

# Comet.Photos: An Interactive Tool for Rapidly Searching and Displaying Rosetta Mission Images by Spatial Location and Other Properties

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DOI: 10.xxxxxx/draft

## Software

- Review
- Repository
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Editor:

Submitted: 20 October 2025  
Published: unpublished

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## Summary

The European Space Agency's Rosetta mission to Comet 67P/Churyumov-Gerasimenko (hereafter, 67P) provided an unparalleled dataset, which has fundamentally reshaped our understanding of comets. Despite these advances, however, the complexity and volume of Rosetta data, coupled with the lack of efficient tools for comprehensive analyses, have hindered its broader utilization. To address this gap, we developed Comet.Photos, an interactive tool for efficiently searching and visualizing images of irregular bodies.

Comet.Photos enables fast, intuitive spatial searches across more than 44,000 high-resolution images of 67P, acquired by the Rosetta Optical, Spectroscopic, and Infrared Remote Imaging System (OSIRIS) Narrow Angle Camera (NAC) and Wide Angle Camera (WAC) (Keller et al., 2007) and the Navigation Camera (NavCam) (Geiger et al., 2021). Users can define a region of interest on a 3D model of the comet by painting with a virtual brush (Figure 1). The application then nearly instantly (< 50 ms in this example) returns a time-ordered list of all images containing that region. Results can be further filtered by pixel scale and viewing geometry (emission, incidence, and phase angle), and displayed either in their original 2D form (Figure 2) or projected onto the 3D shape model (Figure 3).

Designed for both professional researchers and the public, Comet.Photos can be installed locally for fastest performance, but is also accessible in any modern browser without requiring installation. It combines preprocessed data with real-time, client-side filtering to achieve subsecond search speeds, even across large datasets.

## Statement of need

The European Space Agency's (ESA) Rosetta mission to 67P provided the most spatially and temporally comprehensive dataset of any comet to date, enabling a broad range of analyses of its surface. Because each region was imaged multiple times under varying illumination and viewing conditions, studies such as photometric modeling (Fornasier et al., 2023; Oklay et al., 2016), multi-image photoclinometry (Jindal et al., 2024), and surface-change detection (Barrington et al., 2023; Birch et al., 2019; El-Maarry et al., 2017; Fornasier et al., 2017; Groussin et al., 2015; Jindal et al., 2022; Keller et al., 2017) became possible for the first time, substantially advancing our understanding of cometary surface evolution. The sheer number of images, however, makes identifying repeat observations of the same regions a difficult and time-consuming task. This challenge is further compounded by 67P's complex,

highly non-spherical shape and Rosetta's variable orbit, which often caused images of the same region to appear dramatically different from one another.

The scale of this difficulty is illustrated by the fact that it took seven years after the end of the Rosetta mission for the first global catalog of surface changes on 67P to be published (Barrington et al., 2023). Compiling this catalog required manually inspecting more than 20,000 OSIRIS NAC images to identify those showing morphological changes—a monumental task that took over a year to complete! This was followed by an additional year for detailed characterization, involving manual map projection of each image with the shapeViewer software (Vincent, 2018) and then mapping changes in ESRI's ArcGIS software. Given the wide range of imaging geometries, errors in co-registration and change detection were unavoidable. Consequently, despite these efforts, the global change catalog—and, by extension, our understanding of 67P's evolution—knowingly remains incomplete, with numerous consequential surface changes still undocumented, as evidenced by Moser et al. (2025).

Efforts have been made to mitigate these challenges—for example, ESA has introduced an image search capability within their Planetary Science Archive (PSA) (ESA, 2024). This tool, however, remains inadequate (at least for Rosetta), as it (a) is slow, (b) frequently returns incorrect or missing data, and (c) lacks user control over filtering searches by image parameters, a crucial feature for assembling a manageable dataset without wasting time removing irrelevant images. Hence, to fully harness the scientific potential of Rosetta's vast dataset and empower researchers to quickly and accurately identify relevant observations for analysis, an efficient and intuitive tool is needed to streamline image retrieval.

