#### In [2]:

```
%matplotlib inline
import sys
sys.path.insert(0,'..')
from IPython.display import HTML,Image,SVG,YouTubeVideo
```

# **Image restoration**

Image acquisition is rarely perfect, physics or external condition may deform the image being acquired, here are some example of typicall problems:

# **Deformation model**

The original image is f(x, y) undergoes a deformation, given by H, and an additive noise  $\eta(x, y)$  the acquired image is g(x, y).

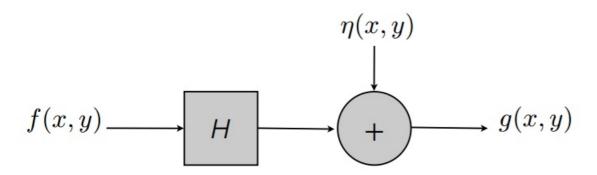
The restoration problem can be stated as follow:

how to recover a good approximation of f(x, y) from g(x, y)?

### In [3]:

Image('http://homepages.ulb.ac.be/~odebeir/data/restauration.png')

#### Out[3]:



Some examples:

$$g(x, y) = H[f(x, y)] + \eta(x, y)$$

H properties:

linear

$$H[k_1f_1(x, y) + k_2f_2(x, y)] = k_1H[f_1(x, y)] + k_2H[f_2(x, y)]$$

additive

$$H[f_1(x, y) + f_2(x, y)] = H[f_1(x, y)] + H[f_2(x, y)]$$

• homogeneous

$$H[k_1f_1(x, y)] = k_1H[f_1(x, y)]$$

• spatially invariant

$$H[f(x-\alpha,y-\beta)]=g(x-\alpha,y-\beta)$$

# **Point Spread Function (PSF)**

We can rewrite f(x, y) as a sum of Dirac function.

$$f(x, y) = \int_{\infty - \infty} \int_{\infty - \infty} f(\alpha, \beta) \delta(x - \alpha, y - \beta) d\alpha d\beta$$

if there is no addiditive noise:

$$g(x,y) = H[f(x,y)] = H\left[\int_{\infty-\infty} \int_{\infty-\infty} f(\alpha,\beta) \delta(x-\alpha,y-\beta) d\alpha d\beta\right]$$

by linearity property:

$$g(x, y) = \int_{\infty - \infty} \int_{\infty - \infty} H[f(\alpha, \beta)\delta(x - \alpha, y - \beta)] d\alpha d\beta$$

 $f(\alpha, \beta)$  is independent of x and y:

$$g(x,y) = \int_{\infty-\infty} \int_{\infty-\infty} f(\alpha,\beta) H[\delta(x-\alpha,y-\beta)] d\alpha d\beta$$

the impulse response of H, also known as point spread function (PSF), is:

$$h(x, \alpha, y, \beta) = H[\delta(x - \alpha, y - \beta)]$$

after substitution we have:

$$g(x, y) = \int_{\infty - \infty} \int_{\infty - \infty} f(\alpha, \beta) h(x, \alpha, y, \beta) d\alpha d\beta$$

this expression means that, if the response H of an impulse is known, the response of any input  $f(\alpha, \beta)$  is known.

if h is spatially invariant (see above):

$$g(x,y) = \int_{\infty - \infty} \int_{\infty - \infty} f(\alpha,\beta) h(x - \alpha, y - \beta) d\alpha d\beta$$

by adding the noise:

$$g(x, y) = \int_{m-m} \int_{m-m} f(\alpha, \beta) h(x - \alpha, y - \beta) d\alpha d\beta + \eta(x, y)$$

see also:

DIP (../00-Preface/06-References.ipynb#[DIP]) p254

#### Restoration

# inverse filtering

If noise is negligeable and PSF is known (in the Fourier domain):

$${}^{\hat{}}\mathsf{F}(\mathsf{u},\mathsf{v}) = \ \mathsf{G}(\mathsf{u},\mathsf{v})\mathsf{H}(\mathsf{u},\mathsf{v}){}^{\hat{}}\mathsf{f}(\mathsf{x},\mathsf{y}) = \ \mathsf{F}^{-1}\big[{}^{\hat{}}\mathsf{F}(\mathsf{u},\mathsf{v})\,\big] = \ \mathsf{F}^{-1}\big[\mathsf{G}(\mathsf{u},\mathsf{v})\mathsf{H}(\mathsf{u},\mathsf{v})\,\big]$$

else:

$$F(u, v) = F(u, v) + N(u, v)H(u, v)$$

and noise is increasing when H is low, restoration is limited where H is big.

if define a restoration transform M(u, v) as:

$$^{\mathsf{F}}(\mathsf{u},\mathsf{v}) = (\mathsf{G}(\mathsf{u},\mathsf{v}) + \mathsf{N}(\mathsf{u},\mathsf{v}))\mathsf{M}(\mathsf{u},\mathsf{v})$$

one can use:

$$M(u, v) = 1/H(u, v), \quad u^2 + v^2 \le w_{20}^2, \quad u^2 + v^2 > w_{20}^2$$

with  $w_0$  being a distance to the origin in the Fourier space.

see also:

• DPP (../00-Preface/06-References.ipynb#[DPP]) p276

# Wiener filtering

To avoid arbitrary setting of a parameter for theinverse transform such as  $w_0$ , one can use the Wiener approach wich consists in minimizing by least square some error function.

Without going into the details, the restoration transform is:

$$\mathsf{M}(\mathsf{u},\mathsf{v}) = \mathsf{H}^{*}(\mathsf{u},\mathsf{v}) \, |\, \mathsf{H}(\mathsf{u},\mathsf{v}) \, |^{\, 2} + \mathsf{S}_{\nu\nu}(\mathsf{u},\mathsf{v}) \mathsf{S}_{gg}(\mathsf{u},\mathsf{v})$$

where  $H^*(u, v)$  is the complex conjugate of H(u, v),

 $S_{\nu\nu}(u,v)$  is the spectral density of the noise and

 $S_{qq}(u,v)$  is the spectral density of the degraded image.

see also:

• <u>IPAMV (../00-Preface/06-References.ipynb#[IPAMV])</u> p107

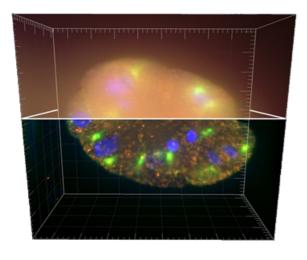
#### **Blind deconvolution**

If PSF in unknown, is has to be estimated. This is called blind deconvolution.

#### In [4]:

Image('http://bigwww.epfl.ch/algorithms/deconvolutionlab/meta/splash.png')

## Out[4]:



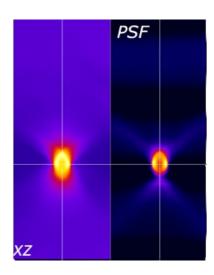


image source (http://bigwww.epfl.ch/algorithms/deconvolutionlab/)

# In [ ]: