

My background and tutorial on diffraction imaging.

Computational Imaging Group (CIG) 2023

Olivier Leblanc.

May 23th, 2023.

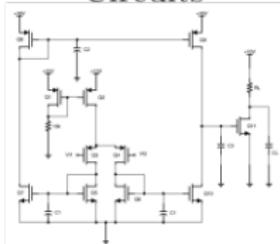


Outline

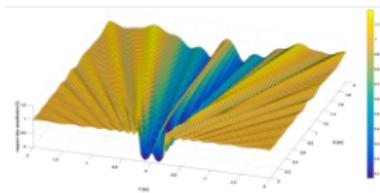
1. Undergrad background
2. PhD - Interferometric Lensless Imaging
3. At WashU: Lippmann-Schwinger for IDT

Undergrad background

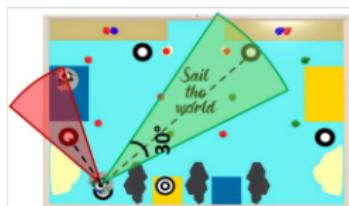
Circuits



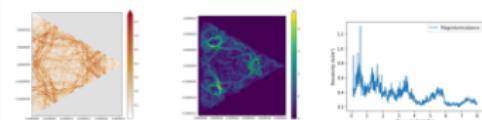
Electromagnetics



Robotics



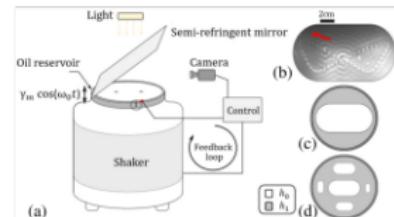
Nanoelectronics



Embedded systems



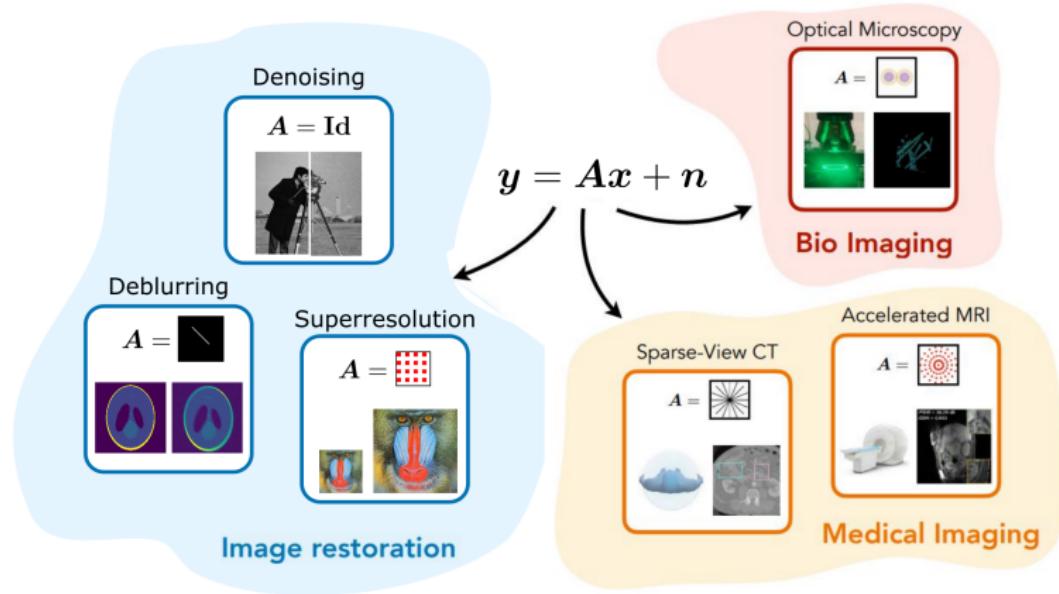
Master thesis



PhD - Interferometric Lensless Imaging

Context - Computational imaging

Many computational imaging problems can be formulated as inverse problems



1

Inverse problem: recover x from the observations y .

¹inspired from Yu Sun

Context - MultiCore Fiber Lensless Imaging (MCF-LI)



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L. Jacques*



M. Hofer†



H. Rigneault†



S. Sivankutty†

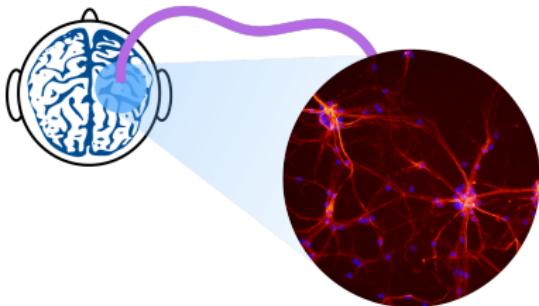
* ISPGGroup, ICTEAM, UCLouvain, Belgium

† Institut Fresnel, Marseille France

‡ CNRS, Université de Lille

Motivation:

- ▶ Deep brain imaging.
- ▶ Limited invasiveness.



Context - MultiCore Fiber Lensless Imaging (MCF-LI)



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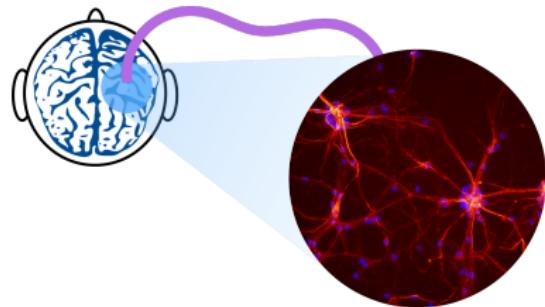
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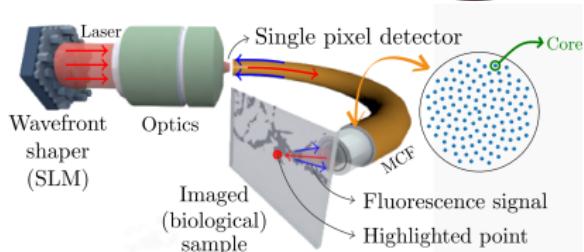
Motivation:

- ▶ Deep brain imaging.
- ▶ Limited invasiveness.



