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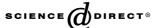
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High-resolution FTIR measurement of the v_4 band of methylene fluoride- d_2

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Abstract

A high-resolution $(0.002\,\mathrm{cm^{-1}})$ infrared absorption spectrum of methylene fluoride-d₂ $(\mathrm{CD_2F_2})$ of the lowest fundamental mode v_4 in the region from 460 to $610\,\mathrm{cm^{-1}}$ has been measured on a Bruker IFS 120-HR Fourier transform infrared spectrometer. More than 3500 transitions have been assigned in this *B*-type band centered at $521.9\,\mathrm{cm^{-1}}$. The data have been combined with upper state pure rotational measurements in a weighted least-squares fit to obtain molecular constants for the upper state resulting in an overall standard deviation of $0.00018\,\mathrm{cm^{-1}}$. Accurate value for the band origin $(521.9578036\,\mathrm{cm^{-1}})$ has been obtained and inclusion of transitions with very high J (≤ 60) and K_a (≤ 34) values has resulted in improved precision for sextic centrifugal distortion constants, in particular D_K , H_{KJ} , and H_K . © 2004 Elsevier Inc. All rights reserved.

Keywords: FTIR; v4 band; CD2F2; B-type band

1. Introduction

There has been obvious interest in the high-resolution infrared spectrum of CH₂F₂ and its isotopomer in particular CD₂F₂ [1-7] due to their importance as an efficient laser medium for generation of far-IR laser emission [1,8–10] optically pumped by CO₂ lasers. Prior studies of the spectrum of CD₂F₂ include the first report of the microwave spectrum by Hirota et al. [11]. Subsequently Hirota and Sahara [12] reinvestigated the microwave spectrum in greater detail up to J = 19 and determined rotational constants and quartic centrifugal distortion constants of the ground state and a few excited vibrational states. The millimeter/sub-millimeter wave spectra of the ground state and the v_4 (v = 1) state has been studied by Deo and Kawaguchi [5] who obtained higher order centrifugal distortion constants by including higher J and K_a transitions.

CD₂F₂ is a near-prolate asymmetric top ($\kappa = -0.89$) and the vibrational assignments ($4a_1 + a_2 + 2b_1 + 2b_2$) of the nine fundamental modes under C_{2V} point group has been reasonably well established from a normal coordinate analysis [13]. The only IR inactive a_2 mode (v_5) does show up as a weak feature in the spectrum and its appearance has been attributed to a possible Coriolis interaction with the neighboring fundamental v_7 level. CD stretching fundamentals (v_1, v_6) appear in the 2200 cm⁻¹ region and its high-resolution study was reported recently [14]. Analysis of the 9–10 µm region has been limited to the work by Goh et al. [6,7].

In the 9–10 µm region, CD_2F_2 has a very crowded and complex spectrum due to the fact that four (v_3, v_5, v_7) and v_9 out of its nine fundamental modes lie in this region. Analysis of the spectrum in this region is complicated due to strong Coriolis interactions among the various vibrational modes as pointed out by Hirota and Sahara [12] in their microwave study and Goh et al. [6,7] in FTIR studies. In addition to the four fundamental modes there is the possibility of complications from the $2v_4$ level and hot band transitions such as $v_9 + v_4 - v_4$,

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 $v_3 + v_4 - v_4$, etc. This is to be expected because the v_4 (v = 1) state is low-lying ($v_0 \sim 521 \, \mathrm{cm}^{-1}$) and has a considerable population of $\sim 7\%$ at room temperature.

Thus a comprehensive study of the v_4 band at high-resolution should be helpful in the analysis of the hot bands in the strongly perturbed 9–10 μ m region.

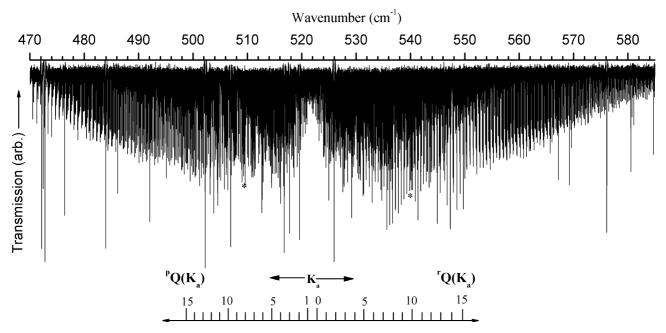


Fig. 1. Overview of the FTIR spectrum of the v_4 Band of CD_2F_2 . Pressure = 80 mTorr, path length = 18 m. The asterisks indicate the sub-bands that are shown under high-resolution in Fig. 2.