Comet.Photos has been developed to fill this critical gap, providing a powerful solution for rapid, spatially targeted image searches and facilitating detailed studies of cometary surface properties and evolution. Users can select a region of interest by interactively painting the desired region on a 3D model of 67P. In a fraction of a second, the application searches through over 44,000 NAC, WAC, and NavCam images to identify all observations of the selected region. Results can then be filtered by key parameters such as pixel scale, emission angle, incidence angle, and phase angle, allowing users to tailor datasets for photometric or temporal analyses. Images matching the search criteria can be viewed in 2D or projected onto the 3D shape model, and users can export a list of matching image IDs for further processing with external tools.

Although designed primarily for researchers, Comet.Photos also provides an accessible interface for educators, students, and the general public to explore Rosetta's imagery of 67P. By making this tool highly efficient, freely available and simple to use, we aim to broaden participation in cometary research and facilitate new scientific insights from this unprecedented dataset. The application can be installed locally for maximum performance, but can also be run in any modern browser without installation, by visiting <https://comet.photos>.

## Notes on the implementation

Comet.Photos achieves its speed through extensive preprocessing of image metadata. For each image (Geiger et al., 2020; Sierks & OSIRIS Team, 2020, 2022), the preprocessing pipeline first simulates the spacecraft camera view and its relative position with respect to the comet's 3D shape model, a SHAP7 model (Preusker et al., 2017) with 100,002 vertices. Vertex visibility across all simulated scenes is stored in a lookup table, which enables, at runtime, fast bitwise operations to compute the fraction of the region of interest (painted vertices) visible in any image. Additional precomputed metadata—such as spacecraft position, solar geometry, and image resolution—allow rapid filtering by geometric or photometric criteria with minimal computational overhead.

To support both high-performance local execution and remote web access, Comet.Photos adopts a client-server architecture. The client is organized into object-based modules (Filter Engine, 3D Scene Manager, Image Browser, GUI Controller, etc.) that communicate through an event-driven system. The user interface is specified via a declarative schema, fully decoupled

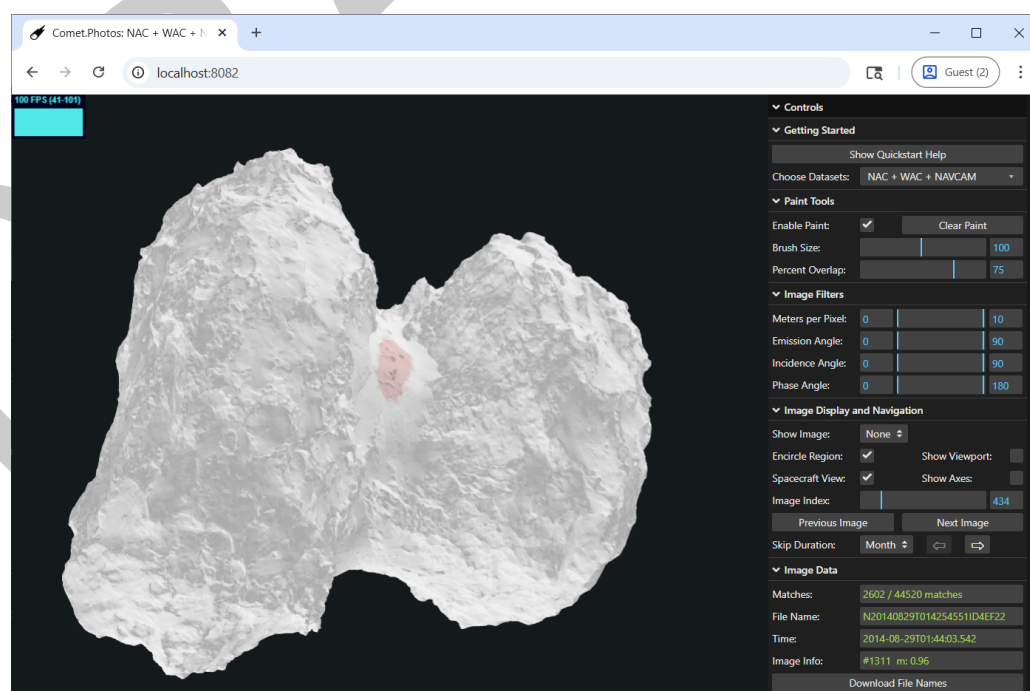
91 from the controller (Filter Engine) and data model. This event-driven runtime architecture  
92 provides the basis for a logging system with automated regression testing. The server component  
93 is a small, modular Node.js application with components for launching the browser when run  
94 locally, loading the platform-specific C library for rapid visibility checks, and distinct modules  
95 that separate preprocessing and runtime event handling. The datasets are separate from the  
96 code, and are loaded dynamically given a dataset catalog.

