429                   | 69.55                   | -299.5063                           |
|                    | C(A)                             | 5/2                      |  | 2.166                   | 1.399                   | 60.53                   | -299.5252                           |
|                    | $C_1(a)$                         | 1/2                      |  | 1.797                   | 1.305                   | 97.05                   | -299.6234                           |
|                    |                                  | 3/2                      |  | 1.883                   | 1.309                   | 115.50                  | -299.5994                           |
|                    | C (1)                            | 5/2                      |  | 1.905                   | 1.323                   | 109.70                  | -299.5315                           |
|                    | $C_1(b)$                         | 1/2                      |  | 1.912                   | 1.327                   | 86.26                   | -299.6746                           |
|                    |                                  | 3/2                      |  | 2.017                   | 1.372                   | 89.81                   | -299.6417                           |
| 01. C              | C                                | 5/2                      | 25   | 2.165                   | 1.377                   | 116.37                  | -299.6075                           |
| RhC <sub>6</sub>   | $C_{\infty V}$                   | 1/2                      | 2 <u>5</u><br>4 <del>5</del>   | 1.871                   | 1.259                   |                         | -337.7612                           |
|                    |                                  | 3/2                      | $^{2}\sum_{\substack{4\sum\\6\sum\\2A_{1}}}$   | 1.824                   | 1.297                   |                         | -337.7853                           |
|                    | C                                | 5/2                      | 2 1  | 1.919                   | 1.312                   | 70.120                  | -337.7082                           |
|                    | $C_{2v}$                         | 1/2<br>3/2               | ${}^{2}\mathrm{A}_{1}$ ${}^{4}\mathrm{A}_{2}$  | 1.885                   | 1.296                   | 70.139                  | -337.6796<br>-337.6754              |
|                    |                                  | 3/2<br>5/2               | 'A <sub>2</sub><br>6 <b>D</b>  | 2.008                   | 1.291                   | 45.053<br>50.637        | -337.6754<br>-337.6162              |
|                    | $C_1$                            | 5/2<br>1/2               | ${}^6\mathrm{B}_1^{2}\mathrm{A}$   | 1.942<br>1.914          | 1.284                   | 59.637                  | -337.6162<br>-337.7488              |
|                    | C <sub>1</sub>                   | 3/2                      |  |                         | 1.311                   | 102.60                  | -337.7488<br>-337.7487              |
|                    |                                  | 4//                      | <sup>4</sup> A   | 1.939<br>1.997          | 1.292<br>1.433          | 95.76<br>91.75          | -337.7487                           |
|                    |                                  | 5/2                      | h A  |                         |                         |                         |                                     |
|                    |                                  | 5/2                      | $^6\mathrm{A}_2$   |                         |                         |                         | -337.7167                           |
|                    | $C_1(a)$                         | 5/2<br>1/2               | $^6\mathrm{A}_2$   | 2.009                   | 1.317                   | 42.27                   | -337.7374                           |
|                    |                                  | 5/2<br>1/2<br>3/2        | $^6\mathrm{A}_2$   | 2.009<br>2.019          | 1.317<br>1.308          | 42.27<br>40.66          | -337.7374 $-337.7004$               |
|                    | $C_1(\mathbf{a})$                | 5/2<br>1/2<br>3/2<br>5/2 | $^6\mathrm{A}_2$   | 2.009<br>2.019<br>2.005 | 1.317<br>1.308<br>1.329 | 42.27<br>40.66<br>40.80 | -337.7374<br>-337.7004<br>-337.6485 |
|                    |                                  | 5/2<br>1/2<br>3/2        | $^6\mathrm{A}_2$   | 2.009<br>2.019          | 1.317<br>1.308          | 42.27<br>40.66          | -337.7374 $-337.7004$               |

<sup>&</sup>lt;sup>a</sup> R1 denote the shortest Rh–C bond lengths. <sup>b</sup> R2 denotes the shortest C–C bond length. <sup>c</sup> α denotes the bond angle of C–Rh–C. <sup>d</sup> E<sub>b</sub> represents the total energy of  $RhC_n$  clusters.