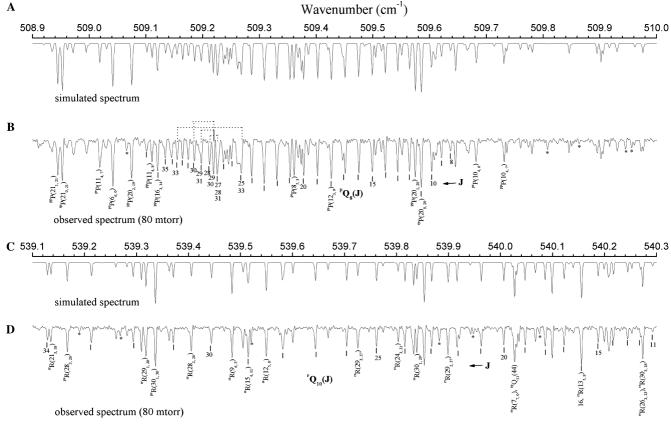


Fig. 2. (A and B) Simulated and observed spectra in the ${}^{p}Q_{8}$ -branch region and (C and D) simulated and observed spectra in the ${}^{r}Q_{10}$ -branch region. The asterisks indicate transitions due to hot bands or impurities.

Table 1 Typical ro-vibrational transitions (in cm $^{-1}$) of the ν_4 band of CD_2F_2 corresponding to the regions shown in Figs. 2 and 3

corresponding to the regions shown in Figs. 2 and 3							
J'	K'_a	K_c'	J''	K_a''	K_c''	Obs. freq	O – C ^a
41	7	35	41	8	34	508.93437	58*
20	0	20	21	1	21	508.94466	-38*
34	2	33	34	3	32	508.94466	50*
20	1	20	21	0	21	508.95253	-05
34	1	33	34	2	32	508.96070	-49*
42	5	38	42	6	37	508.97229	23
40	7	34	40	8	33	508.99539	86*
39	9	30	40	8	33	508.99910	00^{*}
10	3	8	11	4	7	509.01859	42
5	5	1	6	6	0	509.04161	07*
39	7	33	39	8	32	509.04161	70*
38	7	32	38	8	31	509.07465	-98*
19	1	18	20	2	19	509.07465	22
37	7	31	37	8	30	509.10108	-16
10	3	7	11	4	8	509.11047	28
36	7	30	36	8	29	509.12054	45*
15	2	13	16	3	14	509.12054	-46*
35	7	29	35	8	28	509.13392	-30
34	7	28	34	8	27	509.14506	-36
33	7	27	33	8	26	509.15496	-16
32	7	26	32	8	25	509.16439	-08
31	7	25	31	8	24	509.17444	09
30	7	24	30	8	23	509.18533	-01
31	1	31	31	2	30	509.19030	-06*
31	0	31	31	1	30	509.19751	73*
29	7	23	29	8	22	509.19751	-32*
28	7	22	28	8	21	509.21212	11
29	7	22	29	8	21	509.21906	-13*
30	7	23	30	8	22	509.21906	38*
27	7	21	27	8	20	509.22552	-44*
28	7	21	28	8	20	509.22552	08*
31	7	24	31	8	23	509.22552	14*
27	7	20	27	8	19	509.23632	10
32	7	25	32	8	24	509.24073	-43
26	7	20	26	8	19	509.24545	-17
26	7	19	26	8	18	509.25058	-01
36	3	34	36	4	33	509.26001	01*
20	3	18	21	2	19	509.26116	-92*
25	7	19	25	8	18	509.26416	−71*
33	7	26	33	8	25	509.26797	-31*
25	7	18	25	8	17	509.26797	18*
24	7	17	24	8	16	509.28615	-06*
34	7	27	34	8	26	509.30825	-25*
23	7	16	23	8	15	509.30825	-09*
22	7	15	22	8	14	509.33059	-17
21	7 4	14 4	21	8 5	13	509.35396	-16
7			8	5	3	509.36107	06 -09*
37 35	6 7	31 28	38 35	8	34 27	509.36834 509.36834	32*
19	2	18	20	1	19	509.37337	-05
20	7	13	20	8	12	509.37810	-03 -01
41	5	37	41	6	36	509.38612	−01 −78*
36	2	34	36	3	33	509.38612	-78 55*
19	7	12	19	8	11	509.40238	-06
33	2	32	33	3	31	509.42053	-00 -01
11	2	10	12	3	9	509.42662	-01 67*
18	7	11	18	8	10	509.42662	-28*
36	7	29	36	8	28	509.44664	-28 -72*
33	1	32	33	2	31	509.44664	-72 69*
17	7	10	17	8	9	509.45118	-04
16	7	9	16	8	8	509.47526	02
15	7	8	15	8	7	509.49885	08
	•	5		9	,		

