

# Multiple-Time Step Ab Initio Molecular Dynamics Based on Two-Electron Integral Screening

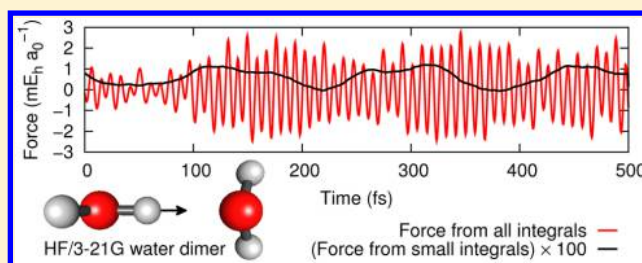
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## Supporting Information

**ABSTRACT:** A multiple-timestep ab initio molecular dynamics scheme based on varying the two-electron integral screening method used in Hartree–Fock or density functional theory calculations is presented. Although screening is motivated by numerical considerations, it is also related to separations in the length- and timescales characterizing forces in a molecular system: Loose thresholds are sufficient to describe fast motions over short distances, while tight thresholds may be employed for larger length scales and longer times, leading to a practical acceleration of ab initio molecular dynamics simulations.

Standard screening approaches can lead, however, to significant discontinuities in (and inconsistencies between) the energy and gradient when the screening threshold is loose, making them inappropriate for use in dynamics. To remedy this problem, a consistent window-screening method that smooths these discontinuities is devised. Further algorithmic improvements reuse electronic-structure information within the dynamics step and enhance efficiency relative to a naïve multiple-timestepping protocol. The resulting scheme is shown to realize meaningful reductions in the cost of Hartree–Fock and B3LYP simulations of a moderately large system, the protonated sarcosine/glycine dipeptide embedded in a 19-water cluster.



## I. INTRODUCTION

Ab initio molecular dynamics (AIMD) techniques combine the locality and efficiency of classical nuclear dynamics with on-the-fly generation of forces from electronic structure theory to generate real-time chemical information at an atomistic level.<sup>1–4</sup> Electronic structure theory can provide significantly more accurate forces than classical force fields, particularly for systems exhibiting strong polarization, charge transfer, and bond rearrangements. The cost of electronic-structure forces is many orders of magnitude larger than that of force fields, however, effectively prohibiting the use of AIMD for many systems and time scales of chemical interest.

Classical, nonpolarizable force fields often partition the potential energy of chemical systems into several bonded and nonbonded contributions,<sup>1</sup>

$$E = \underbrace{E_{\text{bonds}} + E_{\text{angles}} + E_{\text{dihedrals}}}_{E_{\text{bonded}}} + \underbrace{E_{\text{van der Waals}} + E_{\text{electrostatics}}}_{E_{\text{non-bonded}}} \quad (1)$$

each of which is represented by analytic functions which can be differentiated to obtain forces. The bonded terms are typically parametrized to favor the equilibrium bond lengths, bond angles, and dihedral angles of the discrete chemical species in the system; the nonbonded terms primarily describe intermolecular interactions. While the Lennard-Jones potentials

typically used to represent the van der Waals dispersion forces are basically short-range (decaying as  $r^{-6}$ ), the  $(1/r)$  Coulomb potential mediating the electrostatics is inherently long-range. As a result, evaluation of the electrostatic energy and forces is the most time-consuming part of force-field treatments of large systems.<sup>5</sup>

The cost of electrostatics is compounded in MD simulations by the fact that forces must be computed after every timestep. Reliable trajectories are assured by choosing the timestep to be a fraction of the period of the fastest motion in the system, typically the vibration of a hydrogen-containing bond; yet the long-range forces may not change significantly on this time scale. As such, standard MD integration entails unnecessarily frequent reevaluation of these forces. A common strategy for eliminating this redundancy is to treat long-range electrostatics as a periodic correction to a reference system containing the bonded, van der Waals, and short-range electrostatic interactions, such that the reference and correction are effectively integrated with different (“multiple”) timesteps.<sup>5–8</sup>

Multiple-timestep (MTS) molecular dynamics can be derived most naturally from a Liouvillian formulation of classical mechanics,<sup>9</sup> although the MTS idea predates<sup>10–15</sup> the resulting reversible reference-system propagator algorithm (r-RESPA). The basic algorithm is presented in ref 9 and will not be reviewed here. For the purposes of this work, multiple-timestep

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MD constitutes a rigorously time-reversible sequence of  $n$  velocity-Verlet “inner timesteps” using a computationally expedient force, followed by a momentum correction (or “outer timestep”) associated with the remaining contributions to the total force.

In AIMD, analytic potential and force functions are replaced by electronic-structure energies and gradients, which are determined from the full-system Hamiltonian  $H(\mathbf{R})$  and geometry  $\mathbf{R}$  via a nonlinear self-consistent-field (SCF) procedure. This black-box character comes at the cost of separability; in particular, an explicit partitioning of forces is manifestly impossible. It would seem, then, that *ab initio* molecular dynamics simulations must be limited in scope to short time scales.

Multiple-timestep AIMD schemes *can* be formulated, nevertheless. The full-system force from a low-level theory  $\mathbf{F}_{\text{low}}$  is taken as a reference and subsequently corrected with the force *difference* from a higher-level theory,  $\Delta\mathbf{F} \equiv \mathbf{F}_{\text{high}} - \mathbf{F}_{\text{low}}$ . Over the past decade, MTS has been applied to AIMD in this way. For example, an affordable density functional theory (such as PBE) can be corrected with forces from higher-level DFT (such as B3LYP).<sup>16,17</sup> These schemes were treated as *ad hoc* approaches for accelerating dynamics calculations, and general physical justifications for their use remain an open question. Formally justified MTS methods have also been used to propagate electronic information in the context of extended-Lagrangian MD approaches.<sup>18,19</sup>

Recently, an *ab initio* multiple-timestep scheme was presented<sup>20</sup> in which Hartree–Fock theory (HF) was corrected with a correlated wavefunction method, second-order Møller–Plesset perturbation theory (MP2).<sup>21</sup> In the limit of large system size  $N$ , HF/MP2 multiple-timestepping realizes a reduction in the computational cost of an MD simulation which is linear in the number of inner timesteps, as expected from the cost dominance of MP2 over HF [ $O(N^5)$  versus  $O(N^4)$  formal scaling].<sup>22,23</sup> Moreover, this scheme has a solid physical foundation, which is the observed separation between the time scales of variation in the HF forces and in the additional correlation forces obtained from MP2. To the extent that similar separations can be identified among HF, MP2, and more powerful correlated methods (such as coupled-cluster theory), highly accurate dynamics simulations of small systems may soon become feasible.

Of course, correlated wavefunction theories are cost-prohibitive for studying truly large systems, and even Hartree–Fock or DFT may be unappealingly costly. If these latter methods were considered monolithic, then the usefulness of multiple-timestepping as an AIMD acceleration tool would already be exhausted. These basic methods can, however, vary widely in accuracy and computational cost, depending on a range of choices made in advance of a calculation. Some of these choices reflect purely numerical considerations, such as the convergence criterion for the self-consistent-field procedure; others have physical content, such as the basis set or density functional employed. The extent to which variations in each of these (generalized) parameters would support a multiple-timestepping scheme is a subject that remains largely unexplored. The present work is focused on the effect of changes in two-electron integral screening, keeping other parameters fixed.

**A. Motivation for Screening-Based Multiple-Time-stepping.** Screening-based multiple-timestepping is motivated

by the connection between the magnitude of each two-electron integral and the length scale of the interactions it describes. Consider the integral

$$(\mu\nu|\lambda\sigma) = \int d\mathbf{r}_1 \int d\mathbf{r}_2 \phi_\mu^*(\mathbf{r}_1; \mathbf{R}_A) \phi_\nu(\mathbf{r}_1; \mathbf{R}_B) \frac{1}{r_{12}} \phi_\lambda^*(\mathbf{r}_2; \mathbf{R}_C) \phi_\sigma(\mathbf{r}_2; \mathbf{R}_D) \quad (2)$$

where  $\phi_\mu(\mathbf{r}_1; \mathbf{R}_A)$  is a contracted-Gaussian atomic orbital centered on the nucleus at position  $\mathbf{R}_A$  and occupied by an electron with coordinate  $\mathbf{r}_1$ . Orbital pair  $\mu\nu$  forms a charge distribution centered at  $\mathbf{R}_{\mu\nu}$  which decays exponentially as  $S_{\mu\nu} \sim \exp(-\alpha_{\mu\nu} R_{AB}^2)$ , with  $R_{AB} \equiv |\mathbf{R}_A - \mathbf{R}_B|$  the internuclear distance and  $\alpha_{\mu\nu}$  a composite of primitive-Gaussian widths. The asymptotic behavior of the integral is therefore given by<sup>24</sup>

$$(\mu\nu|\lambda\sigma) \sim \frac{S_{\mu\nu} S_{\lambda\sigma}}{R_{\mu\nu\lambda\sigma}} \quad (3)$$

where  $R_{\mu\nu\lambda\sigma} \equiv |\mathbf{R}_{\mu\nu} - \mathbf{R}_{\lambda\sigma}|$  is the distance between the charge centers.

If either charge distribution is highly attenuated—its constituent orbitals are centered on distant nuclei—or if the charge distributions themselves are well-separated, the integral will be small. Such integrals can often be discarded altogether without measurably affecting the accuracy of the calculation. “Screening” is the umbrella term for a constellation of techniques which exclude integrals on this basis, thereby reducing the formal  $O(N^4)$  cost of constructing these four-index quantities to a more modest quadratic (or even linear) scaling.<sup>24</sup>

Equation 3 clearly states that the size of a two-electron integral is inversely proportional to the length scale of the electronic interaction it describes. Thus, the screening threshold is equivalent, in some sense, to a range-separation parameter for the interelectronic repulsion. Given that interactions over small length scales tend to be associated with motions of high frequency, the forces associated with integrals of disparate size should also vary on disparate time scales, which is a scenario tailor-made for multiple-timestepping. This intuition is borne out in the analysis that follows, and it is the foundation of a practical approach for accelerating AIMD simulations.