TABLE 2: Mulliken Atomic Net Populations of  $RhC_n$  (n = 1-6) Clusters at the B3LYP Level Employing LanL2DZ Basis Sets

| cluster | symmetry       | spin | Rh     | $C_1$   | $C_3$   | $C_5$   |
|---------|----------------|------|--------|---------|---------|---------|
| RhC     | $C_{\infty v}$ | 1/2  | 0.0051 |         |         |         |
| $RhC_2$ | $C_{\infty v}$ | 3/2  | 0.3401 | -0.2381 |         |         |
| $RhC_3$ | $C_{\infty v}$ | 1/2  | 0.2095 | 0.0372  | 0.0390  |         |
| $RhC_4$ | $C_{\infty v}$ | 3/2  | 0.3097 | -0.1768 | -0.2545 |         |
| $RhC_5$ | $C_{\infty v}$ | 1/2  | 0.2155 | -0.1023 | 0.1546  | -0.0076 |
| $RhC_6$ | $C_{\infty v}$ | 3/2  | 0.2856 | -0.2058 | 0.0226  | -0.2869 |

optimized results show that the  $C \sim v$  RhC<sub>3</sub> cluster has a preference for the low spin state.

For the calculated total energies of the RhC<sub>3</sub> clusters described above, we calculated that the doublet  $C \infty v$  isomer is the most stable structure; the RhC<sub>3</sub> ( $C \infty v$ , S=1/2) is selected as the ground state with electronic state being  $^2\Sigma$ . The linear ground state for the RhC<sub>3</sub> cluster is similar to the geometries of the AlC<sub>3</sub><sup>52</sup>, NaC<sub>3</sub><sup>53</sup> and PdC<sub>3</sub><sup>25</sup> clusters. However, it significantly differs from those of transition metal carbides, such as LaC<sub>3</sub><sup>28</sup>, NbC<sub>3</sub><sup>29</sup>, YC<sub>3</sub><sup>27</sup>, and ScC<sub>3</sub><sup>30</sup>.

3.1.4.  $RhC_4$ . The RhC<sub>4</sub> clusters with  $C_{\infty\nu}$ ,  $D_{\infty h}$ , and  $C_{3\nu}$  symmetries were selected as the initial geometries to be optimized. The linear  $C_{\infty\nu}$  RhC<sub>4</sub> isomer with the Rh atom localized at the terminal was obtained by adding one C atom to the most stable RhC<sub>3</sub> cluster and optimizing to obtain a stable structure. The linear  $D_{\infty h}$  geometry with the Rh atom located at the central site (Figure 1) was optimized. However, the  $D_{\infty h}$  isomer is less stable than the quartet  $C_{\infty\nu}$  isomer.

After the geometry of the unstable Rh-centered  $C_{3v}$  isomer is relaxed to be the  $C_{2v}$  symmetry, the doublet and quartet  $C_{2v}$  isomers are optimized to be the stable structures while the sextet  $C_{2v}$  isomer converges to be the stable  $C_s$  isomer during the geometry optimization. As seen from Figure 1, the sextet  $C_s$  isomer is analogous to the  $C_{2v}$  isomer in geometry, however, the  $C_s$  isomer with the same spin state is lower by about 0.14 hartree in total energy than the  $C_{2v}$  isomer.

