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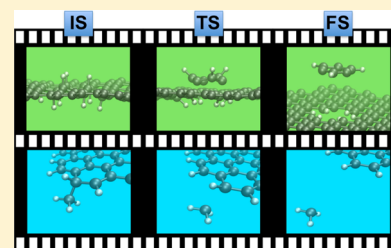
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# Etching of Graphene in a Hydrogen-rich Atmosphere toward the Formation of Hydrocarbons in Circumstellar Clouds

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**ABSTRACT:** We describe a mechanism that explains the formation of hydrocarbons and hydrocarblys from hydrogenated graphene/graphite; hard C–C bonds are weakened and broken by the synergistic effect of chemisorbed hydrogen and high temperature vibrations. Total energies, optimized structures, and transition states are obtained from density functional theory simulations. These values have been used to determine the Boltzman probability for a thermal fluctuation to overcome the kinetic barriers, yielding the time scale for an event to occur. This mechanism can be used to rationalize the possible routes for the creation of small hydrocarbons and hydrocarblys from etched graphene/graphite in stellar regions.



## INTRODUCTION

Recent technological advances providing improved sensitivity in radio-astronomical receivers, and through the usage of large (sub)millimeter radio telescopes, have permitted researchers to find numerous new chemical species in outer space. Several chemical processes lead to the formation of these molecules in the gas-phase in space, e.g., ion–neutral, neutral–neutral, and radical–neutral reactions. In addition, reactions at the surface of the grains can also enhance the gas-phase abundance of these species when the temperature of dust grains increases above a critical value. Heating processes of dust grains can occur due to shocks, UV photon absorption from external sources, or the proximity of a newly formed star. Nevertheless, the presence of complex molecules in photodissociation regions requires a continuous production mechanism that so far is not fully understood, although understanding these chemical processes is crucial to study the chemical evolution of the interstellar gas in galaxies. Therefore, there is an increasing interest in detailed microscopic calculations leading to models that can explain the relative abundance of these molecules.

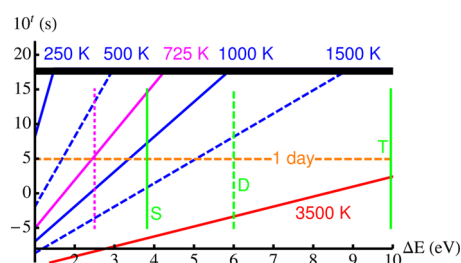
Carbonaceous particles are usually terminated on graphene/graphite-like layers or amorphous carbon, which may transform under heating to other stable forms of carbon. Clouds rich on silicates or carbonaceous dust grains are covered with ices ( $\text{H}_2\text{O}$ ,  $\text{NH}_3$ ,  $\text{CH}_4$ ,  $\text{CH}_3\text{OH}$ ,  $\text{CO}$ , ...) providing support for the formation of different species. A paradigmatic example is the formation of the hydrogen molecule from atomic hydrogen; due to the presence of ices it is produced even at temperatures as low as 10 K, as demonstrated by laboratory experiments on circumstellar grain analogs and by density functional theory (DFT) simulations.<sup>1</sup> On the other hand, in low-density clouds (diffuse) these particles are known to be *clean* of contaminants like ice, and a similar mechanism is lacking since the thermally activated process of molecular hydrogen formation would require temperatures  $T > 250$  K to achieve a reasonable yield in times ranging between a few years to some thousands of years.<sup>2,3</sup>