To the extent that the intuitive correspondence between the screening threshold and a range-separation parameter is valid, the multiple-timestep approach developed here is similar in spirit to the MTS-CASE method of Luehr and co-workers.<sup>25</sup> MTS-CASE is based on an explicit partitioning of the bare Coulomb potential felt by both electrons *and* nuclei,

$$\frac{1}{r} = \frac{\text{erf}(\omega r)}{r} + \frac{\text{erfc}(\omega r)}{r} \quad (4)$$

where  $\text{erf}(\omega r)$  is the (long-range) error function;  $\text{erfc}(\omega r)$  is the (short-range) complementary error function; and the parameter  $\omega$  sets the length scale of the range separation. (The name “MTS-CASE” stems from the fact that neglecting the long-range Coulomb tail in the inner timesteps is equivalent to adopting the Coulomb-attenuated Schrödinger equation, or CASE model.<sup>26</sup>) Luehr et al. noted<sup>25</sup> that MTS-CASE inner timesteps can be made linearly scaling, owing to efficient screening of the CASE two-electron integrals,<sup>27</sup> although this connection has not yet been pursued. Evidence of a rigorous time scale separation between CASE and standard Hartree–

Fock forces also remains to be presented, but such a separation is expected. In the analysis that follows, the range separation implicit in the screening threshold is explicitly shown to be associated with a separation in time scale, as required to justify MTS. Direct comparisons between our method and MTS-CASE are reserved for future work.

**B. Outline of the Present Work.** This manuscript is organized as follows. In Section II, a detailed description of the development of the screening-based multiple-timestep scheme is provided, including a review of the simplest complete sequence of screening methods used in computing ab initio energies and gradients (II A); a demonstration that these classic techniques break down when the screening threshold is significantly loosened (II B); and a remedy for these pathologies (II C). The screening-based MTS protocol is formulated in Section II D, and the method is shown to be phenomenologically justified by a timescale separation in the forces. Instead of proceeding immediately to numerical tests, Section II E highlights the fact that classical and ab initio multiple-timestep schemes are governed by different cost inequalities, such that computational savings are harder to achieve in the ab initio setting. Several approaches to subverting this limitation are described. In Section III, timings are presented for a model biological system, the protonated sarcosine/glycine dipeptide embedded in a 19-water cluster. Section IV concludes with a summary of findings and a discussion of possible future directions.

## II. METHODS DEVELOPMENT

A screening-based multiple-timestep AIMD scheme is developed in this work. Briefly stated, the method consists of several standard MD steps at loose screening thresholds, followed by a correction to the momenta—in the usual r-RESPA fashion<sup>9</sup>—from the difference in forces obtained using loose and tight thresholds. Although this MTS approach is conceptually straightforward, it entails several technical hurdles, which require both an explication of integral screening and an analysis of the pathologies that can arise in standard screening techniques. These pathologies can directly impact the accuracy of single-timestep AIMD as well and, to the authors' knowledge, have not been analyzed in the literature.

**A. Review of Basic Screening Methods.** This section summarizes the simplest complete sequence of screening methods that might be employed in computing energies and gradients using an atomic-orbital-based quantum-chemistry software package.<sup>28–31</sup>

**1. Before SCF: Shell–Pair Economization.** Because screening is based on the asymptotic behavior of the two-electron integrals, it requires a modest computational overhead relative to the cost of computing the integrals outright. In the limit of large system size, however, the number of integrals to be screened continues to grow quartically, even as the number of integrals *retained* grows quadratically. As a result, the screening overhead can eventually dominate the cost of the calculation. One means of reducing the overhead is to eliminate integrals from consideration by some crude (but efficient) prescreening method. Prescreening is *not* the focus of the methodological developments of this work, but its relationship to more rigorous screening merits brief review.

In contracted-Gaussian basis sets, each atomic orbital is composed of a sum of primitive Gaussians. For example, an *s*-type orbital has the form

$$\phi_\mu(\mathbf{r}_i; \mathbf{R}_A) = \sum_m c_{\mu m} g_m(\mathbf{r}_i; \mathbf{R}_A) \quad (5)$$

where  $c_{\mu m}$  is the contraction coefficient of the normalized primitive  $g_m(\mathbf{r}_i; \mathbf{R}_A)$  with width  $\alpha_m$ . Consequently, the charge distribution associated with an *s*-type orbital pair  $\mu\nu$  is a linear combination of primitive Gaussian “shell pairs”  $mn$ ,

$$\phi_\mu^*(\mathbf{r}_i; \mathbf{R}_A) \phi_\nu(\mathbf{r}_j; \mathbf{R}_B) = \sum_{mn} c_{\mu m}^* c_{\nu n} g_m^*(\mathbf{r}_i; \mathbf{R}_A) g_n(\mathbf{r}_j; \mathbf{R}_B) \quad (6)$$

The prescreening approach known as “shell–pair economization” consists of computing the orbital pairs  $\mu\nu$  before the SCF procedure begins, discarding shell pairs if their overlap is smaller than a threshold value  $\epsilon_{\text{shell}}$ .<sup>31</sup> The overlap is given analytically by

$$s_{mn} = \int d\mathbf{r} g_m(\mathbf{r}_i; \mathbf{R}_A) g_n(\mathbf{r}_j; \mathbf{R}_B) \quad (7)$$

$$= \left( \frac{\pi}{\alpha_{mn}} \right)^{3/2} \mathcal{N}_m \mathcal{N}_n e^{-\alpha_m \alpha_n / \alpha_{mn} R_{AB}^2} \quad (8)$$

where  $\alpha_{mn} \equiv \alpha_m + \alpha_n$  and  $\mathcal{N}$  denotes a normalization constant. Shell–pair economization may therefore be written as

$$\begin{aligned} \phi_\mu(\mathbf{r}_i; \mathbf{R}_A) \phi_\nu(\mathbf{r}_j; \mathbf{R}_B) \\ \approx \sum_{mn} c_{\mu m} c_{\nu n} \Theta(s_{mn} - \epsilon_{\text{shell}}) g_m(\mathbf{r}_i; \mathbf{R}_A) g_n(\mathbf{r}_j; \mathbf{R}_B) \end{aligned} \quad (9)$$

with  $\Theta(x)$  the (right-continuous) Heaviside step function,

$$\Theta(x) = \begin{cases} 0, & x < 0 \\ 1, & x \geq 0 \end{cases} \quad (10)$$

If the internuclear distance  $R_{AB}$  or threshold  $\epsilon_{\text{shell}}$  is large, all of the shell pairs constituting  $\mu\nu$  might be discarded by eq 9. All of the integrals involving  $\mu\nu$  would, in turn, be excluded from the calculation.

**2. During SCF: Integral Screening.** After shell–pair economization, the energy is determined via the SCF procedure. Whether the system is treated with Hartree–Fock or Kohn–Sham DFT, each SCF cycle involves the construction of a contribution to the energy from the two-electron integrals,

$$E^{2e-} = \frac{1}{2} \text{Tr} \mathbf{P} \cdot \mathbf{F}^{2e-} = \frac{1}{2} \text{Tr} \mathbf{P} \cdot (\mathbf{J} - c_X \mathbf{K}) = \frac{1}{2} \text{Tr} \mathbf{P} \cdot \mathbf{\Pi} \cdot \mathbf{P} \quad (11)$$

Here,  $\mathbf{P}$  is the density matrix;  $\mathbf{F}^{2e-}$  is the two-electron portion of the Fock matrix, comprising the Coulomb matrix  $\mathbf{J}$  and an appropriate fraction  $c_X$  of the exact Hartree–Fock exchange matrix  $\mathbf{K}$  (but excluding the exchange–correlation matrix); and  $\mathbf{\Pi}$  denotes the antisymmetrized two-electron integrals. (Spin indices and sums have been suppressed.) In the equations that follow,  $c_X$  is taken to be 1, which is the appropriate value for Hartree–Fock. The resulting screening expressions are general; calculations using hybrid functionals ( $0 < c_X \leq 1$ ) use the same screening criteria but scale  $\mathbf{K}$  after it has been constructed. Calculations using pure functionals, by contrast, simply omit the exchange-related terms from any given screening prescription.

Taking permutational symmetry into account, the two-electron contributions to the energy are proportional to



$$E^{2e^-} \propto \sum_{\mu\nu\lambda\sigma} (\mu\nu|\lambda\sigma) \underbrace{(2P_{\mu\nu}P_{\lambda\sigma})}_{\text{Coulomb}} \underbrace{(-P_{\mu\lambda}P_{\nu\sigma} - P_{\mu\sigma}P_{\nu\lambda})}_{\text{exchange}} \quad (12)$$

The overarching goal of screening is to avoid computing integrals unless they contribute materially to the energy. An affordable estimate of the largest contribution from  $(\mu\nu|\lambda\sigma)$  in eq 12 is, therefore, required. The bracket representing the integral has the properties of an inner product, so the Schwarz inequality can be invoked,<sup>29</sup>

$$(\mu\nu|\lambda\sigma) \leq \sqrt{(\mu\nu|\mu\nu)} \sqrt{(\lambda\sigma|\lambda\sigma)} \equiv Q_{\mu\nu} Q_{\lambda\sigma} \quad (13)$$

The Schwarz estimate  $Q_{\mu\nu}Q_{\lambda\sigma}$  provides a rigorous upper bound on  $(\mu\nu|\lambda\sigma)$  in terms of the self-integrals of molecular-orbital pairs, which scale as  $O(N^2)$ . This bound could be tightened further by taking the asymptotic  $R_{\mu\nu\lambda\sigma}^{-1}$  dependence of the exact integral into account,<sup>24,32</sup> but the common practice in atomic-orbital-based quantum chemistry has been to use eq 13 without further refinement.

