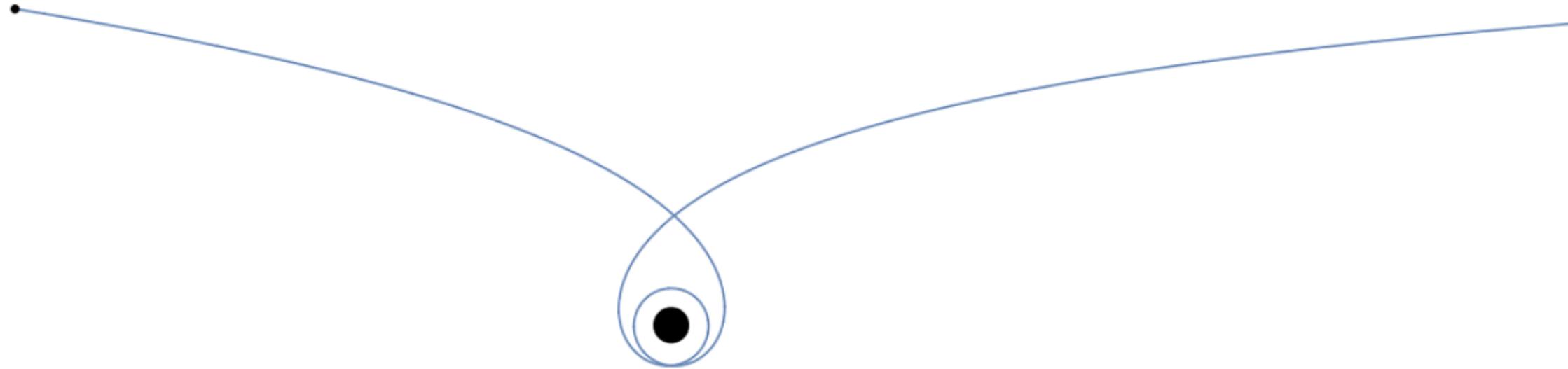


# Towards a self-force calculation of the scatter angle in hyperbolic encounters



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# EMRIs: Expanding in the mass ratio

EMRI: compact object orbiting a supermassive BH with mass ratio  $\eta < 10^{-4}$

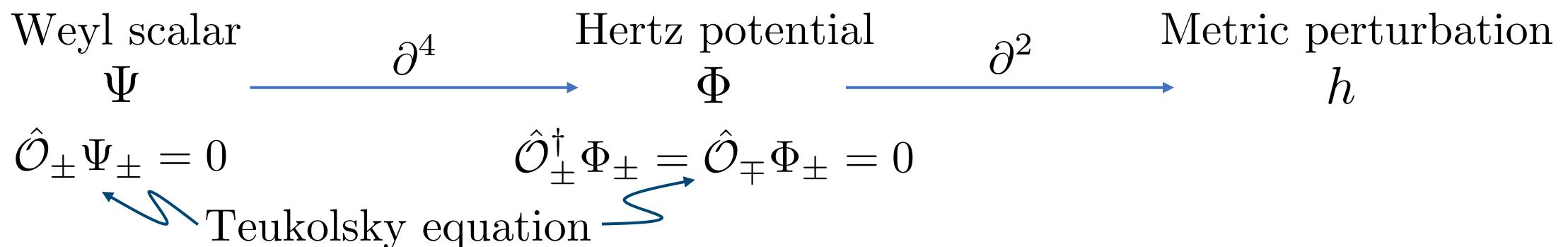
Expand quantities in the mass ratio

$$g_{\alpha\beta} = g_{\alpha\beta}^{(0)} + \eta h_{\alpha\beta}^{(1)} + \eta^2 h_{\alpha\beta}^{(2)} + \dots$$

Schwarzschild/Kerr

1<sup>st</sup> order sufficient for detection but 2<sup>nd</sup> order needed for parameter estimation

Metric reconstruction in vacuum



# 1+1D Teukolsky equation

Master Teukolsky equation

$$\begin{aligned}\hat{\mathcal{O}}_{\pm} \Psi_{\pm} = & -\frac{r^4}{\Delta} \frac{\partial^2 \Psi_s}{\partial t^2} + \Delta^{-s} \frac{\partial}{\partial r} \left( \Delta^{s+1} \frac{\partial \Psi_s}{\partial r} \right) + 2s \left( \frac{Mr^2}{\Delta} - r \right) \frac{\partial \Psi_s}{\partial t} + \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial \Psi_s}{\partial \theta} \right) \\ & + \frac{1}{\sin^2 \theta} \frac{\partial^2 \Psi_s}{\partial \varphi^2} + \frac{2is \cos \theta}{\sin^2 \theta} \frac{\partial \Psi_s}{\partial \varphi} - (s^2 \cot^2 \theta - s) \Psi_s = 0\end{aligned}$$