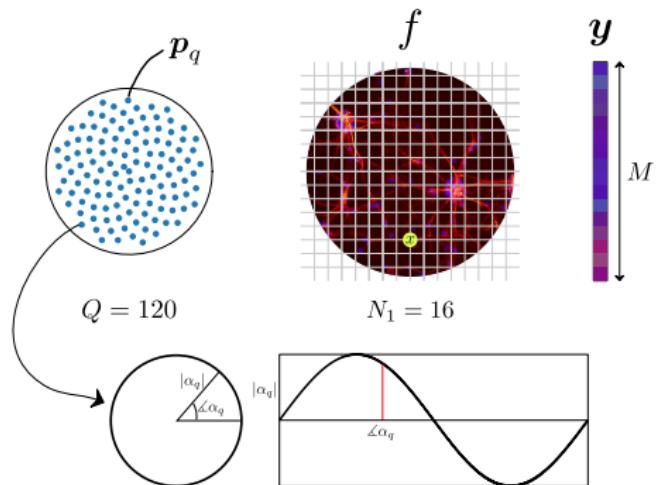
Challenges:

- ▶ Fluorescence imaging.
- ▶ Single-pixel → limited view.
- ▶ Small intensity.



Notations

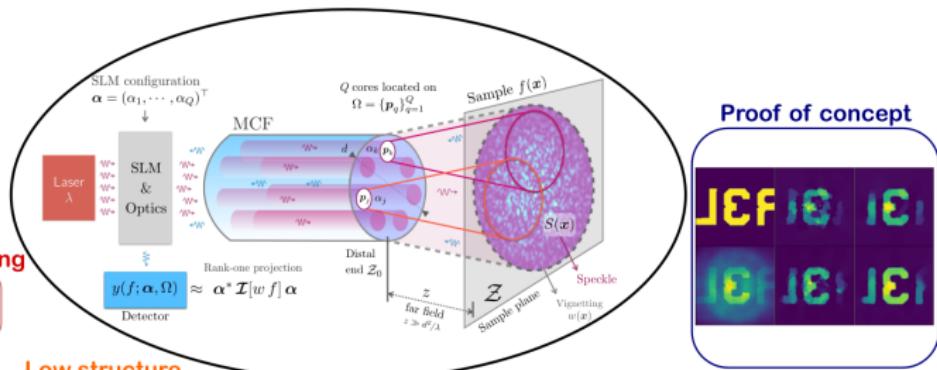
p_q	Position of core q
Q	Number of cores
$\alpha \in \mathbb{C}^Q$	Cores complex amplitudes (tunable!)
$x \in \mathbb{R}^2$	2-D object space
$f(x)$	Object to be imaged
$N = N_1 \times N_1$	Image resolution
M	Number of measurements
$y \in \mathbb{R}_+^M$	Measurement vector



Our contributions

Mathematical modeling

$$\begin{aligned} \mathbf{y} &= \{\alpha_m^* \mathcal{I}_\Omega[wf] \alpha_m\}_{m=1}^M \\ &= \mathcal{A} \circ \mathcal{I}_\Omega[wf] \end{aligned}$$



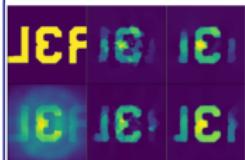
Low structure

Inverse problem
& optimisation

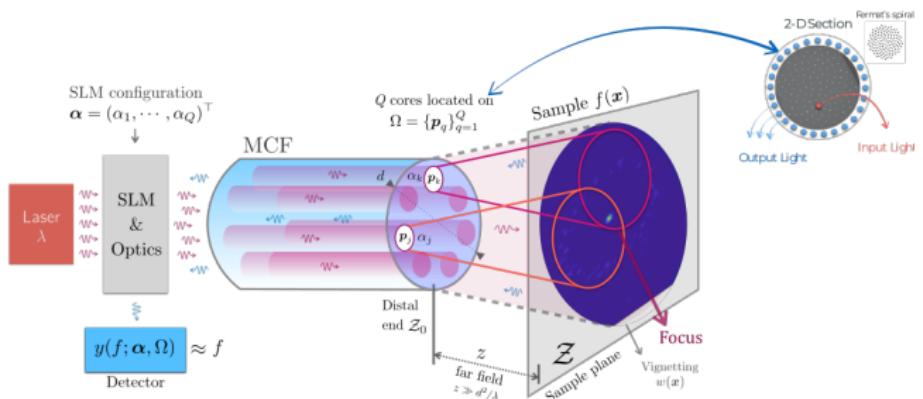
Theoretical guarantees

$$\|f - \hat{f}\|_2 \leq C \frac{\|f - f_K\|_1}{\sqrt{K}} + D \frac{\epsilon}{m}$$

Proof of concept



Raster scanning (RS) mode

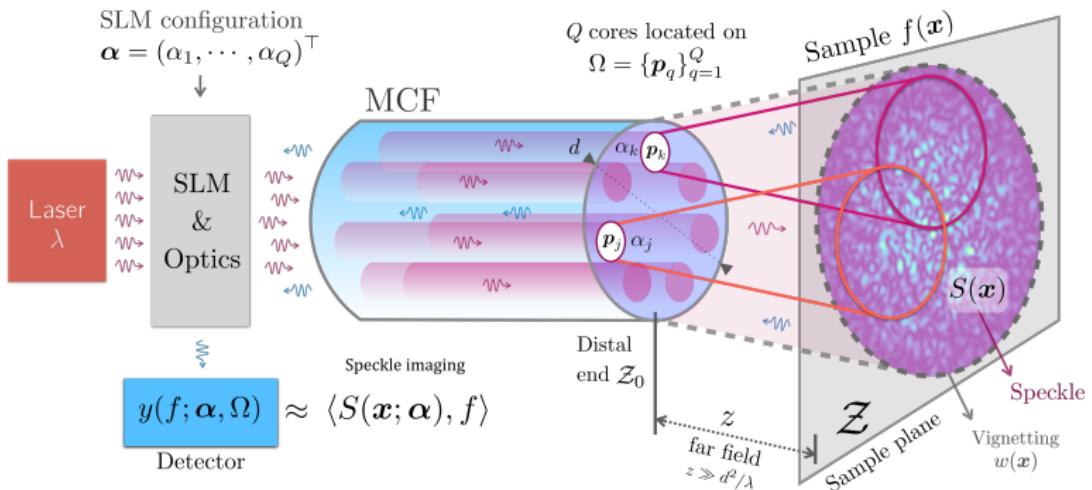


The diagram illustrates the convolution operation in RS mode. It shows three images: a grayscale image x representing the sample, a blue circular mask φ representing the Far field, and a resulting grayscale image y representing the focused output. The operation is represented by the equation $x * \varphi = y$.