Because density matrix elements are restricted in magnitude,  $|P_{\mu\nu}| \in [0,1]$ , eq 12 can be converted into a rigorous estimate for the largest unsigned contribution to the energy from  $(\mu\nu|\lambda\sigma)$ ,

$$\bar{E} \equiv Q_{\mu\nu} Q_{\lambda\sigma} \max(2P_{\mu\nu}P_{\lambda\sigma}, P_{\mu\lambda}P_{\nu\sigma}, P_{\mu\sigma}P_{\nu\lambda}) \quad (14)$$

The screening method suggested by this expression is to discard the integral if  $\bar{E}$  is smaller than a threshold  $\epsilon_{\text{hard}}$ :

$$(\mu\nu|\lambda\sigma) \rightarrow \Theta(\bar{E} - \epsilon_{\text{hard}}) (\mu\nu|\lambda\sigma) \quad (15)$$

In fact, such an approach would be unwise.<sup>28</sup> Although computation of the SCF energy is the aim, steady convergence of the SCF procedure hinges on the accuracy of the two-electron portion of the Fock matrix,  $F^{2e^-} = \mathbf{P} \cdot \mathbf{\Pi}$ , which only includes a single factor of the density. Thus, integrals with negligible estimated contributions to the energy may nevertheless be important in the Fock matrix. Equation 14 can be adapted to reflect this observation by splitting pairs of density matrix elements,

$$\bar{F} \equiv Q_{\mu\nu} Q_{\lambda\sigma} \max(2P_{\mu\nu}, 2P_{\lambda\sigma}, P_{\mu\lambda}, P_{\mu\sigma}, P_{\nu\lambda}, P_{\nu\sigma}) \quad (16)$$

Screening is then equivalent to the replacement

$$(\mu\nu|\lambda\sigma) \rightarrow \Theta(\bar{F} - \epsilon_{\text{hard}}) (\mu\nu|\lambda\sigma) \quad (17)$$

**3. After SCF: Gradient Screening.** With the converged density matrix in-hand, the nuclear gradient of the energy (hereafter referred to simply as “the gradient”) can now be computed. Because the density is obtained variationally, the only explicit contribution to the gradient from the two-electron integrals is the term

$$(\nabla_{\mathbf{R}} E)^{2e^-} = \frac{1}{2} \text{Tr} \mathbf{P} \cdot \mathbf{\Pi}^{[\mathbf{R}]} \cdot \mathbf{P} \quad (18)$$

where the superscript  $[\mathbf{R}]$  is shorthand for the Cartesian gradient over all nuclear coordinates. Any given integral depends on, at most, four nuclear coordinates, so the scaling of this term is still formally  $O(N^4)$ .

A gradient-screening method could potentially be developed directly from eq 18 (or, equivalently, from the gradient of eq 12).<sup>33</sup> A rigorous, practical alternative is to take the gradient of the energy-based screening prescription of eq 15, on the grounds that only those integrals included in the energy should

subsequently contribute to the gradient. The derivative of the Heaviside function is a delta function, so this contribution is given by

$$(\mu\nu|\lambda\sigma)^{[\mathbf{R}]} \rightarrow \Theta(\bar{E} - \epsilon_{\text{hard}}) (\mu\nu|\lambda\sigma)^{[\mathbf{R}]} + \delta(\bar{E} - \epsilon_{\text{hard}}) \bar{E}^{[\mathbf{R}]} (\mu\nu|\lambda\sigma) \quad (19)$$

The delta-function term in this expression has a straightforward meaning; it describes the discontinuity in  $\Theta$ , which admits the integral into the energy calculation when  $\bar{E} = \epsilon_{\text{hard}}$ . This term can be neglected in the gradient, because (1) the delta function goes to infinity when its argument vanishes, which is unphysical, and (2) its effect is restricted to a single value of  $\bar{E}$ , which occurs with vanishingly small probability. Thus, gradient screening is equivalent to the replacement

$$(\mu\nu|\lambda\sigma)^{[\mathbf{R}]} \rightarrow \Theta(\bar{E} - \epsilon_{\text{hard}}) (\mu\nu|\lambda\sigma)^{[\mathbf{R}]} \quad (20)$$

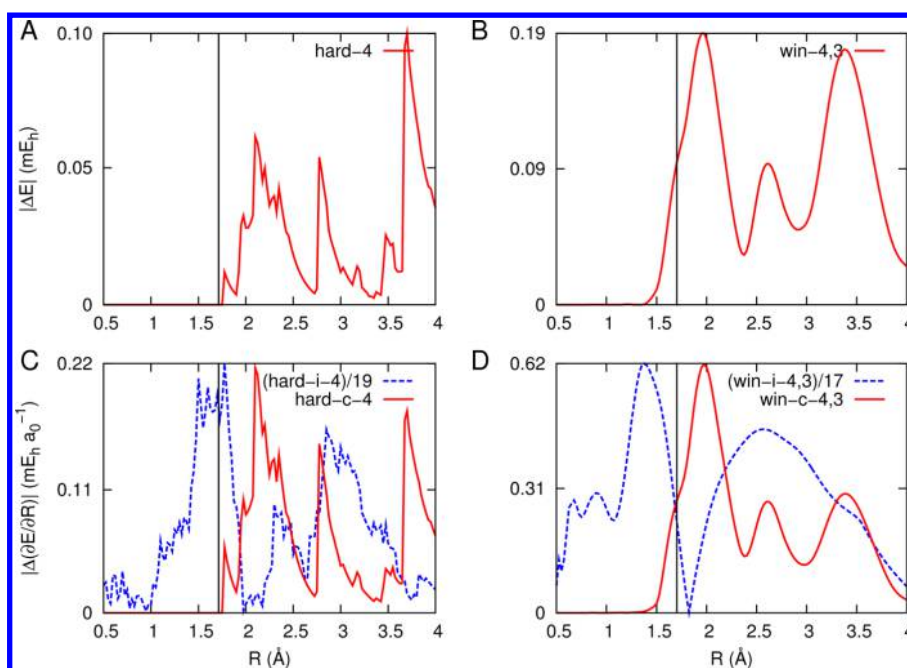
Equations 15 and 17 may admit somewhat different sets of integrals, especially when  $\epsilon_{\text{hard}}$  is large; hence, eq 20 is “inconsistent” gradient screening.

To summarize, a complete prescription for two-electron-integral screening in a Hartree–Fock or DFT calculation of the energy and gradient consists of (1) prescreening shell pairs against a threshold  $\epsilon_{\text{shell}}$  using eq 9, and (2) screening contributions to the Fock matrix and gradient against a threshold  $\epsilon_{\text{hard}}$  using eqs 17 and 20. By analogy to SCF convergence criteria, large thresholds are “loose,” in the sense that the resulting energy will exhibit errors of comparable size, while small thresholds are “tight.”

### B. Loose Thresholds Reveal Flaws in Hard Screening.

A screening-based multiple-timestep scheme will necessarily involve loosening the threshold in the inner timesteps. Instead of varying the thresholds for shell–pair economization and integral screening in the calculations that follow, the prescreening threshold  $\epsilon_{\text{shell}}$  is tightened to the smallest value allowed ( $10^{-14}$ ) by the chosen development environment, Q-CHEM.<sup>34</sup> This choice is motivated by the fact that shell–pair economization is not based on a rigorous integral bound; rather, it is a numerical convenience. Consequently, varying  $\epsilon_{\text{shell}}$  can lead to unpredictable changes in the two-electron integrals. This lack of rigor is in direct conflict with the intention to construct a multiple-timestep scheme based on well-controlled changes in screening. The base method for screening-based multiple-timestepping therefore consists of Fock-matrix-based integral screening (eq 17) and inconsistent gradient screening (eq 20) using threshold  $\epsilon_{\text{hard}} = 10^{-h}$ , where  $h$  is a positive integer. For the sake of brevity, we refer to this method as hard- $i$ - $h$ .

It may be advantageous for  $\epsilon_{\text{hard}}$  to be looser than the chosen SCF convergence criterion,  $10^{-8} E_h$ . In order to establish the properties of hard screening in this unusual regime, HF/6-31G\* potential energy curves and gradients were computed for a series of diatomic molecules—CO, He<sub>2</sub>, HCl, HF, H<sub>2</sub>, LiF, LiH, N<sub>2</sub>, and O<sub>2</sub>—with bond separations  $R \in [0.5, 4.0]$  Å. The increment between points was 0.025 Å, such that each surface was constructed from 141 individual gradient calculations. For each molecule, these calculations were repeated over a range of integer thresholds,  $h \in [2, 14]$ . The effective floor on integral thresholds in Q-CHEM is  $h = 14$ , because the fundamental electron-repulsion integrals are computed by modified Chebyshev interpolation from tabulated values, with a precision set to  $10^{-(h+2)}$ .<sup>35</sup>



**Figure 1.** Unsigned energy and gradient differences with respect to a hard-i-14 reference,  $|\Delta E|$  and  $|\Delta(\partial E/\partial R)|$ , obtained using various screening methods for lithium hydride at the HF/6-31G\* level of theory. Panels A and B show the energy differences for hard-4 and win-4,3, respectively; gradient consistency is immaterial. Panel C shows the gradient differences for hard-i-4 (divided by a factor of 19 for comparison with the consistent result, blue dashed line) and hard-c-4 (red solid line); Panel D, the gradient differences for win-i-4,3 (divided by a factor of 17, blue dashed line) and win-c-4,3 (red solid line). The dropline in each panel indicates the optimum Li–H bond separation,  $R = 1.635749$  Å. See Sections II B and II C for detailed discussion.

