# Application of the Stabilization Method to Temporary Anion States of $\pi$ -Ligand Transition-Metal Carbonyls in Density Functional Theory

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In this paper, density functional theory is used to investigate (benzene)chromium tricarbonyl, (cyclopentadienyl)manganese tricarbonyl, (1,3-butadiene)iron tricarbonyl, and (cyclopentadienyl)cobalt dicarbonyl. For the energies of low-lying temporary anion states, the stabilized Koopmans-based (S-KB) and stabilized Koopmans theorem (S-KT) methods are adopted. Stabilization is accomplished by varying the exponents of appropriate diffuse functions. Results indicate that the calculations of S-KB using PBEPBE and S-KT using CAM-B3LYP are able to yield energies of temporary anion states in good agreement with the experimental values. Furthermore, the ionization potentials can be determined accurately via the Koopmans-based (KB) PBEPBE method.

#### 1. Introduction

The carbonyl-containing transition-metal complexes have aroused both theoretical and experimental interest due to their importance in organic synthesis, catalytic chemistry, and application in potential nonlinear optics. <sup>1–4</sup> The determination of ionization potentials (IPs), electron affinities (EAs), and frontier molecular orbitals are imperative in understanding the bonding, catalytic properties, and nonlinear optical activity of this class of transition-metal complexes. In general, the IPs and EAs can be determined by means of photoelectron spectroscopy (PES)<sup>5</sup> and electron transmission spectroscopy (ETS), <sup>6,7</sup> respectively. However, the assignment of the photoelectron spectra was not accurate in most of the recently studied organometallic systems, as mentioned in the work of Gengeliczki et al. <sup>8</sup> Furthermore, theoretical investigation of the unoccupied orbitals of organometallic complexes is also scarce.

In the theoretical prediction of IPs and EAs, the Koopmans theorem (KT)<sup>9</sup> approximation using Hartree-Fock (HF) or Kohn-Sham (KS) orbital energies is usually adopted. The IPs and EAs in the KT approximation are associated with the negatives of the energies of the filled and unfilled orbitals, respectively. However, it neglects relaxation and correlation effects. Meanwhile, the KT approximation does not mean that any orbital energy may be considered an approximate electron binding energy. Some one-electron equations give physically meaningful orbital energies. Others may not. In using the KT approximation, most density functional theory (DFT)<sup>10</sup> potentials will not yield asymptotic behavior properly. 11,12 Various schemes have been devised for long-range correction (LRC). For instance, Tozer and co-workers have proposed an alternative Koopmansbased (KB) approximation based on the consideration of the integer discontinuity ( $\Delta_{xc}$ ) in the exact exchange-correlation potential. 12-17 However, for species with negative EAs, the temporary anion is unstable with respect to electron detachment. Thus, the unfilled orbitals are prone to collapse onto approximations of continuum functions called orthogonalized discretized continuum (ODC)<sup>18-21</sup> when large basis sets are used. Therefore, the energy calculations of temporary anion states using both the KT and KB approaches are not considered definitive.

The stabilization method proposed by Taylor and co-workers<sup>22–25</sup> can distinguish the temporary anion orbital solutions from the ODC solutions. The vertical attachment energies (AEs) are associated with the energies of the "stabilized" temporary anion states of the neutral molecules. By combining the stabilization method and consideration of LRC, we have applied the S-KB method (i.e., the stabilization method coupled with KB) in the studies of a series of molecules. <sup>26,27</sup> Results have indicated that the S-KB approach has yielded an improvement in predicting both the absolute and relative energies of temporary anion states over other approaches.

In this paper, we continue to study prototypical transition-metal carbonyls such as (benzene)chromium tricarbonyl, (cyclopentadienyl)manganese tricarbonyl, (1,3-butadiene)iron tricarbonyl, and (cyclopentadienyl)cobalt dicarbonyl. These important compounds have been studied previously by the multiple scattering  $X\alpha$  (MS- $X\alpha$ ) method. However, the stability of the temporary anion states was not established for  $\pi$ -ligand transition-metal carbonyls. In the study of the temporary anion resonances, the S-KB method using the local functional will be adopted first. Then, the S-KT method (i.e., the stabilization method coupled with KT)<sup>18,19,29</sup> using the LRC functional will be invoked. As to the filled molecular orbitals (MOs), we will apply the KB and KT approximations to the IPs. Finally, the results will be compared with experimental values.

## 2. Computational Method

The IPs and EAs in the KT approximation can be written as  $IP^{KT} \approx -\epsilon_{OMO}$  and  $EA^{KT} \approx -\epsilon_{VMO}$ , where  $\epsilon_{OMO}$  and  $\epsilon_{VMO}$  denote the occupied and virtual molecular orbital energies, respectively. The asymptotic correction of the Koopmans value in the KB approximation is roughly half the integer discontinuity:

$$\frac{\Delta_{\rm xc}}{2} \approx \varepsilon_{\rm HOMO} + (E_{N-1} - E_N) \tag{1}$$

Notice that  $\varepsilon_{HOMO}$  is the highest-occupied molecular orbital (HOMO) energy determined from a DFT calculation using a

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local exchange-correlation functional on the neutral system. Here,  $E_N$  and  $E_{N-1}$  are the total electronic energies of the neutral and cation, respectively. By applying the correction terms  $\Delta_{xc}/2$  and  $-\Delta_{xc}/2$  to IP<sup>KT</sup> and EA<sup>KT</sup>, respectively, the IP and EA in the KB approximation can be written as

$$IP^{KB} \approx -\varepsilon_{OMO} + [\varepsilon_{HOMO} + (E_{N-1} - E_N)]$$
 (2)

and

$$EA^{KB} \approx -\varepsilon_{VMO} - [\varepsilon_{HOMO} + (E_{N-1} - E_N)]$$
 (3)

The vertical attachment energy (i.e., -EA), can then be represented as

$$AE^{KB} \approx \varepsilon_{VMO} + [\varepsilon_{HOMO} + (E_{N-1} - E_N)]$$
 (4)

The virtual orbital energy associated with temporary anion state is also known as AE. The AE<sup>KB</sup> will be denoted as  $\varepsilon_{VMO}^{KB}$ .

To distinguish the temporary anion state solutions from the virtual ODC ones, the stabilization method is employed. Nine different Gaussian-type basis sets, designated as A1, A2, A3, B1, B2, B4, C1, C2, and C4 are used in our calculations. The convention of designation is explained as follows. For the C and O atoms, the 6-31G+ $\alpha p_1$  basis set A is formed by augmenting the 6-31G basis set with the diffuse  $p_1$  function multiplied by a scale factor  $\alpha$  (denoted by  $\alpha p_1$ ). The  $6-31+G(d)+\alpha p_2$  basis set B is formed by augmenting the 6-31+G(d) basis set with the diffuse  $\alpha p_2$  function. The augcc-pvdz+ $\alpha p_3$  basis set C is formed by augmenting the aug-ccpvdz basis set with the diffuse  $\alpha p_3$  function. The  $p_1$ ,  $p_2$ , and  $p_3$ functions have the exponents of 0.0562, 0.0146, and 0.0135 for the C atom and 0.0900, 0.0282, and 0.0229 for the O atom, respectively. For metal atoms, the following four different Gaussian-type basis sets 1-4 are employed and they are denoted alongside A-C basis sets.

- (1) The Wachters+ $\alpha d_1$  basis set is formed by augmenting the (14s11p6d)/[8s6p4d] all-electron basis set from Wachters<sup>30</sup> with the diffuse  $\alpha d_1$  function. The diffuse  $d_1$  functions have the exponents of 0.0304, 0.0351, 0.0378, and 0.0406 for the Cr, Mn, Fe, and Co atoms, respectively.
- (2) The akr45+ $\alpha d_2$  basis set is formed by augmenting the (13s10p5d)/[5s4p2d] akr45 basis set of Rappe, Smedley, and Goddard<sup>31</sup> with the diffuse  $\alpha d_2$  function. The diffuse  $d_2$  functions have the exponents of 0.0540, 0.0600, 0.0704, and 0.0783 for the Cr, Mn, Fe, and Co atoms, respectively.
- (3) The  $6-31G+\alpha d_3$  basis set is formed by augmenting the 6-31G basis set with the  $\alpha d_3$  diffuse function. The diffuse  $d_3$  functions have the exponents of 0.1335, 0.1539, 0.1681, and 0.1852 for the Cr, Mn, Fe, and Co atoms, respectively.
- (4) The 6-31+G(d)+ $\alpha d_1$  basis set is formed by augmenting the 6-31+G(d) basis set with the  $\alpha d_1$  diffuse function. In summary, the basis set "Xn" represents both the basis set X (X = A, B, or C) for the C and O atoms and n (n = 1, 2, 3, or 4) for the metal atoms. The inclusion of additional diffuse s functions or d polarization functions on C and O and s functions or p functions on metals is found to be unimportant for the energies of resonance states.