Speckle illumination



Steph. Guérit



Classical Compressive Sensing

$$y_m = \mathbf{a}_m^\top \mathbf{x}$$

The diagram illustrates the relationship between observed data y_m , measurement vector \mathbf{a}_m^\top , and signal \mathbf{x} . A horizontal double-headed arrow connects y_m and \mathbf{a}_m^\top . To the right of \mathbf{a}_m^\top is a vertical stack of boxes representing \mathbf{x} . The first few boxes are colored (green, blue, yellow, green, blue, cyan, green, orange), while the rest are white. Ellipses indicate the continuation of the signal \mathbf{x} .

$$y_m = \mathbf{a}_m^\top \mathbf{x} = \langle \mathbf{a}_m, \mathbf{x} \rangle$$

Classical Compressive Sensing

$$\begin{matrix} \mathbf{y} \\ M \end{matrix} = \boxed{\mathbf{A}}_{M \times N} \begin{matrix} \mathbf{x} \\ N \end{matrix}$$

The diagram illustrates the mathematical model of compressive sensing. On the left, a vertical vector \mathbf{y} is shown with a multi-colored bar at its top, followed by M entries. An equals sign follows. In the center, a matrix \mathbf{A} is depicted as a $M \times N$ grid of colored blocks (blue, green, yellow, red). A dashed rectangular box encloses the bottom N rows of \mathbf{A} . To the right, a vertical vector \mathbf{x} is shown with a multi-colored bar at its top, followed by N entries.

Condition: a_{mn} are iid random variables (e.g., $a_{mn} \underset{\text{i.i.d.}}{\sim} \mathcal{N}(0, 1)$).

Speckle illumination: model

The m -th single pixel observation is the integration of the m -th speckle pattern \mathbf{s}_m illuminating the sample \mathbf{f} :

$$y_m = \mathbf{s}_m^\top \mathbf{f} + n_m.$$



Classical compressive sensing model (Guerit et al.)

$$\mathbf{y} = \mathbf{S}\mathbf{f} + \mathbf{n}$$

$$\hat{\mathbf{f}} = \arg \min_{\mathbf{f}} \frac{1}{2} \|\mathbf{y} - \mathbf{S}\mathbf{f}\|_2^2 + \mathcal{R}(\mathbf{f})$$

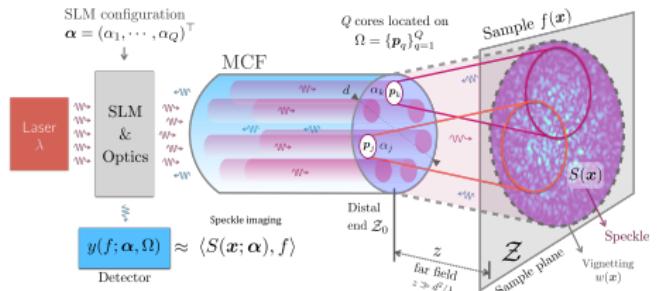
Compressive imaging \Rightarrow number of observations $M \ll N$ image resolution.

Assumption: all coefficients in \mathbf{S} are i.i.d. random variables.

Interferometric imaging

Classical compressive sensing model

$$y_{\alpha} = \int_{\mathbb{R}^2} S_{\alpha}(\mathbf{x}) f(\mathbf{x}) d\mathbf{x} = \langle S_{\alpha}, f \rangle$$



Speckles are interferences (Under far-field approximation)

$$E_q(\mathbf{x}) = \alpha_q \sqrt{w(\mathbf{x})} e^{\frac{2\pi i}{\lambda z} \mathbf{p}_q^\top \mathbf{x}} \quad (\text{Rayleigh-Sommerfeld})$$

$$S(\mathbf{x}; \alpha) \propto w(\mathbf{x}) \left| \sum_{q=1}^Q \alpha_q e^{\frac{2\pi i}{\lambda z} \mathbf{p}_q^\top \mathbf{x}} \right|^2 = \underbrace{w(\mathbf{x})}_{\text{Field of view}} \sum_{j, k=1}^Q \alpha_j \alpha_k^* e^{\frac{2\pi i}{\lambda z} (\mathbf{p}_j - \mathbf{p}_k)^\top \mathbf{x}}$$

Hence (noiseless)

$$\begin{aligned} y_{\alpha} &= \sum_{j, k=1}^Q \alpha_j \alpha_k^* \left[\int_{\mathbb{R}^2} e^{\frac{2\pi i}{\lambda z} (\mathbf{p}_j - \mathbf{p}_k)^\top \mathbf{x}} w(\mathbf{x}) f(\mathbf{x}) d\mathbf{x} \right] \\ &= \alpha^* \mathcal{I}_{\Omega} [wf] \alpha = \langle \alpha \alpha^*, \mathcal{I}_{\Omega} [wf] \rangle \end{aligned}$$

Interferometric imaging

Writing $f^\circ := wf$ we have

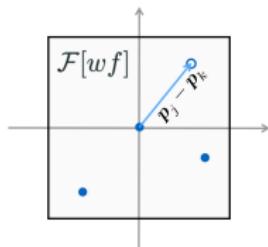
$$y_\alpha = \alpha^* \mathcal{I}_\Omega[f^\circ] \alpha$$

with the interferometric matrix $\mathcal{I}_\Omega[f^\circ] \in \mathbb{C}^{Q \times Q}$ s.t.

$$(\mathcal{I}_\Omega[f^\circ])_{j,k} := \int_{\mathbb{R}^2} e^{\frac{2\pi i}{\lambda z} (\mathbf{p}_j - \mathbf{p}_k)^\top} f^\circ(\mathbf{x}) d\mathbf{x}$$

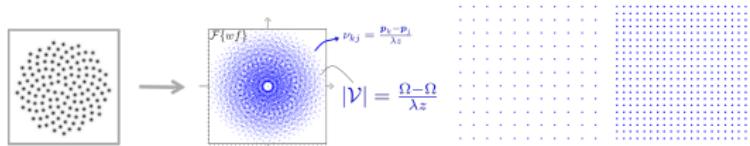
Observation: denser Fourier sampling if

$$|\{\mathbf{p}_j - \mathbf{p}_k : \forall 1 \leq j, k \leq Q\}| \simeq Q^2$$



Fourier plane

- ▶ Lattices are bad core arrangements
- ▶ Fermat's spiral is not bad



$$\begin{aligned}\mathbf{y} &= \{\boldsymbol{\alpha}_m^* \mathcal{I}_\Omega[f^\circ] \boldsymbol{\alpha}_m\}_{m=1}^M \\ &= \mathcal{A} \circ \mathcal{I}_\Omega[f^\circ]\end{aligned}$$

Two-component sensing!