The opposite problem is the effect of atomic hydrogen on carbonaceous particles.<sup>5</sup> It has recently attracted attention because of its potential to explain the formation of hydrocarbon molecules and all sort of radicals in space.<sup>6</sup> These authors have described the high-temperature etching of SiC grains in a hydrogen-rich atmosphere in the laboratory. The main experimental observation is the formation of holes on a graphene surface. This is surprising since the typical energy stored in a C–C bond on graphene is  $\geq 4$  eV, and such a process would only happen on a time-scale of several years in pristine graphene at the temperature of the experiment (1000 K), cf. Figure 1. Indeed, graphene is one of the known toughest materials in nature, with a very large in-plane Young modulus of  $\approx 1$  TPa, which arises from the extraordinary efficiency of the  $\text{sp}^2$  and  $\pi$ -like hybrid orbitals bonding C atoms on a honeycomb topology.<sup>7</sup> Furthermore, bonding between C atoms is very versatile, resulting in a range of strengths from triple to single bonds,  $\approx 10$  to 4 eV per bond, respectively. Characteristic energies for the single (S), double (D), and triple (T) carbon–carbon bond are represented by vertical dashed lines in Figure 1.<sup>8</sup> Graphene is located on an intermediate region with a coordination order of about  $1\frac{1}{3}$  that corresponds to  $\approx 4$ –5 eV per C–C bond. Such a kinetic barrier for the breaking of C–C bonds is too high for the process to take place at 1000 K on the observed experimental time scale, i.e., minutes. Owing to the exponential dependence with temperature of the Boltzmann factors such a process at 500 K would take an amount of time equivalent to the age of the universe and would not happen, while at the same time the solid–liquid phase transition starting around 3500 K is correctly predicted to take place in about milliseconds. We notice that a reduction of the barrier below 3 eV would be enough to allow such a process to happen in about a few minutes. Experiments run on

Received: September 11, 2014

Revised: October 27, 2014





**Figure 1.** Blue lines give typical times ( $10^t$  in s) to overcome an energy barrier ( $\Delta E$  in eV) by picking up a thermal fluctuation at a temperature  $T$  (in K) between 250 and 3500 K. Characteristic energies for the single (S), double (D), and triple (T) C–C bonds are represented by light-green lines. The top thick black horizontal line represents the total time since the big bang ( $\approx 10^{10}$  years), and the middle horizontal orange dashed line represents 1 day. Around the blue 3500 K line the breaking of C–C bonds in the solid–liquid phase transition takes place in about milliseconds. In agreement with experimental evidence,<sup>4</sup> we expect a noticeable amount of defects to appear on the graphene layer after annealing at 725 K for about 1 day (solid and dashed magenta lines).

heavily hydrogenated graphene make it clear that the presence of hydrogen does not eliminate these kinetic barriers,<sup>4</sup> but we find that it acts to reduce their values so it can happen on a reasonable laboratory time scale, between milliseconds and minutes.<sup>6</sup> Here we search for detailed mechanisms that can explain the *catalytic*-like activity of atomic hydrogen leading to the reported formation of holes in the graphene layer, as well as the nature of the fragments formed at the same time, i.e., hydrocarbons and radicals of various sizes.

## METHODOLOGY

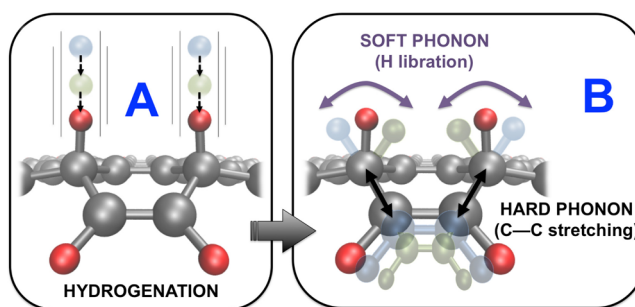
In the *ab initio* atomistic simulations, total energies, forces, and stresses were minimized by using DFT as implemented in the plane-waves package QUANTUM ESPRESSO.<sup>9</sup> A perturbative van der Waals (vdW) correction was used to check for the effect of long-range interactions on the different configurations analyzed. For this purpose, we have used an empirical vdW  $R^{-6}$  correction to add dispersive forces to conventional density functionals within the DFT+D formalism.<sup>10–12</sup> This vdW contribution to the total energy is almost negligible around the transition states; it only becomes significant ( $\sim 15\%$ ) when distances between the expelled fragments and graphene increase by  $>4$  Å, which only matters for a good description of energies at asymptotic distances. The local density approximation (LDA) has been used for the exchange and correlation potential. It describes fairly well both the C–C and C–H bonds, and at the same time offers a most simple and clear conceptual frame. Similar conclusions have been reached by checking a few selected cases with generalized gradient approximations (GGA) to the exchange and correlation potential. Norm-conserving scalar-relativistic pseudopotentials have been used to model the ion–electron interaction.<sup>13</sup>