At the loosest thresholds ( $h = 2$  and  $3$ ), the calculations unsurprisingly exhibited poor convergence behavior, often failing to converge even after hundreds of SCF cycles using DIIS extrapolation.<sup>36,37</sup> These failures occur due to the exclusion of large integrals which are necessary to obtain a well-converged solution to the Roothaan equations. For the remaining thresholds,  $h \in [4, 14]$ , all of the potential curves and gradients were visually indistinguishable on the scale of the plots. Subtle features in the curves for each method were revealed by plotting unsigned energy and gradient differences with respect to hard-i-14, which is treated as reference data.

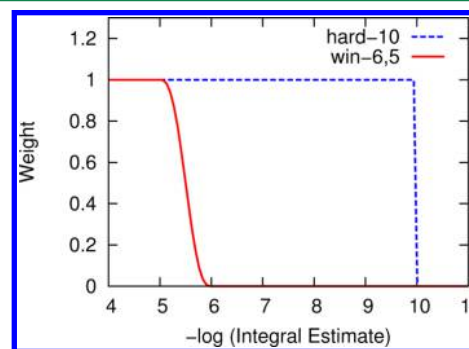
As a concrete example, energy and gradient differences for lithium hydride, treated with hard-i-4 screening, are plotted in Figures 1A and C. Both the energy and the gradient exhibit discontinuities with respect to the reference curve, some of which lie very close to (or even on top of) the potential minimum, which is the region most relevant for dynamics. Discontinuities in the energy are of the same order as  $\epsilon_{\text{hard}}$ , with concomitantly large discontinuities in the gradient; further, large spikes in the energy are not always matched by corresponding features in the gradient (and vice versa), confirming the expectation that inconsistent gradient screening might be problematic at loose thresholds.

Any discontinuity in the energy and gradient will have deleterious effects in a molecular dynamics simulation because the system will experience unphysical, instantaneous impulses. Inconsistencies in the gradient are especially troubling because they imply that the forces used to integrate the trajectory will be fundamentally incorrect. The solution to this latter problem is straightforward: The gradients should be screened using the Fock estimate of eq 16 rather than the energy estimate of eq 14,

$$(\mu\nu|\lambda\sigma)^{[R]} \rightarrow \Theta(\bar{F} - \epsilon_{\text{hard}}) (\mu\nu|\lambda\sigma)^{[R]} \quad (21)$$

When the Fock-based integral screening of eq 17 is paired with this consistent gradient screening, the combined method is referred to as hard-c- $h$ . The gradient results in Figure 1C show that, while still discontinuous, the hard-c-4 gradient is fully consistent with the energy. Note that hard-i-14 results remain a good reference—they are indistinguishable from hard-c-14, which only serves to emphasize that screening in the loose-threshold regime is qualitatively different from tighter screening.

**C. Window Screening.** Remedying the remnant discontinuities that appear in hard energies and gradients at loose thresholds requires an understanding of their origin. Hard screening effectively assigns a binary weight to each integral, as dictated by the Heaviside weight function  $w_{\text{hard}}(\bar{F}; \epsilon_{\text{hard}}) = \Theta(\bar{F} - \epsilon_{\text{hard}})$  plotted in Figure 2. Now, consider an integral with estimated Fock-matrix contribution comparable to the thresh-



**Figure 2.** Weight functions used in integral screening as a function of the logarithm of the integral estimate. The figure depicts a hard-10 weight function (blue dashed line) and a win-6,5 window-screening function (red solid line), as described in Section II C.

old ( $\bar{F} \approx \epsilon_{\text{hard}}$ ) when the system is in some geometry  $\mathbf{R}$ . Whether that integral will be retained or discarded after a small perturbation to the geometry,  $\mathbf{R} + d\mathbf{R}$ , will be determined by changes in the Schwarz estimates and density-matrix elements constituting  $\bar{F}$ . Depending on the sign and magnitude of these changes, therefore, the integral can “blink” on or off. When  $\epsilon_{\text{hard}}$  is very tight, blinking integrals will have an insignificant effect on the energy and gradient, but when the threshold is sufficiently loose, these integrals will lead to the observed discontinuities.

If screening-based multiple-timestepping is to be viable, a screening method must be developed which is less sensitive to small changes in the integral estimate, such that the corresponding potential energy curves and gradients will be continuous. To accomplish this aim, the hard cutoff  $\epsilon_{\text{hard}}$  is paired with a looser “soft” threshold  $\epsilon_{\text{soft}} = 10^{-5}$  to form a window spanning  $\Omega \equiv h - s$  orders of magnitude. This approach is hereafter termed window screening.

Within the window, each integral is assigned a fractional weight from a smoothly varying function satisfying two conditions: First, integrals with estimates close to  $\epsilon_{\text{hard}}$  should be assigned very small weights, such that blinking is a more modest source of error. Second, integrals with estimates close to  $\epsilon_{\text{soft}}$  should be rescaled only very delicately, with integrals above the soft threshold carrying full weight. Sigmoid functions of the *logarithm* of the estimate,  $\bar{f} \equiv -\log \bar{F}$ , have just these properties; a quintic spline with vanishing first and second derivatives at the boundaries of the window is chosen here, in direct analogy to smoothing functions used in force-field-based multiple-timestep approaches.<sup>38,39</sup>

The window-screening approach for the energy is therefore defined by the replacement

$$(\mu\nu|\lambda\sigma) \rightarrow w(\bar{f}; h, s) (\mu\nu|\lambda\sigma) \quad (22)$$

where the weight function  $w$  is

$$w(\bar{f}; h, s) = \begin{cases} 1, & \bar{f} \leq s \\ 1 + u^3(15u - 6u^2 - 10), & s < \bar{f} < h \\ 0, & \bar{f} \geq h \end{cases} \quad (23)$$

with spline variable  $u$  given by

$$u(\bar{f}; h, s) = \frac{\bar{f} - s}{\Omega} \quad (24)$$

An example of this weight function is plotted in Figure 2. As may be apparent from the figure,  $w = u = (1/2)$  at the midpoint of the window,  $\bar{f} = s + (\Omega/2)$ . Equivalently, integrals  $(\sqrt{10})^\Omega$  times smaller than the soft threshold—a bit more than a factor of 3, for an integer window—will be assigned half weight or less.

Because eq 23 recovers the Heaviside function when  $h = s$ , window screening can be understood as a hard screening generalization that yields the same set of two-electron integrals but is not as susceptible to blinking. Additional error is incurred because the integrals have been rescaled—thereby interpolating between hard- $h$  and hard- $s$  results and costs, with a small additional overhead from computing and applying the integral weights—but Figure 1B indicates that the largest errors only increase by a factor of 2. Importantly, these errors are associated with smooth features in the potential-energy surface.

In practice, an integer window has been found to achieve the best compromise between accuracy and smoothing.

Of course, the discontinuities that appear in the gradient must also be smoothed. As before, the gradient of eq 22 is taken to obtain

$$(\mu\nu|\lambda\sigma)^{[\mathbf{R}]} \rightarrow w(\bar{f}; h, s) (\mu\nu|\lambda\sigma)^{[\mathbf{R}]} + \frac{\partial w}{\partial \bar{f}} \bar{f}^{[\mathbf{R}]} (\mu\nu|\lambda\sigma) \quad (25)$$

The quintic spline  $w$  was chosen because its derivatives vanish at the boundaries of the window, which significantly limits the effect of the weight-function derivative in eq 25. However, unlike the delta function that initially appeared in the hard gradient (eq 19), this term is neither patently unphysical nor restricted in effect to a single value of  $\bar{f}$ :

$$\frac{\partial w}{\partial \bar{f}} = \begin{cases} 0, & \bar{f} \leq s \\ -\frac{30}{\Omega} u^2(u-1)^2, & s < \bar{f} < h \\ 0 & \bar{f} \geq h \end{cases} \quad (26)$$

The derivative smoothly increases from zero at either boundary to a modest value of  $(\partial w/\partial \bar{f}) = -(15/8\Omega)$  at the center of window. The gradient of  $\bar{f}$  is given by

$$\bar{f}^{[\mathbf{R}]} = -(\log \bar{F})^{[\mathbf{R}]} = -\left(\frac{\ln \bar{F}}{\ln 10}\right)^{[\mathbf{R}]} = -\frac{\bar{F}^{[\mathbf{R}]}}{\bar{F} \ln 10} \quad (27)$$

Evaluating the gradient of the Fock-matrix estimate (eq 16) and canceling terms,

$$\frac{\partial w}{\partial \bar{f}} \bar{f}^{[\mathbf{R}]} (\mu\nu|\lambda\sigma) = \frac{30u^2(u-1)^2}{\Omega \ln 10} \left( \frac{Q_{\mu\nu}^{[\mathbf{R}]}}{Q_{\mu\nu}} + \frac{Q_{\lambda\sigma}^{[\mathbf{R}]}}{Q_{\lambda\sigma}} + \frac{P_{\max}^{[\mathbf{R}]}}{P_{\max}} \right) (\mu\nu|\lambda\sigma) \quad (28)$$

where  $P_{\max}$  is shorthand for the largest of the density-matrix elements in eq 16.