\mathcal{A} : Symmetric rank-one projections of a matrix.

\mathcal{I}_Ω : Partial Fourier sensing with replacement.

Objective

Show

$$\hat{\mathbf{f}} \in \arg \min_{\mathbf{u}} \|\mathbf{u}\|_1 \quad \text{s.t. } \|\mathbf{y} - \mathcal{A}(\mathcal{I}_{\Omega}[\mathbf{u}])\|_1 \leq \epsilon \quad (\text{BPDN}_{\ell_1})$$

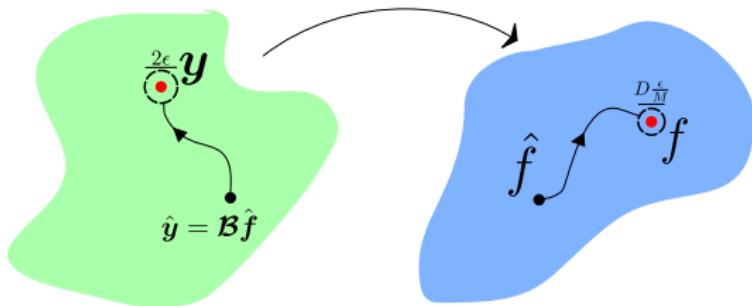
satisfies

$$\|\mathbf{f} - \hat{\mathbf{f}}\|_2 \leq C \frac{\|\mathbf{f} - \mathbf{f}_K\|_1}{\sqrt{K}} + D \frac{\epsilon}{m}.$$

Prove the RIP- ℓ_2/ℓ_1

$$\tilde{A}\|\mathbf{u}\| \leq \frac{1}{M}\|\mathcal{A}(\mathcal{I}_{\Omega}[\mathbf{u}])\|_1 \leq \tilde{B}\|\mathbf{u}\|$$

Measurement space **Object space**



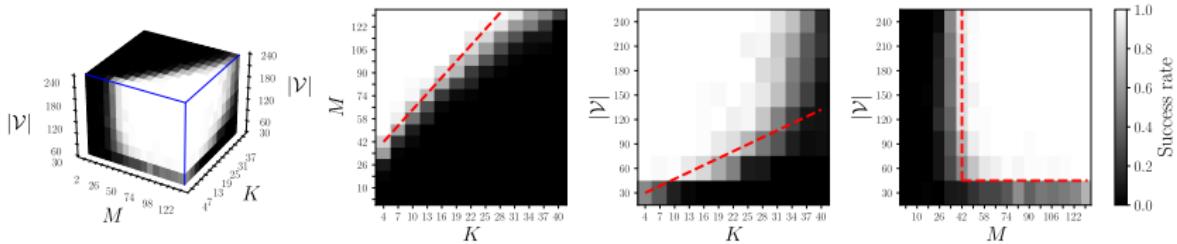
Simulation results

K = sparsity of f , $f \in \Sigma_K$.

$|\mathcal{V}|$ = cardinality of the visibility set in the Fourier space, grows with Q .

M = number of measurements.

$$\hat{\mathbf{f}} = \arg \min_{\mathbf{u}} \|\mathbf{y} - \mathcal{B}\mathbf{u}\|_2 \quad \text{s.t.} \quad \|\mathbf{u}\|_1 \leq K \quad (\text{LASSO})$$

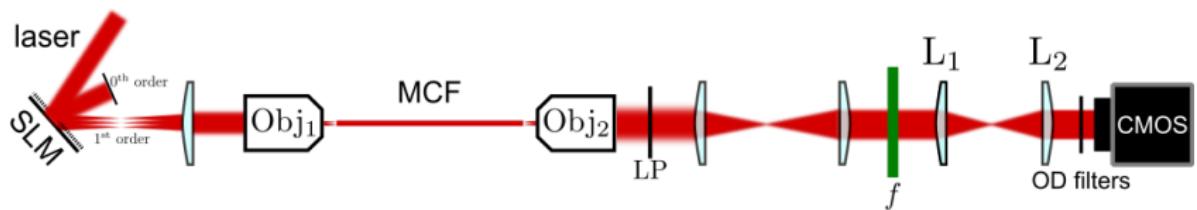
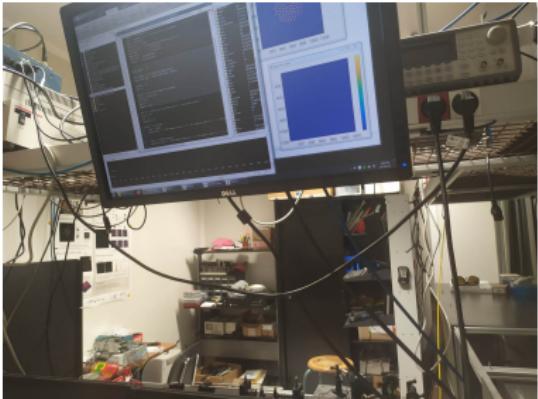
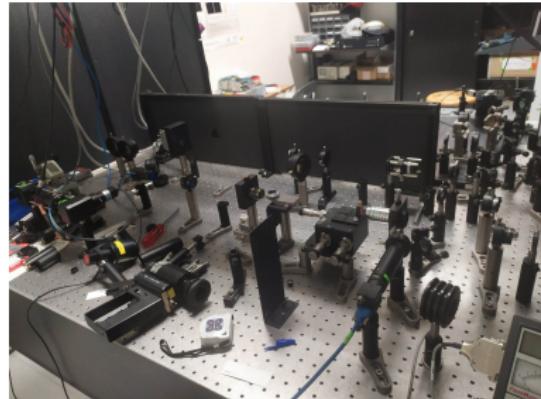


$$M > CK \log \left(\frac{12eN}{K} \right)$$

$$|\mathcal{V}_0| = Q(Q-1) \geq \delta^{-2} K \text{plog}(N, K, \delta)$$

Success if SNR > 40dB.