As for the kitelike  $C_{2\nu}(a)$  RhC<sub>4</sub> structure, it was found that the terminal carbon atoms are attached to the C ring with the rhodium atom embedded. Like the  $C_{2\nu}(a)$  isomer, the total energy of the fanlike  $C_s(a)$  isomer increases monotonically, together with a drop of stability, when the spin increases from S=1/2, to 5/2. It should be mentioned that the fanlike  $C_s(a)$  RhC<sub>4</sub> geometry is lower in total energy than the  $D_{\infty h}$ ,  $C_{2\nu}$ ,  $C_{2\nu}(a)$ , and  $C_s$  isomers. However, it is higher by about 0.01 hartree in total energy than the  $C_{\infty \nu}$  isomer with the same spin state. Finally, the stable scooplike  $C_1$  isomer, which is described as a distortion of the  $C_{2\nu}(a)$  RhC<sub>4</sub> isomer above, was considered. We found that its total energy is slightly higher than those of the  $C_s(a)$  and  $C_{\infty \nu}$  RhC<sub>4</sub> clusters.

On the basis of the calculated total energies of the RhC<sub>4</sub> isomers, the linear RhC<sub>4</sub> ( $C_{\infty\nu}$ , S=3/2) cluster is the most stable structure and its electronic ground state is  $^4\Sigma$ . The most stable  $C_{\infty\nu}$  RhC<sub>4</sub> structure is similar to that of the PdC<sub>4</sub> isomer<sup>25</sup> and deviates strongly from those of the TMC<sub>4</sub> (TM = Y, Nb, La, Sc, and Cr)<sup>26-30</sup> clusters in geometrical structures.

3.1.5.  $RhC_5$ . Guided by the calculated results on the RhC<sub>3</sub> and RhC<sub>4</sub> clusters, the RhC<sub>5</sub> clusters with respective  $C_{\infty v}$ ,  $C_{4v}$ , and  $C_{5v}$  symmetries were considered. However, the Rh-capped tetragonal bipyramidal  $C_{4v}$  isomer and the Rh-capped pentagonal pyramidal  $C_{5v}$  isomer turned out not to be stable isomers. After a slight distortion of the  $C_{5v}$  isomer, stable doublet and sextet Rh-capped  $C_1$  RhC<sub>5</sub> structures were finally obtained.

Ringlike  $C_{2v}$  and kitelike  $C_{2v}(a)$  isomers were optimized to be stable geometries, whereas the quartet  $C_{2v}$  isomer corresponds

to a transition state. A relaxation of the quartet  $C_{2\nu}$  geometry along the coordinates of the imaginary vibrational mode was carried out until a true local minimum  $C_1(b)$  isomer was reached. As seen from Figure 1, the  $C_1(b)$  geometry is analogous to the  $C_{2\nu}$  geometry. For the kitelike  $C_{2\nu}(a)$  isomer, the doublet and sextet  $C_{2\nu}(a)$  isomers are actually formed to be the stable states. However, the quartet isomer corresponds to a transition state.

As the most likely ground state candidate, the linear  $C_{\infty \nu}$  structure was also optimized. The doublet  $C_{\infty \nu}$  isomer is actually the ground state of the RhC<sub>5</sub> unit. In addition, a bent  $C_1(a)$  isomer, which can be viewed as the Rh atom substituted for a C atom on the center site of the linear  $C_{\infty \nu}$  RhC<sub>5</sub> cluster, is also optimized. The calculated results indicate that the  $C_1(a)$  isomer is the stable geometry. However, its total energy is higher than that of the linear  $C_{\infty \nu}$  isomer.

According to the investigated RhC<sub>5</sub> clusters, it is found that the Rh atom in the most stable RhC<sub>5</sub> isomer localizes at the terminal of the linear structure and the corresponding electronic state is  $^2\Sigma$ . The most stable geometry is different from those of TMC<sub>5</sub> (TM = Y, Nb, La, and Sc)<sup>27–30</sup> isomers. However, it is consistent with those of the CrC<sub>5</sub><sup>26</sup> and PdC<sub>5</sub><sup>25</sup> clusters.