Models having from tens to hundreds of atoms have been considered to construct both finite-size clusters (labeled as C) and extended periodic-cell models (labeled as c). A  $k$ -space mesh of  $\Delta k \leq 0.01$  Å<sup>-1</sup>, and an energy cutoff of 500 eV, yielded total energies with an accuracy of  $\Delta E \approx \pm 0.05$  eV (converged to a precision better than  $10^{-6}$  eV). Optimized geometries were obtained with residual forces lower than 0.01 eV/Å, and stresses below 0.1 GPa. Besides enthalpies of formation,  $H$ , for a particular process the all important feature to ascertain its

feasibility is the height of the barrier at the transition state (TS),  $\Delta E$ . These TSs have been investigated within the climbing-image nudge elastic band (CI–NEB) approach<sup>14–16</sup> implemented in the QUANTUM ESPRESSO package,<sup>9</sup> where the initial, the final, and all the intermediate image-states were free to fully relax. At this point, it is important to remark that the underlying *chemistry* of these systems is well-represented by all the approximations adopted here, and we expect a good correlation with experimental data, as it is usually attained.<sup>17–19</sup>

## RESULTS AND DISCUSSION

A plausible mechanism that would allow the breaking of the strong in-plane C–C bonds should involve a transition state low enough to get from the initial to the final state in a meaningful time on the experiment time scale. Our simulations show that atomic hydrogen plays a dual role in this job. On one hand, it chemisorbs on the graphene layer with an adsorption energy of  $\approx 1$  eV, and a small initial sticking barrier of  $\approx 0.15$  eV. This small barrier appears as a result of the induced deformation upon adsorption of H on the stiff planar graphene layer, where the  $sp^2$  acquires a partial  $sp^3$  character.<sup>20,21</sup> On the other hand, hydrogen can adsorb on the edges of the newly created holes saturating the dangling bonds very effectively. That kind of C–H bond can store as much as 5 eV, and it is characterized by vibrational frequencies that match the phonons of the graphene layer.<sup>7</sup> Therefore, the C–H bond can effectively compete with the C–C bond making it possible to reduce transition barriers for various processes to reasonable amounts (see Figure 2). Figure 1 gives in a logarithmic scale the



**Figure 2.** Schematic mechanism to detach a  $C_2H_2$  molecule from the edge of a hydrogenated graphene nanoribbon. In step A hydrogen atoms are chemisorbed on nearby positions to progress toward a transition state under the combination of two phonons (step B). Once the H atoms have been inserted near the transition state the C–C bonds are weakened enough to be broken by thermal activation after annealing.

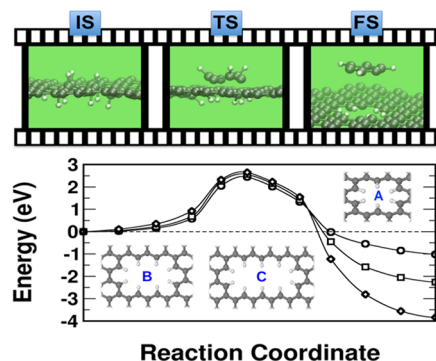
typical time in seconds of having a process with a transition barrier of  $\Delta E$  (eV) at a temperature  $T$  (K). At the temperatures of interest ( $T > 500$  K) we can safely neglect tunneling contributions, and we assume that the probability for a thermal fluctuation to overcome the TS barrier is determined by the canonical Boltzmann factor<sup>22</sup>

$$\Gamma = \Gamma_0 e^{-\Delta E/k_B T} \quad (1)$$

Since the exponential dominates the behavior of this probability as a function of  $T$  we take the prefactor as a temperature-independent value. This prefactor  $\Gamma_0$  is indeed determined by the combined action of a soft and a hard vibrational modes. We have calculated typical frequencies for these modes: the H libration (left panel in Figure 2) comes around 300 cm<sup>-1</sup>, while

