

- MPI-Rockstar: a Hybrid MPI and OpenMP Parallel
- 2 Implementation of the Rockstar Halo finder
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# Summary

According to the concordance cosmological theory, structure formation and evolution proceeds hierarchically in the Universe. Smaller-scale dark matter structures gravitationally collapse first everywhere in the Universe, then merge into larger-scale structures. Such dense gravitationally bound structures of dark matter are called halos. Halos can host smaller halos, so-called subhalos (or substructures). Cosmological N-body simulations are vital in understanding the formation and evolution of halos and subhalos. Halo/subhalo finders are a post-processing step to identify those dense structures in the particle dataset of cosmological N-body simulations.

MPI-Rockstar is a massively parallel halo finder based on the Rockstar phase-space temporal halo finder code (Behroozi et al., 2013), which is one of the most extensively used halo finding codes. Compared to the original code, parallelized by a primitive socket communication library, we parallelized it in a hybrid way using MPI and OpenMP, which is suitable for analysis on the hybrid shared and distributed memory environments of modern supercomputers. This implementation can easily handle the analysis of more than a trillion particles on more than 100,000 parallel processes, enabling the production of a huge dataset for the next generation of cosmological surveys.

## Statement of need

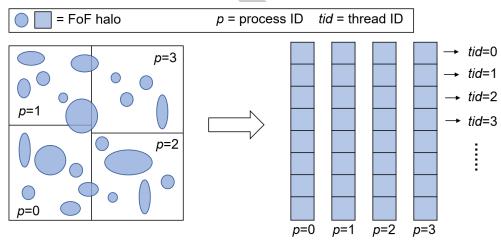
Owing to the advance of supercomputing power and highly scalable parallel gravitational N-body codes (Garrison et al., 2021; Ishiyama et al., 2009, 2012; Potter et al., 2017; Wang et al., 2018), the number of particles in recent massive cosmological simulations exceeds a trillion (Ishiyama et al., 2021; Potter et al., 2017; Wang et al., 2022), posing significant challenges for halo finding. Several halo finding algorithms have been suggested (Knebe et al., 2013), and some of their implementation are publicly available (Behroozi et al., 2013; Elahi et al., 2019; Knollmann & Knebe, 2009; Springel et al., 2021). The computational performance of these implementations differs substantially, and they have not been uniformly tested yet on the large-scale hybrid shared and distributed memory environments of modern supercomputers. MPI-Rockstar addresses these issues and is designed to run on more than 100,000 parallel processes in a hybrid way using MPI and OpenMP. As new functions to the original Rockstar code, MPI-Rockstar supports HDF5 as an output format and can output additional halo properties such as the inertia tensor.



## Parallelization

As a process parallelization, the original Rockstar divides a simulation box by the number of parallel processes and assigns each sub-box to each process. Then, each process performs 3D Friends-of-Friends (FoF) to find overdense regions, and FoF halos across processes are linked by communicating boundary regions. Rockstar then performs the subhalo finding for each FoF halo using 6D phase space information. Data communications between multiple processes are performed by one-to-one communications using sockets. As a result, many sockets (file descriptors) are issued simultaneously in the case of analysis with many processes, complicating analysis on modern supercomputers because the number of file descriptors issued simultaneously is normally limited.

In MPI-Rockstar, we replaced all socket communications in the original Rockstar with Message Passing Interface (MPI) communications, while maintaining compatibility with the analysis results. Rather than simply using MPI one-to-one communication, we changed the order of communication and computation to utilize collective communications and to run efficiently on large supercomputers. Furthermore, we parallelized MPI-Rockstar in a hybrid way, where thread parallelization is implemented within each process using OpenMP. The subhalo finding is parallelized not only on a process level but also on a thread level, improving the overall performance of MPI-Rockstar. This hybrid parallel design also reduces the risk of per-process out-of-memory compared with a flat-MPI configuration. The below figure illustrates this parallelization storategy.



Performing 3D friends-of-friends by process parallelization

Performing subhalo finding for FoF halos in each process by thread parallelization

Figure 1: Parallelization storategy

The below figure shows a strong scaling of MPI-Rockstar using up to 1,024 nodes (48 CPU cores per node) on supercomputer Fugaku. The horizontal and vertical axes represent the number of computational nodes and the time taken for the halo and subhalo finding of one snapshot, respectively. The blue and green curves show the strong scaling for simulations with 40963 particles in a 2 Gpc/h box and with 25603 particles in a 400 Mpc/h box, respectively The redshift of the snapshot analyzed was 2.0. We measured the code's performance using 2 MPI processes per node and 24 OpenMP threads per process. This choice gives an optimal configulation for these snapshots, considering the balance between the computation and I/O time. The parallel efficiency is excellent, ~90% for both the 40963 box from 256 to 1024 nodes and the 25603 box from 256 to 1024 nodes. Thanks to the communication optimization



- and hybrid parallelization, MPI-Rockstar could run up to three times faster than the original
- 74 Rockstar when compared in the same execution environment. We confirm that MPI-Rockstar
- <sub>75</sub> can analyze 2 trillion particle simulations on 16,384 nodes (786,432 CPU cores) of Fugaku.