**3.1.6.** RhC<sub>6</sub>. On the basis of the investigations on the RhC<sub>n</sub> (n = 2-5) clusters above, a variety of possible RhC<sub>6</sub> isomers with respective  $C_{\infty\nu}$ ,  $D_{6h}$ , and  $C_{6\nu}$  symmetries as the ground state candidates were investigated. The optimized results show that the Rh-capped  $C_{6\nu}$  isomer and the  $D_{6h}$  isomer are the unstable structures. After an adjustment of the  $D_{6h}$  geometry, a bicyclic fused stable  $C_1(a)$  isomer, which contains a triangular C-Rh-C ring, is obtained. A cyclic stable  $C_1$  structure was also obtained by changing the location of the Rh atom and the tailing C atom. (Figure 1,  $C_1$ ). When spin goes from S = 1/2, to 3/2, and to 5/2, the total energies of the  $C_1$  and  $C_1(a)$  isomers increase. The  $C_1$  isomer tends to have lower total energy as compared to the  $C_1(a)$  isomer with the same spin value. Consequently, the cyclic doublet  $C_1$  structure with electronic state of  ${}^2A$  is more stable than the doublet  $C_1(a)$  isomer. A kitelike stable  $C_{2v}$  isomer is found. The stable doublet  $C_{2\nu}$  isomer is higher by about 0.07 hartree in total energy than the doublet  $C_1$  isomer.

Two RhC<sub>6</sub> isomers are expected to be the most likely ground state candidates: one is a linear  $C_{\infty v}$  structure; The other is a distorted linear  $C_1(b)$  isomer which can obtained by an exchange of one C atom for one Rh atom in the linear  $C_{\infty y}$  structure; The calculated results show that the  $C_{\infty \nu}$  and  $C_1(b)$  RhC<sub>6</sub> isomers are the stable structures. In analoge to the  $C_{\infty \nu}$  RhC<sub>4</sub> isomer, it is surprisingly found that the slight deviated linear doublet  $C_1$ -(b) RhC<sub>6</sub> isomer is higher by about 0.06 hartree in total energy than the linear quartet  $C_{\infty v}$  isomer, the quartet  $C_{\infty v}$  isomer is lower in total energy than the doublet and sextet isomers. Thus, the linear quartet  $C_{\infty \nu}$  isomer is the most stable structure with electronic state of  $^4\Sigma.$  The Rh atom in the lowest-energy RhC  $_6$ geometry occupies the terminal site and Rh acts as only an impurity atom for the carbon chain. It should be pointed out that the equilibrium geometry of the most stable  $C_{\infty \nu}$  RhC<sub>6</sub> cluster is similar to those of the TMC<sub>6</sub> (TM= Cr, Y, Pd, and Nb) $^{25-27,30}$  clusters.

The available experimental data indicate that the linear pure carbon  $C_n$  chains with  $2 \le n \le 9$  are the most stable structures.<sup>54</sup> As far as the RhC<sub>n</sub> clusters are concerned, the linear isomer with the Rh atom acting as the terminal is the lowest-energy form for each unit of the RhC<sub>n</sub> ( $n \le 1-6$ ) series. The terminal Rh in the linear chain structure does not directly bind with all carbon atoms and acts as only an impurity atom without changing the framework of the most stable  $C_n$  structure. Moreover, the electronic ground state of RhC<sub>n</sub> clusters has a

TABLE 3: Natural Populations and Natural Electron Configurations of the Most Stable RhC<sub>n</sub> (n = 1-6) Clusters