the C–C stretching (right panel in Figure 2) comes around  $1000\text{ cm}^{-1}$ . These modes crucially determine the prefactor in such a way that if we imagine these frequencies tending to zero,  $\Gamma_0$  vanishes and the effect we are introducing would not be possible. On the other hand, the variation with the temperature of these values, or the difference between the soft and the hard modes, is totally negligible compared with the exponential factor that changes by many orders of magnitude with small variations of  $T$ , while the prefactor mostly changes by 1 or 2 orders of magnitude at most. Therefore, the vibrational mechanism we have identified is at the same time crucial to understand the effect, but irrelevant to describe its variation with  $T$ . According to the aforementioned details, the prefactor is here taken as a typical phonon  $\Gamma_0 = 1000\text{ cm}^{-1}$ . By propagation of errors, we find for typical temperatures of  $\approx 1000\text{ K}$ , and typical energy barriers of  $\approx 1\text{ eV}$ , that uncertainties in the calculation of barriers of  $\approx \pm 0.05\text{ eV}$ , and in the values of temperatures of  $\approx \pm 10\text{ K}$ , result in Figure 1 in fractional errors  $< 35\%$ . These error bars, or even much higher values, are quite acceptable to draw conclusions since the rates change so quickly with the values of the barriers.

First, we discuss how a defective hole on a graphene layer can be created, as it has been observed in UHV-STM and AFM experimental images.<sup>6</sup> We start by adsorbing a few H atoms on both sides of the graphene layer surrounding the region where the hole may appear and producing a  $\text{sp}^3$ -like deformation on the layer. The role of chemisorbed atoms is to weaken first, and then to passivate the broken C–C bonds. Take as an example the extraction of a six C ring (see Figure 3). It would be



**Figure 3.** (Top) Pictorial sketch of the initial (IS), transition (TS), and final (FS) states for the  $\text{C}_6\text{H}_6$  detachment. (Bottom) Energy barriers to create holes of different sizes on an extended periodic graphene layer: (○) corresponds to the creation of a  $\text{C}_6$  hole, (□) to a  $\text{C}_{10}$  hole, and (◇) to a  $\text{C}_{14}$  hole.

desirable from the point of view of the formation enthalpy to saturate this ring with H atoms to directly obtain benzene. However, this process involves too high barriers due to steric impediments from hydrogens in the extracted fragment and other hydrogens left in the layer. Instead, we find it is optimal to extract  $\text{C}_6\text{H}_6$ . The process starts by breaking the three C–C bonds that are left unsaturated in the fragment, and subsequently the three remaining ones. In the absence of chemisorbed hydrogen the formation of such a six-C ring hole has a barrier of  $+7.12\text{ eV}$ , which at  $1000\text{ K}$  simply would not happen even in typical stellar times (see Figure 1). Adsorption of atomic hydrogen drastically modifies this picture; values in Table 1 show how the transition state to extract  $\text{C}_6\text{H}_6$  is reduced down to  $+2.43\text{ eV}$ , a process that now may take place

**Table 1.** Perimetral Length,  $P$  (Å), Enthalpies of Formation,  $H$  (eV), and Transition State Barriers,  $\Delta E$  (eV), for Different Sizes of the Hole (Created by Detaching  $\text{C}_6\text{H}_6$ ,  $\text{C}_{10}\text{H}_4$ , and  $\text{C}_{14}\text{H}_5$ )<sup>a</sup>

species	$P$ (Å)	$H$ (eV)		$\Delta E$ (eV)	
		C	c	C	c
$\text{C}_6\text{H}_6$	9	−0.82	−1.12	2.81	2.43
$\text{C}_{10}\text{H}_4$	15		−2.35		2.56
$\text{C}_{14}\text{H}_5$	21		−3.96		2.65

<sup>a</sup>C and c indicate finite-size clusters and periodic-cell models, respectively.