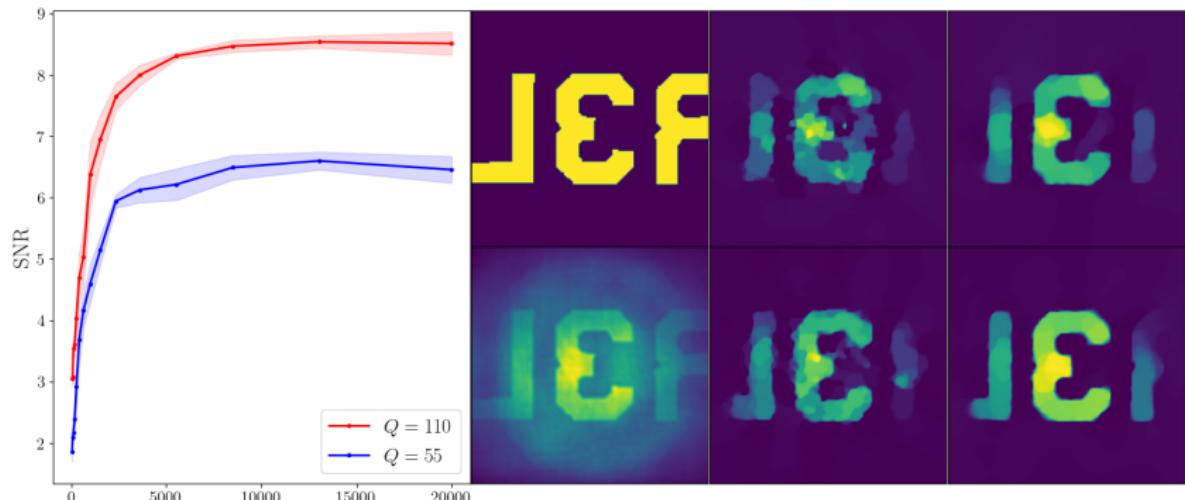
Experimental results



Experimental results



$$\hat{\mathbf{f}} = \arg \min_{\mathbf{u}} \frac{1}{2} \left\| \mathbf{y} - \hat{\mathcal{B}}(\mathbf{u}) \right\|_2 + \rho \|\mathbf{u}\|_{\text{TV}} \quad \text{s.t. } \mathbf{u} \geq 0, \quad (1)$$



At WashU: Lippmann-Schwinger
for IDT

Outline

1. Undergrad background
2. PhD - Interferometric Lensless Imaging
3. At WashU: Lippmann-Schwinger for IDT
 - 3.1 From Maxwell to Wave equation
 - 3.2 From Wave to Lippmann-Schwinger equation
 - 3.3 Discretization and inverse problem

Context - Intensity Diffraction Tomography (IDT)

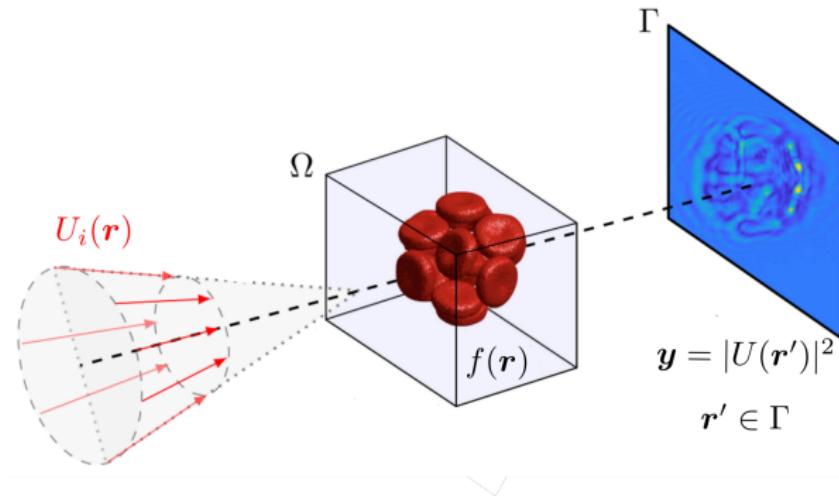


Figure 1: Inspired from T. Pham.

Can we get $y = h(U_i, f)$?

From Maxwell to Wave equation

$$\nabla \cdot \mathbf{D} = \rho, \quad (\text{Coulomb})$$

$$\nabla \cdot \mathbf{B} = 0, \quad (\text{Gauss})$$

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}, \quad (\text{Ampère})$$

$$\nabla \times \mathbf{H} = \frac{\partial \mathbf{D}}{\partial t} + \mathcal{J}, \quad (\text{Faraday})$$

► \mathbf{H} is the magnetic excitation.

► \mathbf{B} is the magnetic field.

► \mathbf{D} is the displacement field.

► \mathbf{E} is the electric field.

□: simplifying assumption.

□: rewriting.

From Maxwell to Wave equation

$$\nabla \cdot \mathbf{D} = \rho, \quad (\text{Coulomb})$$

$$\nabla \cdot \mathbf{B} = 0, \quad (\text{Gauss})$$

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}, \quad (\text{Ampère})$$

$$\nabla \times \mathbf{H} = \frac{\partial \mathbf{D}}{\partial t} + \mathbf{0}, \quad (\text{Faraday})$$

No current density (no moving charged particle): $\mathcal{J} = 0$

From Maxwell to Wave equation

$$\nabla \cdot \mathbf{D} = \rho$$

$$\nabla \cdot \mathbf{B} = 0$$

$$\nabla \times \left(\frac{1}{\mu} \nabla \times \mathbf{E} \right) = -\nabla \times \left(\frac{1}{\mu} \frac{\partial \mathbf{B}}{\partial t} \right)$$

$$\frac{\partial}{\partial t} (\nabla \times \mathbf{H}) = \frac{\partial}{\partial t} \left(\frac{\partial \mathbf{D}}{\partial t} \right)$$