|         | -              |      |      |                    |   |
|---------|----------------|------|------|--------------------|---|
| system  | symmetry       | spin | atom | natural population | natural electron configuration  |
| RhC     | $C_{\infty v}$ | 1/2  | Rh   | -0.1379            | [core]5s <sup>0.69</sup> 4d <sup>8.45</sup> 5p <sup>0.01</sup>                    |
|         |                |      | C    | 0.1379             | $[\text{core}]2\text{s}^{1.89}2\text{p}^{1.96}3\text{s}^{0.01}$                   |
| $RhC_2$ | $C_{\infty v}$ | 3/2  | Rh   | 0.4099             | [core]5s <sup>0.35</sup> 4d <sup>8.26</sup>                                       |
|         |                |      | C(1) | -0.4007            | [core] $2s^{1.34}2p^{3.03}3s^{0.02}3p^{0.02}$                                     |
|         |                |      | C(2) | -0.0092            | [core] $2s^{1.63}2p^{2.35}3s^{0.02}3p^{0.01}$                                     |
| $RhC_3$ | $C_{\infty v}$ | 1/2  | Rh   | 0.2392             | [core]5s <sup>0.31</sup> 4d <sup>8.47</sup>                                       |
|         |                |      | C(1) | 0.1316             | [core] $2s^{1.27}2p^{2.58}3s^{0.01}3p^{0.01}$                                     |
|         |                |      | C(2) | -0.5211            | [core] $2s^{1.27}2p^{2.58}3s^{0.01}3p^{0.01}$                                     |
|         |                |      | C(3) | 0.1504             | [core] $2s^{1.62}2p^{2.22}3s^{0.01}3p^{0.01}$                                     |
| $RhC_4$ | $C_{\infty v}$ | 3/2  | Rh   | 0.3914             | [core]5s <sup>0.28</sup> 4d <sup>8.34</sup>                                       |
|         |                |      | C(1) | -0.1422            | [core]2s <sup>1.25</sup> 2p <sup>2.87</sup> 3s <sup>0.01</sup> 3p <sup>0.01</sup> |
|         |                |      | C(2) | -0.0454            | $[\text{core}]2\text{s}^{0.94}2\text{p}^{3.09}3\text{p}^{0.01}$                   |
|         |                |      | C(3) | -0.2770            | $[\text{core}]2\text{s}^{0.99}2\text{p}^{3.28}3\text{p}^{0.02}$                   |
|         |                |      | C(4) | 0.0732             | $[\text{core}]2\text{s}^{1.63}2\text{p}^{2.28}3\text{s}^{0.01}3\text{p}^{0.01}$   |
| $RhC_5$ | $C_{\infty v}$ | 1/2  | Rh   | 0.2572             | [core]5s <sup>0.30</sup> 4d <sup>8.46</sup>                                       |
|         |                |      | C(1) | 0.0595             | $[\text{core}]2\text{s}^{1.27}2\text{p}^{2.66}3\text{s}^{0.01}3\text{p}^{0.01}$   |
|         |                |      | C(2) | -0.2615            | [core] $2s^{0.91}2p^{3.34}3p^{0.01}$  |
|         |                |      | C(3) | 0.2615             | $[\text{core}]2\text{s}^{0.92}2\text{p}^{2.86}3\text{p}^{0.01}$                   |
|         |                |      | C(4) | -0.3949            | [core]2s <sup>0.98</sup> 2p <sup>3.32</sup> 3p <sup>0.02</sup>                    |
|         |                |      | C(5) | 0.1233             | [core] $2s^{1.62}2p^{2.25}3s^{0.01}3p^{0.01}$                                     |
| $RhC_6$ | $C_{\infty v}$ | 3/2  | Rh   | 0.3652             | $[\text{core}]5S^{0.28}4d^{8.36}5p^{0.01}$  |
|         |                |      | C(1) | -0.1102            | [core] $2S^{1.26}2p^{2.83}3S^{0.01}3p^{0.01}$                                     |
|         |                |      | C(2) | -0.1068            | [core] $2S^{0.92}2p^{3.17}3p^{0.01}$  |
|         |                |      | C(3) | -0.0094            | [core] $2S^{0.89}2p^{3.11}3p^{0.01}$  |
|         |                |      | C(4) | 0.0901             | [core] $2S^{0.90}2p^{3.00}3p^{0.01}$  |
|         |                |      | C(5) | -0.3094            | [core] $2S^{0.97}2p^{3.32}3p^{0.02}$  |
|         |                |      | C(6) | 0.0805             | [core] $2S^{1.63}2p^{2.28}3S^{0.01}3p^{0.01}$                                     |