where  $\Delta := r^2 - 2Mr$  and  $s = \pm 2$

$s = \pm 2$  fields modally decomposed into  $s = \pm 2$  harmonics

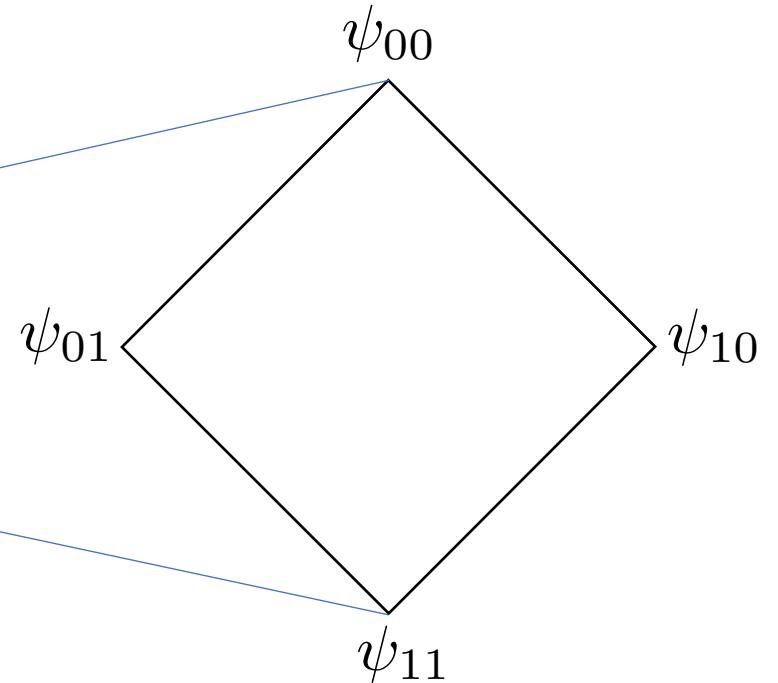
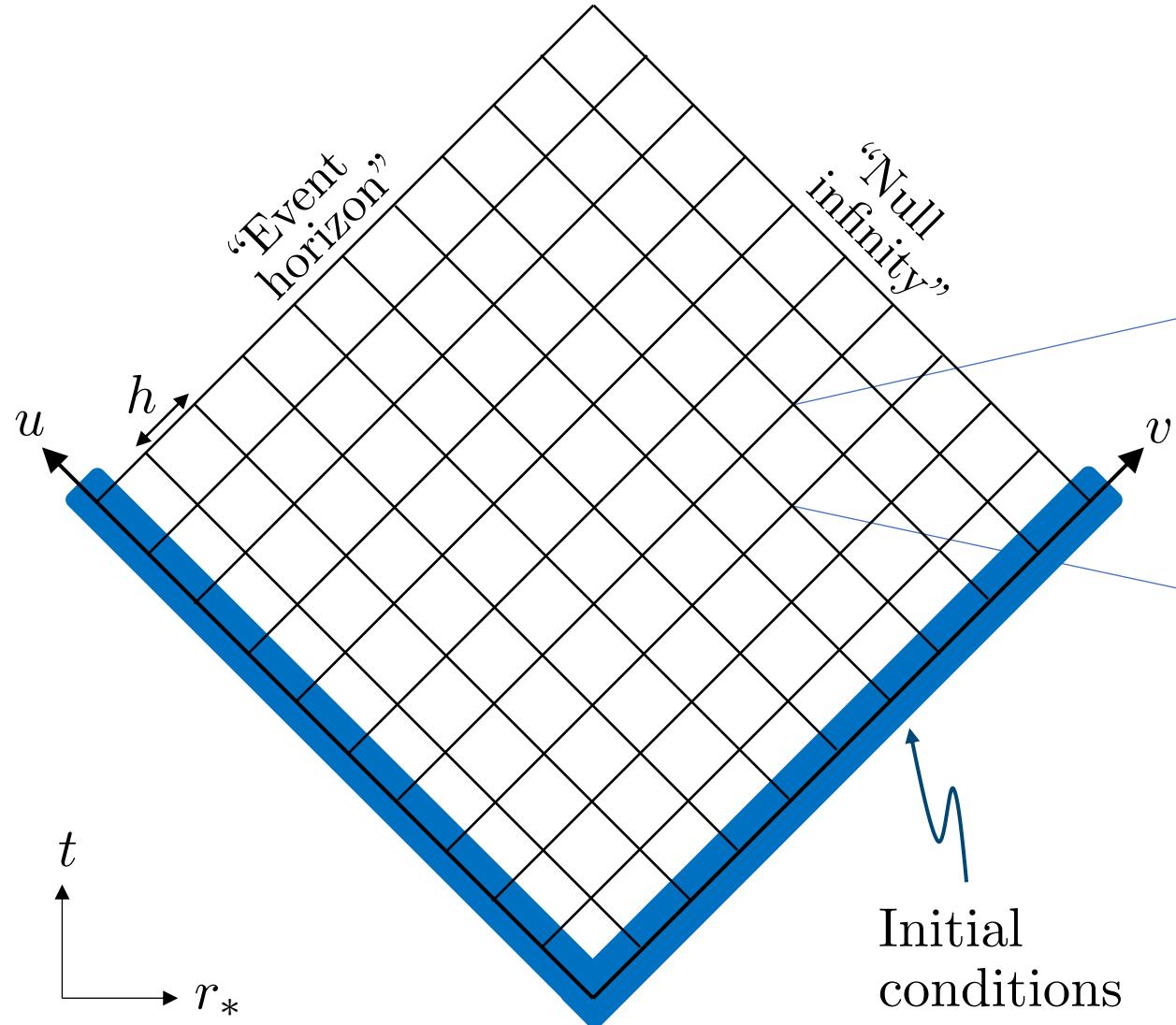
$$\Psi_{\pm} = \sum_{l=2}^{\infty} \sum_{m=-l}^l {}_{\pm 2} \psi_{lm}(t, r) {}_{\pm 2} Y_{lm}(\theta, \varphi) \quad \Phi_{\pm} = \sum_{l=2}^{\infty} \sum_{m=-l}^l {}_{\pm 2} \phi_{lm}(t, r) {}_{\pm 2} Y_{lm}(\theta, \varphi)$$

Teukolsky equation split in modes

$$\psi_{,uv} + U(r)\psi_{,u} + V(r)\psi_{,v} + W(l; r)\psi = 0$$

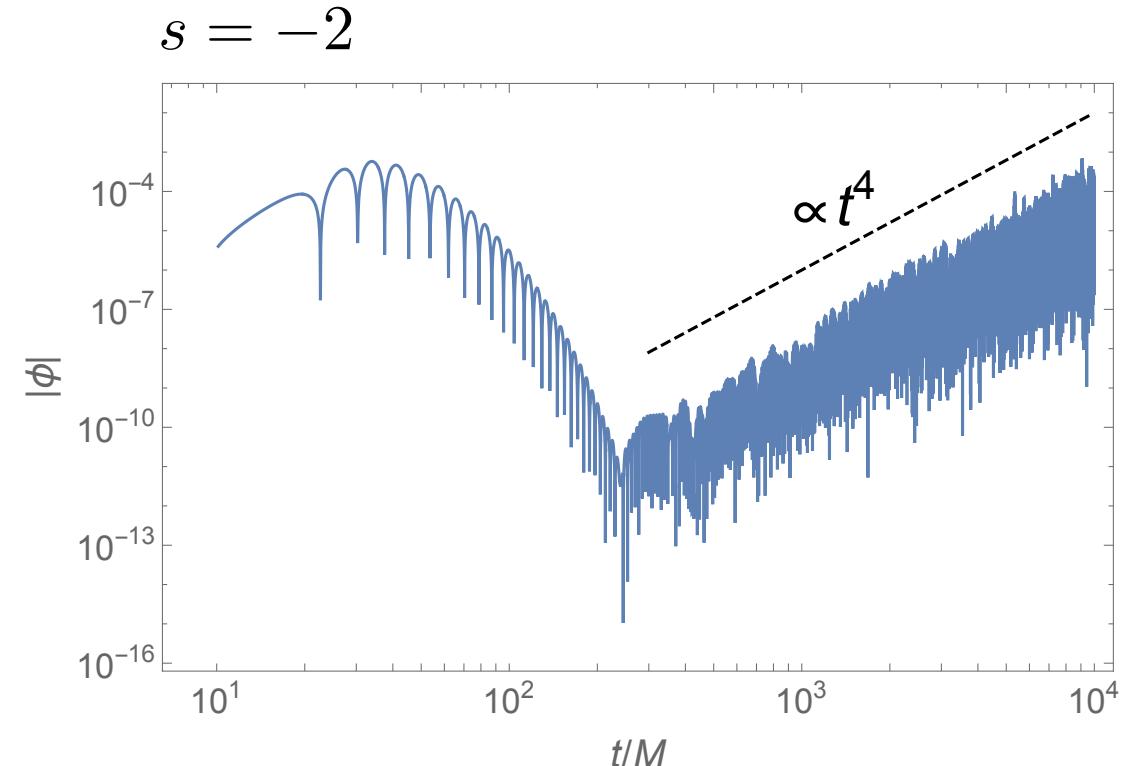
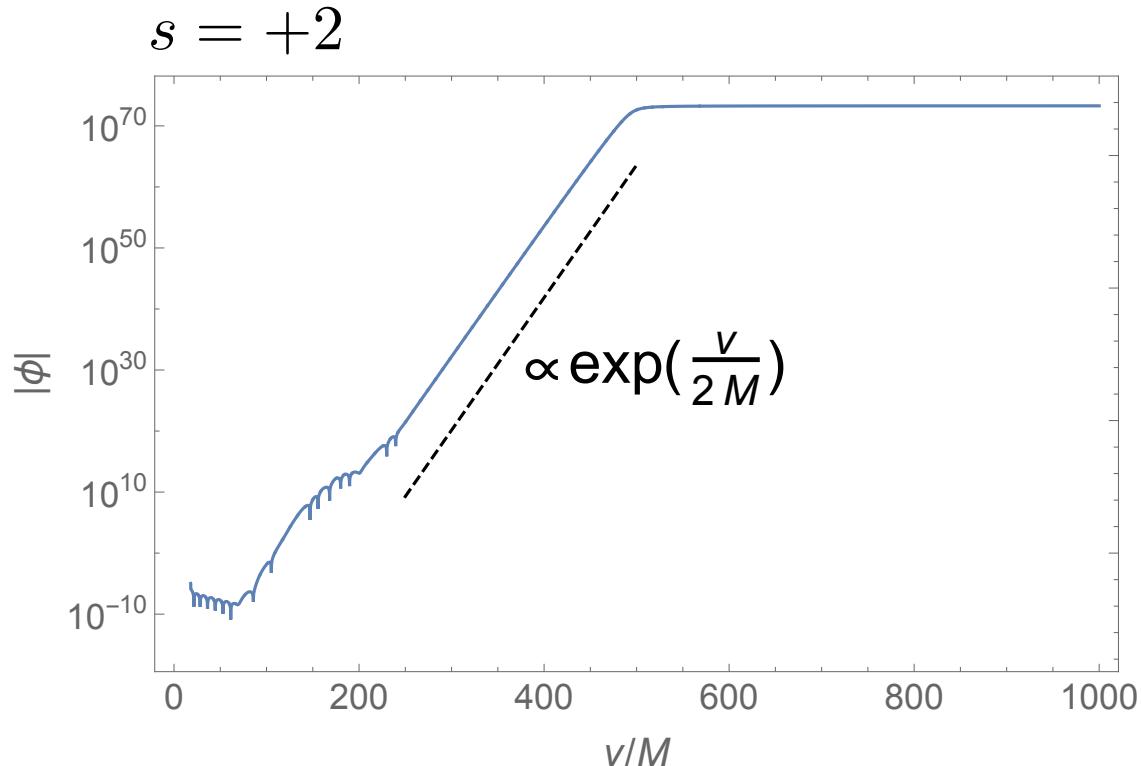
Eddington-Finkelstein coordinates  $u = t - r_*$  and  $v = t + r_*$