As  $\alpha$  increases, the ODC solutions may approach the temporary anion state orbital solutions in energy and lead to avoided crossing between the two types of solutions. The stabilization graphs are obtained by plotting the calculated energies  $(\epsilon_{VMO}^{KB})$  as a function of the scale factor  $\alpha.$  When the avoided crossing between temporary anion and ODC solutions occurs at their point of closest approach  $\alpha_{ac}$ , the energy of the anion shape resonance is taken as the mean value of these two eigenvalues.  $^{32}$ 

In the present study, we adopt the (1) local functional via the S-KB method and (2) LRC functional via the S-KT method for the temporary anion states. The local functional chosen is the PBEPBE,<sup>33</sup> which utilizes pure generalized gradient approximation (GGA) functional due to its superior calculated AEs.<sup>26,27</sup> For the LRC functional, we use CAM-B3LYP,<sup>34</sup> wB97XD,<sup>35</sup> and LC-wPBE<sup>36</sup> functionals. All calculations are performed using the Gaussian 09 program.<sup>37</sup> The geometry of  $(\eta^6\text{-}C_6\text{H}_6)\text{Cr(CO)}_3$  is optimized under  $C_{3\nu}$ , and the geometries

TABLE 1: Calculated IPs (eV) for  $(\eta^6-C_6H_6)Cr(CO)_3$  and  $(\eta^5-C_5H_5)Mn(CO)_3$ 

			$(\eta^6 - C_6 H_6)$	$_6$ )Cr(CO) $_3$				$(\eta^5-C_5H$	$H_5$ )Mn(CO) $_3$		
method	basis set	$a_1$	e	e	d <sup>b</sup> /eV	a'	a"	a'	a"	a'	d <sup>b</sup> /eV
KBPBEPBE	A1	7.38	7.40	10.52	0.08	8.05	8.06	8.35	9.88	9.96	0.08
	A2	7.45	7.46	10.54	0.08	8.13	8.13	8.44	9.88	9.96	0.11
	A3	7.43	7.44	10.56	0.06	8.05	8.06	8.32	9.94	10.03	0.08
	B1	7.57	7.59	10.64	0.13	8.18	8.19	8.48	9.93	10.01	0.13
	B2	7.58	7.59	10.60	0.14	8.22	8.23	8.53	9.92	10.00	0.16
	B4	7.60	7.64	10.69	0.14	8.19	8.19	8.48	9.94	10.03	0.13
	C1	7.56	7.58	10.62	0.13	8.16	8.17	8.45	9.90	9.98	0.12
	C2	7.57	7.60	10.61	0.14	8.19	8.20	8.49	9.90	9.97	0.14
	C4	7.55	7.59	10.63	0.12	8.17	8.17	8.45	9.91	9.98	0.12
KT <sup>CAM-B3LYP</sup>	A1	7.33	7.00	10.43	0.26	7.94	7.93	8.53	9.38	9.47	0.34
	B1	8.05	7.75	10.51	0.38	8.13	8.12	8.71	9.49	9.58	0.32
	C1	8.02	7.73	10.49	0.37	8.13	8.12	8.68	9.48	9.56	0.32
KTwB97XD	A1	7.89	7.59	11.05	0.33	8.51	8.51	9.09	10.00	10.09	0.38
	B1	8.05	7.75	11.08	0.45	8.65	8.64	9.21	10.06	10.15	0.46
	C1	8.02	7.73	11.05	0.42	8.63	8.62	9.17	10.03	10.12	0.44
KT <sup>LC-wPBE</sup>	A1	8.45	8.13	11.83	0.96	9.17	9.16	9.72	10.73	10.84	0.99
	B1	8.64	8.32	11.87	1.10	9.33	9.32	9.87	10.81	10.91	1.11
	C1	8.61	7.30	11.83	0.81	9.31	9.30	9.83	10.77	10.87	1.08
$X\alpha^a$		7.3	7.4	10.8	_	8.0	8.0	8.2	10.1	9.9	_
$expt^a$		7.42	10.70			8.05	8.40	9.90	10.29		

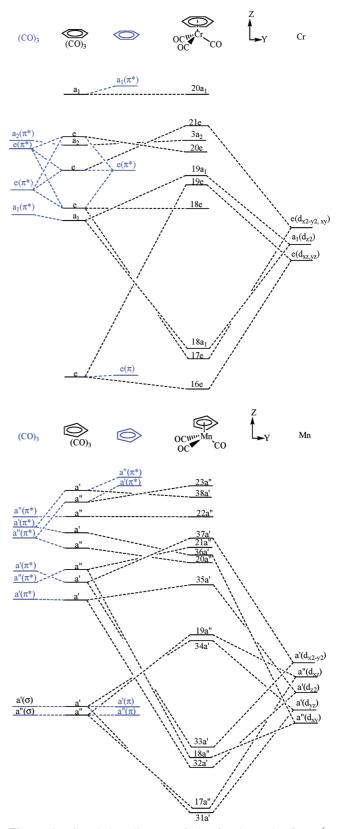
<sup>&</sup>lt;sup>a</sup> The IPs are obtained from previous studies. <sup>28</sup> <sup>b</sup> Here d denotes the mean error relative to experimental IP data.

of  $(\eta^5-C_5H_5)Mn(CO)_3$ ,  $(\eta^4-C_4H_6)Fe(CO)_3$ , and  $(\eta^5-C_5H_5)Co$  $(CO)_2$  are optimized under  $C_S$  symmetry constraints at the B3LYP/6-31+G(d) level. The overall optimized results of structural parameters are close to the crystallographic data.<sup>38–40</sup> For  $(\eta^6-C_6H_6)Cr(CO)_3$ , the optimized bond lengths of Cr–CO and C-O in Cr(CO)<sub>3</sub> and the optimized distance between Cr and the benzene ring are 1.86, 1.16, and 1.74 Å, respectively. For  $(\eta^5-C_5H_5)Mn(CO)_3$ , the optimized bond lengths of Mn–CO and C-O and the distance of Mn-cyclopentadienyl (Mn-Cp) are 1.81, 1.16, and 1.81 Å, respectively. The optimized bond distances of Co-CO, C-O, and Co-Cp are 1.75, 1.16, and 1.73 Å for  $(\eta^5-C_5H_5)Co(CO)_2$ , respectively. Finally, for the optimized structure of (butadiene)Fe(CO)<sub>3</sub>, one unique CO group is aligned with the open side of the butadiene. The bond lengths of the unique Fe-CO and the other two Fe-CO are 1.79 and 1.80 Å, respectively. The distances between Fe and two inner C atoms of the butadiene and those of Fe and two outer C atoms are 2.08 and 2.13 Å, respectively.

### 3. Results and Discussion

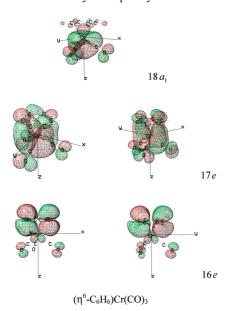
We perform KB calculations using PBEPBE (KBPBEPBE) and KT calculations using CAM-B3LYP, wB97XD, and LC-wPBE (KT<sup>CAM-B3LYP</sup>, KT<sup>wB97XD</sup>, and KT<sup>LC-wPBE</sup>) methods on the filled orbitals for transition-metal carbonyls. The calculated IPs for  $(\eta^6-C_6H_6)Cr(CO)_3$  and  $(\eta^5-C_5H_5)Mn(CO)_3$  are tabulated along with the experimental values in Table 1. In the KBPBEPBE method, the first IPs are due to ionization from the a<sub>1</sub> (denoted as 18a<sub>1</sub>) for  $(\eta^6-C_6H_6)Cr(CO)_3$  and a' (denoted as 33a') for  $(\eta^5-$ C<sub>5</sub>H<sub>5</sub>)Mn(CO)<sub>3</sub>, respectively. The increasing order of IPs of filled MOs are  $a_1 \approx e < e$  for  $(\eta^6 - C_6 H_6) Cr(CO)_3$  and  $a' \approx a'' < e$ a' < a'' < a' for  $(\eta^5 - C_5 H_5) Mn(CO)_3$ . As can be seen from Table 1, the calculated IPs using the KBPBEPBE method are in very good agreement with the experimental values. The range of errors is within 0.2 eV for all basis sets. As to the KT calculations, the order of IPs for  $(\eta^6-C_6H_6)Cr(CO)$ ,  $e < a_1 < e$ , is not conformable with that obtained from the KBPBEPBE method. The errors for IPs of the KTCAM-B3LYP, KTWB97XD, and  $KT^{LC\text{-wPBE}}$  methods are 0.26-0.38, 0.33-0.46, and 0.81-1.11 eV, respectively. They are larger than 0.08-0.13 eV of the KBPBEPBE method when the same basis sets are used. Consequently, the KB<sup>PBEPBE</sup> approach has better prediction in IPs than the KT methods. In the  $X\alpha$  method, the predicted values are close to those of the KBPBEPBE method but the orders of the HOMO-4 (a") and HOMO-5 (a') orbitals for  $(\eta^5-C_5H_5)$ -Mn(CO)<sub>3</sub> are not conformable with ours.