From Maxwell to Wave equation

$$\begin{aligned}\nabla \cdot (\epsilon \mathbf{E}) &= \rho \\ \nabla \cdot (\mu \mathbf{H}) &= 0 \\ \nabla \times \left(\frac{1}{\mu} \nabla \times \mathbf{E} \right) &= -\nabla \times \left(\frac{1}{\mu} \frac{\partial \mu \mathbf{H}}{\partial t} \right) \\ \frac{\partial}{\partial t} (\nabla \times \mathbf{H}) &= \frac{\partial}{\partial t} \left(\frac{\partial \epsilon \mathbf{E}}{\partial t} \right)\end{aligned}$$

Linear optics: $\mathbf{B} = \mu \mathbf{H}$, $\mathbf{D} = \epsilon \mathbf{E}$

From Maxwell to Wave equation

$$\begin{aligned}\nabla \cdot (\epsilon \mathbf{E}) &= \rho \\ \nabla \cdot (\mu \mathbf{H}) &= 0 \\ \nabla \times \left(\frac{1}{\mu} \nabla \times \mathbf{E} \right) &= -\nabla \times \left(\mathbf{H} \frac{\partial \mathbf{H}}{\partial t} \right) \\ \nabla \times \left(\frac{\partial \mathbf{H}}{\partial t} \right) &= \epsilon \frac{\partial^2 \mathbf{E}}{\partial t^2}\end{aligned}$$

$$\frac{\partial}{\partial t} (\nabla \times \mathbf{B}) = \nabla \times \left(\frac{\partial \mathbf{B}}{\partial t} \right), \quad \mu, \epsilon \perp\!\!\!\perp t$$

From Maxwell to Wave equation

$$\nabla \cdot (\epsilon \mathbf{E}) = \rho$$

~~$$\nabla \cdot (\mu \mathbf{H}) = 0$$~~

$$\nabla \times \left(\frac{1}{\mu} \nabla \times \mathbf{E} \right) = -\nabla \times \left(\frac{\partial \mathbf{H}}{\partial t} \right)$$

$$\nabla \times \left(\frac{\partial \mathbf{H}}{\partial t} \right) = \epsilon \ddot{\mathbf{E}}$$

From Maxwell to Wave equation

$$\nabla \cdot (\epsilon \mathbf{E}) = \rho$$

$$\nabla \cdot (\mu \mathbf{H}) = 0$$

$$\nabla \times \left(\frac{1}{\mu} \nabla \times \mathbf{E} \right) = -\epsilon \ddot{\mathbf{E}}$$

From Maxwell to Wave equation

$$\nabla \cdot (\epsilon \mathbf{E}) = \rho$$

$$\nabla \times \left(\frac{1}{\mu} \nabla \times \mathbf{E} \right) = -\epsilon \ddot{\mathbf{E}}$$

From Maxwell to Wave equation

$$\boxed{\nabla \cdot (\epsilon \mathbf{E}) = \rho}$$

$$\nabla \times \left(\frac{1}{\mu} \nabla \times \mathbf{E} \right) = -\epsilon \ddot{\mathbf{E}}$$



stationary conditions \rightarrow get rid of time:

$$\mathbf{E}(\mathbf{r}, t) = \frac{1}{\sqrt{\epsilon}} \mathbf{E}(\mathbf{r}) e^{-i\omega t} \implies \ddot{\mathbf{E}}(\mathbf{r}, t) = -\omega^2 \mathbf{E}(\mathbf{r}, t)$$



$$\nabla^2 \mathbf{E} + k^2 n^2 \mathbf{E} + \nabla \log(\mu) \times (\nabla \times \mathbf{E}) + \nabla \left(\frac{\rho}{\epsilon} \right) + \nabla (\mathbf{E} \cdot \nabla (\log(\epsilon))) = 0$$

From Maxwell to Wave equation

$$\nabla^2 \mathbf{E} + k^2 n^2 \mathbf{E} + \nabla \log(\mu) \times (\nabla \times \mathbf{E}) + \nabla \left(\frac{\rho}{\epsilon} \right) + \nabla (\mathbf{E} \cdot \nabla (\log(\epsilon))) = 0$$

From Maxwell to Wave equation

$$\nabla^2 \mathbf{E} + k^2 n^2 \mathbf{E} + \nabla \log(\mu) \times (\nabla \times \mathbf{E}) + \cancel{\frac{\nabla \rho}{\epsilon}} + \nabla (\mathbf{E} \cdot \nabla (\log(\epsilon))) = 0$$

No charge density: $\rho = 0$

From Maxwell to Wave equation

$$\nabla^2 \mathbf{E} + k^2 n^2 \mathbf{E} + \cancel{\nabla \log(\mu) \times (\nabla \times \mathbf{E})} + \nabla \left(\frac{\rho}{\epsilon} \right) + \nabla (\mathbf{E} \cdot \nabla (\log(\epsilon))) = 0$$

No charge density: $\rho = 0$

Non magnetic medium: $\mu(r) = \mu_0 \perp r$

From Maxwell to Wave equation

$$\nabla^2 \mathbf{E} + k^2 n^2 \mathbf{E} + \cancel{\nabla \log(\mu) \times (\nabla \times \mathbf{E})} + \cancel{\nabla \left(\frac{\rho}{\epsilon} \right)} + \underline{\nabla (\mathbf{E} \cdot \nabla (\log(\epsilon)))} = 0$$

No charge density: $\rho = 0$

Non magnetic medium: $\mu(r) = \mu_0 \perp r$

Slow spatial variations of ϵ : $\nabla (\mathbf{E} \cdot \nabla (\log(\epsilon))) \ll \nabla^2 \mathbf{E}$

From Maxwell to Wave equation

$$\boxed{\nabla^2 \mathbf{E} + k^2 n^2 \mathbf{E} = 0} \quad (\text{Wave eq.})$$

From Maxwell to Wave equation

$$\boxed{\nabla^2 \mathbf{E} + k^2 n^2 \mathbf{E} = 0} \quad (\text{Wave eq.})$$

With $\mathbf{E} = (E_x, E_y, E_z)$

$$\begin{cases} \nabla^2 E_x + k^2 n^2 E_x = 0 \\ \nabla^2 E_y + k^2 n^2 E_y = 0 \\ \nabla^2 E_z + k^2 n^2 E_z = 0. \end{cases}$$

Write U for any component E_x, E_y or E_z .

$$\boxed{\nabla^2 U + k^2 n^2 U = 0} \quad (\text{Scalar Wave eq.})$$

From Maxwell to Wave equation

$$\nabla^2 U + k^2 n^2 U = 0$$

(Scalar Wave eq.)