TABLE 4: HOMO-LUMO Gap of the Most Stable RhC<sub>n</sub> Clusters (unit: Hartree)

| system  | symmetry       | spin | α-НОМО   | α-LUMO   | $\beta$ -HOMO | $\beta$ -LUMO | gap (eV) |
|---------|----------------|------|----------|----------|---------------|---------------|----------|
| RhC     | $C_{\infty v}$ | 1/2  | -0.19193 | -0.09815 | -0.24543      | -0.11965      | 1.96602  |
| $RhC_2$ | $C_{\infty v}$ | 3/2  | -0.24724 | -0.10300 | -0.24698      | -0.15865      | 2.40258  |
| $RhC_3$ | $C_{\infty v}$ | 1/2  | -0.24768 | -0.12028 | -0.23793      | -0.14476      | 2.53422  |
| $RhC_4$ | $C_{\infty v}$ | 3/2  | -0.25134 | -0.10697 | -0.24342      | -0.16633      | 2.09685  |
| $RhC_5$ | $C_{\infty v}$ | 1/2  | -0.25096 | -0.13944 | -0.23780      | -0.14564      | 2.50675  |
| $RhC_6$ | $C_{\infty v}$ | 3/2  | -0.24234 | -0.10555 | -0.24116      | -0.17140      | 1.89747  |

strong odd—even alternation as a function of the number of carbon atoms: the doublet isomers with electronic state of  $^2\Sigma$  are associated with n-odd members of series while the quartet isomers with electronic state of  $^4\Sigma$  are associated with the n-even members. This even—odd effect is similar to those occurring in the  $C_n$ ,  $^{54}$  Pd $C_n$ ,  $^{25}$  Ca $C_n$ ,  $^{14}$  Nb $C_n$ ,  $^{29}$  and  $Zr_n$ ,  $^{55}$  clusters.

**3.2. Binding Energies and Fragmentation Energies.** To predict the relative stability of the most stable  $RhC_n$  (n=1-6) clusters, it is significant to calculate the averaged atomic binding energies ( $E_b(n)$ ) with respect to the isolated atoms and fragmentation energies (D(n, n-1)) with respect to removal of one C atom from the RhCn clusters. The averaged atomic binding energies and fragmentation energies of the RhCn (n=1-6) clusters are defined as<sup>45</sup>

$$E_{b}(n) = \frac{nE_{T}(C) + E_{T}(Rh) - E_{T}(RhC_{n})}{n+1}$$
  
$$D(n, n-1) = E_{T}(RhC_{n-1}) + E_{T}(C) - E_{T}(RhC_{n})$$

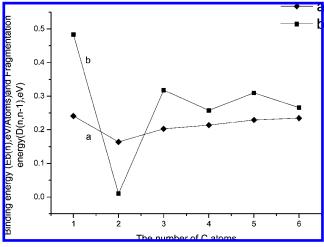
Where  $E_T(RhC_{n-1})$ ,  $E_T(C)$ ,  $E_T(Rh)$ , and  $E_T(RhC_n)$  represent the total energies of the most stable  $RhC_{n-1}$ , C, Rh, and  $RhC_n$  clusters, respectively.

The calculated results are plotted in Figure 2. Curve a in Figure 2 shows the size dependence of the average atomic binding energies  $(E_b(n))$  of RhC<sub>n</sub> clusters; curve b in Figure 2 shows the size dependence of the fragmentation energies D(n, n-1) of RhC<sub>n</sub> clusters. The features of the size evolution are obviously displayed, and the peaks of the curves correspond to those clusters having enhanced relative stabilities. It can be seen