97 We are currently using Comet.Photos as part of a broader toolset (Jindal et al., 2025) to  
98 discover and map new surface changes on 67P using Rosetta data. With the appropriate  
99 preprocessing, the same framework could support other Rosetta instruments, such as VIRTIS  
100 (Coradini et al., 2007), MIRO (Gulkis et al., 2007), and ALICE (Stern et al., 2007); it could  
101 also be adapted to other small-body missions, including Lucy (Levison et al., 2021), Hera  
102 (Michel et al., 2022), OSIRIS-REx (Lauretta et al., 2017), OSIRIS-APEX (DellaGiustina et al.,  
103 2023), and Hayabusa2 (Watanabe et al., 2017).

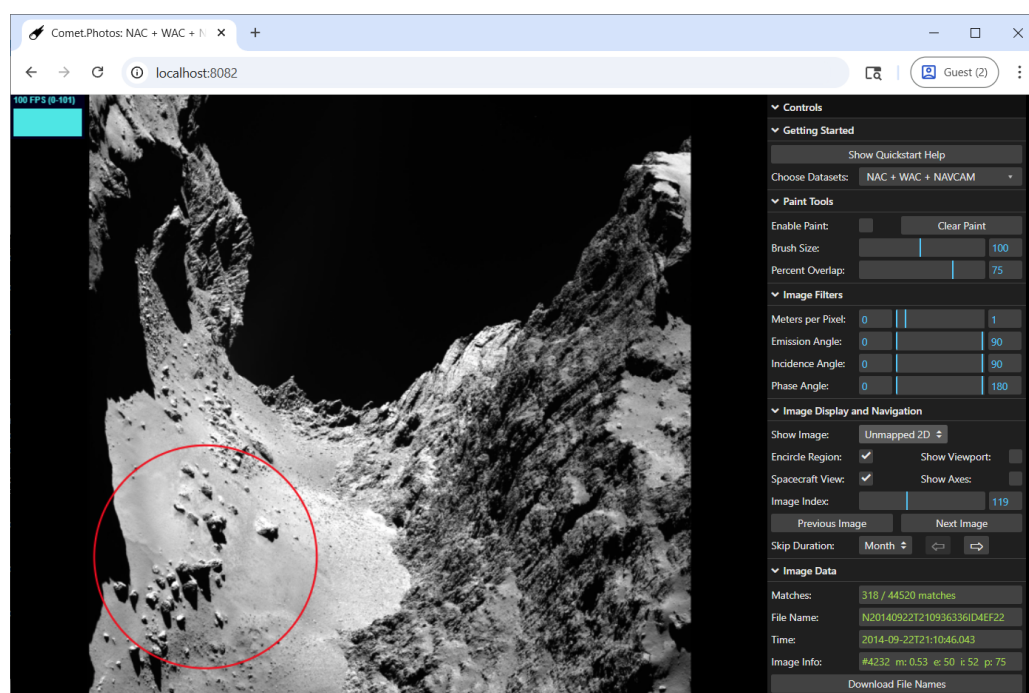
104 The Comet.Photos GitHub repository (Kurlander, 2025) provides the [full source code](#), a  
105 detailed [user manual](#), [an example workflow](#), [local installation](#) and [testing instructions](#), and  
106 documentation on [design](#), [architecture](#), [implementation](#), and [performance](#).