# 1+1D vacuum evolution code



$$\psi_{00} = \sum_{i,j=0}^1 H_{ij}(h, l; r) \psi_{ij}$$

# Problem of divergent modes



There exist non-physical asymptotic solutions:

$$\psi_{\mathcal{H}^+} \sim C_1(u) \exp(v s/4M) + C_2(v)$$

$$\psi_{\mathcal{I}^+} \sim (t + r)^{-2s} + C_4(u)$$

Non-physical terms usually removed by physical boundary conditions

Both terms originate from the  $\phi_{,t}$  term of the Teukolsky equation

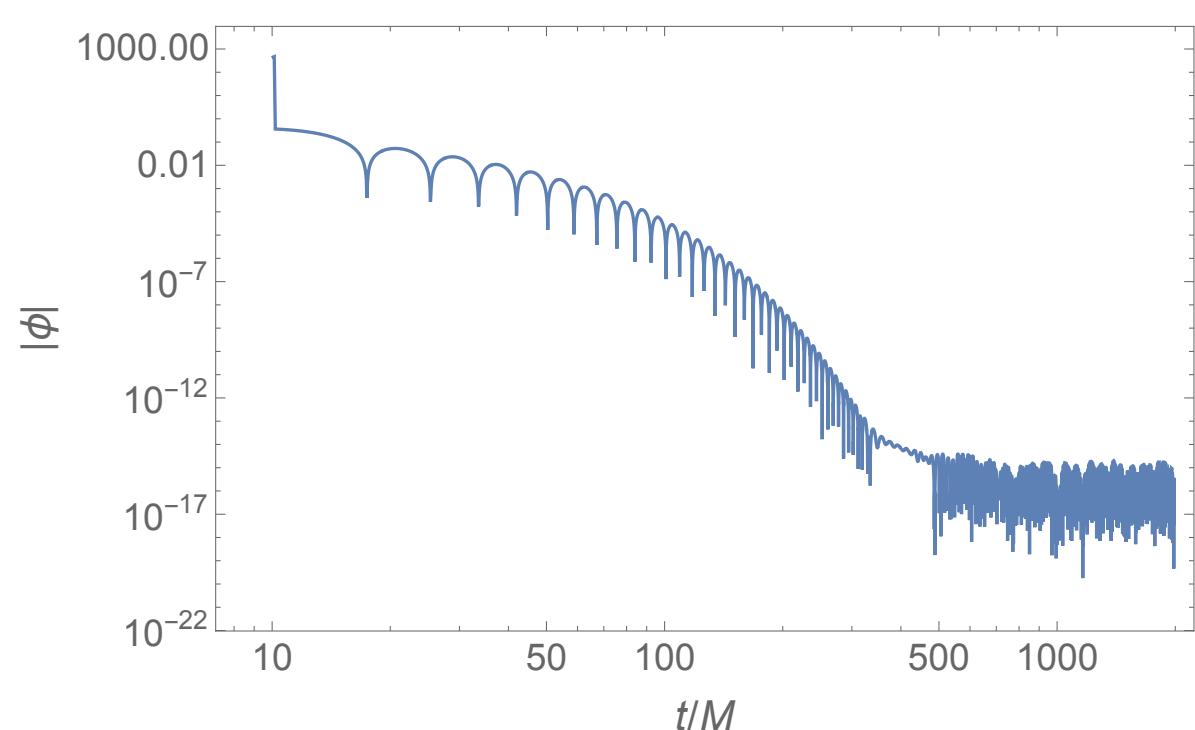
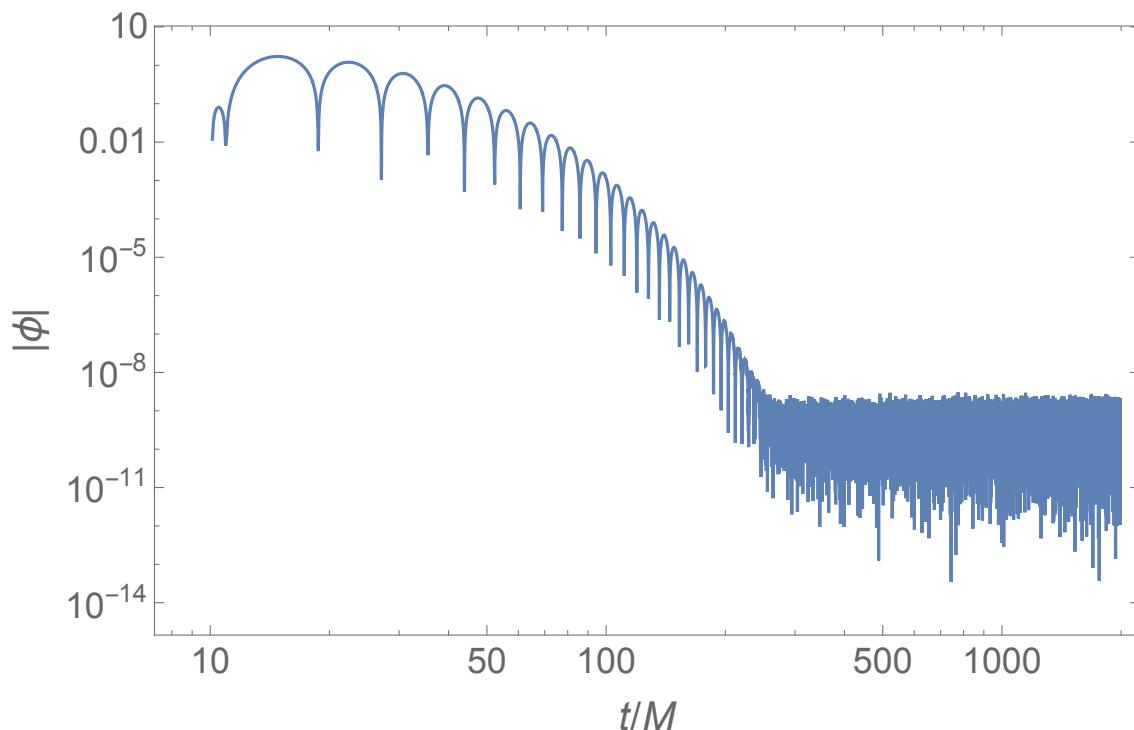
# Transformation to Regge-Wheeler form