We then examine the characteristics of MOs using the KBPBEPBE method. The correlation diagrams of the frontier MOs of  $(\eta^6-C_6H_6)Cr(CO)_3$  and  $(\eta^5-C_5H_5)Mn(CO)_3$  are illustrated in Figure 1. Our symmetry in the labeling of the orbital is based on that the z axis is perpendicular to the arene ring. In Figure 1, the ordering of MOs for metals, benzene (Bz), Cp, (CO)<sub>3</sub>, Bz(CO)<sub>3</sub>, and Cp(CO)<sub>3</sub> are estimated using the  $(\eta^6$ -C<sub>6</sub>H<sub>6</sub>)Cr(CO)<sub>3</sub>/  $(\eta^5-C_5H_5)Mn(CO)_3$  structures by setting up the other atoms as ghost atoms<sup>37</sup> via KT. For  $(\eta^6-C_6H_6)Cr(CO)_3$ , the 3d orbitals of chromium yield  $a_1(d_{z^2})$ ,  $e(d_{x^2-y^2,xy})$ , and  $e(d_{xz,yz})$  orbitals in the  $C_{3v}$  point group. The  $\pi^*$  orbitals of (CO)<sub>3</sub> are of  $a_1$ , e, e, and  $a_2$ symmetries, and those of benzene are of  $a_1$  and e symmetries. For  $(\alpha^5-C_5H_5)Mn(CO)_3$ , the yz plane is defined as the plane of symmetry. The 3d orbitals of manganese yield a'  $(d_{r^2})$ , a'  $(d_{r^2-v^2})$ ,  $a'(d_{yz})$ ,  $a''(d_{xy})$ , and  $a''(d_{xz})$  orbitals in the  $C_S$  point group. The  $\pi^*$  orbitals of (CO)<sub>3</sub> are of a', a", a', a", a', and a" symmetries, and those of Cp are of a' and a" symmetries. For both metal carbonyls, 4s or 4p orbitals of metals are mainly core-excited resonances.41



**Figure 1.** Correlation diagram of the frontier MOs for  $(\eta^6-C_6H_6)Cr(CO)_3$  and  $(\eta^5-C_5H_5)Mn(CO)_3$ .

 $(\eta^6\text{-}\mathrm{C}_6\mathrm{H}_6)\mathrm{Cr}(\mathrm{CO})_3$  has 18 a<sub>1</sub>, 17 e, and 2 a<sub>2</sub> occupied MOs. Its frontier filled MOs using basis set A1 are illustrated in Figure 2. As indicated in Figure 2, the 18a<sub>1</sub> orbital is the HOMO. This orbital results from Cr  $(\mathrm{d}_{z^2})$  and  $(\mathrm{CO})_3$   $\pi^*$ . The 17e orbitals result from the mixing of the benzene and  $(\mathrm{CO})_3$   $\pi^*$  orbitals with Cr  $(\mathrm{d}_{x^2-y^2,xy})$ . The 16e orbitals are essentially from benzene e  $\pi$ 



**Figure 2.** Plots of the frontier filled MOs for  $(\eta^6-C_6H_6)Cr(CO)_3$ . The isosurface values are chosen to be 0.02 for all the MO plots.

orbitals. As for  $(\eta^5-C_5H_5)Mn(CO)_3$ , it has 33 a' and 18 a'' occupied MOs. The HOMO 33a' results from the mixing of Mn  $(d_{x^2-y^2})$  with  $(CO)_3 \pi^*$ . The 18a" and 32a' orbitals result from the mixing of Mn ( $d_{xy}$ ) and Mn ( $d_{z^2}$ ) with (CO)<sub>3</sub>  $\pi^*$  orbitals, respectively. The 17a" and 31a' orbitals are essentially from the mixing of Cp  $\pi$  orbitals with Mn ( $d_{xz}$ ) and Mn ( $d_{vz}$ ), respectively.

Table 2 lists the calculated IPs using the representative results of A1, B1, and C1 basis sets for  $(\eta^4-C_4H_6)Fe(CO)_3$  and  $(\eta^5 C_5H_5$ )Co(CO)<sub>2</sub> along with the experimental values. For  $(\eta^4$ -C<sub>4</sub>H<sub>6</sub>)Fe(CO)<sub>3</sub>, the increasing order of IPs of filled MOs is a' <  $a' \approx a'' < a' < a'' < a'$  for all methods. For  $(\eta^5 - C_5 H_5) Co(CO)_2$ , the increasing orders of IPs of filled MOs are a' < a' < a'' < a''< a' < a'' for KB and a' < a' < a'' < a'' < a' < a' for KT calculations, respectively. According to Table 2, the calculated IPs using the KBPBEPBE method are in good agreement with the experimental values. The range of errors as compared with experimental values is within 0.2 eV. For the KT calculations, the errors for IPs are 0.25-0.30, 0.30-0.51, and 1.03-1.27 eV for the KT<sup>CAM-B3LYP</sup>, KT<sup>wB97XD</sup>, and KT<sup>LC-wPBE</sup> methods, respectively. If the X $\alpha$  method is used for  $(\eta^5-C_5H_5)Co(CO)_2$ , the first IP is due to ionization from the a" orbital. It is clear that the KBPBEPBE approach so far has the best prediction in IPs. One of the main reasons why the order of KS PBEPBE orbital energies is likely to be correct is because the present approach contains exchange-correlation potential.<sup>42</sup>

Figure 3 shows the correlation diagrams of the frontier MOs of  $(\eta^4-C_4H_6)Fe(CO)_3$  and  $(\eta^5-C_5H_5)Co(CO)_2$ . Similar to  $(\eta^5-C_5H_5)Co(CO)_2$ .  $C_5H_5$ )Mn(CO)<sub>3</sub>, the 3d orbitals of metals yield a'  $(d_{r^2})$ , a'  $(d_{r^2-v^2})$ ,  $a'(d_{yz})$ ,  $a''(d_{xy})$ , and  $a''(d_{xz})$  orbitals. The  $\pi^*$  orbitals of (CO)<sub>3</sub> are of a', a', a" a', a", and a" symmetries, and those of (CO)<sub>2</sub> are of a', a', a", and a" symmetries. The  $\pi^*$  orbitals of  $(\eta^4$ C<sub>4</sub>H<sub>6</sub>) and Cp are both of a' and a" symmetries. There are 31 a' and 18 a" occupied MOs for ( $\eta^4$ -C<sub>4</sub>H<sub>6</sub>)Fe(CO)<sub>3</sub>, and 28 a' and 17 a" occupied MOs for  $(\eta^5-C_5H_5)Co(CO)_2$ . According to our analysis of the nature of orbitals for  $(\eta^4-C_4H_6)Fe(CO)_3$ , the HOMO (31a') results from Fe ( $d_{z^2}$ ) mixed with ( $\eta^4$ -C<sub>4</sub>H<sub>6</sub>)(CO)<sub>3</sub>  $\pi^*$ . The 30a' orbital is essentially from Fe  $(d_{r^2-v^2})$  and  $(\eta^4$  $C_4H_6$ (CO)<sub>3</sub>  $\pi^*$ . MO 18a" results from Fe (d<sub>xv</sub>) mixed with ( $\eta^4$ - $C_4H_6$ (CO)<sub>3</sub>  $\pi^*$ , and MO 29a' results from Fe ( $d_{yz}$ ) mixed with  $(\eta^4\text{-C}_4\text{H}_6)(\text{CO})_3 \pi^*$ . The 17a" orbital is mainly from  $(\eta^4\text{-C}_4\text{H}_6)$  $\pi$  and Fe (d<sub>xz</sub>). MO 28a' is mainly from ( $\eta^4$ -C<sub>4</sub>H<sub>6</sub>)  $\pi$ . As for  $(\eta^5-C_5H_5)C_0(CO)_2$ , the 28a' orbital results from  $(CO)_2$   $\pi^*$  mixed with Co ( $d_{yz}$ ). The 27a' orbital is essentially from the Co ( $d_{x^2-y^2}$ ) orbital. The 17a" and 26a' orbitals are mainly from Co  $(d_{xz})$ and Co ( $d_{z^2}$ ) mixed with (CO)<sub>2</sub>  $\pi^*$  orbitals, respectively. The 25a' orbital corresponds to the  $\pi$ -bonding interaction between Cp  $\pi$  and Co (d<sub>vz</sub>) orbitals, and 16a" corresponds to the  $\pi$ -bonding interaction between Cp  $\pi$  and Co ( $d_{xz}$ ) orbitals. The 24a' orbital is mainly from Cp  $\pi$ .