Table 1 (continued)

rable i	(conti	nuea)					
J'	K'_a	K'_c	J''	K_a''	K_c''	Obs. freq	O – C ^a
14	2	12	15	3	13	509.50457	-16
42	6	37	42	7	36	509.51460	-06
14	7	7	14	8	6	509.52160	-01
13	7	6	13	8	5	509.54357	-07
37	7	30	37	8	29	509.55141	14
12	7	5	12	8	4	509.56448	-20
19	0	19	20 20	1	20	509.57459	-09 -58*
19 11	1 7	19 4	20 11	0 8	20 3	509.58527 509.58527	-58° 65*
10	7	4	10	8	3	509.60362	30*
39	3	36	39	4	35	509.60808	-46
22	4	19	23	3	20	509.61167	48
9	7	2	9	8	1	509.62070	02
8	7	2	8	8	1	509.63678	18
18	1	17	19	2	18	509.64546	06
30	0	30	30	1	29	509.66867	-30*
9	3	7	10	4	6	509.68180	39
38	7	31	38	8	30	509.68317	00*
9	3	6	10	4	7	509.73132	07
35 41	3 6	33 36	35 41	4 7	32 35	509.73621 509.76032	-16 13
40	5	36	40	6	35	509.78043	02
39	7	32	39	8	31	509.84575	04
32	2	31	32	3	30	509.89527	-14
13	2	11	14	3	12	509.90191	-16
35	2	33	35	3	32	509.90555	05
32	1	31	32	2	30	509.93078	-25
10	1	10	11	2	9	509.96070	-03*
40	6	35	40	7	34	509.97650	19
34	11	24	34	10	25	539.12900	18
22	4	19	21	3	18	539.13516	10
29	2	27	28	3	26	539.16691	-05
33	11	22	33	10	23	539.21322	12
32	4	28	31	5	27	539.26041	06
17	3	14	16	2	15	539.28156	-14
32	11	22	32	10	23	539.29352	-13
30	1	29	29	2	28	539.30943	09
30	2	29	29	1 1	28	539.31831 539.33646	03
31 23	4	31 20	30 22	3	30 19	539.33646	03 11
31	11	20	31	10	21	539.30204	02
29	3	27	28	2	26	539.40529	-06
30	11	20	30	10	21	539.44389	02
10	7	4	9	6	3	539.48362	-01
30	3	27	29	4	26	539.50465	-40
16	5	12	15	4	11	539.51456	-65*
29	11	19	29	10	20	539.51456	76*
13	6	7	12	5	8	539.54966	-16
28	11	18	28	10	19	539.58046	02*
24 16	4 5	21 11	23 15	3 4	20 12	539.58266 539.60071	-14*
27	11	17	27	10	18	539.60071	04 10
18	4	14	17	3	15	539.66836	-05
26	11	16	26	10	17	539.70423	02
30	2	28	29	3	27	539.72550	05
25	11	14	25	10	15	539.76158	02
25	4	22	24	3	21	539.80303	-18
24	11	14	24	10	15	539.81599	00
31	1	30	30	2	29	539.83286	02
31	2	30	30	1	29	539.83890	-07
32	0	32	31	1	31	539.85325	01*
32 23	1 11	32 13	31 23	0 10	31 14	539.85325 539.86760	-07^{*} 00
23	11	13	23	10	14	JJ7.8U/0U	υU

Table 1 (continued)