in less than a second at  $1000\text{ K}$ . The mechanism involves C–C stretching so a wagging mode of a nearby chemisorbed H can place the H closer to the middle of the C–C bond and can help to break it while forming at the same time a new C–H bond, cf. Figure 2. The formation of a favorable configuration of chemisorbed hydrogen atoms depends on the available density of atoms impinging on the surface; the reference experiments have been performed under conditions where atomic H has been admitted in the UHV chamber up to a  $10^{-6}\text{ mbar}$  ( $\approx 10^{10}\text{ atoms/cm}^3$ ), and temperature has been raised up to  $1000\text{ K}$ .<sup>6</sup> These conditions imply hydrogen saturation conditions, making it likely to reach a favorable configuration, that then determines the shape of the hole. These experiments have been designed to mimic regions near red-giant stars between 3 and 15 stellar radii, where temperature ranges from  $1500$  to  $200\text{ K}$ , and hydrogen densities take values between  $10^{11}$  to  $10^6\text{ atoms/cm}^3$ .<sup>23</sup> We remark that the enthalpy of formation for this particular  $\text{C}_6$  hole is negative ( $-1.12\text{ eV}$ ), therefore making the hole saturated with hydrogen and the fragment,  $\text{C}_6\text{H}_6$ , stable. A similar pattern follows for a large cluster labeled C in Table 1 with values of  $+2.81$  and  $-0.82\text{ eV}$  for the barrier and the enthalpy of formation, respectively. The main difference between the periodic and finite-size simulations resides in the barriers to stick the hydrogen initially near the region to be etched. This is mainly due to the different elastic constants for a continuous layer and a small chunk of material that may be supported by some substrate or not. Finally, we analyze, for the case of the periodic system, the upstart of the scaling with the size of the hole formed, i.e., with the number of C atoms to be extracted, by considering 10 and 14 C atoms. These holes have been labeled in Figure 3 with A, B, and C, and in the continuous lines marked with empty circles, squares, and diamonds, respectively. Enthalpies of formation grow slowly with the size of the fragment tending to the saturation value to be found on very large systems. Interestingly enough, however, the transition state barriers tend to be nearly the same, indicating that the main difficulty to extract large fragments resides with the configurational statistical complexity more than with the chemistry. This can be fixed by introducing an abundant amount of hydrogen favoring the appearance of many different configurations; the holes observed would then correspond to the particular configuration realized on the graphene layer.

According to the hole-creation mechanism proposed, the surrounding adsorbed H atoms helping to weaken C–C bonds and reducing the barrier to extract small fragments not involving any aromatic ring (such as a C atom,  $\text{C}_2$ ,  $\text{C}_3$ , etc.) suffer steric impediments to be accommodated during the extraction of the fragment. As an example, if we apply the proposed mechanism to the extraction of a single C atom, or a



CH functional group, from pristine extended graphene to create a hole, the barriers predicted by theory are in excess above 8 eV. This may be explained in terms that the hydrogen-saturated hole is left in a very stressed structural situation, with no reduction of the kinetic extraction barrier, in such a way that the event simply does not occur. This rationalization can be easily extended to the case of any aliphatic fragment leaving hydrogen-saturated holes in very structurally stressed situations. Besides, the fact that the  $C_6H_3$  fragment is found to be the smallest optimal fragment to be extracted at a reasonable kinetic barrier can be justified following the same rationalization. If one analyzes within the proposed mechanism the case of the  $C_6H_6$  (benzene) fragment, it is physically unfeasible because of another arising steric effect, and the impossibility of accommodating the H atoms (adsorbed on both graphene basal planes) during the extraction process. This occurs in the same way in attempting to analyze the extraction of the  $C_6H_5$  and  $C_6H_4$  fragments. It is in the extraction process of the  $C_6H_3$  fragment where the mentioned steric impediment is not a problem anymore, since the  $C_6$  skeleton is able to attach the 3 of the hydrogens weakening the C–C bonds from the side it is detaching from. On the other hand, the extraction of the  $C_6H_2$  and  $C_6H$  fragments yields values for the extraction kinetic barriers in excess above 5.5 eV; once again, these events simply would not occur.