For the temporary anion shape resonance, we perform S-KB calculations on the unfilled orbitals to distinguish them from the ODC solutions for  $\alpha$  between 0.0 and 3.0 via the PBEPBE (S-KBPBEPBE), and then the S-KT calculations via CAM-B3LYP, wB97XD, and LC-wPBE (S-KTCAM-B3LYP, S-KTWB97XD, and  $S\text{-}KT^{\text{LC-wPBE}})$  methods. First, we will present the results of S-KB<sup>PBEPBE</sup> for  $(\eta^6$ -C<sub>6</sub>H<sub>6</sub>)Cr(CO)<sub>3</sub> and  $(\eta^5$ -C<sub>5</sub>H<sub>5</sub>)Mn(CO)<sub>3</sub>. Figure 4a shows the energies of the discretized continuum  $(DC)^{19-23,43}$  solutions as a function of the scale factor  $\alpha$  for the e,  $a_1$ , and  $a_2$  virtual orbitals of  $(\eta^6-C_6H_6)Cr(CO)_3$  using the basis set A1 located at the appropriate nuclear positions. The energies of the DC solutions are obtained by solving the Kohn-Sham equation for a single electron in the absence of any potential. The stabilization graphs of the energies as a function of  $\alpha$  for the e,  $a_1$ , and  $a_2$  virtual states of  $(\eta^6-C_6H_6)Cr(CO)_3$  using basis set A1 for the S-KBPBEPBE calculations are shown in Figure 4b.

TABLE 2: Calculated IPs (eV) for  $(\eta^4-C_4H_6)$ Fe(CO)<sub>3</sub> and  $(\eta^5-C_5H_5)$ Co(CO)<sub>2</sub>

			$(\eta^4$ -C <sub>4</sub> H <sub>6</sub> )Fe(CO) <sub>3</sub>							$(\eta^5\text{-C}_5\text{H}_5)\text{Co}(\text{CO})_2$						
method	basis set	a'	a′	a"	a′	a"	a′	d <sup>b</sup> /eV	a'	a'	a"	a′	a′	a"	d <sup>b</sup> /eV	
KBPBEPBE	A1	8.13	8.58	8.66	8.84	9.73	11.04	0.24	7.54	7.66	8.41	9.37	9.87	9.97	0.12	
	B1	8.28	8.75	8.83	9.00	9.83	11.09	0.13	7.63	7.74	8.52	9.48	9.91	10.00	0.10	
	C1	8.26	8.73	8.81	8.99	9.80	11.06	0.14	7.65	7.76	8.53	9.48	9.91	9.99	0.10	
$KT^{CAM-B3LYP}$	A1	7.77	8.61	8.76	9.15	9.39	11.05	0.30	7.10	8.15	8.73	9.64	9.78	9.28	0.29	
	B1	7.95	8.83	8.98	9.38	9.54	11.16	0.25	7.25	8.28	8.90	9.76	9.95	9.39	0.25	
	C1	7.94	8.83	8.98	9.38	9.53	11.13	0.26	7.27	8.29	8.92	9.77	9.95	9.38	0.25	
$KT^{wB97XD}$	A1	8.36	9.18	9.34	9.72	9.98	11.63	0.30	7.66	8.69	9.28	10.22	10.35	9.88	0.42	
	B1	8.48	9.34	9.49	9.89	10.07	11.68	0.42	7.77	8.77	9.40	10.30	10.46	9.94	0.51	
	C1	8.45	9.32	9.48	9.87	10.05	11.64	0.40	7.78	8.77	9.41	10.30	10.44	9.93	0.51	
$KT^{LC\text{-wPBE}}$	A1	9.07	9.86	10.01	10.32	10.76	12.55	1.03	8.35	9.44	10.01	10.98	11.06	10.69	1.16	
	B1	9.21	10.05	10.19	10.51	10.86	12.61	1.17	8.47	9.52	10.15	11.12	11.15	10.76	1.27	
	C1	9.18	10.03	10.18	10.50	10.83	12.55	1.14	8.48	9.52	10.15	11.10	11.13	10.73	1.26	
$X\alpha^a$		7.7	8.4	8.5	8.8	9.8	11.7	0.31	$7.7^{c}$	8.5	9.5	8.9	9.8	10.2	0.21	
$expt^a$		8.23	8.82	9.09	9.93	11.52			7.59	7.95	8.51	9.41	9.87	10.23		

<sup>&</sup>lt;sup>a</sup> The IPs are obtained from previous studies. <sup>28</sup> <sup>b</sup> Here d denotes the mean error relative to experimental IP data. <sup>c</sup> Value for a".

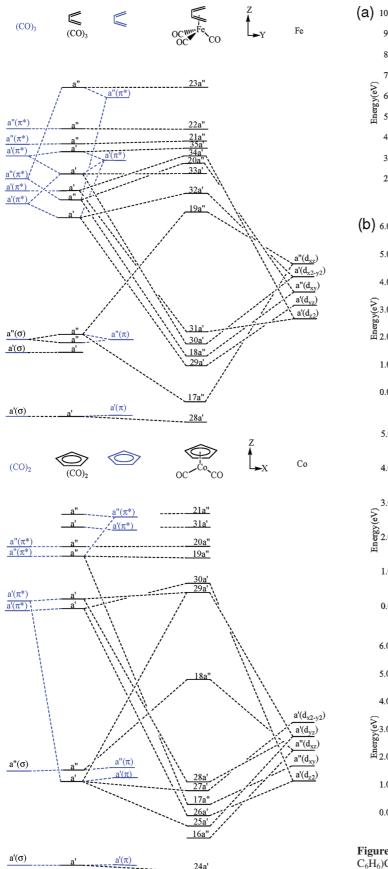
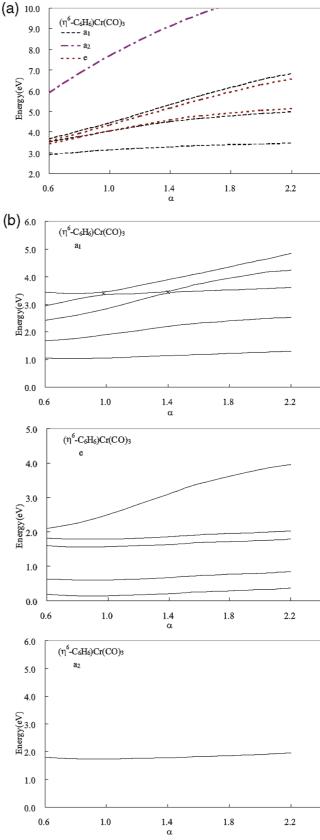


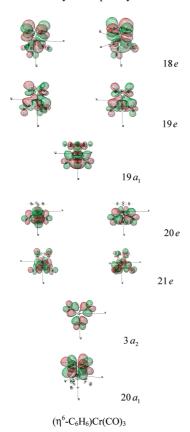
Figure 3. Correlation diagram of the frontier MOs for  $(\eta^4\text{-}C_4H_6)Fe(CO)_3$  and  $(\eta^5\text{-}C_5H_5)Co(CO)_2.$ 

There are two types of energies for virtual orbital solutions in the S-KB calculations. One is the unfilled orbital solution



**Figure 4.** (a) Energies of e,  $a_1$ , and  $a_2$  virtual orbitals of  $(\eta^6-C_6H_6)Cr(CO)_3$  as a function of the scaling factor α for a free electron in the absence of potentials. (b) Stabilization graphs for  $(\eta^6-C_6H_6)Cr(CO)_3$  via S-KB<sup>PBEPBE</sup> method. Energies of e,  $a_1$ , and  $a_2$  virtual orbitals as a function of α. The location of  $\alpha_{ac}$  is marked with  $\times$ .

and the other is the ODC virtual orbital solution. The unfilled orbital solution and the ODC solutions are readily distinguished