•	J'	K'_a	K'_c	J''	K_a''	K_c''	Obs. freq	O – Ca
	30	3	28	29	2	27	539.89886	04
	22	11	12	22	10	13	539.91641	-05
	36	5	31	35	6	30	539.94187	-04
	21	11	10	21	10	11	539.96266	00
	20	11	9	20	10	10	540.00633	07
	44	12	33	44	11	34	540.02223	-55*
	8	8	1	7	7	0	540.02699	-02
	26	4	23	25	3	22	540.03131	07
	44 19	7 11	37 8	43 19	8 10	36 9	540.03519 540.04726	-26 -09
	17	5	13	16	4	12	540.04720	18
	58	2	56	58	1	57	540.07447	-44
	18	11	8	18	10	9	540.08589	-07
	11	7	4	10	6	5	540.09922	-07
	17	11	7	17	10	8	540.12223	03
	40	6	34	39	7	33	540.13231	32
	43	12	31	43	11	32	540.13940	-19*
	14	6	9	13	5	8	540.15544	57*
	16	11	6	16	10	7	540.15544	-65*
	15	11	5	15	10	6	540.18771	01
	33 17	4 5	29 12	32 16	5 4	28 13	540.19931 540.20754	-05 17
	31	3	28	30	4	27	540.20734	03
	14	11	4	14	10	5	540.21699	-10
	13	11	3	13	10	4	540.24437	09
	42	12	31	42	11	32	540.25260	08
	12	11	1	12	10	2	540.26911	-25
	27	4	24	26	3	23	540.27356	-51*
	31	2	29	30	3	28	540.27356	17*
	11	11	1	11	10	2	540.29222	-12
	34	22	13	34	21	14	560.03900	03
	23	14	10	22	13	9	560.07156	-01
	36	10	27	35	9	26	560.08723	19
	33	22	11	33	21	12	560.10237	15
	20	15	6	19	14	5	560.15125	-01
	40	9	32	39	8	31	560.16245	42*
	17	16	1	16	15	2	560.22228	10*
	31 33	22 11	9 22	31 32	21 10	10 23	560.22228 560.29902	-56*
	29	22	8	29	21	9	560.33549	06 -21
	30	12	19	29	11	18	560.44727	-05
	26	22	5	26	21	6	560.49077	22
	25	22	4	25	21	5	560.53837	02
	27	13	15	26	12	14	560.56394	-06
	24	22	3	24	21	4	560.58442	16
	41	9	33	40	8	32	560.62343	34*
	37	10	28	36	9	27	560.62665	17
	37	10	27	36	9	28	560.62940	-15
	24 50	14 23	11 28	23 50	13 22	10 29	560.66227 560.71035	-09 -21
	21	15	28 7	20	14	6	560.74870	-21 07
	18	16	3	17	15	2	560.82570	03
	34	11	24	33	10	23	560.85831	00
	31	12	19	30	11	20	561.01916	02
	42	9	34	41	8	33	561.06206	-10
	28	13	16	27	12	15	561.14497	-07
	38	10	29	37	9	28	561.16074	06
	38	10	28	37	9	29	561.16564	01
	25	14	12	24	13	11	561.25095	-04 07
	40 43	6 23	34 20	39 43	5 22	35 21	561.30965 561.33118	07 -39
	22	23 15	20 7	21	14	8	561.33118	-39 07
	42	9	33	41	8	34	561.35305	09

Table 1 (continued)