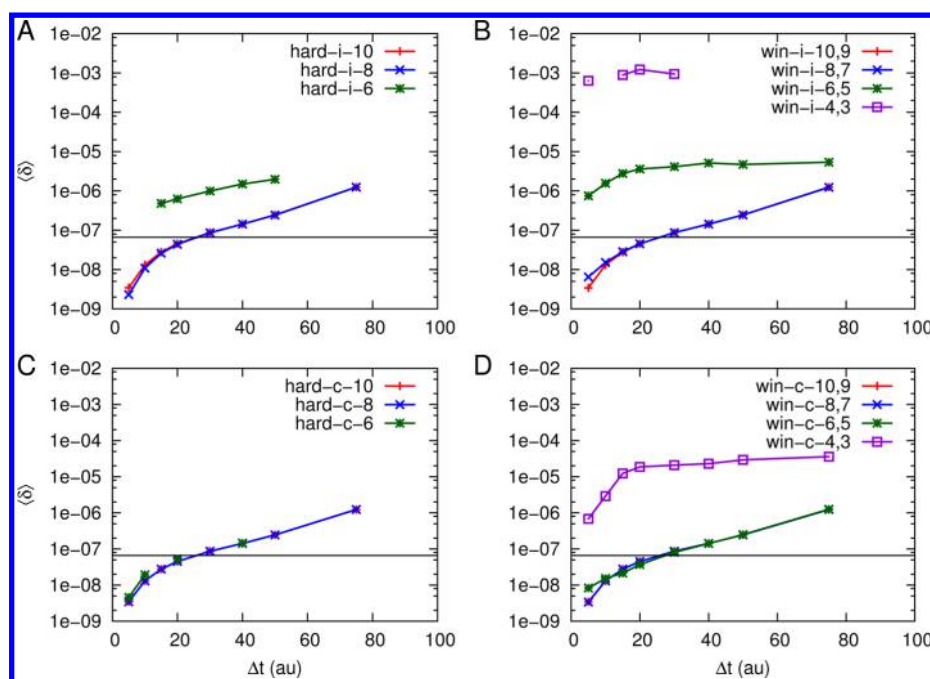
Each Schwarz-gradient term in eq 28 will vanish identically for all nuclear coordinates but those of the (one or two) centers forming the relevant orbital pair. For those coordinates, the overall asymptotic behavior will be, for example (cf. eq 3),

$$\frac{Q_{\mu\nu}^{[\mathbf{R}_A]}}{Q_{\mu\nu}} (\mu\nu|\lambda\sigma) \sim R_{AB} \frac{S_{\mu\nu} S_{\lambda\sigma}}{R_{\mu\nu\lambda\sigma}} \quad (29)$$

which tends to zero in the limit of small or large internuclear separations. These terms may, therefore, be discarded with little or no effect.

The density gradient  $P_{\max}^{[\mathbf{R}]}$  will be largest when perturbations to the geometry strongly mix occupied and virtual orbitals, such as when the HOMO–LUMO gap is small. To the extent that Hartree–Fock and DFT are unreliable in the limit of vanishing fundamental gap, this term is discarded, too. There is an additional practical reason for doing so: To evaluate  $P_{\max}^{[\mathbf{R}]}$  exactly would require that the coupled-perturbed Hartree–Fock or Kohn–Sham equations be solved, at significant additional cost.<sup>40–44</sup> As shown below, neglecting this gradient has no serious consequences for MD; in fact, the statistical properties of the dynamics improve when window-screened gradients are used.





**Figure 3.** Mean energy fluctuations  $\langle \delta \rangle$  (eq 31) in completed HF/3-21G water-dimer trajectories as a function of timestep and screening method. Panels A and B show fluctuations for hard and win screening methods with inconsistent gradients, respectively; Panels C and D, with consistent gradients. The horizontal line in each panel indicates the threshold below which the magnitude of the mean energy fluctuation will be smaller than  $10 \mu E_h$ . See Section II D for detailed discussion.

In light of the above discussion, the consistent window screening method win-c- $h,s$  is implemented for gradients as

$$(\mu\nu|\lambda\sigma)^{[R]} \rightarrow w(\bar{f}; h, s) (\mu\nu|\lambda\sigma)^{[R]} \quad (30)$$

Inconsistent screening (win-i- $h,s$ ) can be performed instead by replacing  $\bar{f}$  in eq 30 with the negative logarithm of the energy estimate  $\bar{E}$  (eq 14).

Lithium hydride gradients obtained using these methods are shown in Figure 1D. As was observed for the energy, the size of the errors in the gradient increases by a factor of 2 or 3; the associated discontinuities are fully smoothed. Moreover, the behavior of the win-c gradient is consistent with the behavior of the energy.

**D. Formulation and Validation of Screening-Based Multiple-Timestepping.** Having established that window-screening methods smooth the discontinuities engendered in potential energy surfaces and gradients by standard screening approaches, this section examines the effect of screening method on the statistical properties of single-timestep ab initio molecular dynamics trajectories in the microcanonical ensemble. To probe the effect of screening, the water dimer was simulated with HF/3-21G. The initial configuration of the system was obtained by reoptimizing the minimum-energy structure from the Cambridge Cluster Database<sup>45,46</sup> using hard-c-14 screening; initial velocities were identical in all simulations and were sampled from the Maxwell–Boltzmann distribution at 298 K.

Trajectories were computed using a range of screening methods—hard and win, with and without gradient consistency; with hard cutoffs  $h \in \{4, 6, 8, 10, 12, 14\}$  and soft thresholds  $s = h - 1$ , and with timesteps  $\Delta t \in \{5, 10, 15, 20, 30, 40, 50, 75, 100\}$  au. Each trajectory was integrated for a maximum of  $\tau_{\text{steps}} = 10\,000$  steps or to the first instance of failed SCF convergence. In this way, the longest viable timestep was found to be  $\Delta t = 75$  au; when  $\Delta t = 100$  au, one of the

water molecules invariably autoionized, followed closely by SCF convergence failure. While this result is entirely dependent on the choice of SCF convergence algorithm, the success or failure of a series of trajectories—each employing the *same* convergence algorithm—is indicative of the robustness of the MD method. This result also highlights the fact that AIMD functions as a reactive force field and that pathologies in the forces or the MD integrator can lead to qualitatively incorrect behavior.

While symplectic integrators are guaranteed to yield energy-conserving trajectories for small enough timesteps,<sup>47</sup> step-to-step energy fluctuations are not guaranteed to vanish, and these fluctuations are indicative of the quality of the integration. Energy conservation of all complete trajectories was assessed using the time-average of these energy fluctuations relative to the initial energy,  $E_0$ :

$$\langle \delta \rangle = \left\langle \left| \frac{E_\tau - E_0}{E_0} \right| \right\rangle = \frac{1}{\tau_{\text{steps}}} \sum_{\tau=1}^{\tau_{\text{steps}}} \left| \frac{E_\tau - E_0}{E_0} \right| \quad (31)$$

The smaller the value of  $\langle \delta \rangle$ , the more closely the trajectory conserves energy.

The fluctuation data are presented in Figure 3. Curves for methods with  $h \geq 10$  are indistinguishable, and only  $h = 10$  curves are shown. Several conclusions can be drawn from these data:

- Window screening makes trajectories with loose thresholds more robust against SCF convergence failures, in exchange for a modest increase in the mean fluctuation. In particular, the win-i-6,5 curve is complete, while the hard-i-6 curve is missing several points. Further, some of the win-i-4,3 trajectories ran to completion, unlike those using hard-i-4.
- Gradient consistency has a qualitative effect on energy conservation in loosely screened trajectories. Fluctua-

tions in hard-c-6 were reduced by an order of magnitude relative to hard-i-6, while fluctuations in win-c-6,5 and win-c-4,3 were reduced by a factor of 100 relative to their inconsistently screened counterparts.

- Window screening with a consistent gradient is surprisingly stable—all of the win-c-4,3 trajectories ran to completion.

The magnitude of the mean energy fluctuation provides a rational basis for selecting the inner timestep (denoted  $\delta t$ ) in screening-based MTS:  $\delta t$  is chosen such that the mean energy fluctuation will be smaller than  $10 \mu E_h$ , equivalent to a 1% fluctuation in the effective temperature of the microcanonical system. Keeping in mind that the total energy of this test system is roughly  $151 E_h$ , a timestep is desired with relative fluctuations smaller than about  $7 \times 10^{-8}$ , as indicated by the horizontal lines in Figure 3. Examination of the data shows that the largest timestep meeting this criterion is  $\delta t = 20$  au (0.484 fs), which is a commonly used step size for hydrogen-containing systems. Indeed, to the extent that a given screening method yields an accurate description of the water-dimer system, the inner timestep should be determined *entirely* by its fastest motions (O–H bond vibrations).

For the inner- and outer-timestep screening protocol, a pair of methods should be identified that are expected to provide stable, physically sensible trajectories, even for large systems. Because the SCF procedure is converged strictly numerically, and because screened energies and gradients are inexact (independent of any considerations related to windowing), even symplectic integrators can exhibit long-time energy drift. (This property is sometimes addressed via extended-Lagrangian MD methods,<sup>48</sup> which directly enforce energy conservation, at the expense of slightly perturbed dynamics.) Consequently, the estimated energy drift over the course of the trajectory is a relevant figure of merit.

To establish whether a given trajectory drifted appreciably, the fluctuation data were fit to a line,  $E_t = mt + E_0$ , using the Levenberg–Marquardt algorithm implemented in GNUPLLOT.<sup>49,50</sup> The estimated unsigned drift over the length of the trajectory was then computed as a fraction of the mean energy fluctuation,

$$\zeta = \frac{|\text{mlt}_{\text{max}}|}{E_1 \langle \delta \rangle} \quad (32)$$

Perfect energy conservation corresponds to vanishing  $m$ , or  $\zeta = 0$ , while  $\zeta > 1$  indicates a drift exceeding the typical fluctuations. Estimated drifts for the methods depicted in Figure 3 with  $\Delta t = 20$  au are listed in Table 1. Note that consistent screening of the gradient leads to significant reductions in the drift, as expected.