From Maxwell to Wave equation

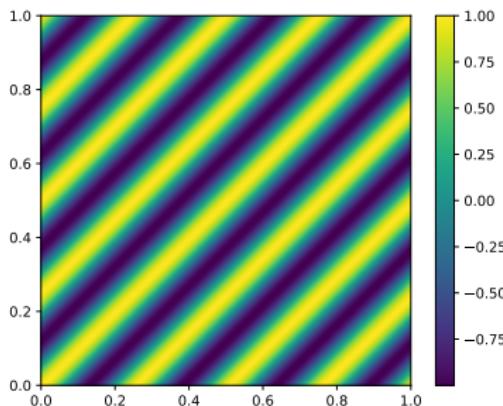
$$(\nabla^2 + k^2 n(\mathbf{r})^2) U(\mathbf{r}) = 0$$

(Scalar Wave eq.)

Example: plane wave

(with $n(\mathbf{r}) = 1$, $\forall \mathbf{r} \in \mathbb{R}^3$):

$$U(\mathbf{r}) = e^{i\mathbf{k}^\top \mathbf{r}}$$



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From Wave to Lippmann-Schwinger equation

$$\boxed{(\nabla^2 + k^2 n(\mathbf{r})^2) U(\mathbf{r}) = 0}$$

k is the wavenumber in vacuum.

$k_m := kn_m$ is the wavenumber in medium m with RI n_m .

$$\left(\nabla^2 + k_m^2 \left(\frac{n(\mathbf{r})}{n_m} \right)^2 \right) U(\mathbf{r}) = 0$$

$$(\nabla^2 + k_m^2)U(\mathbf{r}) = -f(\mathbf{r})U(\mathbf{r})$$

with the scattering potential

$$f(\mathbf{r}) := k_m^2 \left[\left(\frac{n(\mathbf{r})}{n_m} \right)^2 - 1 \right].$$

From Wave to Lippmann-Schwinger equation

$$(\nabla^2 + k_m^2)U(\mathbf{r}) = -f(\mathbf{r})U(\mathbf{r})$$

Separation into homogeneous and inhomogeneous equations. $U(\mathbf{r}) = U_i(\mathbf{r}) + U_s(\mathbf{r})$ with:

$$(\nabla^2 + k_m^2)U_i(\mathbf{r}) = 0$$

$$(\nabla^2 + k_m^2)U_s(\mathbf{r}) = -f(\mathbf{r})U(\mathbf{r})$$

U_i is the incident field.

U_s is the scattered field.

From Wave to Lippmann-Schwinger equation

$$(\nabla^2 + k_m^2) U_s(\mathbf{r}) = -f(\mathbf{r})U(\mathbf{r})$$

The Green's function is defined as:

$$(\nabla^2 + k_m^2) G(\mathbf{r}) = -\delta(\mathbf{r})$$

Now

$$\int (\nabla^2 + k_m^2) G(\mathbf{r} - \mathbf{r}') f(\mathbf{r}') U(\mathbf{r}') d\mathbf{r}' = \int -\delta(\mathbf{r} - \mathbf{r}') f(\mathbf{r}') U(\mathbf{r}') d\mathbf{r}'$$

so that

$$U_s(\mathbf{r}) = \int_{\Omega} f(\mathbf{r}') U(\mathbf{r}') G(\mathbf{r} - \mathbf{r}') d\mathbf{r}'$$

Finally

$$U(\mathbf{r}) = U_i(\mathbf{r}) + \int_{\Omega} f(\mathbf{r}') U(\mathbf{r}') G(\mathbf{r} - \mathbf{r}') d\mathbf{r}' \quad \text{(Lippmann-Schwinger)}$$

Born approximation and interpretation

$$U(\mathbf{r}) = U_i(\mathbf{r}) + \int_{\Omega} f(\mathbf{r}') U(\mathbf{r}') G(\mathbf{r} - \mathbf{r}') d\mathbf{r}'$$

(Lippmann-Schwinger)

Born approximation and interpretation

$$U(\mathbf{r}) = U_i(\mathbf{r}) + \int_{\Omega} f(\mathbf{r}') U(\mathbf{r}') G(\mathbf{r} - \mathbf{r}') d\mathbf{r}'$$

(Lippmann-Schwinger)

$$\mathbf{U}_1(\mathbf{r}) = U_i(\mathbf{r}) + \int_{\Omega} f(\mathbf{r}') \mathbf{U}_i(\mathbf{r}') G(\mathbf{r} - \mathbf{r}') d\mathbf{r}'$$

(1st Born approx.)

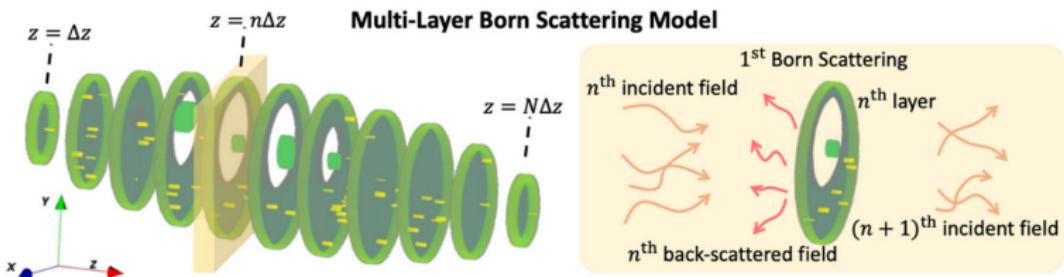
$$\mathbf{U}_{n+1}(\mathbf{r}) = U_i(\mathbf{r}) + \int_{\Omega} f(\mathbf{r}') \mathbf{U}_n(\mathbf{r}') G(\mathbf{r} - \mathbf{r}') d\mathbf{r}'$$

(n-th Born approx.)

$$\mathbf{U}_{n+1}(\mathbf{r}) = \mathbf{U}_n(\mathbf{r}) + \int_{\Omega} f(\mathbf{r}') \mathbf{U}_n(\mathbf{r}') G(\mathbf{r} - \mathbf{r}') d\mathbf{r}'$$

(multi-layer Born approx.)