## 107 Acknowledgements

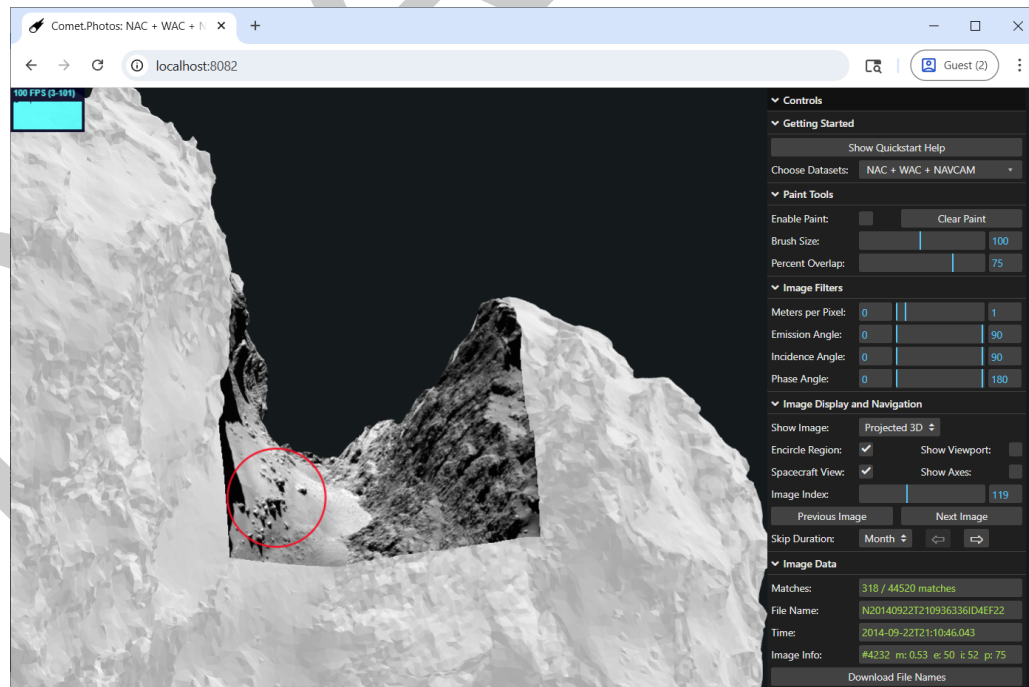
108 This research was supported by the NASA Discovery Data Analysis Program (grant  
109 80NSSC22K1399 supported D.A.K., J.M.S., and J.K.S., and grant 80NSSC24K0060 supported  
110 A.S.J. and S.P.D.B.). We gratefully acknowledge Jean-Baptiste Vincent, discussions with  
111 whom made this software possible as he helped us navigate Rosetta's dataset. A 2023  
112 MIT Open Data Prize for an earlier version of this work also provided recognition and  
113 encouragement to continue developing Comet.Photos (Fay, 2023). Lastly, we thank all of the  
114 early users of the program for feedback that led to numerous improvements.



**Figure 1:** In this example, the user paints a region of interest on the 3D comet shape model. Almost immediately, the application finds 2,602 images from the 44,520 images in the dataset that contain the painted region.



**Figure 2:** In this example, to find only those images with a high level of surface detail, the user sets the pixel scale slider to a maximum of 1 meter per pixel. This results in 318 image matches, one of which is displayed here in its original 2D form.



**Figure 3:** To examine the image from Figure 2 in context, the user projects it onto the 3D model.



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