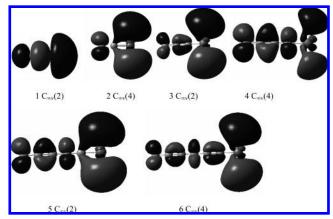
from curve a that the average atomic binding energies generally fall as the number of carbon atoms increases from n = 1 to n= 2. Thereafter, the average atomic binding energies of the clusters begin to increase with increasing n. This regularity is in good agreement with the observed regularity of the IrSi<sub>n</sub> (n = 2-6) clusters. 40 According to the calculated fragmentation energies shown in curve b, three marked peaks are found respectively at n = 1, n = 3, and n = 5; indicating that the corresponding clusters have enhanced stabilities and have high abundances in mass spectrum as compared to the corresponding neighbors. Consequently, the  $RhC_n$  clusters with odd-numbers of carbon atoms are relatively more stable than the neighboring ones with even-numbers of carbon atoms. The general behavior of the RhC<sub>n</sub> clusters is very similar to that of the  $C_n^{54}$  and  $ScC_n^{30}$ clusters; however, it is opposite to that of PbC<sub>n</sub>, <sup>25</sup> CaC<sub>n</sub>, <sup>14</sup> and  $MgC_n^{11}$  clusters.

**3.3. Population Analyses.** Our calculated results of Mulliken atomic net population on RhC<sub>n</sub> (n=1-6) clusters obtained at the B3LYP/LanL2DZ level are summarized at Table 2. We find that the Mulliken atomic net population of Rh in the most stable RhC<sub>n</sub> (n=1-6) clusters is positive, and that the nearest neighbor C of the Rh atom is negative (except in RhC<sub>3</sub>) is in good agreement with what is found in NbC<sub>n</sub> (n=3-8),<sup>29</sup> YC<sub>n</sub> (n=2-6),<sup>27</sup> and LaC<sub>n</sub> (n=2-6)<sup>28</sup> clusters.

Mulliken populations were sometimes found to be misleading for  $TMSi_n$  clusters,<sup>56</sup> whereas natural populations and natural electron configurations are a reasonable explanation of the charge transfer within the clusters.<sup>40,45,57</sup> For this reason, we calculated the natural population and natural electron configurations are listed in Table 3 for the  $RhC_n$  (n = 1-6) clusters.



**Figure 2.** Size dependence of the average atomic binding energies and the fragmentation energies of the most stable  $RhC_n$  clusters.



**Figure 3.** Contour maps of the HOMOs of the most stable  $RhC_n$  clusters; For example, the 6  $C_{\infty v}(4)$  represents the most stable quartet  $RhC_n$  cluster.

From Table 3, one finds that the natural population of the Rh atom for RhC is negative, and that it is positive for the stable  $RhC_n$  (n = 2-6) clusters. On the basis of the calculated natural electron configurations, we show that more than 8.3 electrons occupy the 4d subshell of Rh in the RhC<sub>n</sub> clusters, and that the 4d orbitals of the Rh atom in RhC<sub>n</sub> clusters do not behave as core orbitals but take an active role in chemical bonding. The charges in the RhC<sub>n</sub> (n = 2-6) clusters mainly transfer from the 5s orbitals of the Rh atom and the 2s orbitals of the C atoms to the 4d orbitals of the Rh atom and the 2p orbitals of the C atoms. Consequently, the Rh atom in the most stable  $RhC_n$ clusters acts as an electron acceptor, influencing the chemical bonding properties of the  $C_n$  framework. Natural population of Rh atom in the RhC<sub>n</sub>(n = 1-6) units also varies with the n: the natural populations decrease with n in even-n clusters, whereas the natural populations increase with n in odd-n clusters.

**3.4 HOMO–LUMO Gap.** The electronic properties of RhC<sub>n</sub> clusters are characterized by the energy gap between the highest occupied molecular orbital(HOMO) and the lowest unoccupied molecular orbital(LUMO). The calculated HOMO–LUMO gaps of RhC<sub>n</sub> clusters are listed in Table 4. The calculated results indicate that the HOMO–LUMO gaps of RhC<sub>3</sub> (2.53 eV) and RhC<sub>5</sub> (2.51 eV) clusters are bigger than those of the other RhC<sub>n</sub> clusters. Therefore, the chemical stabilities of these two clusters are improved dramatically as compared to the others, and the clusters with odd numbers of carbon atoms are more stable than those with even-numbers of carbon atoms. These findings are

in good agreement with those of the calculated fragmentation energies above.