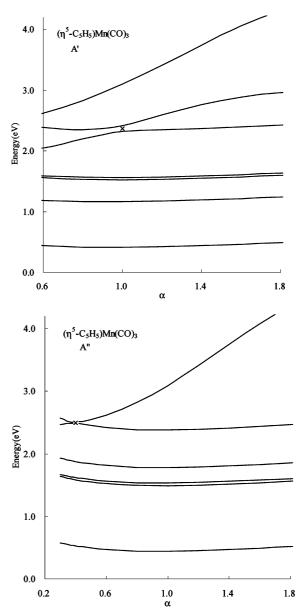


**Figure 5.** Plots of the frontier unfilled MOs at  $\alpha = 2.0$  for  $(\eta^6\text{-C}_6\text{H}_6)\text{Cr}(\text{CO})_3$ . The isosurface values are chosen to be 0.02 for all the MO plots.

by examining how their energies vary with a. As shown in Figure 4b for the a<sub>1</sub> virtual orbitals, the first solution which remains stabilized with  $\alpha$  is the  $19a_1$  orbital solution and the stabilized energy value is 1.14 eV.44 The second solution corresponds to the first ODC solution. The avoided crossings at  $\alpha_{ac} = 1.0$  and 1.4 are due to the coupling between the  $20a_1$ orbital solution with the second and third ODC solutions. The energies of the 20a<sub>1</sub> orbital can be extracted from each twostate avoided crossing region. From Figure 4b, the energy values of the  $20a_1$  orbital are both 3.40 eV at  $\alpha_{ac}=1.0$  and 1.4. Thus, the energy of the 20a<sub>1</sub> orbital is 3.40 eV. As for the e virtual orbitals, the first to fourth solutions which remain stabilized with  $\alpha$  are the 18e, 19e, 20e, and 21e orbital solutions and their stabilized energy values are 0.19, 0.65, 1.61, and 1.84 eV, respectively. The next four solutions correspond to the ODC solutions. For the a<sub>2</sub> virtual orbital, the first solution is the 3a<sub>2</sub> orbital solution and the stabilized energy value is 1.79 eV.

In Figure 5, the first and third a<sub>1</sub> virtual orbitals, the first to fourth e virtual orbitals, and the first a2 virtual orbital that correspond to the resonance solutions of  $(\eta^6-C_6H_6)Cr(CO)_3$  for  $\alpha = 2.0$  are displayed. As can be seen in Figures 1 and 5, the 18e orbitals essentially result from Bz(CO)<sub>3</sub>  $\pi^*$  orbitals. The 19e orbitals result from the  $\pi$ -antibonding interaction between Bz(CO)<sub>3</sub>  $\pi$  orbitals and Cr (d<sub>xz</sub>, d<sub>yz</sub>). The 20e orbitals are essentially derived from the Bz(CO)<sub>3</sub>  $\pi^*$  orbitals. The 19a<sub>1</sub> orbital results from the antibonding interaction between Bz(CO)<sub>3</sub>  $\pi^*$  orbitals and Cr (d<sub>7</sub>2). The 21e orbitals result from the antibonding interaction between Bz(CO)<sub>3</sub>  $\pi^*$  orbitals and Cr  $(d_{x^2-y^2,xy})$ . The  $3a_2$  and 22e orbitals are mainly from the Bz(CO)<sub>3</sub>  $\pi^*$  orbitals. The 20a<sub>1</sub> orbital is essentially from the benzene  $\pi^*$ 

Next, for  $(\eta^5-C_5H_5)Mn(CO)_3$ , the stabilization graphs of the energies as a function of  $\alpha$  for the a' and a" virtual states using



**Figure 6.** Stabilization graphs for  $(\eta^5-C_5H_5)Mn(CO)_3$  via S-KB<sup>PBEPBE</sup> method. Energies of a' and a" virtual orbitals as a function of  $\alpha$ . The location of  $\alpha_{ac}$  is marked with  $\times$ .

basis set A1 are shown in Figure 6. As shown in Figure 6, the first to fourth solutions which remain stabilized with  $\alpha$  are the 34a', 35a', 36a', and 37a' orbital solutions and their stabilized energy values are 0.44, 1.19, 1.55, and 1.58 eV, respectively. The avoided crossing at  $\alpha_{ac} = 1.0$  is due to the coupling between the fifth a' (38a') unfilled orbital solution and first ODC solution. In Figure 6, the energy value of 38a' orbital is 2.37 eV at  $\alpha_{ac}$ = 1.0. The seventh solution corresponds to the second ODC solution. For the a" virtual states, the first to fourth solutions which remain stabilized with  $\alpha$  are the 19a", 20a", 21a", and 22a" orbital solutions and their stabilized energy values are 0.47, 1.51, 1.56, and 1.81 eV, respectively. The avoided crossing at  $\alpha_{ac} = 0.4$  is due to the coupling between the 23a" orbital solution and the first ODC solution. The energy value of the 23a" orbital is 2.50 eV at  $\alpha_{ac} = 0.4$ . According to Figures 1 and 6, the LUMO 34a' corresponds to the antibonding interaction between  $Cp(CO)_3 \pi^*$  orbitals and Mn ( $d_{vz}$ ). The 19a" orbital corresponds to the antibonding interaction between  $Cp(CO)_3 \pi^*$  orbitals and Mn ( $d_{xz}$ ). The 35a' orbital is essentially from (CO)<sub>3</sub>  $\pi^*$  orbitals and Mn (d<sub>7</sub><sup>2</sup>). The 20–22a" orbitals are essentially from (CO)<sub>3</sub>

TABLE 3: Calculated AEs<sup>a</sup> (eV) for  $(\eta^6-C_6H_6)Cr(CO)_3$  and  $(\eta^5-C_5H_5)Mn(CO)_3$  via S-KB<sup>PBEPBE</sup> Method

					basis set							
	A1	A2	A3	B1	B2	B4	C1	C2	C4	2	$X\alpha^b$	$expt^b$
					(η	6-C <sub>6</sub> H <sub>6</sub> )Cr(0	CO) <sub>3</sub>					
$a_1$	3.40	3.51	3.52	3.52	3.51	3.58	3.55	3.56	3.54	$a_2$	3.2	3.63
e	1.84	1.84	1.93	2.10	2.06	2.09	2.08	2.08	2.08	$a_2$	1.9	2.4
$a_2$	1.79	1.88	1.86	2.12	2.12	2.09	2.08	2.08	2.08	e	1.8	2.4
e	1.61	1.63	1.68	1.92	1.89	1.89	1.90	1.89	1.90	e	1.7	1.82
$a_1$	1.14	1.14	1.21	1.28	1.25	1.26	1.27	1.27	1.27	$a_1$	1.2	0.96
e	0.65	0.66	0.71	0.89	0.83	0.86	0.88	0.86	0.88	e	1.0	0.96
e	0.19	0.25	0.26	0.39	0.38	0.36	0.36	0.37	0.35	$a_1$	0.9	c
										e	0.1	
										e	-0.1	
					$(\eta^5$	$^{5}$ - $C_{5}H_{5}$ )Mn( $^{6}$	CO) <sub>3</sub>					
a"	2.50	2.46	2.26	2.38	2.40	2.37	2.34	2.34	2.33	a"	2.0	2.36
a'	2.37	2.38	2.16	2.33	2.35	2.32	2.29	2.29	2.28	a"	1.7	2.36
a"	1.81	1.83	1.64	1.94	1.94	1.92	1.92	1.91	1.90	a"	1.6	1.8
a'	1.58	1.61	1.43	1.69	1.68	1.67	1.69	1.68	1.67	a'	1.6	1.8
a"	1.56	1.57	1.40	1.67	1.66	1.65	1.67	1.65	1.65	a'	1.5	1.8
a'	1.55	1.57	1.39	1.66	1.66	1.63	1.65	1.64	1.63	a'	1.1	1.8
a"	1.51	1.56	1.36	1.62	1.63	1.59	1.62	1.61	1.59	a'	1.0	1.8
a'	1.19	1.20	1.03	1.26	1.26	1.25	1.27	1.26	1.26	a"	1.0	1.22
a"	0.47	0.48	0.29	0.49	0.46	0.47	0.49	0.48	0.48	a"	0.8	0.2
a'	0.44	0.45	0.27	0.46	0.44	0.44	0.47	0.45	0.45	a'	0.8	0.2
										a'	0.4	
										a"	0.2	
										a'	0.1	

<sup>&</sup>lt;sup>a</sup> The energies of the HOMO ( $\varepsilon_{HOMO}$ ) in eq 4 are calculated for each value of α even though the variations of  $\varepsilon_{HOMO}$  values are within 0.1 eV. <sup>b</sup> The AEs are obtained from previous studies. <sup>28 c</sup> The calculated energy value is outside the observed range.