Regge-Wheeler equation

$$X_{,uv} + \mathcal{W}(l; r)X = 0$$

$$\phi_{s=+2} = r^4 \left( X_{,vv} + \frac{r - 3M}{r^2} X_{,v} \right)$$

$$\phi_{s=-2} = \frac{r^4}{\Delta^2} \left( X_{,uu} - \frac{r - 3M}{r^2} X_{,u} \right)$$

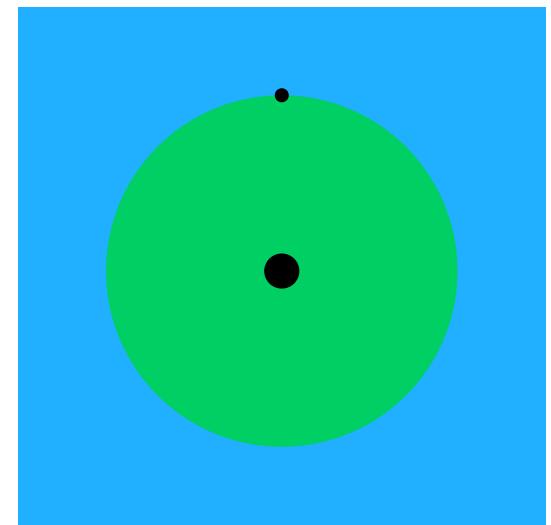
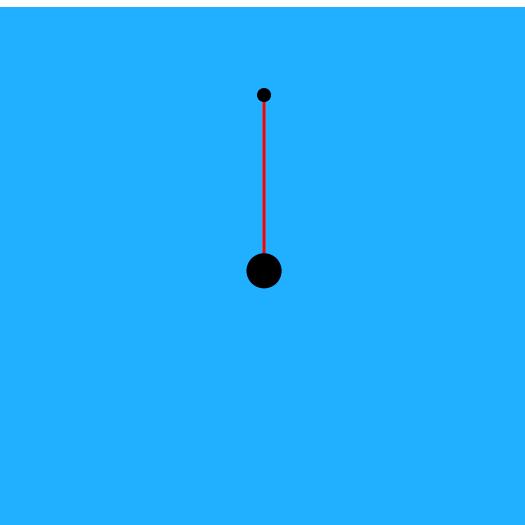
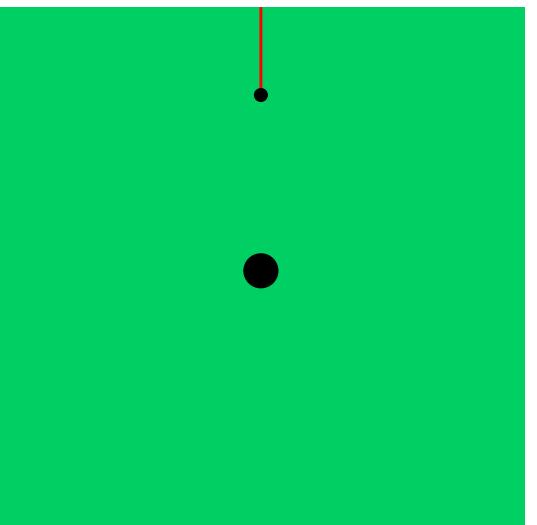
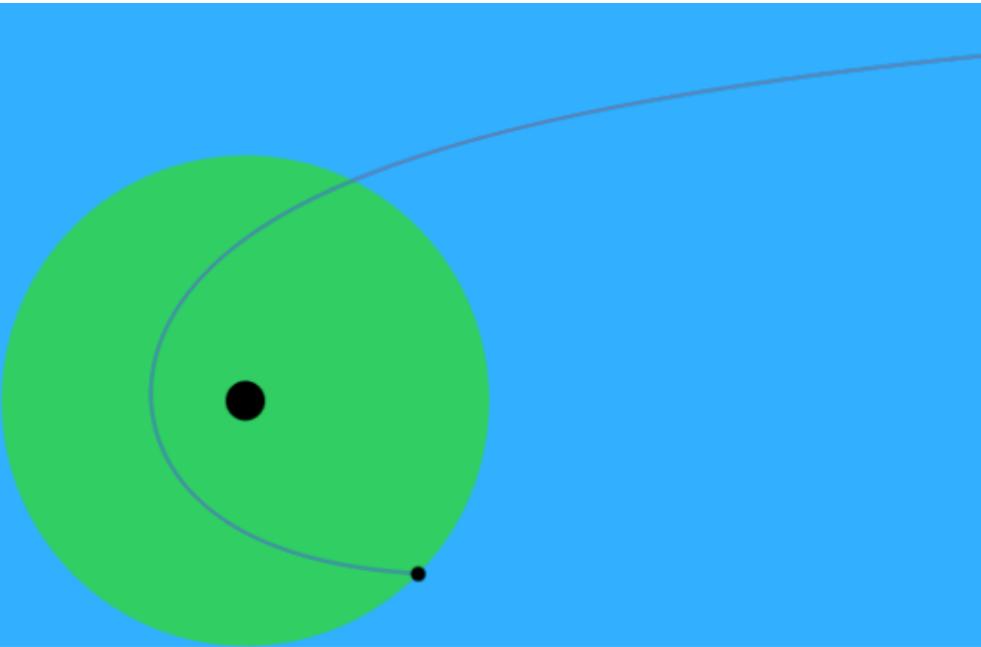


# Point-particle metric reconstruction

Sourced Teukolsky equation

$$\hat{\mathcal{O}}_{\pm} \psi_{\pm} = T_{\pm} \propto \delta(r - r_P)$$

Point particle solutions for  $\phi$  and  $h_{\alpha\beta}$



# Jumps in the field

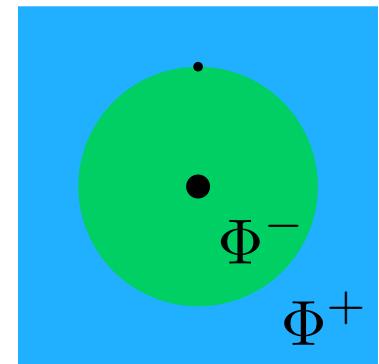
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Jump in a field

$$[\Phi] := \lim_{\epsilon \rightarrow 0} [\Phi^+(\tau, r_P(\tau) + \epsilon) - \Phi^-(\tau, r_P(\tau) - \epsilon)]$$