J' K'a K'c J'' K'a K'c Obs. freq O - Ca	Γable 1	Table 1 (continued)							
19	J'	K'_a	K'_c	J''	K_a''	K_c''	Obs. freq	O – C ^a	
19	44	8	36	43	7	37	561.40336	17	
43 9 35 42 8 34 561.47482 21 32 12 21 31 11 20 561.58817 02 39 10 30 38 9 29 561.68928 07 39 10 29 38 9 30 561.69714 07 29 13 17 28 12 16 561.78528 -17* 37 23 15 37 22 16 561.78528 -05* 43 7 36 42 6 37 561.879158 -05* 43 7 36 42 6 37 561.85490 -25* 43 9 34 42 8 35 561.85490 -25* 43 9 34 42 8 35 561.8648 -01 49 22 28 48 21 27 590.0842 -01	35	11	24	34	10	25	561.41430	04	
32 12 21 31 11 20 561.58817 02 39 23 17 39 22 18 561.64193 -03 39 10 30 38 9 29 561.68928 07 39 10 29 38 9 30 561.78356 -10 31 2 29 30 1 30 561.78528 -17* 37 23 15 37 22 16 561.78528 -05 43 7 36 42 6 37 561.79158 -05 26 14 13 25 13 12 561.83752 11 44 9 36 43 8 35 561.83790 -25 43 9 34 42 8 35 561.83490 -25 43 9 34 25 10 590.18471 -0	19	16	4	18	15	3	561.42724	04	
39 23 17 39 22 18 561.64193 -03 39 10 30 38 9 29 561.68928 07 39 10 29 38 9 30 561.69714 07 29 13 17 28 12 16 561.78528 -17* 37 23 15 37 22 16 561.78528 -0* 43 7 36 42 6 37 561.79158 -05 26 14 13 25 13 12 561.85490 -25 43 9 34 42 8 35 561.88415 06 23 15 9 22 14 8 561.96648 -01 49 22 28 48 21 27 590.01847 17* 28 28 1 27 27 0 590.02625 10	43	9	35	42	8	34	561.47482	21	
39	32	12	21	31	11	20	561.58817	02	
39 10 29 38 9 30 561.69714 07 29 13 17 28 12 16 561.78526 -10 31 2 29 30 1 30 561.78528 -17* 37 23 15 37 22 16 561.78528 -05 43 7 36 42 6 37 561.8752 11 44 9 36 43 8 35 561.88490 -25 43 9 34 42 8 35 561.86489 -25 43 9 34 42 8 35 561.86489 -25 43 9 34 42 8 35 561.86489 -25 43 9 34 42 8 31 27 590.01847 17* 28 28 1 27 27 0 590.02625 10 <td>39</td> <td>23</td> <td>17</td> <td>39</td> <td>22</td> <td>18</td> <td>561.64193</td> <td>-03</td>	39	23	17	39	22	18	561.64193	-03	
39 10 29 38 9 30 561.69714 07 29 13 17 28 12 16 561.78528 -10* 31 2 29 30 1 30 561.78528 -17* 37 23 15 37 22 16 561.78528 -05* 43 7 36 42 6 37 561.8752 11 44 9 36 43 8 35 561.85490 -25 43 9 34 42 8 35 561.86489 -25 43 9 34 42 8 35 561.86489 -25 43 9 34 42 8 35 561.86489 -25 43 9 34 42 8 31 26 9 34 25 10 39 11 36 26 9 34 25 <	39	10	30	38	9	29	561.68928	07	
29 13 17 28 12 16 561.78528 -10* 31 2 29 30 1 30 561.78528 -17* 37 23 15 37 22 16 561.78528 02* 43 7 36 42 6 37 561.78528 -05 26 14 13 25 13 12 561.83490 -25 43 9 34 42 8 35 561.88415 06 23 15 9 22 14 8 561.96648 -01 49 22 28 48 21 27 590.01847 17* 28 28 1 27 27 0 590.2625 10 35 26 9 34 25 10 590.10016 -11* 42 24 19 41 23 18 590.10016 -11* <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td>561.69714</td> <td>07</td>							561.69714	07	
31 2 29 30 1 30 561.78528 -17* 37 23 15 37 22 16 561.78528 02* 43 7 36 42 6 37 561.79158 -05 26 14 13 25 13 12 561.83490 -25 44 9 36 43 8 35 561.88415 06 23 15 9 22 14 8 561.93715 -05 36 11 26 35 10 25 561.96648 -01 49 22 28 48 21 27 590.01847 17* 28 28 1 27 27 0 590.2625 10 35 26 9 34 25 10 590.10433 -33 46 23 23 45 22 24 590.20433 -33			17	28			561.72356	-10	
37 23 15 37 22 16 561.78528 02* 43 7 36 42 6 37 561.79158 -05 26 14 13 25 13 12 561.83752 11 44 9 36 43 8 35 561.86415 06 23 15 9 22 14 8 561.93715 -05 36 11 26 35 10 25 561.96648 -01 49 22 28 48 21 27 590.01847 17* 28 28 1 27 27 0 590.02625 10 35 26 9 34 25 10 590.10016 -11* 53 21 33 52 20 32 590.20433 -33 46 23 23 45 22 24 590.34949 0		2	29	30	1	30	561.78528	-17*	
43 7 36 42 6 37 561.79158 -05 26 14 13 25 13 12 561.83752 11 44 9 36 43 8 35 561.85490 -25 43 9 34 42 8 35 561.864815 06 23 15 9 22 14 8 561.93715 -05 36 11 26 35 10 25 561.96648 -01 49 22 28 48 21 27 590.01847 17* 28 28 1 27 27 0 590.02625 10 35 26 9 34 25 10 590.10016 14* 42 24 19 41 23 18 590.10016 -11* 53 21 33 52 20 32 590.20433 -33									
26			36	42	6	37	561.79158	-05	
44 9 36 43 8 35 561.85490 -25 43 9 34 42 8 35 561.88415 06 23 15 9 22 14 8 561.93715 -05 36 11 26 35 10 25 561.96648 -01 49 22 28 48 21 27 590.01847 17* 28 28 1 27 27 0 590.02625 10 35 26 9 34 25 10 590.10016 -11* 42 24 19 41 23 18 590.10016 -11* 53 21 33 52 20 32 590.20433 -33 46 23 23 45 22 24 590.34920 -30 32 27 6 31 26 590.36051 09 <		14					561.83752		
43 9 34 42 8 35 561.88415 06 23 15 9 22 14 8 561.93715 -05 36 11 26 35 10 25 561.96648 -01 49 22 28 48 21 27 590.01847 17* 28 28 1 27 27 0 590.02625 10 35 26 9 34 25 10 590.10016 -11* 53 21 33 52 20 32 590.20433 -33 46 23 23 45 22 24 590.34920 -30 32 27 6 31 26 5 590.34920 -30 32 27 6 31 26 5 590.34920 -30 32 27 6 31 26 5 590.34920 -30									
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In the present work, the FTIR spectrum of the v_4 band has been recorded under high-resolution and analyzed in order to obtain the band origin and other