In previous literature, Elias et al.<sup>4</sup> reports on a reversible process of charge and discharge of H on graphene, where the discharge is achieved by annealing up to a maximum of 723 K for 24 h. Higher temperatures resulted in damage of the graphene layer. As Figure 1 (magenta line for 725 K) shows, in combination with our values in Table 1, our simulations predict the formation of holes in the graphene layer in a time-scale of approximately 24 h near  $T = 725$  K, which is in good agreement with the observation of Raman spectra by Elias et al.,<sup>4</sup> and the STM experiments beyond that critical temperature by Merino et al.<sup>6</sup>

Finally, we notice that in an elegant experiment Xie et al. have recently measured the etching of single graphene layer edges by hydrogen plasma at a rate of approximately 0.3 nm/min.<sup>24</sup> According to our model, to create a  $C_6$  hole with a perimetral length of approximately 9 Å at the measured rate it would require a barrier of  $\approx 2.85$  eV, which is remarkably close to the values in Table 1 taking into account that our focus is on the general mechanism using arguments based in the order of magnitude of rates, not on the detailed values that we cannot expect to obtain with such accuracy.

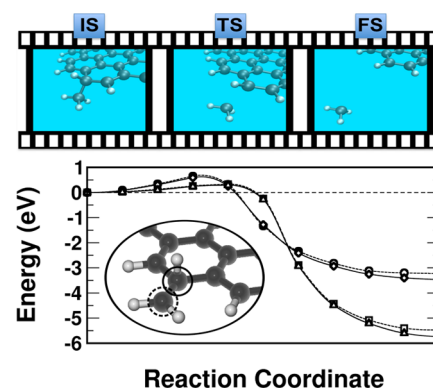
Second, we focus on the small molecules formed on the internal edges of the holes left on the surface, and in the external edges of the ejected fragments following the creation of a large defect on the layer. Table 2 shows a comparison between enthalpies and transition state barriers for a variety of small molecules. Formation of  $CH_2$  and  $CH_3$  involves the lowest barriers and the most stable systems; e.g., see discontinuous lines in Figure 3. These two molecules can be formed on the external edges of graphene flakes, or on the internal edges of the hole (c). As an example of a straightforward mechanism on the edge of a hydrogen saturated nanoribbon we draw an external C atom (surrounded by a dashed circle) attached to two H atoms and a third H chemisorbed on a top position in the nearby C (surrounded by a continuous circle). Barriers to form both  $CH_2$  and  $CH_3$  are similar since the important point here is the local geometry giving rise to the breaking of a C–C bond by insertion of a

**Table 2.** Enthalpies of Formation,  $H$  (eV), and Transition State Barriers,  $\Delta E$  (eV), To Obtain Different Fragments from the Inner Part of the Hole in the Graphene Layer<sup>a</sup>

molecule	$H$ (eV)		$\Delta E$ (eV)	
	C	c	C	c
$CH_2$	−3.22	−3.46	0.65	0.59
$CH_3$	−5.46	−5.73	0.33	0.29
$C_2H_2$	+1.52	+1.33	2.35	2.18
$C_2H_4$	+0.81	+0.74	1.93	1.79
$C_2H_6$	−0.61	−0.82	1.48	1.67

<sup>a</sup>C and c indicate finite-size clusters and periodic-cell models, respectively.

chemisorbed H (Figure 4). The more saturated with hydrogen the fragment is, the lower are the barriers, and the formation



**Figure 4.** (Top) Pictorial sketch of the initial (IS), transition (TS), and final (FS) states for the  $CH_3$  detachment. (Bottom) Energy barriers for the proposed mechanism of breaking a C–C bond in the edge of a H saturated graphene nanoribbon. A chemisorbed H on top of a C atom (continuous circle, referred to in the text as the first C) moves toward a transition state located in between the second C (dashed circle) to create fragments like  $CH_2$  (inset) or  $CH_3$  (upper panel). (○, □)  $CH_2$  and  $CH_3$  from a H saturated edge; (◇, Δ)  $CH_2$  and  $CH_3$  from an H saturated hole.