$[\psi]$  obtained from the sourced Teukolsky equation

Jumps in the Hertz potential from ODE



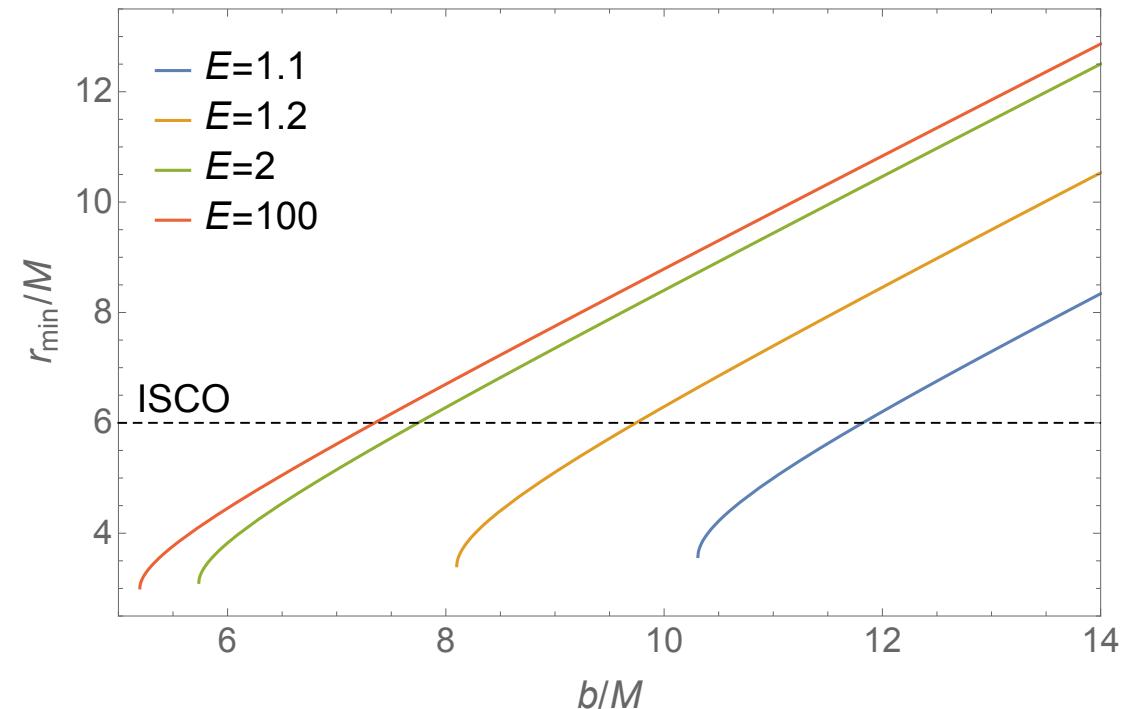
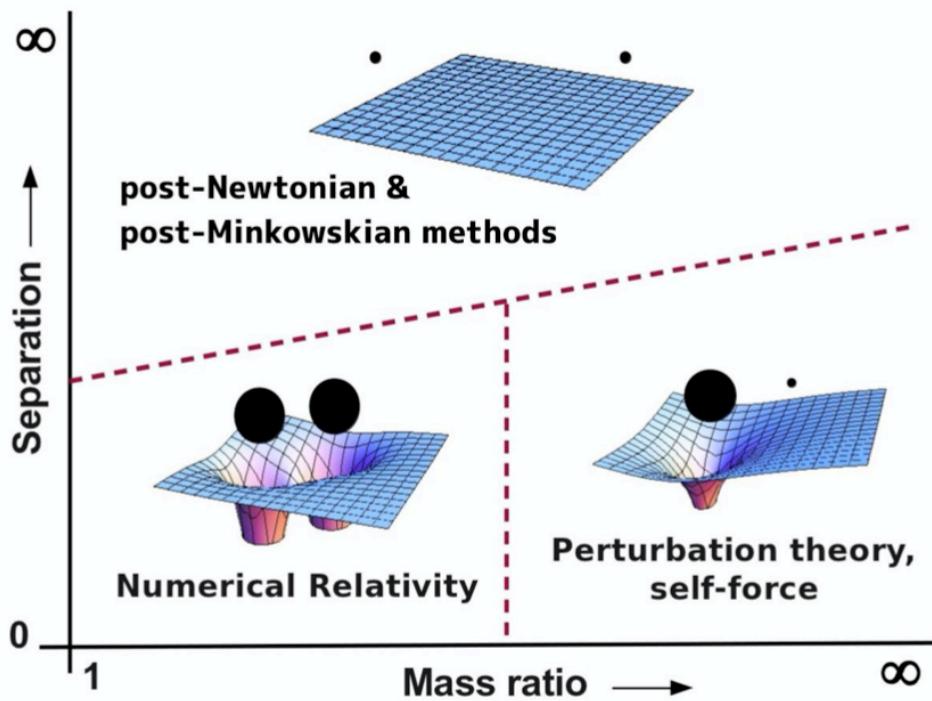
$$\dot{[\phi]} + \mathcal{A}(\tau)[\phi] = \mathcal{B}(\tau, [\psi])$$

Jumps in the Regge-Wheeler variable from ODE

$$\dot{[X]} + \mathcal{C}(\tau)[X] = \mathcal{D}(\tau, [\phi])$$

# Motivation: EMRI scatter orbits

- Exact post-Minkowskian calculations [Damour '19; Bini, Damour, Geralico '20]
  - 1<sup>st</sup> order self-force **exactly** determines 2 body Hamiltonian up to 4PM
- Comparison with quantum-like scatter amplitude calculations
- Calibration of Effective One Body (EOB) **exact** at  $O(\eta)$  (no PN/PM expansions)