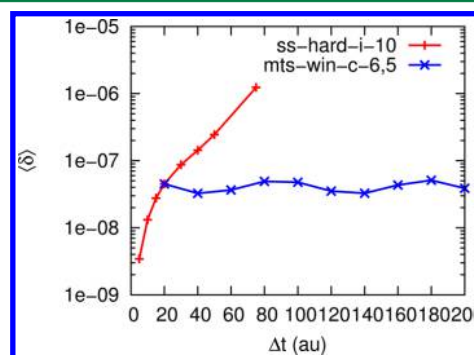
Selection of the outer-timestep screening method should be guided by the necessity of describing the system reliably, while the inner-timestep screening should be as loose as possible while maintaining similar quality. The criterion for a high-quality trajectory is that the drift in Table 1 should be no larger than 1% of the mean energy fluctuation. Consequently, the hard-i-10 method was chosen as the screening method for the outer timestep, and win-c-6,5 was chosen for the inner timestep. Because the statistical properties of these methods are similar, the corresponding screening-based MTS scheme is expected to yield stable trajectories; the effect of looser screening in the inner timestep will be explored in Section III B.

**Table 1.** Estimated Energy Drifts  $\zeta$  (eq 32) as a Percentage of Mean Energy Fluctuations for HF/3-21G Water-Dimer Trajectories Integrated with Timestep  $\Delta t = 20$  au<sup>a</sup>

screening	threshold(s)	$\zeta$ (%)
hard-i	10	0.00
	8	6.41
	6	235.43
hard-c	10	0.00
	8	0.00
	6	15.05
win-i	10, 9	0.00
	8, 7	7.73
	6, 5	205.53
	4, 3	71.01
win-c	10, 9	0.00
	8, 7	0.00
	6, 5	0.27
	4, 3	296.69

<sup>a</sup>Drift percentages are not strictly comparable between the hard and win methods; mean energy fluctuations are generally larger in the latter case, as shown in Figure 3.

To test this expectation, the HF/3-21G water dimer was simulated for 4.84 ps, using between 1 and 10 inner timesteps of fixed size  $\delta t = 20$  au. (While not strictly an MTS method, using a single inner timestep is equivalent to performing sequential momentum updates using the partitioned forces and closely tracks the single-timestep hard-i-10 trajectory.) As Figure 4 shows, fluctuations in the resulting trajectories are



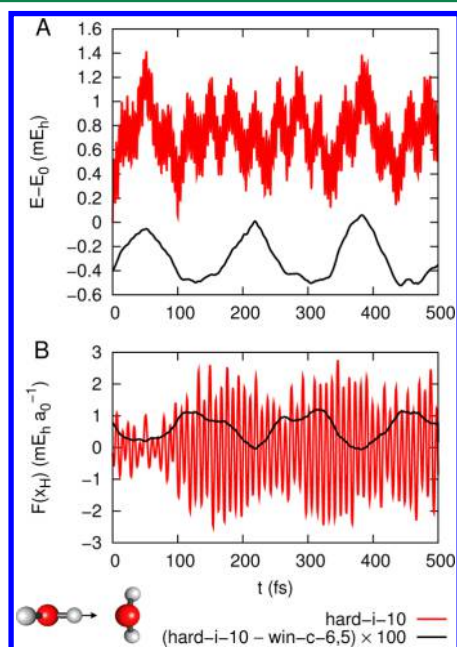
**Figure 4.** Mean energy fluctuations relative to the initial energy  $E_0$  in HF/3-21G water-dimer multiple-timestep trajectories as a function of outer timestep. The fluctuations are largely determined by the inner-timestep method.

effectively constant—the subtle oscillations are a small-system artifact—and are narrowly distributed around the value observed for the single-timestep method. This encouraging result is not as mysterious as it may at first appear: The win-c-6,5 inner-timestep method was chosen to exhibit comparable fluctuations to hard-i-10, and this behavior is reproduced in the MTS data. The drift is consistent, as well, remaining smaller than  $\zeta = 1\%$  until  $\Delta t = 140$  au and increasing to no more than a few percent ( $\zeta \approx 3.11\%$ ) at  $\Delta t = 200$  au.

The screening-based MTS scheme is now complete— $n$  inner timesteps of length  $\delta t = 20$  au are integrated using win-c-6,5 screening, followed by an outer-timestep momentum correction using hard-i-10 screening—and it has been shown to be stable. Now, this MTS scheme will be justified on physical (rather than purely numerical) grounds, on the basis of an associated time



scale separation. To this end, a single-timestep hard-i-10 trajectory was integrated with  $\Delta t = 20$  au; overall rotations were removed from the initial velocities, as were translations, so that trivial oscillations would not appear in the energies and forces. Gradient calculations were then performed on the corresponding geometries using win-c-6,5. The difference between the hard-i-10 and win-c-6,5 results is shown for the energy and for a single force component in Figure 5; this difference yields the



**Figure 5.** Phenomenological timescale separation in the energy and a force component of the HF/3-21G water dimer. Panel A shows the energy relative to the initial energy  $E_0$  for the hard-i-10 method (red or gray line) and the energy difference between hard-i-10 and the same trajectory reevaluated using win-c-6,5 (scaled up by a factor of 100, black line). Panel B makes a similar comparison for the force component associated with the Cartesian  $x$  coordinate of one of the hydrogens, as illustrated at lower left. See Section II D for detailed discussion.

effective contribution to these quantities from integrals with estimated magnitude between  $10^{-6}$  and  $10^{-10}$ , with fractional contributions from integrals as large as  $10^{-5}$ . Thus, any higher-frequency features in the hard-i-10 energy and force must be associated with integrals of larger magnitude. These results lead to the conclusion that a separation in time scale between larger and smaller integrals does indeed exist, consistent with the original motivation for this approach. Similar time scale separations were observed for the water octamer  $(\text{H}_2\text{O})_8$  treated with HF/3-21G and HF/6-31G\*\*—see Section S1 of the Supporting Information for more details—indicating that the time scale separation is not simply an artifact associated with a small system and basis set. As such, screening-based MTS is more than just an ad hoc method.

**E. Computational Cost Considerations for Ab Initio MTS.** With the results of Section II D in-hand, MTS could immediately be applied to realistic chemical systems. In doing so, however, an important distinction between classical and ab initio MTS would be neglected. Suppose that the inner timestep  $\delta t$  is appropriate for integrating the dynamics of a system with single-timestep integration methods. The cost of incrementing the system by an outer timestep  $\Delta t = n\delta t$  is then

cost of increment

$$= \begin{cases} nS, & \text{single timestep integrator} \\ nI + O, & \text{multiple timestep integrator} \end{cases} \quad (33)$$

where  $I$  is the composite cost of the energy and gradient in the inner timestep,  $O$  is the cost of the outer timestep, and  $S$  is the equivalent cost for the single-timestep method.

In classical MTS,  $S$  is simply the cost of evaluating the complete set of analytic energy and gradient functions describing the system; in the ab initio case,  $S$  is the cost of computing the energy and gradient at the same level of theory as in the outer timestep. Thus,

$$S = \begin{cases} I + O, & \text{classical MTS} \\ O, & \text{ab initio MTS} \end{cases} \quad (34)$$

Equations 33 and 34 imply that classical MTS will be cheaper than the single-timestep equivalent whenever  $I < S$ ; any nontrivial partitioning of the system is guaranteed to meet this requirement. The ab initio case is qualitatively different: MTS with  $n$  inner timesteps will only be cheaper than the single-timestep method when

$$I < \frac{n-1}{n}O \quad (35)$$

In other words, the repeated work at the outer ab initio timestep leads to an overhead, not present in classical MD, that sets a lower bound on the number of inner timesteps that will lead to computational savings. Choosing  $n$  to exceed this bound may not be practicable, due to resonance considerations<sup>51–55</sup> or to long physical time scales in a given system. The approach taken here is to dig further into the workings of the electronic-structure methodology in order to minimize this overhead. In particular, by carrying forward information from the inner timesteps, much of the repeated outer-timestep effort can be eliminated. As a result, eq 35 may be rewritten as

$$I < \frac{n-\kappa}{n}O \quad (36)$$

with  $\kappa < 1$ , subverting what appeared to be a hard inequality. The smaller the value of  $\kappa$ , the more favorable MTS will be.

The most effective means of reducing the outer-timestep cost are also the simplest: In the trivial case that the method used to screen the energies does not change—as is true when only switching gradient consistency, or for correlation-based MTS<sup>20</sup>—the inner- and outer-timestep energies will be identical. Thus, the SCF procedure can be skipped altogether, and the gradient can be computed using the inner-timestep density,  $P_i$ . Of course, the present hard-i-10/win-c-6,5 MTS scheme is not so simple; this possibility is mentioned strictly for completeness.

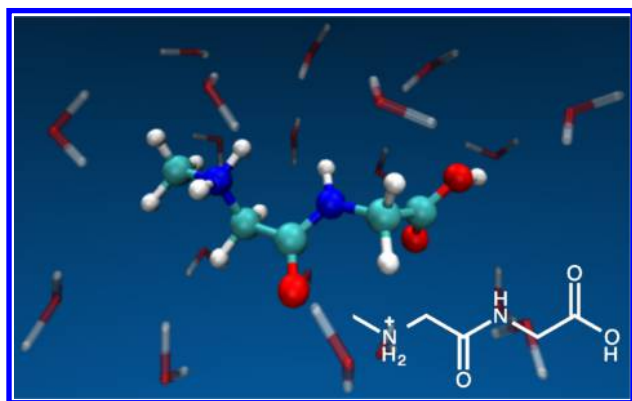
A more immediately relevant (and general) improvement involves the initial SCF guess.  $P_i$  is an approximation to the desired outer-timestep density,  $P_o$ ; it can, therefore, be used as the initial guess for  $P_o$ . Although the initial Fock matrix will be more expensive to construct than that associated with a sparse guess formed by superposition of atomic densities (Q-CHEM's default), the total number of SCF cycles at the higher level of theory will be appreciably reduced.