enthalpies are more negative, since the C–C bond becomes weaker. This mechanism does not result in the formation of CH nor  $CH_4$ , although for opposite reasons. In the first case, the H chemisorbed on top position (continuous circle) finds it too favorable to bond to the second C (dashed circle) and cannot participate in the breaking of the C–C bond. In the second case, the second carbon atom cannot support the interaction with four H plus the first carbon atom. Nevertheless,  $CH_2$  and  $CH_3$  fragments detached from hydrogenated graphene/graphite will capture environmental H atoms to spontaneously form  $CH_4$ .<sup>25</sup> Other fragments with higher C content, such as  $C_2H_2$ ,  $C_2H_4$ , and  $C_2H_6$ , are not exothermic, but display reasonable formation barriers too. The inverse process, where the fragment joins the edge, is hindered by entropic orientational considerations and does not play an important role. The transition barriers for these molecules having two C atoms are higher than for those having just one C atom, but still reasonable if compared with typical times in the circumstellar medium. To complement the scenario, since these particles may be bombarded by charged particles in a realistic situation located near a star, we have considered the breaking of  $C_2H_2$  on a charged surface, leading to the anion  $C_2H_2^-$ . The barrier for this case is of +2.20 eV, i.e., very similar to the one

found on the inner side of the hole for the periodic system. The presence of the extra charge on the surface does not significantly alter the mechanism of breaking the relevant C–C bonds, and indicates that the abundance of anions and neutral species should be similar. Since the experiments have been performed on graphene formed on a SiC substrate we have additionally tested the influence of the layer immediately below the surface, the so-called buffer layer. For  $C_2H_2$  we see a negligible change of the height of the barrier, from 2.35 to 2.33 eV. This finding manifests that the weak interaction between layers does not interfere with the strong chemical-like bonds broken and formed that explains the mechanism for etching of the layer, and assigns to the buffer layer a purely physical confining role.<sup>20,26</sup> Since in most cases graphene layers grown on metallic surfaces keep also a weak interaction with the support, we expect the same principles to apply for those cases.

## CONCLUSIONS

Summarizing, in this Article we propose a mechanism underlying the formation of prototypical hydrocarbons, hydrocarbyls, and other polyaromatics from hydrogenated graphene. We have studied the formation of large holes on graphene upon atomic hydrogen adsorption and high- $T$  annealing, as observed in experiments performed in the laboratory. These have been rationalized in terms of the adsorption of H atoms around the region where the hole is expected to appear. Chemisorbed atoms act to passivate broken bonds in synergy with phonons simultaneously weakening the hard C–C bonds. This is a temperature-activated process exhibiting typical barriers of around 3 eV, and occurs at 1000 K in a few minutes. Computed transition state energies are used to determine the Boltzmann probability for a thermal fluctuation to overcome such a barrier, and the time scale for an event to occur, which is of paramount importance from the practical point of view. The proposed mechanism also sheds light into possible routes for the creation in the circumstellar medium of small hydrocarbons and hydrocarbyls, such as  $CH_2$ ,  $CH_3$ ,  $C_2H_2$ ,  $C_2H_2^-$ ,  $C_2H_4$ , and  $C_2H_6$ .

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### Notes

The authors declare no competing financial interest.

## ACKNOWLEDGMENTS

We acknowledge funding from the Spanish MINECO (Grants MAT2011-26534 and CSD2009-00038), the EU (FP7 Program Grant 604391 GRAPHENE FLAGSHIP, and ERC Synergy Grant ERC-2013-SYG-610256 NANOCOSMOS), and computing resources from CTI-CSIC. J.L.M. acknowledges a CSIC-JaeDoc Fellowship, cofunded by ESF.

## REFERENCES

- (1) Hornekaer, L.; Baurichter, A.; Petrunin, V. V.; Field, D.; Luntz, A. C. *Science* **2003**, *302*, 1943–1946.
- (2) McKay, H.; Wales, D. J.; Jenkins, S. J.; Verges, J. A.; de Andres, P. L. *Phys. Rev. B* **2010**, *81*, 075425.
- (3) Casolo, S.; Martinazzo, R.; Bonfanti, M.; Tantardini, G. F. *J. Phys. Chem. A* **2009**, *113*, 14545–14553.