 $^{^{}a}$ O – C = (observed – calculated) \times 10⁵ cm⁻¹. * The symbol indicates lines not included in the fit.

molecular parameters accurately. Also since transitions with high J (≤ 60) and K_a (≤ 34) have been observed the precision of the J and K_a dependent parameters has greatly improved compared to previous reports [5].

2. Experiment

Fourier transform absorption spectra of CD₂F₂ between 470 and 610 cm⁻¹ were measured at an apodized spectral resolution of 0.002 cm⁻¹ on a Bruker IFS 120HR FTIR spectrometer using a globar infrared source, KBr beamsplitter and an appropriate optical filter. Spectra were recorded with two detectors, liquid helium cooled Ge-bolometer and liquid nitrogen cooled HgCdTe (MCT) for the region 470-550 and 520-610 cm⁻¹, respectively. The spectra were obtained with a multiple pass cell set at a path length of 18 m and at sample pressures of 0.08, 0.20, 0.66, and 1.8 Torr in order to measure low as well as high J and K_a transitions. For each set, about 180 scans were co-added for a total integration time of 12 h. The reference spectra of the empty cell at lower resolution was recorded for a duration of 5h, which was, after zero fillings, used to reduce water vapour lines and to get good base line. Residual water vapour lines were used for calibration. The accuracy of the measurement is expected to be better than 0.00025 cm⁻¹. The observed overall highresolution features are shown in Fig. 1, in a condensed scale.

3. Analysis, results, and discussions

 $\mathrm{CD}_2\mathrm{F}_2$ is a near-prolate asymmetric rotor with $\kappa = -0.89$ and belongs to the C_{2V} point group. The v_4 mode ascribes to the CF_2 scissoring motion [13] which produces a dipole moment change parallel to the intermediate axis of moment of inertia, thus giving rise to a B-type band. The strong transitions are determined by the selection rules

$$\Delta K_a = \pm 1$$
, $\Delta K_c = \pm 1$, with $\Delta J = 0, \pm 1$

for a *B*-type band the allowed transitions are of the type $ee \leftrightarrow oo$ and $eo \leftrightarrow oe$.

Transitions originating from oe or eo levels have a nuclear spin statistical weight of 21 compared to 15 for transitions from other levels. Transitions from corresponding levels therefore show an intensity alternation with a ratio 21:15. This feature is useful in checking line assignments. The v_4 mode being the lowest vibrational state above the ground state is well isolated from other vibrational levels, hence it is free from perturbations. However, the observed band may be complicated due to the presence of overlapping hot bands and the presence

of impurities such as CHDF₂ ($\sim 0.5\%$) and 13 CD₂F₂ (natural abundance $\sim 1\%$).

Since CD_2F_2 is a near-prolate asymmetric top with an asymmetric parameter, $\kappa = -0.89$, the *B*-type band is expected to bear a close resemblance to the perpendicular band of a symmetric top molecule. Spectra are

Table 2A Molecular parameters (in cm $^{-1}$) for the v_4 (v=1) band of ${\rm CD_2F_2}^a$ S-reduction; I^r representation