**3.5. Chemical Bonding Properties.** As seen from the HOMO of RhC<sub>n</sub> clusters (Figure 3), the  $\sigma$ -type bond of RhC dimer is formed between the Rh and C atoms. However, when one Rh atom is capped on the  $C_2$  molecule, the  $\sigma$ -type bond between Rh and C atoms is changed to be the  $\pi$ -type bond. As observed from the HOMO of RhC<sub>3</sub> isomer, the  $\pi$ -type orbitals extend over the 2C and 3C atoms (The order of carbon atoms is named from left to right). However, the shape of the HOMO and bonding property for the most stable RhC<sub>3</sub> cluster are obviously different from those of RhC2 cluster. For the RhC4, its HOMO obviously corresponds to a  $\pi$ -type bond between the 2C and 3C atoms, the Rh and 4C atoms formed the weak delocalized  $\pi$ -type bond, which is similar to that of RhC<sub>2</sub> cluster. However, for the RhC<sub>5</sub>, its HOMO corresponds to the delocalized  $\pi$ -type bonds on 2C-3C and Rh-5C bonds. As far as the RhC<sub>6</sub> isomer is concerned, the 2C-3C and 5C-6C bonds are  $\pi$ -type bonds; The 1C, 4C, and Rh atoms appear the localized  $\pi$ -type bonds.

#### 4. Conclusions

The RhC<sub>n</sub> (n = 1-6) clusters with different spin states have been studied employing the B3LYP method with the Lan2DZ basis sets. The total energies, equilibrium geometries, and stabilities of RhC<sub>n</sub> (n = 1-6) clusters, together with fragmentation energies and averaged atomic binding energies, were presented and discussed.

According to the calculated results, we find that the most stable Rh-terminated linear  $RhC_n$  (n=1-6) geometries are similar to those of the ground states of the  $C_{n+1}$  frameworks. Therefore, linear structures are the most interesting possible targets for an experimental search of rhodium-doped carbon clusters. Furthermore, the electronic ground state has a strong even—odd alternation: the n-odd  $RhC_n$  clusters have a doublet ground state ( $^2\Sigma$ ), whereas the n-even isomers have a quartet ground state ( $^4\Sigma$ ). This feature is also found in  $C_n$ ,  $^{54}$   $PdC_n$ ,  $^{25}$   $CaC_n$ ,  $^{14}$  and  $NbC_n$  clusters.

The relative stability of the different clusters has been discussed in term of the calculated average atomic binding energies and fragmentation energies. The calculated results exhibit an even—odd alternation trend in the cluster's stability: the n-odd RhC $_n$  clusters are more stable than the corresponding n-1 and n+1 ones. The relative stability of the RhC $_n$  (n=1-6) clusters agrees with findings for  $C_n^{54}$  and  $ScC_n^{30}$  clusters, but it is opposite to what is found for PbC $_n^{25}$  CaC $_n^{14}$  and MgC $_n^{11}$  clusters.

The population analysis indicates that the charges in the RhC<sub>n</sub> (n = 2-6) clusters mainly transfer from the 5s orbitals of Rh atom and the 2s orbitals of C atoms to the 4d orbitals of Rh atom and the 2p orbitals of C atoms. The Rh atom in the most stable RhC<sub>n</sub> clusters acts as an electron absorber. Generally, the charges are transferred from the Rh atom to the C<sub>n</sub> in the most stable RhC<sub>n</sub> (n = 1-6) clusters.

Finally, the HOMO-LUMO gaps exhibit size and species dependence. These findings are in good agreement with those of the calculated fragmentation energy. In addition, the chemical bonding properties of  $RhC_n$  clusters is also discussed in this article.

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