

Problem Set 4

Physics 180

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8.9 Three nuclei A, B, C are radioactive isotopes of the same element and all decay via α -emission. The half-lives $t_{1/2}$ and ranges r of the alpha particles for A and B are:

$$A (t_{1/2} = 10^3 \text{ yr}, r = 3 \text{ cm}), \quad B (t_{1/2} = 10^2 \text{ days}, r = 4 \text{ cm}).$$

If the range of the α particles from the decay of C is $r = 6 \text{ cm}$, estimate its half-life.

We will be using the Geiger–Nuttall relation to get the half-life of the C isotope:

$$\log_{10} \lambda = c + d \log_{10} r \quad (1)$$

where $\lambda = \frac{\ln 2}{t_{1/2}}$. We begin with finding the values of c and d using the given for the A and B isotopes with:

$$\log_{10} \left(\frac{\ln 2}{t_{1/2}} \right) = c + d \log_{10} r \quad (2)$$

converting them to similar units, we have the system of equations:

$$\log_{10} \left(\frac{\ln 2}{525\,948\,766 \text{ min}} \right) = c + d \log_{10}(3 \text{ cm}) \quad (3)$$

$$\log_{10} \left(\frac{\ln 2}{144\,000 \text{ min}} \right) = c + d \log_{10}(4 \text{ cm}) \quad (4)$$

$$(5)$$

we can solve these using substitution:

$$c = \log_{10} \left(\frac{\ln 2}{525\,948\,766 \text{ min}} \right) - d \log_{10}(3 \text{ cm}) \quad (6)$$

$$\log_{10} \left(\frac{\ln 2}{144\,000 \text{ min}} \right) = \log_{10} \left(\frac{\ln 2}{525\,948\,766 \text{ min}} \right) - d \log_{10}(3 \text{ cm}) + d \log_{10}(4 \text{ cm}) \quad (7)$$

$$d = \left[\log_{10} \left(\frac{\ln 2}{525\,948\,766 \text{ min}} \right) - \log_{10} \left(\frac{\ln 2}{144\,000 \text{ min}} \right) \right] \left[\frac{1}{\log_{10}(3 \text{ cm}) - \log_{10}(4 \text{ cm})} \right] \quad (8)$$

from here we get the following values:

$$c = -22.49 \tag{9}$$

$$d = 28.51 \tag{10}$$

getting the half-life for the C isotope:

$$10^{\log_{10}(\ln 2) - c - d \log_{10} r} = t_{1/2} \tag{11}$$

$$10^{\log_{10}(\ln 2) + 22.49 - 28.51 \log_{10}(6)} = t_{1/2} \tag{12}$$

$$\boxed{t_{1/2} = 1.37 \text{ min}} \tag{13}$$

8.15 Use the Weisskopf formulas (8.88a) and (8.88b) to calculate the radiative width $\Gamma_\gamma(\text{E3})$ expressed in a form analogous to (8.89).

The emission rate $T_{fi}^{E,M}$ is given by:

$$T_{fi}^{E,M} = \frac{1}{4\pi\epsilon_0} \frac{8\pi(L+1)}{L[(2L+1)!!]^2} \frac{1}{\hbar} \left(\frac{E_\gamma}{\hbar c} \right)^{2L+1} B_{fi}^{E,M}(L) \quad (14)$$

in this case, we only need $B_{fi}^{E,M}(L) = B^E(L)$ for the E3 multipole transition. The reduced transition probability for electric radiation ($B^E(L)$) is given by:

$$B^E(L) = \frac{e^2}{4\pi} \left(\frac{3}{L+3} \right)^2 (R_0)^{2L} A^{2L/3} \quad (15)$$

$$(16)$$

For E3 multipole transition ($L = 3$), we then have $B^E(L)$ as:

$$B^E(L=3) = \frac{e^2}{4\pi} \left(\frac{1}{2} \right)^2 (R_0)^6 A^2 \quad (17)$$

$$B^E(3) = \frac{e^2}{16\pi} R_0^6 A^2 \quad (18)$$

using this for $T_{fi}^{E,M}$:

$$T^E = \frac{1}{4\pi\epsilon_0} \frac{8\pi(4)}{3[(7)!!]^2} \frac{1}{\hbar} \left(\frac{E_\gamma}{\hbar c} \right)^7 B^E(3) \quad (19)$$

$$T^E = \frac{1}{\epsilon_0} \frac{8}{3[(7)!!]^2} \frac{1}{\hbar} \left(\frac{E_\gamma}{\hbar c} \right)^7 B^E(3) \quad (20)$$

we note that for the double factorial for odd integers, we have:

$$(2n+1)!! = \frac{(2n+1)!}{2^n n!} \quad (21)$$

$$(7)!! = \frac{(7)!}{2^3(3!)} = 105 \quad (22)$$

$$(23)$$

going back to the emission rate, we have:

$$T^E = \frac{1}{\epsilon_0} \frac{8}{3[105]^2} \frac{1}{\hbar} \left(\frac{E_\gamma}{\hbar c} \right)^7 \left(\frac{e^2}{16\pi} R_0^6 A^2 \right) \quad (24)$$

$$= \frac{e^2 R_0^6}{6\pi[105]^2 \epsilon_0 \hbar^8 c^7} E_\gamma^7 A^2 \quad (25)$$

$$= \frac{e^2 (1.21 \text{ fm})^6}{6\pi[105]^2 (55.263 \text{ e}^2 \text{ GeV}^{-1} \text{ fm}^{-1}) (6.582 \times 10^{25} \text{ GeV s})^8 (3 \times 10^{23} \text{ fm/s})^7} E_\gamma^7 A^2 \quad (26)$$

$$(27)$$

to get the radiative width Γ , we use:

$$\Gamma = \hbar T^E \quad (28)$$

$$= \frac{e^2 (1.21 \text{ fm})^6}{6\pi[105]^2 (55.263 \text{ e}^2 \text{ GeV}^{-1} \text{ fm}^{-1}) (6.582 \times 10^{25} \text{ GeV s})^7 (3 \times 10^{23} \text{ fm/s})^7} E_\gamma^7 A^2 \quad (29)$$

$$(30)$$

$$\boxed{\Gamma = (2.346 \times 10^{16}) E_\gamma^7 A^2 \text{ eV}} \quad (31)$$