- (4) Elias, D. C.; Nair, R. R.; Mohiuddin, T. M. G.; Morozov, S. V.; Blake, P.; Halsall, M. P.; Ferrari, A. C.; Boukhvalov, D. W.; Katsnelson, M. I.; Geim, A. K.; Novoselov, K. S. *Science* **2009**, *323*, 610–613.
- (5) Guisinger, N. P.; Rutter, G. M.; Crain, J. N.; First, P. N.; Stroscio, J. A. *Nano Lett.* **2009**, *9*, 1462–1466.
- (6) Merino, P.; Svec, M.; Martinez, J. I.; Jelinek, P.; Lacovig, P.; Dalmiglio, M.; Lizzit, S.; Soukiassian, P.; Chernicharo, J.; Martin-Gago, J. A. *Nat. Commun.* **2014**, *5*, 3054.
- (7) de Andres, P. L.; Guinea, F.; Katsnelson, M. I. *Phys. Rev. B* **2012**, *86*, 245409.
- (8) Langhoff, S. R.; Bauschlicher, C. W. *J. Chem. Phys. Lett.* **1991**, *180*, 88–94.
- (9) Baroni, S.; Corso, A. D.; de Gironcoli, S.; Giannozzi, P. *QUANTUMESPRESSO Package*; 2005.
- (10) Grimme, S. *J. Comput. Chem.* **2006**, *27*, 1787–1799.
- (11) Elstner, M.; Hobza, P.; Frauenheim, T.; Suhai, S.; Kaxiras, E. *J. Chem. Phys.* **2001**, *114*, 5149–5155.
- (12) Dunitz, J. D.; Gavezotti, A. *Acc. Chem. Res.* **1999**, *32*, 677–684.
- (13) Vanderbilt, D. *Phys. Rev. B* **1990**, *41*, 7892–7895.
- (14) *Classical and Quantum Dynamics in Condensed Phase Simulations*; Berne, B. J., Cicotti, G., Coker, D. F., Eds.; World Scientific Publishing Company: Singapore, 1998.
- (15) Henkelman, G.; Jónsson, H. *J. Chem. Phys.* **2000**, *113*, 9978–9985.
- (16) Henkelman, G.; Uberuaga, B. P.; Jónsson, H. *J. Chem. Phys.* **2000**, *113*, 9901–9904.
- (17) Balog, R.; Jorgensen, B.; Wells, J.; Laegsgaard, E.; Hofmann, P.; Besenbacher, F.; Hornekaer, L. *J. Am. Chem. Soc.* **2009**, *131*, 8744–8745.
- (18) Schwabe, T.; Grimme, S. *Phys. Chem. Chem. Phys.* **2006**, *8*, 4398–4401.
- (19) Chen, L.; Cooper, A. C.; Pez, G. P.; Cheng, H. *J. Phys. Chem. C* **2007**, *111*, 18995–19000.
- (20) Balgar, T.; Kim, H.; Hasselbrink, E. *J. Phys. Chem. Lett.* **2013**, *4*, 2094–2098.
- (21) Johns, J. E.; Hersam, M. C. *Acc. Chem. Res.* **2013**, *46*, 77–86.
- (22) de Andres, P.; Flores, F.; Cerda, J. R.; Echenique, P. *J. Phys.: Condens. Matter* **1993**, *5*, A411.
- (23) Agundez, M.; Cernicharo, J. *Astrophys. J.* **2006**, *650*, 374.
- (24) Xie, L.; Jiao, L.; Dai, H. *J. Am. Chem. Soc.* **2010**, *132*, 14751–14753.
- (25) Wood, B. J.; Wise, H. *J. Phys. Chem.* **1969**, *73*, 1348–1351.
- (26) Seyller, T.; Graupner, R.; Sieber, N.; Emtsev, K. V.; Ley, L.; Tadich, A.; Riley, J. D.; Leckey, R. C. G. *Phys. Rev. B* **2005**, *71*, 245333.