TABLE 4: Calculated AEs<sup>a</sup> (eV) for ( $\eta^4$ -C<sub>4</sub>H<sub>6</sub>)Fe(CO)<sub>3</sub> and ( $\eta^5$ -C<sub>5</sub>H<sub>5</sub>)Co(CO)<sub>2</sub> via S-KB<sup>PBEPBE</sup> Method

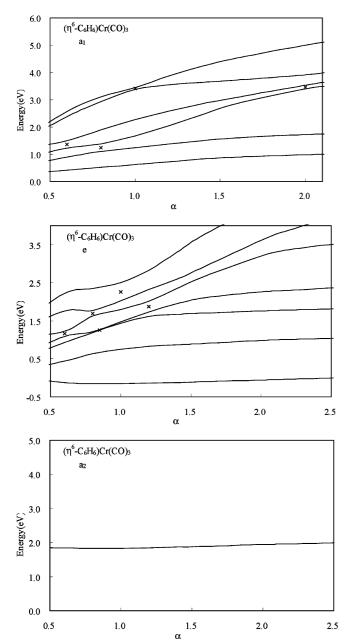
					basis set <sup>a</sup>							
	A1	A2	A3	B1	B2	B4	C1	C2	C4	2	$X\alpha^b$	expt <sup>l</sup>
					$(\eta^4$	-C <sub>4</sub> H <sub>6</sub> )Fe(C0	D) <sub>3</sub>					
a"	2.24	2.27	2.21	2.17	2.20	2.15	2.12	2.11	2.09	a"	2.0	2.21
a"	1.48	1.50	1.44	1.62	1.66	1.60	1.59	1.58	1.55	a"	1.7	1.26
a"	1.27	1.30	1.24	1.39	1.43	1.36	1.38	1.36	1.34	a'	1.3	1.26
a'	1.26	1.27	1.22	1.37	1.41	1.35	1.35	1.32	1.30	a"	1.2	1.26
a"	1.18	1.17	1.15	1.28	1.29	1.27	1.27	1.22	1.23	a'	1.1	1.26
a′	1.20	1.19	1.17	1.26	1.26	1.24	1.25	1.21	1.22	a'	1.0	1.26
a′	0.85	0.84	0.81	0.92	0.93	0.90	0.91	0.87	0.87	a"	1.0	1.26
a'	0.67	0.66	0.64	0.66	0.65	0.64	0.64	0.59	0.60	a"	0.8	c
a"	0.34	0.31	0.29	0.31	0.27	0.29	0.30	0.23	0.27	a'	0.1	c
										a"	0.0	
										a'	0.0	
										a'	-0.2	
					$(\eta^5)$	-C <sub>5</sub> H <sub>5</sub> )Co(Co	$O)_2$					
a"	2.62	2.63	2.61	2.49	2.51	2.48	2.45	2.45	2.39	a"	1.7	2.05
a′	2.58	2.58	2.64	2.50	2.52	2.48	2.45	2.45	2.40	a'	1.6	2.05
a"	1.78	1.72	1.76	1.87	1.85	1.84	1.87	1.83	1.81	a'	1.5	2.05
a"	1.72	1.64	1.70	1.79	1.75	1.77	1.79	1.74	1.73	a'	1.0	2.05
a′	1.02	0.93	0.99	1.06	1.02	1.04	1.06	1.02	1.01	a"	1.0	0.88
a′	0.97	0.89	0.95	0.99	0.95	0.96	0.98	0.94	0.92	a''	0.9	0.88
a"	0.00	-0.08	-0.04	-0.05	-0.12	-0.08	-0.05	-0.11	-0.10	a'	0.9	c
										a'	0.8	
										a'	0.0	

<sup>&</sup>lt;sup>a</sup> The energies of the HOMO ( $\varepsilon_{\text{HOMO}}$ ) in eq 4 are calculated for each value of α even though the variations of  $\varepsilon_{\text{HOMO}}$  values are within 0.1 eV. <sup>b</sup> The AEs are obtained from previous studies. <sup>28 c</sup> The calculated energy value is outside the observed range.

 $\pi^*$ . The 36a' and 37a' orbitals are mainly from the mixing of Mn (d<sub>xy</sub>) and (d<sub>x²-y²</sub>) with (CO)<sub>3</sub>  $\pi^*$  orbitals, respectively. The 38a' and 23a" unfilled orbitals are derived from Cp  $\pi^*$  orbitals.

Table 3 illustrates the results of AEs using the S-KB methods for  $(\eta^6\text{-}C_6H_6)\text{Cr}(\text{CO})_3$  and  $(\eta^5\text{-}C_5H_5)\text{Mn}(\text{CO})_3$  along with the experimental values. The increasing orders of AEs of unfilled MOs are  $e < e < a_1 < e < a_2 \approx e < a_1$  for  $(\eta^6\text{-}C_6H_6)\text{Cr}(\text{CO})_3$  and  $a' \approx a'' < a'' < a'' \approx a' \approx a'' \approx a' < a'' <$ 

 $a' \approx a''$  for  $(\eta^5 - C_5 H_5) Mn(CO)_3$ . Notice that, if the  $X\alpha$  method is used, the order of AEs is not conformable with ours. The additional anion states found in the  $X\alpha$  calculations may represent ODC solutions. As can be seen from Table 3, the calculated AEs using the S-KB<sup>PBEPBE</sup> methods are better than those using the  $X\alpha$  method. One possible reason for the  $X\alpha$  assignments to be worse is that it does not treat the  $\sigma$  and  $\pi$  orbitals on an equal footing when using the muffin-tin



**Figure 7.** Stabilization graphs for  $(\eta^6\text{-C}_6\text{H}_6)\text{Cr}(\text{CO})_3$  via S-KT<sup>CAM-B3LYP</sup> method. Energies of e, a<sub>1</sub>, and a<sub>2</sub> virtual orbitals as a function of α. The location of  $\alpha_{ac}$  is marked with  $\times$ .

approximation (MTA). The AEs obtained are also quite sensitive to the sphere overlap used for MTA.

Our assignments for the shape resonance observed in the ET spectrum are slightly different from those of the  $X\alpha$  method. According to our S-KB calculations for  $(\eta^6-C_6H_6)Cr(CO)_3$ , the 3.63 eV feature in the ET spectrum is ascribed to electron capture into the empty 20a<sub>1</sub> orbital. The resonances at 2.4 and 1.82 eV are associated with the capture into the 21e, 3a<sub>2</sub>, and 20e orbitals. The resonance at 0.96 eV is associated with the capture into the 19e and 19a<sub>1</sub> orbitals. The electron addition to 18e orbitals may not be observed in the ET spectrum. As for  $(\eta^5-C_5H_5)Mn(CO)_3$ , the electron capture into the 38a' and 23a" orbitals corresponds to the resonance at 2.36 eV. The 1.8 eV feature can be ascribed to the electron capture into the 36a', 37a', and 20a"-22a" orbitals. The capture into the 35a' orbital can be associated with the resonance at 1.22 eV. Finally, the electron addition to the 34a' and 19a" MOs can be associated with the resonance at 0.2 eV.

Table 4 lists the calculated AEs via the S-KBPBEPBE method for  $(\eta^5-C_4H_6)Fe(CO)_3$  and  $(\eta^4-C_5H_5)Co(CO)_2$ . As can be seen for  $(\eta^5-C_4H_6)$ Fe(CO), the increasing order of AEs of unfilled orbitals is  $a'' < a' < a' < a' \approx a'' < a' \approx a'' < a'' < a''$ . For  $(\eta^5-C_5H_5)Co(CO)_2$ , the increasing order of AEs of unfilled orbitals is  $a'' < a' \approx a' < a'' \approx a'' < a' \approx a''$ . According to Tables 3 and 4, the calculated AEs are in good agreement with the experimental data. Both the inherent experimental errors for the ETS structures and the errors associated with determination of the resonance energies from the stabilization graphs could be as large as 0.1 eV. 45,46 Consequently, the S-KBPBEPBE calculations are able to yield accurate energies of temporary anion states of transition-metal carbonyls.