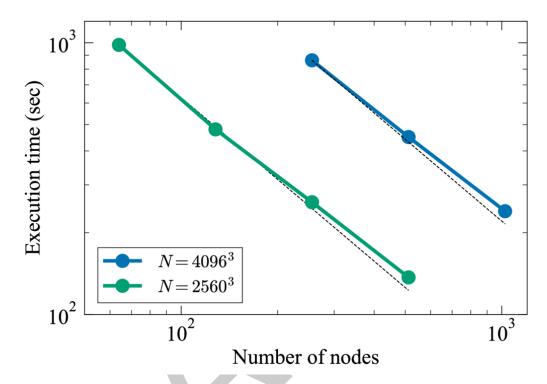


Figure 2: Strong scaling of MPI-Rockstar

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# 4 References

- Behroozi, P. S., Wechsler, R. H., & Wu, H.-Y. (2013). The ROCKSTAR Phase-space Temporal Halo Finder and the Velocity Offsets of Cluster Cores. 762, 109. https://doi.org/10.1088/0004-637X/762/2/109
- Elahi, P. J., Cañas, R., Poulton, R. J. J., Tobar, R. J., Willis, J. S., Lagos, C. del P., Power, C., & Robotham, A. S. G. (2019). Hunting for galaxies and halos in simulations with VELOCIraptor. 36, e021. https://doi.org/10.1017/pasa.2019.12
- Garrison, L. H., Eisenstein, D. J., Ferrer, D., Maksimova, N. A., & Pinto, P. A. (2021). The
  ABACUS cosmological N-body code. 508(1), 575–596. https://doi.org/10.1093/mnras/
  stab2482
- Ishiyama, T., Fukushige, T., & Makino, J. (2009). GreeM: Massively Parallel TreePM Code



- for Large Cosmological N -body Simulations. 61, 1319. https://doi.org/10.1093/pasj/61.
   6.1319
- Ishiyama, T., Nitadori, K., & Makino, J. (2012). 4.45 pflops astrophysical n-body simulation on k computer: The gravitational trillion-body problem. Proc. Int. Conf. High Performance Computing, Networking, Storage and Analysis, SC'12 (Los Alamitos, CA: IEEE Computer Society Press), 5:, (arXiv:1211.4406). https://doi.org/10.1109/SC.2012.3
- Ishiyama, T., Prada, F., Klypin, A. A., Sinha, M., Metcalf, R. B., Jullo, E., Altieri, B., Cora, S.
   A., Croton, D., de la Torre, S., Millán-Calero, D. E., Oogi, T., Ruedas, J., & Vega-Martínez,
   C. A. (2021). The Uchuu simulations: Data Release 1 and dark matter halo concentrations.
   506(3), 4210–4231. https://doi.org/10.1093/mnras/stab1755
- Knebe, A., Pearce, F. R., Lux, H., Ascasibar, Y., Behroozi, P., Casado, J., Moran, C. C.,
   Diemand, J., Dolag, K., Dominguez-Tenreiro, R., Elahi, P., Falck, B., Gottlöber, S., Han,
   J., Klypin, A., Lukić, Z., Maciejewski, M., McBride, C. K., Merchán, M. E., ... Zemp,
   M. (2013). Structure finding in cosmological simulations: the state of affairs. 435(2),
   1618–1658. https://doi.org/10.1093/mnras/stt1403
- Knollmann, S. R., & Knebe, A. (2009). AHF: Amiga's Halo Finder. 182(2), 608–624.
   https://doi.org/10.1088/0067-0049/182/2/608
- Potter, D., Stadel, J., & Teyssier, R. (2017). PKDGRAV3: beyond trillion particle cosmological simulations for the next era of galaxy surveys. *Computational Astrophysics and Cosmology*, 4(1), 2. https://doi.org/10.1186/s40668-017-0021-1
- Springel, V., Pakmor, R., Zier, O., & Reinecke, M. (2021). Simulating cosmic structure formation with the GADGET-4 code. 506(2), 2871-2949. https://doi.org/10.1093/mnras/stab1855
- Wang, Q., Cao, Z.-Y., Gao, L., Chi, X.-B., Meng, C., Wang, J., & Wang, L. (2018). PHoToNs A parallel heterogeneous and threads oriented code for cosmological N-body simulation.
   Research in Astronomy and Astrophysics, 18(6), 062. https://doi.org/10.1088/1674-4527/
   18/6/62
- Wang, Q., Gao, L., & Meng, C. (2022). The ultramarine simulation: properties of dark matter haloes before redshift 5.5. *517*(4), 6004–6012. https://doi.org/10.1093/mnras/stac3072