Interpretation:



Outline

1. Undergrad background
2. PhD - Interferometric Lensless Imaging
3. At WashU: Lippmann-Schwinger for IDT
 - 3.1 From Maxwell to Wave equation
 - 3.2 From Wave to Lippmann-Schwinger equation
 - 3.3 Discretization and inverse problem

Discretizing the 1st Born approx.

$$U(\mathbf{r}) = U_i(\mathbf{r}) + \int_{\Omega} f(\mathbf{r}') U_i(\mathbf{r}') G(\mathbf{r} - \mathbf{r}') d\mathbf{r}' \quad (1\text{st Born approx.})$$

$$\mathbf{U} = \mathbf{U}_i + \mathbf{G} \operatorname{diag}(\mathbf{U}_i) \mathbf{f}$$

$$\mathbf{y} = |\mathbf{U}|^2 = |\mathbf{U}_i|^2 + 2\Re\{\mathbf{U}_i^* \mathbf{G} \operatorname{diag}(\mathbf{U}_i) \mathbf{f}\} + \text{negl.}$$

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$$\mathbf{y}' = |\mathbf{U}|^2 - |\mathbf{U}_i|^2 = 2\Re\{\mathbf{U}_i^* \mathbf{G} \operatorname{diag}(\mathbf{U}_i) \mathbf{f}\} + \text{negl.}$$

Remove background intensity

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Remove background intensity

Remind

$$f(\mathbf{r}) = k_m^2 [n^2(\mathbf{r}) - n_m^2] = k_m^2 [\epsilon(\mathbf{r}) - \epsilon_m] \mu_0 = k_m^2 \mu_0 \Delta \epsilon(\mathbf{r})$$

Hence,

$$\mathbf{y}' = \mathbf{A} \Delta \epsilon$$

\mathbf{A} is nice and fast \rightarrow FFT implementation.

Inverse problem

Try to estimate the volume of RI as

$$\widetilde{\Delta \epsilon} = \arg \min_{\boldsymbol{x}} \frac{1}{2} \|\boldsymbol{y}' - \boldsymbol{A}\boldsymbol{x}\|_2^2 + \mathcal{R}(\boldsymbol{x})$$

Inverse problem

Try to estimate the volume of RI as

$$\widetilde{\Delta\epsilon} = \arg \min_{\mathbf{x}} \frac{1}{2} \|\mathbf{y}' - \mathbf{Ax}\|_2^2 + \mathcal{R}(\mathbf{x})$$

Eventually use Neural Fields to get a continuous representation $\Delta\epsilon = \mathcal{M}_\phi(\mathbf{r})$ and solve

$$\tilde{\phi} = \arg \min_{\phi} \frac{1}{2} \|\mathbf{y}' - \mathbf{Ax}\|_2^2 + \mathcal{R}(\mathbf{x}) \text{ s.t. } \mathbf{x} = \mathcal{M}_\phi(\mathbf{r}).$$

Inverse problem

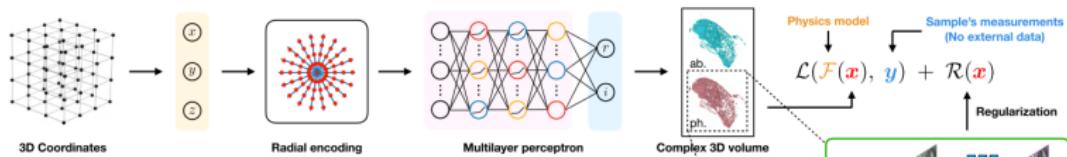
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This is the approach used in DeCaf¹.



¹R. Liu, Y. Sun, J. Zhu, L. Tian, U. Kamilov, "Recovery of Continuous 3D Refractive Index Maps from Discrete Intensity-Only Measurements using Neural Fields," *Nature Mach Intell* 4, 781-791 (2022)

Discretizing Lippmann-Schwinger

$$U(\mathbf{r}) = U_i(\mathbf{r}) + \int_{\Omega} f(\mathbf{r}') U(\mathbf{r}') G(\mathbf{r} - \mathbf{r}') d\mathbf{r}'$$

(Lippmann-Schwinger)

$$\mathbf{U} = \mathbf{U}_i + \mathbf{G} \operatorname{diag}(\mathbf{f}) \mathbf{U}$$

Discretizing Lippmann-Schwinger

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$$\mathbf{U} = \mathbf{U}_i + \mathbf{G} \operatorname{diag}(\mathbf{f}) \mathbf{U}$$

$$\mathbf{U} = (\mathbf{I} - \mathbf{G} \operatorname{diag}(\mathbf{f}))^{-1} \mathbf{U}_i$$

Discretizing Lippmann-Schwinger

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$$\mathbf{y} = \mathcal{A}(\Delta\epsilon)$$

\mathcal{A} is less nice. Accurate but slow...

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IDEA 1 Keep Neural Fields approach and solve

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IDEA 2 Can we get a loss differentiable in ϕ with the continuous model?

$$\mathbf{y} = \mathcal{B}(\Delta\epsilon)$$

\mathcal{B} is a continuous-to-discrete operator.

$$\tilde{\phi} = \arg \min_{\phi} \frac{1}{2} \|\mathbf{y}' - \mathcal{B}(\mathbf{x})\|_2^2 + \mathcal{R}(\mathbf{x}) \text{ s.t. } \mathbf{x} = \mathcal{M}_{\phi}(\mathbf{r}).$$

$$U(\mathbf{r}) = U_i(\mathbf{r}) + \int_{\Omega} f(\mathbf{r}') U(\mathbf{r}') G(\mathbf{r} - \mathbf{r}') d\mathbf{r}' \quad (\text{Lippmann-Schwinger})$$

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IDEA 3 Could we train a neural field \mathcal{M}_{ϕ} for multiple RI volumes?

Conclusion

Conclusion

Today, you learned:

- ▶ a little more about me.
- ▶ some diffraction theory.
- ▶ my contribution perspectives.

From now on

- ▶ Your tips? Ideas?
- ▶ Please help me with \mathcal{M}_ϕ :).

Thank you!