According to analysis of the nature of orbitals for  $(\eta^4$ C<sub>4</sub>H<sub>6</sub>)Fe(CO)<sub>3</sub>, the LUMO (19a") corresponds to the antibonding interaction between  $(\eta^4-C_4H_6)$   $\pi$  orbitals and Fe  $(d_{xz})$ . The 20a" orbital is mainly from  $(\eta^4\text{-C}_4\text{H}_6)(\text{CO})_3$  a"  $\pi^*$  orbitals and Fe  $(d_{xy})$ . The 21a" and 22a" orbitals are essentially from (CO)<sub>3</sub> a"  $\pi^*$ . The 23a" orbital is derived from the  $(\eta^4$ -C<sub>4</sub>H<sub>6</sub>)  $\pi^*$  orbital. Orbital 32a' is mainly from  $(\eta^4-C_4H_6)(CO)_3$  a'  $\pi^*$  orbitals and Fe  $(d_{x^2-y^2})$ . The 33a' and 35a' orbitals are essentially from  $(\eta^4$  $C_4H_6)(CO)_3$  a'  $\pi^*$  orbitals. The 34a' orbital is mainly from ( $\eta^4$ - $C_4H_6)(CO)_3$  a'  $\pi^*$  orbitals and Fe  $(d_{z^2})$ . For  $(\eta^5-C_5H_5)Co(CO)_2$ , the 18a" orbital corresponds to the antibonding interaction between  $(\eta^5-C_5H_5)$   $\pi$  orbitals and Co  $(d_{rz})$ . The 29a' orbital corresponds to antibonding interaction between ( $\eta^5$ -C<sub>5</sub>H<sub>5</sub>)  $\pi$ orbitals and Co ( $d_{yz}$ ). Orbital 30a' is mainly from (CO)<sub>2</sub> a'  $\pi^*$ orbitals and Co (d<sub>z²</sub>). The 19a" and 20a" unfilled orbitals are essentially from (CO)<sub>2</sub> a"  $\pi$ \*. The 31a' and 21a" orbitals are derived from Cp  $\pi^*$  orbitals.

For the S-KT calculations, the stabilization graphs of the energies as a function of  $\alpha$  for the e, a<sub>1</sub>, and a<sub>2</sub> virtual states of  $(\eta^6\text{-C}_6\text{H}_6)\text{Cr}(\text{CO})_3$  via the S-KT<sup>CAM-B3LYP</sup> method using the A1 basis set are shown in Figure 7. The obtained energy values of the  $19a_1$  orbital are 1.36~eV at  $\alpha_{ac}=0.6$  and 1.25~eV at  $\alpha_{ac}=$ 0.8. The energy values of the  $20a_1$  orbital are 3.42 eV at  $\alpha_{ac}$  = 1.0 and 3.47 eV at  $\alpha_{ac} = 2.0$ . The lowest value from each set will be defined as the energy of the temporary anion state.<sup>26</sup> Thus, the energies of the 19a<sub>1</sub> and 20a<sub>1</sub> orbitals are 1.25 and

TABLE 5: Calculated AEs (eV) for  $(\eta^6-C_6H_6)Cr(CO)_3$  via S-KT Methods for Basis Sets A1, B1, and C1

		S-KT <sup>CAM-B3LYP</sup>		S-KT <sup>wB97XD</sup>						
	A1	B1	C1	A1	B1	C1	A1	B1	C1	$expt^a$
a <sub>1</sub>	3.42	3.59	3.55	3.94	4.00	3.90	4.56	4.35	4.57	3.63
e	2.26	2.56	2.71	3.01	3.15	3.16	3.76	3.78	3.76	2.4
$a_2$	1.94	2.05	2.02	2.52	2.58	2.62	3.17	3.15	3.19	1.82
e	1.69	1.89	1.95	2.36	2.57	2.54	3.03	3.16	3.22	
$a_1$	1.25	1.19	1.21	1.98	2.00	2.00	2.48	2.44	2.61	0.96
e	1.16	1.10	1.19	1.64	1.81	1.77	2.23	2.40	2.38	
e	-0.06	-0.07	-0.10	0.54	0.58	0.57	0.97	0.88	0.92	

<sup>&</sup>lt;sup>a</sup> The AEs are obtained from previous studies.<sup>28</sup>

TABLE 6: Calculated AEs (eV) for  $(\eta^5-C_5H_5)Mn(CO)_3$ ,  $(\eta^4-C_4H_6)$ Fe(CO)<sub>3</sub>, and  $(\eta^5-C_5H_5)$ Co(CO)<sub>2</sub> via S-KT Methods

		method		
	S-KT <sup>CAM-B3LYP</sup>	S-KT <sup>wB97XD</sup>	S-KT <sup>LC-wPBE</sup>	$expt^a$
	(	$(\eta^5-C_5H_5)Mn(CO)$	)3	
a"	2.09	2.73	3.32	2.36
a'	2.07	2.77	3.22	
a''	1.41	2.01	2.53	1.8
a'	1.30	1.82	2.46	1.22
a"	1.29	1.89	2.40	
a'	1.06	1.69	2.31	
a''	1.04	1.69	2.30	
a'	0.54	1.14	1.81	0.2
a"	0.28	0.95	1.41	
a'	0.26	0.94	1.29	
		$(\eta^4$ -C <sub>4</sub> H <sub>6</sub> )Fe(CO)	3	
a"	2.00	2.68	3.21	2.21
a''	1.12	1.77	2.33	1.26
a''	1.06	1.73	2.11	
a'	0.89	1.62	2.18	
a"	0.66	1.24	1.96	
a'	0.81	1.50	2.08	
a'	0.29	0.96	1.57	
a'	0.21	0.90	1.42	
a"	0.34	1.04	1.48	
		$(\eta^5-C_5H_5)Co(CO)$	)2	
a"	2.19	2.80	3.38	2.05
a'	2.13	2.74	3.35	
a''	1.39	2.06	2.54	
a"	0.99	1.68	2.07	0.88
a'	0.31	0.98	1.60	
a'	0.16	0.76	1.32	
a"	-0.07	0.64	0.93	

<sup>&</sup>lt;sup>a</sup> The AEs are obtained from previous studies. <sup>28</sup>

3.42 eV, respectively. As for the e virtual orbitals, the stabilized energy value for the 18e orbital is -0.06 eV. The energy values of the 19e orbitals are 1.16 eV at  $\alpha_{ac} = 0.6$  and 1.27 eV at  $\alpha_{ac}$ = 0.9, respectively. The energy values of the 20e orbitals are 1.69 eV at  $\alpha_{ac} = 0.8$  and 1.87 eV at  $\alpha_{ac} = 1.2$ , respectively. Accordingly, the energies of the 19e and 20e anion states are 1.16 and 1.69 eV, respectively. The energy value for the 21e orbitals is 2.26 eV at  $\alpha_{ac} = 1.0$ . For the  $a_2$  virtual orbitals, the stabilized energy value for the 3a<sub>2</sub> virtual orbital is 1.94 eV.

The calculated AEs of (η<sup>6</sup>-C<sub>6</sub>H<sub>6</sub>)Cr(CO)<sub>3</sub> for various S-KT via CAM-B3LYP, wB97XD, and LC-wPBE methods are summarized in Table 5. The S-KT<sup>LC-wPBE</sup> and S-KT<sup>wB97XD</sup> approaches overestimate AEs when compared with the experimental values. The AEs obtained from the S-KT<sup>LC-wPBE</sup> and S-KTwB97XD methods are larger than those obtained from the S-KT<sup>CAM-B3LYP</sup> method (about 0.4 and 0.9 eV). Table 6 illustrates the calculated AEs of S-KT<sup>CAM-B3LYP</sup>, S-KT<sup>wB97XD</sup>, and S-KT<sup>LC-wPBE</sup> approaches using the representative A1 basis set for  $(\eta^5-C_5H_5)Mn(CO)_3$ ,  $(\eta^4-C_4H_6)Fe(CO)_3$ , and  $(\eta^5-C_5H_5)Co$ (CO)2. In Table 6, the calculated AEs using the S-KT<sup>CAM-B3LYP</sup> method are in good agreement with the experimental values for three carbonyls. Nevertheless, the S-KT<sup>LC-wPBE</sup> and S-KT<sup>wB97XD</sup> approaches overestimate AEs when compared with the experimental values. The AEs obtained from the S-KTwB97XD and S-KT<sup>LC-wPBE</sup> methods are larger than those obtained from the S-KT<sup>CAM-B3LYP</sup> method (about 0.7 and 1.2 eV). Possible reasons for the discrepancy are due to the different considerations of the exchange-correlation potential, self-interaction effect, and Coulomb contributions at large electron-molecule distance among these methods. 47,48 For instance, the LC-wPBE functional may be overcorrected by inclusion of a too large fraction of HF exchange, and the wB97XD functional may suffer from some self-interaction at short range. On the other hand, in the method of CAM-B3LYP, two extra parameters are used in the range separated Coulomb operator. To sum up, as compared with experimental values in Tables 3-6, the S-KTCAM-B3LYP and S-KBPBEPBE methods generally yield better AEs than those of the S-KTwB97XD and S-KBLC-wPBE methods for the transition-metal carbonyls.