# Scatter geodesics in Schwarzschild

Energy and angular momentum

$$E > 1$$

$$L > L_{\text{crit}}(E)$$

Eccentricity and semi-latus rectum

$$e \geq 1$$

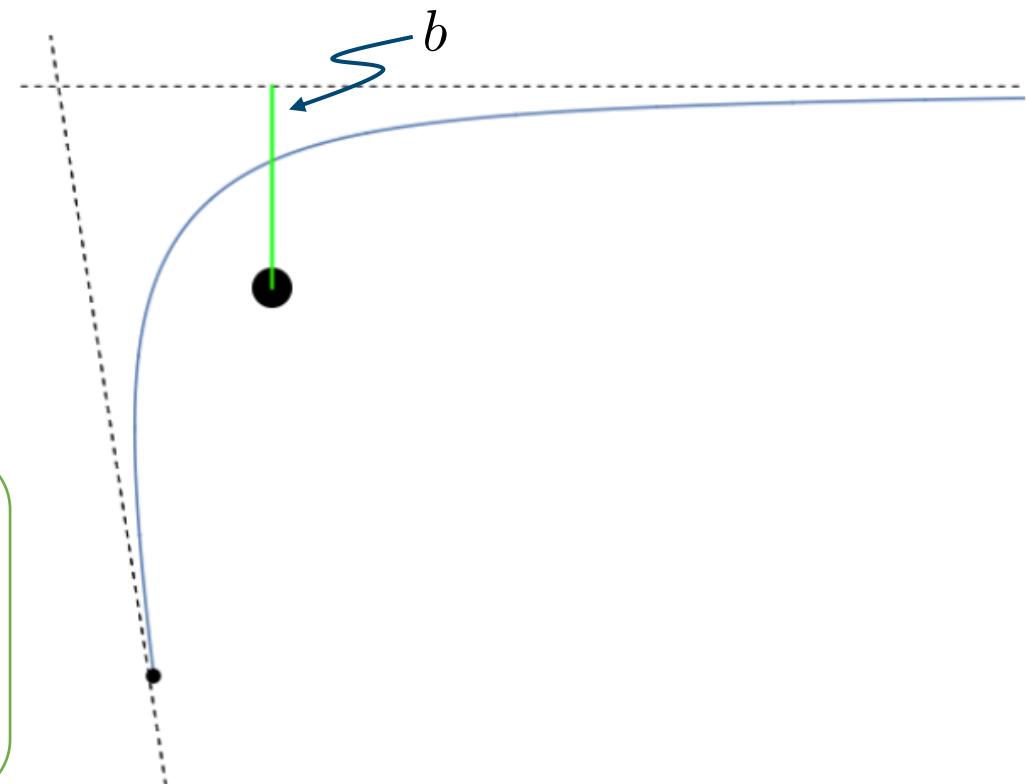
$$p > p_{\text{crit}}(e)$$

$$r = \frac{Mp}{1 + e \cos \chi}$$

$$-\chi_\infty \leq \chi \leq \chi_\infty \text{ where } \chi_\infty := \arccos(-1/e)$$

Velocity at infinity and impact parameter

$$v_\infty := \left. \frac{dr}{dt} \right|_{r \rightarrow \infty} \quad b := \lim_{r \rightarrow \infty} r \sin |\varphi(r) - \varphi(\infty)|$$



# Scatter angle

Definition

$$\delta\varphi := \int_{-\chi_\infty}^{\chi_\infty} \frac{d\varphi}{d\chi} d\chi - \pi = \delta\varphi^{(0)}(v_\infty, b) + \eta \delta\varphi^{(1)}(v_\infty, b)$$

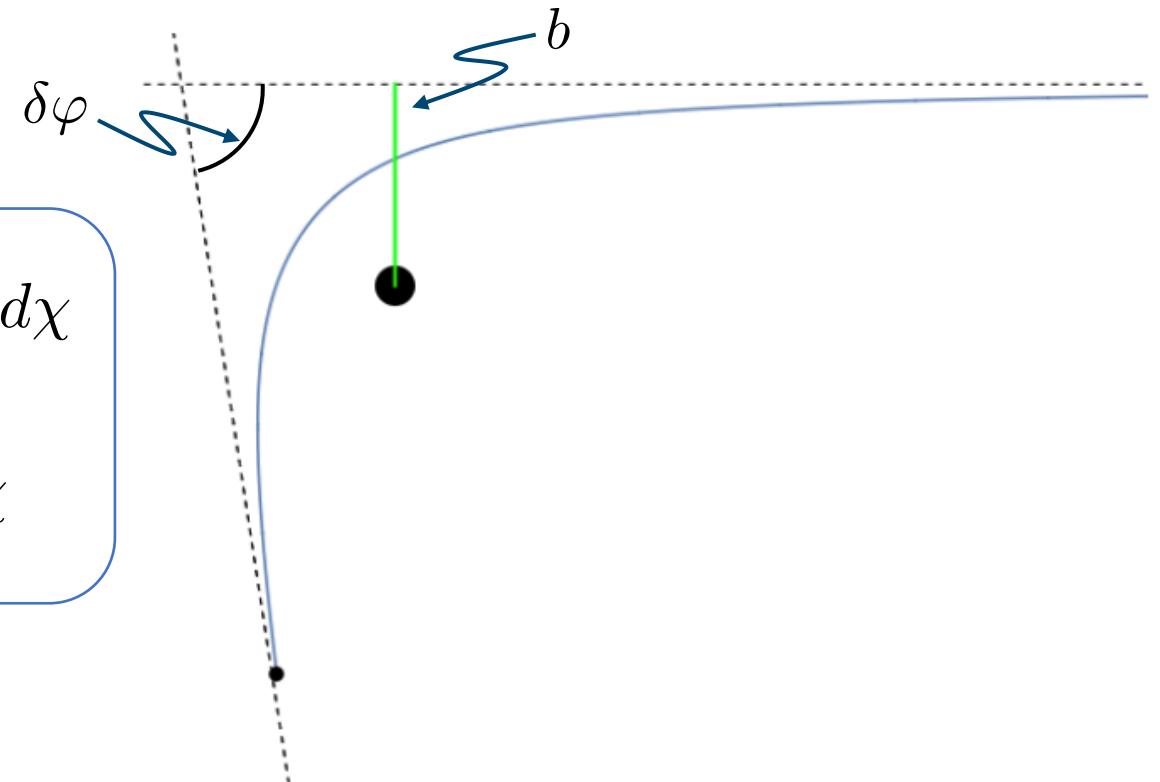
geodesic

Solve 1<sup>st</sup> order self-force equations  
of motion at fixed  $v_\infty$  and  $b$