Parameter	Ground state ^b	$v_4 \ (v=1)$ state	
		Present	Previous ^b
v_0	_	521.9578036(39)	_
A	1.15897681(1)	1.16501523(1)	1.16501530(3)
B	0.341618053(5)	0.340867516(5)	0.340867505(8)
C	0.294597299(5)	0.293470927(5)	0.293470917(8)
$D_J imes 10^6$	0.3111105(52)	0.3076311(57)	0.307607(15)
$D_{JK} imes 10^6$	-0.986060(22)	-0.971075(23)	-0.971171(54)
$D_K imes 10^5$	0.7809106(55)	0.8000726(97)	0.80029(10)
$\delta_J imes 10^7$	-0.667962(11)	-0.664413(14)	-0.664413(20)
$\delta_K imes 10^8$	-0.380616(37)	-0.404174(50)	-0.404108(74)
$H_J \times 10^{12}$	0.3329(17)	0.3189(19)	0.3069(69)
$H_{JK} imes 10^{12}$	0.3209(87)	0.371(12)	0.373(20)
$H_{KJ} imes 10^{10}$	-0.46809(85)	-0.47843(48)	-0.4905(60)
$H_K \times 10^9$	0.18331(14)	0.191236(84)	0.209(11)
$h_J \times 10^{12}$	0.17696(48)	0.17100(65)	0.17108(98)
$h_{JK} \times 10^{13}$	0.2878(30)	0.3067(40)	0.3045(61)
$h_K \times 10^{14}$	0.3856(61)	0.4232(93)	0.415(13)
Standard deviation		$0.000176\mathrm{cm^{-1}}$	

 $^{^{\}rm a}$ Standard deviations given in parentheses are 1σ in terms of the last digit quoted.

Table 2B Molecular parameters (in cm⁻¹) for the v_4 (v=1) band of CD₂F₂^a A-reduction; I^r representation

Parameter	Ground state ^b	$v_4 \ (v=1)$ state	
		Present	Previous ^b
v_0	_	521.9578036(39)	_
A	1.15897676(1)	1.16501519(1)	1.16501525(3)
B	0.341619157(5)	0.340868687(5)	0.340868678(8)
C	0.294596222(5)	0.293469784(5)	0.293469775(8)
$\Delta_J \times 10^6$	0.3187199(53)	0.3157099(58)	0.315687(15)
$\Delta_{J\!K} imes 10^5$	-0.1031775(23)	-0.1019614(23)	-0.1019692(53)
$\Delta_K \times 10^5$	0.7847184(55)	0.8041177(96)	0.80433(10)
$d_J \times 10^7$	0.667974(10)	0.664416(13)	0.664421(19)
$d_K \times 10^6$	0.544954(50)	0.578834(66)	0.57871(10)
$arPhi_J imes 10^{12}$	0.3895(18)	0.3783(20)	0.3668(71)
$\Phi_{JK} imes 10^{11}$	0.2185(41)	0.2401(57)	0.2386(81)
$arPhi_{KJ} imes 10^{10}$	-0.5406(14)	-0.5571(19)	-0.5687(60)
$\Phi_K imes 10^9$	0.18869(16)	0.19707(15)	0.215(10)
$\varphi_J imes 10^{12}$	0.18107(52)	0.17505(68)	0.1754(10)
$\varphi_{J\!K} imes 10^{12}$	0.411(44)	0.498(59)	0.439(89)
$\varphi_K imes 10^{10}$	0.5161(86)	0.555(13)	0.547(18)
Standard deviation		$0.000176\mathrm{cm^{-1}}$	

^a Standard deviations given in parentheses are 1σ in terms of the last digit quoted.

^b Values taken from [5].

^b Values taken from [5].

therefore characterized by prominent Q branches with a separation of $\sim 2(A-\bar{B})$ where $\bar{B}=(B+C)/2$. In this case, $2(A - \bar{B})$ is about 1.7 cm⁻¹. Since for CD₂F₂ $A > \bar{B}$, the separation of successive Q branches $2(A - \bar{B})$ is much greater than the separation of the P and R lines in each sub-band given by $\sim 2\bar{B}$ ($\sim 0.6 \, \mathrm{cm}^{-1}$). Thus it is seen that Q branches for different K_a values spread out (Fig. 1) whereas within a K_a sub-band lines with different J cluster together. Two such Q branches, one in the Pbranch and another in the R-branch region are shown in Fig. 2. Since the rotational constants for the ground as well as for the v_4 state were available from recent pure rotational spectra in the millimeter and sub-millimeter wave [5] region, spectra were simulated in the region of the predicted band center near 521.7 cm⁻¹ from lowresolution studies [13]. The value of the band origin was then adjusted until the simulated spectra resembled the characteristic *Q*-branch features in the observed spectra. The assignment of Q-branch features with low K_a values was complicated due to large asymmetry splittings. Subbands with $K_a > 5$ give rise to characteristic features free from asymmetry splittings for moderate J values and could be more readily identified. Assignment of transitions was done via a bootstrap procedure, initially using the method of combination differences and later by predicted line positions from the refined molecular constants. Data were analyzed with a least-squares fit program using Watsons' A-reduced Hamiltonian [15] in the I^r representation. The assignments were gradually extended to include higher J and K_a transitions. In this manner a total of about 3500 ro-vibrational transitions have been assigned. In the final fitting, about 3100 FTIR data were combined with microwave [12] and millimeter wave [5] data of the upper state with appropriate weighting factors. Lines with deviation $> 0.0007 \,\mathrm{cm}^{-1}$ were excluded from the fit and lines not split by asymmetry were included only once. Standard deviations of the independent fits of various data sets (σ_i) were used as weights $(1/\sigma_i^2)$ in the combined weighted least-squares analysis. The overall standard deviation of the FTIR

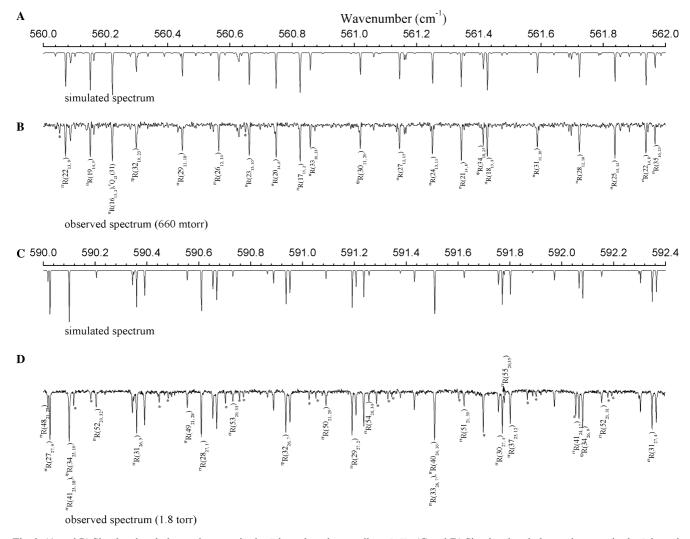


Fig. 3. (A and B) Simulated and observed spectra in the R-branch region, medium J, K_a . (C and D) Simulated and observed spectra in the R-branch region, high J, K_a . The asterisks indicate transitions due to hot bands or impurities.

data was found to be 0.00018 cm⁻¹ which is well within the experimental uncertainty. Since the FTIR data set is very large only a few ro-vibrational transitions have been listed in Table 1 as a representative of the quality of the fit. Complete set of data can be obtained on request from the authors.

The S-reduction in the I^r representation is the obvious choice for $\mathrm{CD}_2\mathrm{F}_2$ ($\kappa=-0.89$) but for comparison with previous reported results we have carried out the calculation in the S-reduction as well as in the A-reduction. The molecular parameters obtained are listed in Tables 2A and 2B along with previously reported values for comparison. It can be seen that the precision of parameters in particular D_k , H_{KJ} , H_K , etc. has improved considerably. This is because higher K_a transitions were included in the present study, which helped in improving the K dependent parameters.

As a final confirmation of the assignments computer simulation of the spectra was carried out with the molecular parameters reported in Table 2A. The simulated spectra are compared with the observed spectra in Figs. 2 and 3. Fig. 2 shows the simulated spectra along with the experimentally observed spectra of ${}^{p}Q_{8}$ and ${}^{r}Q_{10}$ in the P- and R-branch region, respectively. Fig. 3 shows Rbranch transitions in the region of medium and higher K_a transitions. It can be seen that there is excellent agreement between the observed and simulated spectra with respect to line positions as well as relative intensities of spectral lines. A few extra weak features appearing in the observed spectra could be attributed to either hot band transitions $(2v_4 \leftarrow v_4)$ or due to the presence of CHDF₂ ($\sim 0.5\%$) as impurity and $^{13}\text{CD}_2\text{F}_2$ which has about 1% natural abundance in the sample. Hot band transition $2v_4 \leftarrow v_4$ is expected as v_4 (v=1)state has about 7% population at room temperature.

In conclusion, more than 3500 lines in the high-resolution FTIR absorption spectrum of the v_4 band of CD_2F_2 were assigned. Out of these 3100 lines were combined with all other available data to obtain a set of very precisely determined molecular parameters for the

 v_4 (v=1) state including the band origin. The availability of precise parameters for the v_4 state may serve as useful input in analyzing the spectrum in the 9–10 µm region where hot band transitions involving v_4 could occur.

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