

¹spherical: A Comprehensive Database and Automated Pipeline for VLT/SPHERE High-Contrast Imaging

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Software

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⁵ Summary

⁶ The Spectro-Polarimetric High-contrast Exoplanet REsearch instrument [SPHERE; Beuzit et
⁷ al. (2019)] at the Very Large Telescope (VLT) is a leading facility for coronagraphic imaging
⁸ of exoplanets and circumstellar disks in the optical and near-infrared. Over the last decade,
⁹ SPHERE has contributed to hundreds of publications and major legacy surveys [e.g., SHINE;
¹⁰ Chauvin et al. (2017); Chomez et al. (2025)].

¹¹ spherical streamlines the end-to-end analysis of SPHERE data by integrating a curated,
¹² searchable database with an automated Python-based pipeline. The software enables users to
¹³ filter the complete observation history by target properties or observing conditions, automatically
¹⁴ download raw datasets from the ESO archive, and reduce Integral Field Spectrograph (IFS)
¹⁵ data using the adapted CHARIS pipeline. Beyond basic reduction, the workflow handles
¹⁶ astrometric calibration and post-processing via TRAP for companion detection, while also
¹⁷ facilitating the discovery and retrieval of IRDIS dual-band and polarimetric imaging data for
¹⁸ analysis with community tools.

¹⁹ As of May 2025, the database includes approximately 6000 IRDIS dual-band imaging (DBI)
²⁰ observations, ~1000 IRDIS dual-beam polarimetric imaging [DPI; de Boer et al. (2020); van
²¹ Holstein et al. (2020a)] observations, and ~4500 IFS observations. SPHERE is undergoing the
²² SPHERE+ upgrade (Boccaletti et al., 2022, 2020), including a second-stage adaptive optics
²³ system [SAXO+; Stadler et al. (2022)], ensuring long-term scientific relevance and providing
²⁴ a pathfinder for the ELT's Planetary Camera and Spectrograph [PCS; Kasper et al. (2021)].

²⁵ Statement of Need

²⁶ The ESO VLT/SPHERE archive constitutes the world's largest collection of high-contrast
²⁷ imaging data to detect exoplanets, substellar companions, and circumstellar disks. However,
²⁸ utilizing this rich dataset for end-to-end research remains problematic. Researchers face
²⁹ challenges through need of manual identification and acquisition of available data and the
³⁰ lack of integration between various reduction pipelines. These barriers disproportionately
³¹ affect researchers attempting to assemble large homogeneous samples for population studies,
³² homogenous extraction of exoplanet spectra, and survey teams requiring rapid follow-up.

³³ State of the Field

³⁴ Several existing tools address isolated components of the SPHERE data workflow. The **High**
³⁵ **Contrast Data Center (DC)** (Delorme et al., 2017) provides Java-based access to reduce
³⁶ datasets using ESO's internal pipeline but offers limited capabilities for programmatic interaction
³⁷ or custom batch processing. **vlt-sphere** (Vigan, 2020) offers Python wrappers around the ESO
³⁸ pipeline for user-provided raw data, yet it lacks automated archival download and integrated

39 post-processing features. For polarimetry, **IRDAP** (van Holstein et al., 2020b) serves as a
40 robust automated pipeline, though it leaves dataset discovery and retrieval as manual user
41 tasks.

42 spherical was build to fill the gap that a missing high level overview of what data SPHERE has
43 taken over its lifetime leaves. It creates a unified workflow for identification of data, download,
44 reduction and post-processing (via TRAP). From identification of observations of interest,
45 all the way to the identification of exoplanets and extraction of their spectra spherical is
46 an end-to-end framework. spherical can also be used as infrastructure to download and
47 pre-process data for alternative post-processing ecosystems such as VIP (Christiaens et al.,
48 2023; Gomez Gonzalez et al., 2017), pyKLIP (Wang et al., 2015), and IRDAP.

49 Software Design

50 The architecture of spherical is designed as a high-level abstraction layer over the ESO
51 raw archive, providing a systematic interface that currently does not exist for high-contrast
52 imaging data. A central design choice was the decoupling of metadata curation from data
53 reduction. This is achieved through a multi-stage process that ingests science headers from
54 the ESO archive and cross-matches them with the Gaia catalog based on telescope pointings.
55 To handle the inherent noise in archival metadata, the system employs a proximity, brightness,
56 and object type-based resolution logic to resolve target ambiguities, accounting for epoch of
57 observation of proper motion of stellar targets, ensuring the resulting database is both clean
58 and astronomically accurate.

59 The core of the package is a structured, searchable observation table—archived on Zenodo
60 ([DOI: 10.5281/zenodo.15147730](https://doi.org/10.5281/zenodo.15147730))—which aggregates information such as observing mode,
61 total exposure time, parallactic angle coverage, and atmospheric conditions. By choosing a
62 local, curated table over live archive queries, spherical enables rapid, complex filtering of the
63 entire instrument history that would otherwise be difficult.

64 Technically, spherical adopts a wrapper-based architecture. Instead of reimplementing
65 reduction algorithms, it provides a Pythonic interface to the Samland et al. (2022) IFS
66 instrument pipeline. This design trade-off prioritizes scientific continuity and maintainability; by
67 “gluing” the complex CHARIS instrument-based SPHERE IFS pipeline and TRAP high-contrast
68 imaging post-processing workflows into an object-oriented framework, spherical transforms
69 specialized code into a user-friendly end-to-end workflow without sacrificing the precision of
70 the original algorithms. This hybrid approach allows the software to act as a bridge between
71 archival discovery and specialized community tools. Likewise, spherical can be used as an
72 initial stage for post-processing frameworks like VIP and IRDAP.

73 spherical relies purely on open-source software: Astropy (Astropy Collaboration et al., 2013,
74 2018, 2022), astroquery for ESO archive and catalog access [Ginsburg et al. (2019), NumPy
75 (Harris et al., 2020) for numerical operations, and pandas (The pandas development team,
76 n.d.) for tabular data handling. The IFS reduction uses the adapted CHARIS pipeline (Brandt
77 et al., 2017; Samland et al., 2022), with calibration routines derived in part from Vigan (2020)
78 (see the vlt-sphere repository for individual contributors). Post-processing employs TRAP
79 (Samland et al., 2021).

80 Research Impact Statement

81 spherical has been used in several recently published peer-reviewed studies to analyse SPHERE
82 IFS data (Franson et al., 2023; Hammond et al., 2025; Stolker et al., 2025), improving the
83 extraction of spectra of exoplanetary atmospheres. At the time these works were published, the
84 code had not yet been citable via an existing DOI. Since the database of existing observations
85 has been made available on Zenodo, it has been downloaded more than 400 times. New

86 exoplanet candidates identified through spherical have been accepted for follow-up observation
87 in an ESO P117 proposal (117.2A06.001), awarding 28.5 hours—more than 60% of all time
88 awarded in P117 to SPHERE observations.

89 The research impact of spherical is centered on enabling large-scale, reproducible science that
90 was previously inhibited by the heterogeneous nature of the SPHERE archive. By lowering the
91 technical barrier from raw archival to science-ready data products, the software facilitates the
92 construction of homogeneous samples essential for statistically significant exoplanet occurrence
93 rate and population studies. This systematic approach allows researchers to apply consistent
94 algorithms across the entire instrument lifetime.

95 Furthermore, the integration of high-performance post-processing tools like TRAP within an
96 automated framework enables the “mining” of archival data for low-mass companions that were
97 missed by earlier, less sensitive reduction techniques. This capability is particularly timely as the
98 astronomical community prepares for the SPHERE+ upgrade and future ELT-era observations;
99 spherical provides the necessary infrastructure to benchmark current performance and develop
100 the automated workflows required to handle the next generation of high-contrast imaging data.
101 By shifting the focus from manual data curation to atmospheric characterization and spectral
102 extraction, the package accelerates the transition from exoplanet discovery to detailed physical
103 understanding.

104 Future Work

105 spherical is built with extensibility as a core tenet. Future development will focus on expanding
106 database coverage and streamlining data retrieval for additional observing modes, specifically
107 ZIMPOL, IRDIS LSS, and SAM, while integrating their respective community reduction
108 pipelines into the established script-driven framework. To enhance software reliability and
109 facilitate community contributions, we intend to provide a public test dataset for the continuous
110 integration of selected pipeline stages. Furthermore, the package will evolve alongside the
111 SPHERE+ upgrade; calibration and post-processing modules will be adapted to account for
112 the instrument’s enhanced performance and the new capabilities of the SAXO+ adaptive optics
113 system.

114 AI Usage Disclosure

115 The authors acknowledge the use of generative AI (Large Language Models) in the development
116 of spherical and the preparation of this manuscript. Specifically, AI tools were utilized to
117 generate source code docstrings, draft the project README and release notes, and assist in
118 refactoring existing code to improve maintainability. During the writing process, generative AI
119 was used for language editing and to refine the manuscript’s conciseness.

120 The authors certify that generative AI was not involved in the original conceptual design,
121 the software architecture, or the underlying scientific methodology. All AI-generated outputs,
122 including code and text, were manually reviewed, verified, and edited by the authors to ensure
123 technical accuracy and integrity.

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126 Contributors are listed at the project repository’s [contributors page](#). We acknowledge ESO for
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