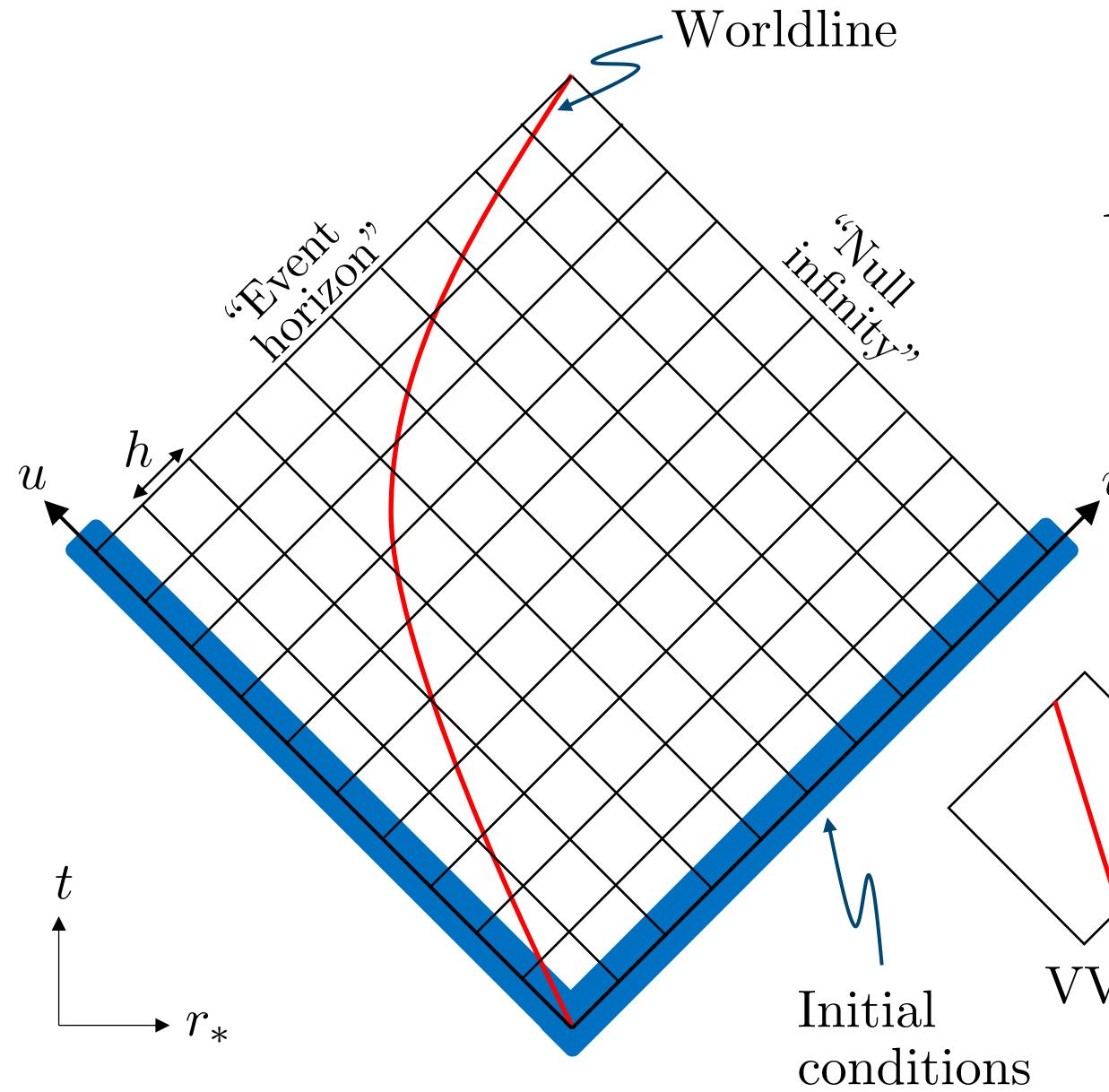
$$\begin{aligned}\delta\varphi^{(1)} &= \mathcal{A}(e, p) \int_{-\chi_\infty}^{\chi_\infty} F_t d\chi + \mathcal{B}(e, p) \int_{-\chi_\infty}^{\chi_\infty} F_\varphi d\chi \\ &+ \int_{-\chi_\infty}^{\chi_\infty} [a(\chi; e, p)F_t + b(\chi; e, p)F_\varphi] d\chi\end{aligned}$$