Additional efficiencies have been identified in the outer timestep, beyond simple improvements in the SCF guess, as described in Section S2 of the Supporting Information. These

efficiencies include reuse of information about the DFT quadrature grid<sup>56,57</sup> and a “bandpass screening” method that reduces the number of integral contractions required in the initial SCF step by considering changes in the weights of two-electron integrals. Bandpass screening is developed in a similar spirit to standard Fock-matrix increments.<sup>58,59</sup>

### III. RESULTS

With the formulation, validation, and cost analysis of the screening-based MTS scheme completed, application of the method to a challenging chemical system is now possible. Here, the MTS scheme—as implemented in a development version of Q-CHEM<sup>34</sup>—is applied to a biological model, the protonated sarcosine-glycine dipeptide (SarGlyH<sup>+</sup>, cf. Figure 6 inset)



**Figure 6.** SarGlyH<sup>+</sup> (structure inset at lower right) embedded in a 19-water cluster.

embedded in a 19-water cluster. Although this choice of model was inspired by recent experiments using gas-phase SarGlyH<sup>+</sup> as a testbed for understanding the structural and dynamical effects of peptide methylation,<sup>60,61</sup> these phenomena are not addressed in this work. Rather, a simple demonstration is made that screening-based MTS realizes significant computational savings for a reasonably large system containing 29 heavy atoms and 268 electrons.

The ground-state dynamics of SarGlyH<sup>+</sup>(H<sub>2</sub>O)<sub>19</sub> were simulated at both the Hartree–Fock and B3LYP levels of theory using the 6-31G\*\* basis set (corresponding to a total of 680 basis functions) and a serial implementation of the underlying electronic-structure methodology. The initial configuration of the system, shown in Figure 6, was obtained by building SarGlyH<sup>+</sup> in IQmol;<sup>62</sup> optimizing the structure with the MMFF94s force field;<sup>63</sup> inserting the optimized dipeptide in the HF/6-31G\*\* minimum-energy structure of (H<sub>2</sub>O)<sub>19</sub> from the Cambridge Cluster Database;<sup>45,46</sup> and reoptimizing with MMFF94s. Initial velocities were sampled from the Maxwell–Boltzmann distribution at 298 K. As before, the base inner timestep for the MTS trajectories was  $\delta t = 20$  au. Timesteps were screened using hard-i-10 (outer) and win-c-6,5 (inner) methods. All trajectories were propagated for at least 480 au (11.6 fs); for purposes of timing comparison, single-timestep simulations using hard-i-10 and win-c-6,5 were also performed.

**A. Outer-Time Step Efficiencies Are Large.** In order to assess the usefulness of the improved outer-timestep SCF guess proposed in Section II E—as well as the additional efficiencies discussed in depth in Section S2 of the Supporting Information—a series of simulations was performed in which

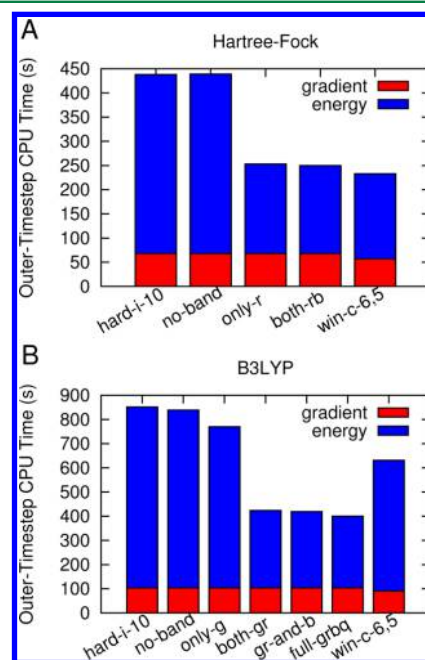
these algorithmic improvements were added hierarchically. Consequently, three types of Hartree–Fock calculation were performed:

1. no-band, a Hartree–Fock calculation with no improvements.
2. only-r, the inner-timestep density  $P_i$  is used as a guess.
3. both-rb, bandpass screening is applied, in addition to only-r.

After quadrature considerations were taken into account, there were five types of B3LYP calculation:

1. no-band, a B3LYP calculation with no improvements.
2. only-g, quadrature-grid information is reused.
3. both-gr,  $P_i$  is used as a guess, in addition to only-g.
4. gr-and-b, bandpass screening is applied, in addition to both-gr.
5. full-grbq, DFT quadrature results are reused, in addition to gr-and-b.

As Figure 7 indicates, reusing the density overwhelms all other efficiencies, reducing the cost of the outer timestep



**Figure 7.** Average CPU seconds required to compute the hard-i-10 outer-timestep energy and gradient in AIMD simulations of the SarGlyH<sup>+</sup>(H<sub>2</sub>O)<sub>19</sub> system using Hartree–Fock and B3LYP in 6-31G\*\* basis set. hard-i-10 and win-c-6,5 denote costs from single-timestep trajectories, while the other labels indicate the cost of the outer timestep in screening-based MTS trajectories exploiting efficiencies to differing degrees. See Section III A for detailed discussion.

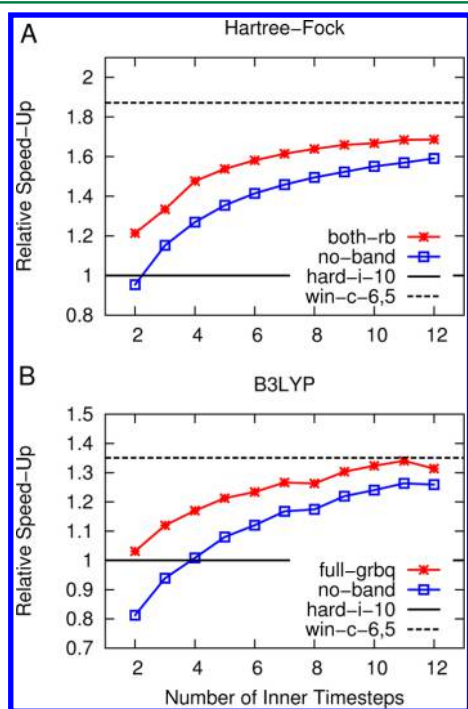
roughly by half for both Hartree–Fock and B3LYP. Additional savings accrue in B3LYP calculations when quadrature information is reused, reducing the cost by a further  $\approx 12\%$ . While bandpass screening produces more modest savings, the aggregate effect of a few seconds saved per outer timestep can be significant for long trajectories.

Figure 7 shows that an HF/both-rb outer timestep is comparable in cost to an inner timestep. Further, B3LYP/full-grbq outer timesteps are actually *cheaper*, such that the inner timesteps become the computational bottleneck. In both cases, these steep reductions stem from the fact that much of the

effort required to converge the hard-i-10 calculation has already been done in the preceding inner timestep. Therefore, the goal of addressing outer-timestep “overhead,” as articulated in Section II E, has largely been met.

**B. Minimal Multiple-Timestepping Realizes Significant Savings.** Table S1 of the Supporting Information lists the average CPU minutes required to simulate the SarGlyH<sup>+</sup>(H<sub>2</sub>O)<sub>19</sub> system for a femtosecond as a function of the number of inner timesteps,  $n$ . The ratio between single-timestep timings for hard-i-10 and win-c-6,5 provides an estimate of the maximum theoretically achievable acceleration in these calculations—roughly 87% for Hartree–Fock and 35% for B3LYP. The smaller percentage for B3LYP reflects the fact that DFT quadrature adds what amounts to a fixed cost in each and every SCF cycle (apart from the initial Fock build in the outer timestep). This penalty for DFT relative to HF should diminish with increasing system size, because the cost of quadrature grows more slowly than the cost of constructing the Coulomb and exchange matrices.

Figure 8 presents the same data graphically, clearly showing that roughly a fifth of the achievable speed-up for Hartree–



**Figure 8.** Speed-up of screening-based MTS simulations of the SarGlyH<sup>+</sup>(H<sub>2</sub>O)<sub>19</sub> system in the 6-31G\*\* basis set. Red stars indicate the fully bandpassed results (HF/both-rb or B3LYP/full-grbq, cf. Table S1 of the Supporting Information), while blue boxes are data for simulations without any use of efficiencies in the outer timestep. The solid and dashed black lines indicate the single-timestep cost of the hard-i-10 and win-c-6,5 methods, respectively. Results below the black line indicate cases where the MTS simulation is more expensive than the single-timestep method.

Fock is realized with a minimal MTS scheme ( $n = 2$ ), and more than half is realized when  $n = 4$ . (The same trend holds for B3LYP.) These results are due entirely to the overhead-trimming interventions in the outer timestep, as evidenced by naïve MTS calculations costing more than a single-timestep calculation until  $n = 3$  for HF or  $n = 4$  for B3LYP. Reusing information in the outer timestep has effectively skirted the cost

inequality that governs ab initio MTS (eq 35); the effective values of the constant  $\kappa$  in eq 36 derived from our simulations are  $\kappa_{\text{HF}} \approx 0.57$  and  $\kappa_{\text{B3LYP}} \approx 0.47$ . These constants are not necessarily transferable between systems, but future analysis will establish whether they are typical.