#### 4. Conclusion

The energies of filled and unfilled orbitals in various transition-metal carbonyls have been systematically studied by various approaches. The present investigation has demonstrated that the KBPBEPBE calculations can yield very accurate results for IPs. In addition, the S-KBPBEPBE and S-KTCAM-B3LYP methods can yield good energy results for temporary anion states. Hence, it is believed that these two methods can be very useful in the studies of temporary anion states for transition-metal complexes.

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# References and Notes

- (1) (a) Muetterties, E. L.; Bleeke, J. R.; Wucherer, E. J.; Albright, T. A. Chem. Rev. 1982, 82, 499-525. (b) Albright, T. A. Acc. Chem. Res. 1982, 15, 149-155.
- (2) Merlic, C. A.; Walsh, J. C.; Tantillo, D. J.; Houk, K. N. J. Am. Chem. Soc. 1999, 121, 3596.
  - (3) Seyferth, D. Organometallics 2003, 22 (1), 2.
- (4) Frazier, C. C.; Harvey, M. A.; Cockeriam, M. P.; Hand, H. M.;
- Chauchard, E. A.; Lee, C. H. J. Phys. Chem. 1986, 90, 5703.(5) Rabalais, J. W. Principles of Ultraviolet Photoelectron Spectroscopy; John Wiley and Sons: New York, 1977.
  - (6) Sanche, L.; Schulz, G. J. Phys. Rev. A 1972, 5, 1672.
  - (7) Jordan, K. D.; Burrow, P. D. Chem. Rev. 1987, 87, 557.
- (8) Gengeliczki, Z.; Pongor, C. I.; Sztáray, B. Organometallics 2006, 25, 2553.
  - (9) Koopmans, T. Physica 1934, 1, 104.
  - (10) Kohn, W.; Sham, L. J. Phys. Rev. A 1965, 140, 1133.
  - (11) Simons, J. J. Phys. Chem. A 2008, 112, 6401.
  - (12) Tozer, D. J.; De Proft, F. J. Phys. Chem. A 2005, 109, 8923.
  - (13) Tozer, D. J.; De Proft, F. J. Chem. Phys. 2007, 127, 034108.
- (14) De Proft, F.; Sablon, N.; Tozer, D. J.; Geerlings, P. Faraday Discuss. 2007, 135, 151.
- (15) Sablon, N.; De Proft, F.; Geerlingsa, P.; Tozer, D. J. Phys. Chem. Chem. Phys. 2007, 9, 5880.
- (16) Teale, A. M.; De Proft, F.; Tozer, D. J. J. Chem. Phys. 2008, 129, 044110.
- (17) Hajgató, B.; Deleuze, M. S.; Tozer, D. J.; De Proft, F. J. Chem. Phys. 2008, 129, 084308.
  - (18) Falcetta, M. F.; Jordan, K. D. J. Phys. Chem. 1990, 94, 5666.
  - (19) Falcetta, M. F.; Jordan, K. D. J. Am. Chem. Soc. 1991, 113, 2903.
- (20) Burrow, P. D.; Howard, A. E.; Johnston, A. R.; Jordan, K. D. J. Phys. Chem. 1992, 96, 7570.
  - (21) Juang, C.-Y.; Chao, J. S.-Y. J. Phys. Chem. 1994, 98, 13506.
  - (22) Hazi, A. U.; Taylor, H. S. Phys. Rev. A 1970, 1, 1109.
  - (23) Taylor, H. S. Adv. Chem. Phys. 1970, 18, 91.
  - (24) Fels, M. F.; Hazi, A. U. Phys. Rev. A 1972, 5, 1236.
  - (25) Taylor, H. S.; Hazi, A. U. Phys. Rev. A 1976, 14, 2071.
- (26) Cheng, H.-Y.; Shih, C.-C. J. Phys. Chem. A 2009, 113, 1548. (27) Cheng, H.-Y.; Shih, C.-C.; Chang, J.-T. J. Phys. Chem. A 2009,
- (28) Modelli, A.; Distefano, G.; Guerra, M.; Jones, D. J. Am. Chem. Soc. 1987, 109, 4440.
  - (29) Wei, Y.-H.; Cheng, H.-Y. J. Phys. Chem. A 1998, 102, 3560.
  - (30) Wachters, A. J. H. J. Chem. Phys. 1970, 52, 1033.
- (31) Rappe, A. K.; Smedley, T. A.; Goddared, W. A., III. J. Phys. Chem. **1981**, 85, 2607.
- (32) We use the simplest "midpoint method" adopted by Burrow et al. to extract resonance energy (ref 20).

- (33) Perdew, J. P.; Burke, K.; Ernzerhof, M. Phys. Rev. Lett. 1996, 77, 3865.
- (34) Yanai, T.; Tew, D. P.; Handy, N. C. Chem. Phys. Lett. **2004**, 393, 51
- (35) Chai, J.-D.; Head-Gordon, M. Phys. Chem. Chem. Phys. 2008, 10, 6615.
- (36) Vydrov, O. A.; Scuseria, G. E.; Perdew, J. P. J. Chem. Phys. 2007, 126, 154109.
- (37) Frisch, M. J.; Trucks, G. W.; Schlegel, H. B.; Scuseria, G. E.; Robb, M. A.; Cheeseman, J. R.; Scalmani, G.; Barone, V.; Mennucci, B.; Petersson, G. A.; Nakatsuji, H.; Caricato, M.; Li, X.; Hratchian, H. P.; Izmaylov, A. F.; Bloino, J.; Zheng, G.; Sonnenberg, J. L.; Hada, M.; Ehara, M.; Toyota, K.; Fukuda, R.; Hasegawa, J.; Ishida, M.; Nakajima, T.; Honda, Y.; Kitao, O.; Nakai, H.; Vreven, T.; Montgomery, J. A., Jr.; Peralta, J. E.; Ogliaro, F.; Bearpark, M.; Heyd, J. J.; Brothers, E.; Kudin, K. N.; Staroverov, V. N.; Kobayashi, R.; Normand, J.; Raghavachari, K.; Rendell, A.; Burant, J. C.; Iyengar, S. S.; Tomasi, J.; Cossi, M.; Rega, N.; Millam, J. M.; Klene, M.; Knox, J. E.; Cross, J. B.; Bakken, V.; Adamo, C.; Jaramillo, J.; Gomperts, R.; Stratmann, R. E.; Yazyev, O.; Austin, A. J.; Cammi, R.; Pomelli, C.; Ochterski, J. W.; Martin, R. L.; Morokuma, K.; Zakrzewski, V. G.; Vorh, G. A.; Salvador, P.; Dannenberg, J. J.; Dapprich, S.; Daniels, A. D.; Farkas, O.; Foresman, J. B.; Ortiz, J. V.; Cioslowski, J.; Fox, D. J. Gaussian 09, revision A.02; Gaussian, Inc.: Wallingford, CT, 2009.
  - (38) Simion, D. V.; Sorensen, T. S. J. Am. Chem. Soc. 1996, 118, 7345.

- (39) Berndt, A. F.; Marsh, R. F. Acta Crystallogr. 1963, 16, 118.
- (40) Byers, L. R.; Dahl, L. F. Inorg. Chem. 1980, 19, 277.
- (41) Burrow, P. D.; Modelli, A.; Guerra, M.; Jordan, K. D. Chem. Phys. Lett. 1985, 118, 328.
- (42) (a) Baerends, E. J.; Gritsenko, O. V. J. Phys. Chem. A 1997, 101,
  5383. (b) Baerends, E. J. Theor. Chem. Acc. 2000, 103, 265. (c) Chong,
  D. P.; Gritsenko, O. V.; Baerends, E. J. J. Chem. Phys. 2002, 116, 1760.
  (43) Falcetta, M. F.; Chui, Y.; Jordan, K. D. J. Phys. Chem. A 2000,

1044, 9605.

- (44) (a) When the energies of ODC solutions are much higher than those of the resonance states, no avoided crossings will be found in the stabilization graphs. (b) In typical stabilization graphs, the lowest states correspond to ODC solutions if basis sets with sufficiently low exponents are used in HF or post-HF methods. However, the lowest states are typically the resonance states when using the DFT method. The reasons for the discrepancy are perhaps caused by different considerations of exchange-correlation potential and self-interaction effect.
- (45) Jordan, K. D.; Michejda, J. A.; Burrow, P. D. J. Am. Chem. Soc. 1976, 98, 7189.
  - (46) Chao, J. S.-Y.; Jordan, K. D. J. Phys. Chem. 1987, 91, 5578.
  - (47) Vydrov, O. A.; Scuseria, G. E. J. Chem. Phys. 2005, 122, 184107.
  - (48) Dutoi, A. D.; Head-Cordan, M. Chem. Phys. Lett. 2006, 422, 230.

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