$$\frac{dE}{d\tau} = -\eta F_t$$

$$\frac{dL}{d\tau} = \eta F_\varphi$$

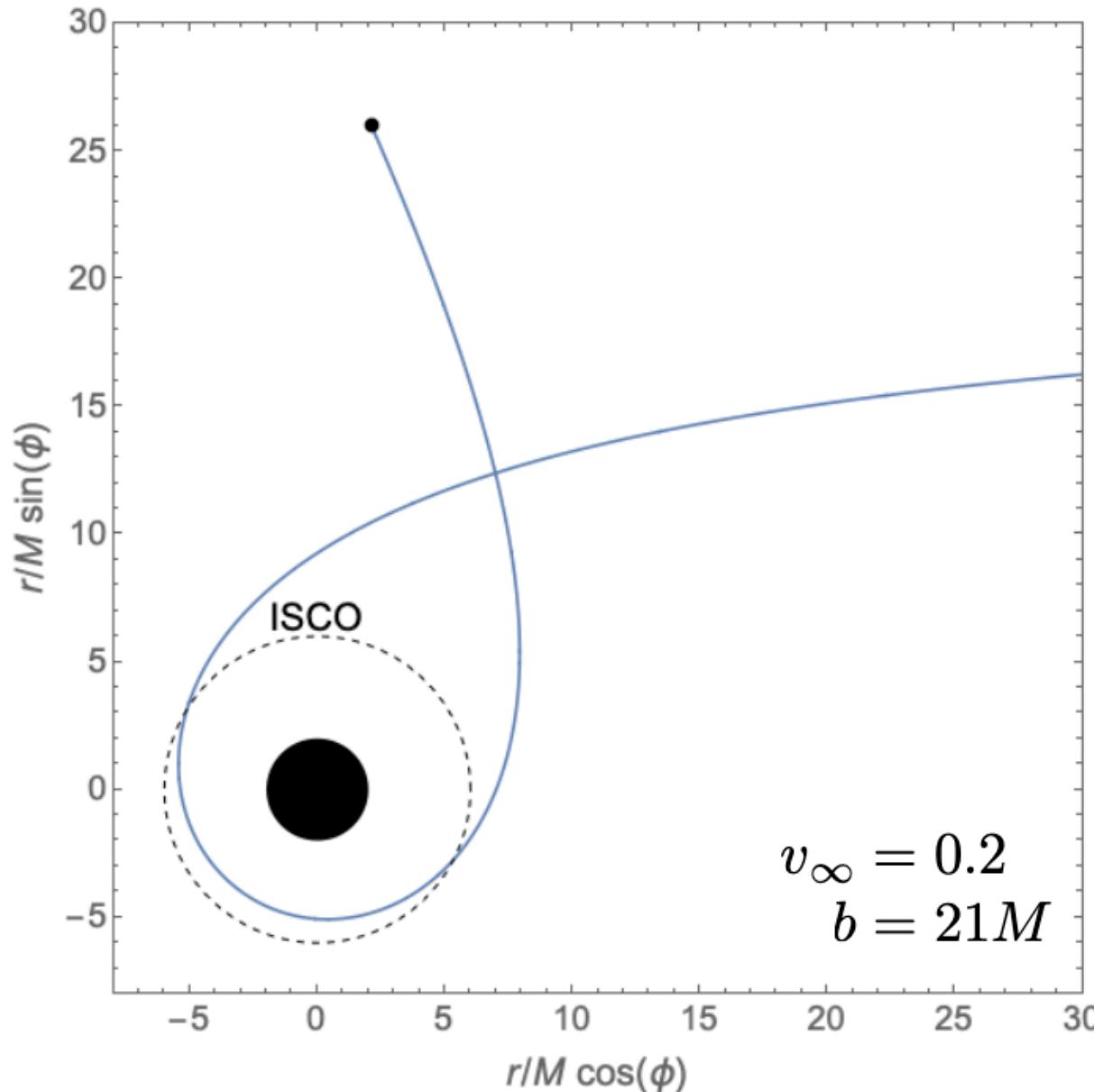


# 1+1D scatter evolution code

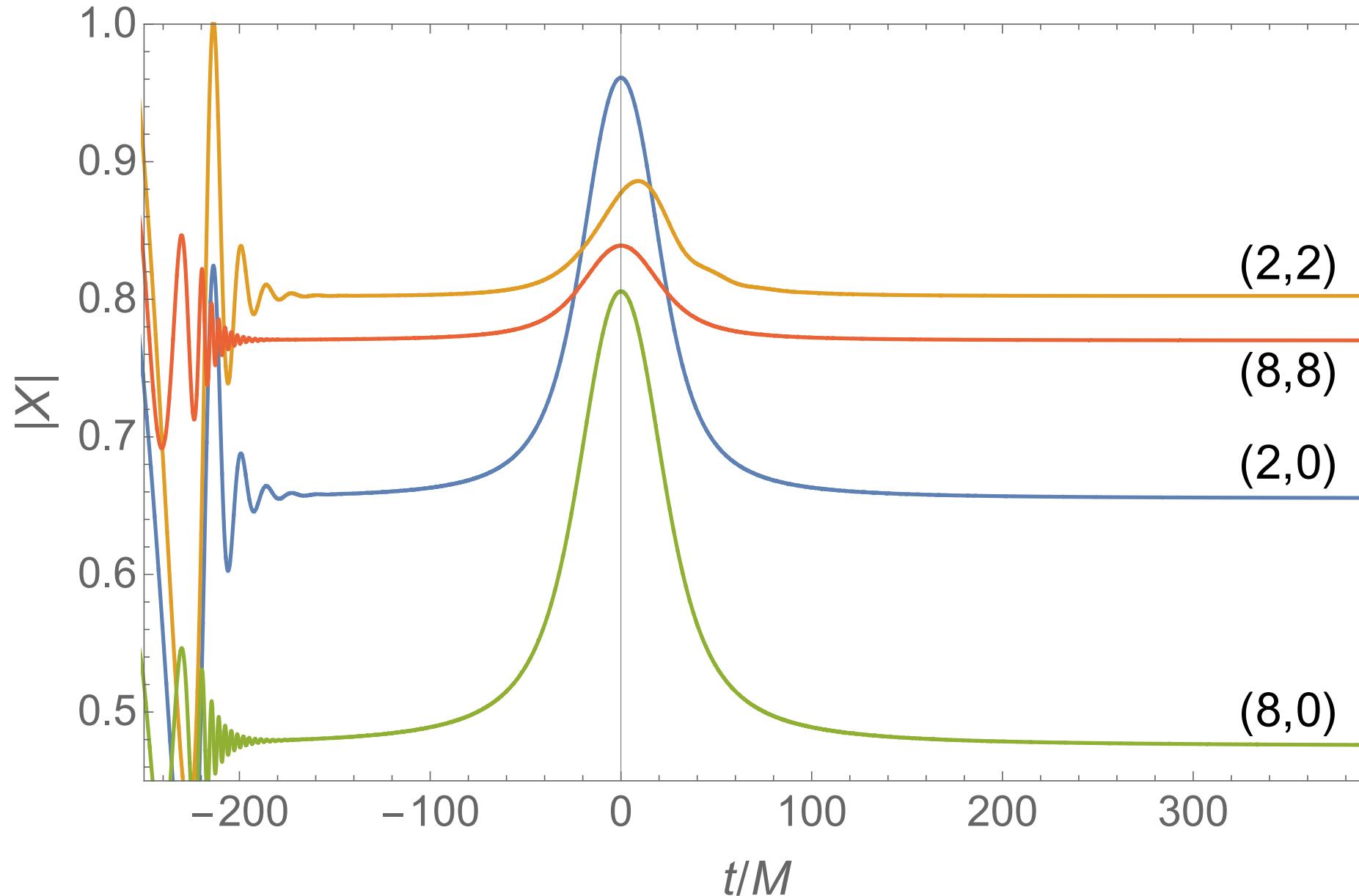


$$X_{00} = \sum_{i,j=0}^1 H_{ij}(h, l; r) X_{ij} + \mathcal{J}([X], h, l; r)$$

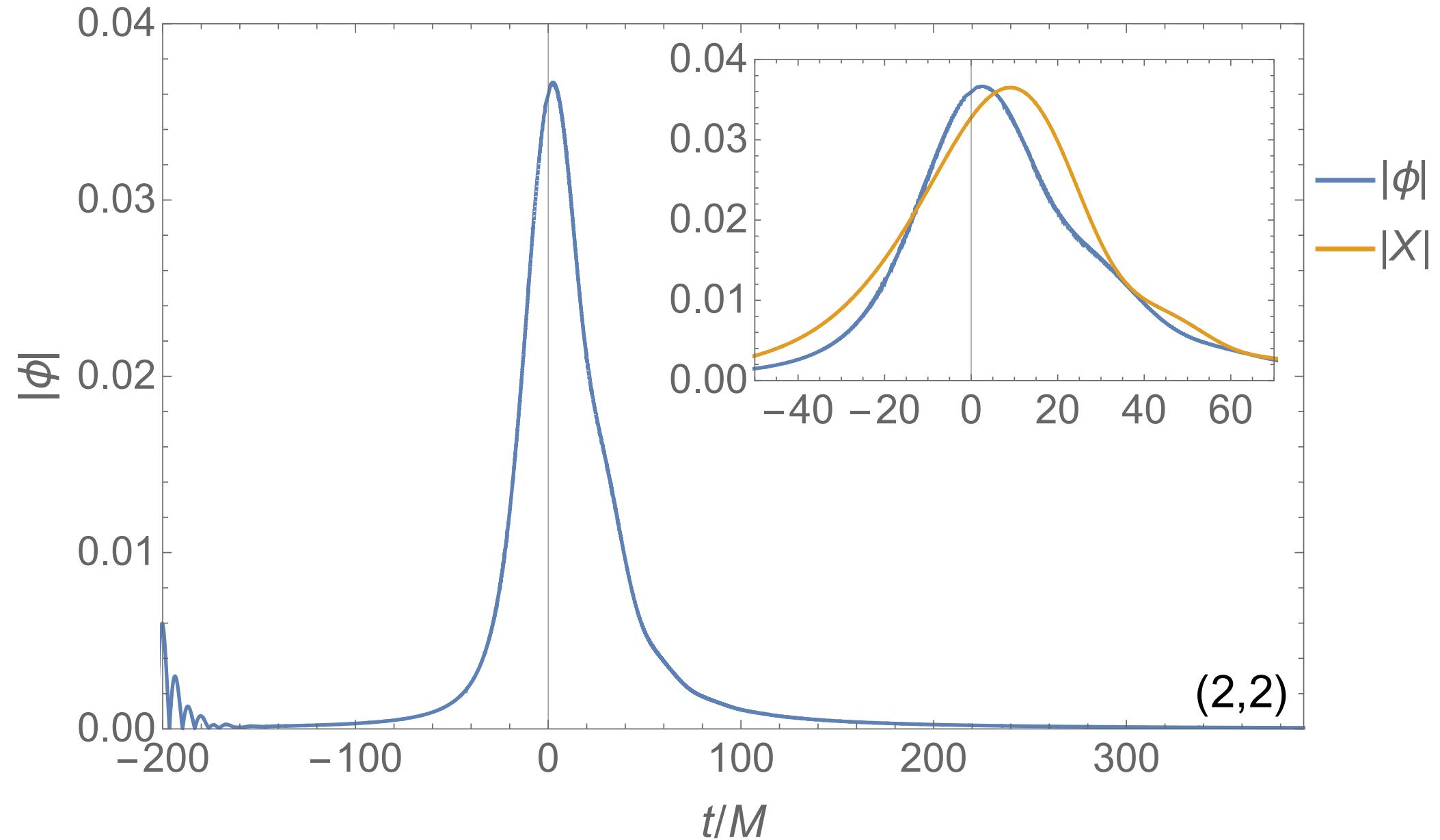
# Hyperbolic orbit



# Regge-Wheeler field along the worldline



# Hertz potential along the worldline



# Summary

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- The 1+1D Teukolsky equation is numerically unstable without boundary conditions
- Transforming to Regge-Wheeler removes the numerical instability
- Self-force scattering results can be used to inform other modelling methods outside of the EMRI regime
- Future work: calculate metric perturbation and obtain self-force data