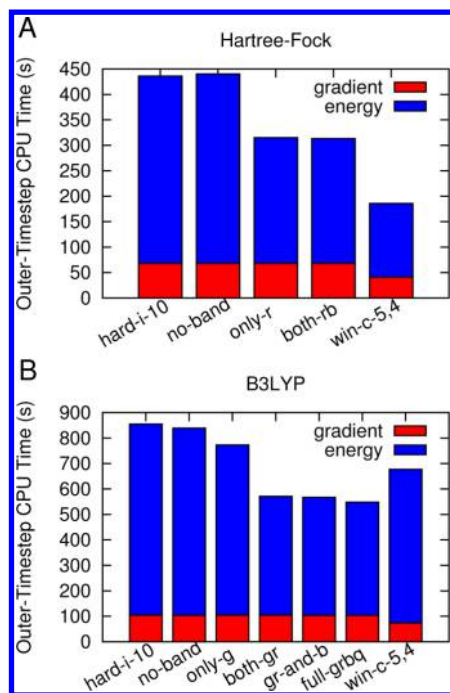
**C. Looser Inner-Time Step Screening Can Provide Further Savings.** The hard-i-10/win-c-6,5 MTS scheme was constructed to satisfy fairly stringent criteria for statistical properties of the corresponding single-timestep trajectories on a picosecond time scale. As noted in Section II D, looser screening in the inner timestep may still yield quality results to the extent that inaccuracies in the inner timestep are corrected by frequent-enough outer timesteps. Here, the degree to which loosening the threshold affects the savings achievable for SarGlyH<sup>+</sup>(H<sub>2</sub>O)<sub>19</sub> is examined.

Looser screening reduces the time required for each SCF cycle, but the errors introduced into the density can make it harder to achieve SCF convergence, such that the average number of cycles increases. For any given system, then, a regime may exist in which loosening the threshold makes the calculation more expensive. As a case in point, the SCF did not converge when a win-c-4,3 was used for SarGlyH<sup>+</sup>(H<sub>2</sub>O)<sub>19</sub>. Therefore, narrower windows were tried— $s \in \{3.12, 3.3, 3.6\}$ , corresponding roughly to  $2.5 \times 10^{-4}$  steps in integral magnitude between the hard cutoff and soft threshold. (A narrower window reduces errors associated with integral rescaling, which could restore convergence, but it also reduces robustness against blinking.) The narrow-window B3LYP calculations never converged, while the HF calculations were costlier than win-c-6,5, owing to a sharp increase in the number of SCF cycles required (as much as a factor of 6.5). The aggressive win-c-4,3 method was ultimately abandoned for the present application, and win-c-5,4 was tested instead. While both HF and B3LYP calculations converged, only the HF trajectories were cheaper than the corresponding win-c-6,5 calculations—by about 20%, as shown in Table S2 of the Supporting Information.

Another consequence of looser screening in the inner timestep is that the various efficiencies discussed in Section II E can be less useful, as shown in Figure 9. In particular, because the density from the last inner timestep is not as good of a guess for the outer-timestep density, the number of SCF cycles required is not reduced as sharply. These more modest savings—equivalent to  $\kappa_{\text{HF}} \approx 0.67$  and  $\kappa_{\text{B3LYP}} \approx 0.59$ —are compensated in the case of Hartree–Fock by the fact that the inner timestep cost is significantly reduced. As a result, win-c-5,4/hard-i-10 MTS always involves some speed-up, as shown in Figure 10A, and about half of the roughly 135% theoretically achievable speed-up is realized when  $n = 4$ . By contrast,  $n = 4$  is the minimum number of inner timesteps required for B3LYP to achieve even a modest speed-up, as shown in Figure 10B. In light of these results, a wise practice would be to compare timings for a small number of possible inner-timestep methods and single increments of the corresponding MTS protocol before committing to a specific scheme.

Finally, it is worth noting that these timing assessments are fairly conservative, due to the fact that screening-based multiple-timestepping has been evaluated with all other parameters fixed. Algorithmic improvements in the outer timestep led to a situation in which the inner timesteps are cost-dominant—an inversion of the usual relationship. Because a much looser threshold is employed when screening these inner steps, a looser SCF convergence criterion or coarser DFT





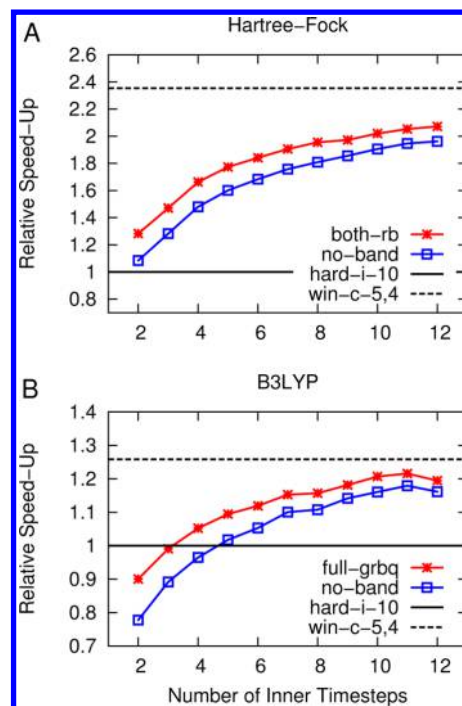
**Figure 9.** Average CPU seconds required to compute the hard-i-10 outer-timestep energy and gradient in AIMD simulations of the SarGlyH<sup>+</sup>(H<sub>2</sub>O)<sub>19</sub> system using Hartree–Fock and B3LYP in the 6-31G\*\* basis set. hard-i-10 and win-c-5,4 denote costs from single-timestep trajectories, while the other labels indicate the cost of the outer timestep in screening-based MTS trajectories exploiting efficiencies to differing degrees. See Section III A for detailed discussion.

quadrature grids could likely be employed as well, reducing the inner-timestep cost and increasing the theoretically achievable speed-up. Established AIMD extrapolation techniques (such as Fock-matrix extrapolation<sup>64,65</sup>) could also potentially be applied to reduce the cost of the inner timesteps. Finally, the shorter spatial extent of the inner-timestep thresholding suggests that linearly scaling SCF algorithms would likely become effective sooner, with respect to system size, than the analogous tight-threshold outer step. In this large-system regime, relative timings would likely improve further.

#### IV. CONCLUDING REMARKS

This work addressed the question of whether multiple-timestep methods could be applied to Hartree–Fock or DFT molecular dynamics on a firmer-than-ad-hoc basis. This question was answered affirmatively by showing that two-electron integrals of disparate magnitude are associated with forces that vary on disparate time scales. In the process, a multiple-timestep scheme was formulated using classic hard-cutoff screening techniques and a window-screening method of our own devising. Once efficiencies in the outer-timestep were identified and exploited, this multiple-timestep protocol realized computational savings with minimal or small numbers of inner timesteps ( $n = 2–4$ ).

The Schwarz-based integral estimates in this work highlight only one possibility for screening-based multiple-timestep AIMD; the same principles should apply to more sophisticated screening protocols, such as those based on multipole expansions.<sup>24</sup> In particular, screening methods that explicitly account for the asymptotic decay of interactions between  $\mu\nu$  and  $\lambda\sigma$  orbital pairs are likely to introduce additional



**Figure 10.** Speed-up of screening-based MTS simulations of the SarGlyH<sup>+</sup>(H<sub>2</sub>O)<sub>19</sub> system in the 6-31G\*\* basis set. Red stars indicate the fully bandpassed results (HF/both-rb or B3LYP/full-grbq, cf. Table S2 of the Supporting Information), while blue boxes are data for simulations without any use of efficiencies in the outer timestep. The solid and dashed black lines indicate the single-timestep cost of the hard-i-10 and win-c-5,4 methods, respectively. Results below the black line indicate cases where the MTS simulation is more expensive than the single-timestep method.

opportunities for dynamics acceleration: Because the corresponding inequalities provide tighter bounds on integral magnitudes, inner timesteps may be significantly cheaper; at the same time, the corresponding range and time scale separations may be more pronounced.

The field of multiple-timestep Born–Oppenheimer molecular dynamics is still quite young, and many possible correspondences between electronic-structure parameters and physical timescale separations remain to be tested. Varying these parameters is expected to lead to a new appreciation of hidden pathologies in commonly used AIMD protocols, just as varying the threshold highlighted blinking integrals and gradient inconsistency. For example, the SCF procedure is not iterated to exact convergence; as a result, the dynamics do not strictly preserve phase-space area, leading to drift.<sup>3,66</sup> Varying the SCF convergence criterion is likely to exacerbate this problem, and stable trajectories may only be achievable within the rigorously time-reversible approach of Niklasson et al.<sup>66</sup>

As the number of validated resources for multiple-timestep Hartree–Fock and DFT increases, combining these methods directly may become useful. Combined methods may be constructed by varying several parameters simultaneously or in hierarchical fashion, with inner timesteps subject to further Trotter splitting. The latter scenario requires careful implementation, because the MTS framework must be flexible enough to combine different protocols to an arbitrary hierarchical depth while also allowing for the use of

thermostats, which are fully compatible with the r-RESPA formalism.<sup>9,67,68</sup>

Omnibus methods of either type may be viable for outer timesteps well above the onset of resonance ( $\Delta t \gtrsim 160$  au  $\approx 4$  fs).<sup>51–55</sup> Indeed, Figure 4 indicates that screening-based multiple-timestepping already fits this description, although any resonance-related drift remains modest. Translating existing resonance-resistant MTS methods<sup>69–78</sup> into the ab initio context may open the way to efficient, accurate, genuinely long-time scale ab initio molecular dynamics.

## ■ ASSOCIATED CONTENT

### ■ Supporting Information

Evidence for time scale separation in the dynamics of the water octamer; details of additional efficiencies beyond SCF that reduce the cost of the outer timestep, including reuse of DFT quadrature-grid information and a bandpass screening method introduced in a similar spirit to traditional Fock increments; and timing data for SarGlyH<sup>+</sup>(H<sub>2</sub>O)<sub>19</sub>. This material is available free of charge via the Internet at <http://pubs.acs.org>.

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### Notes

The authors declare no competing financial interest.

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