[Template:About](/wiki/Template:About" \o "Template:About) [Template:Pp-semi](/wiki/Template:Pp-semi) [Template:Pp-move-indef](/wiki/Template:Pp-move-indef) [Template:Use dmy dates](/wiki/Template:Use_dmy_dates) [Template:Infobox planet](/wiki/Template:Infobox_planet)

**Saturn** is the sixth [planet](/wiki/Planet) from the [Sun](/wiki/Sun) and the second-largest in the [Solar System](/wiki/Solar_System), after [Jupiter](/wiki/Jupiter). It is a [gas giant](/wiki/Gas_giant) with an average radius about nine times that of [Earth](/wiki/Earth).[[1]](#cite_note-1)[[2]](#cite_note-2) Although only one-eighth the average density of Earth, with its larger volume Saturn is just over 95 times more massive.[[3]](#cite_note-3)[[4]](#cite_note-4)[[5]](#cite_note-5) Saturn is named after the [Roman](/wiki/Roman_mythology) [god of agriculture](/wiki/Saturn_(mythology)); its [astronomical symbol](/wiki/Astronomical_symbol) (♄) represents the god's [sickle](/wiki/Sickle).

Saturn's interior is probably composed of a core of [iron–nickel](/wiki/Iron–nickel_alloy) and rock ([silicon](/wiki/Silicon) and [oxygen](/wiki/Oxygen) compounds). This core is surrounded by a deep layer of [metallic hydrogen](/wiki/Metallic_hydrogen), an intermediate layer of [liquid hydrogen](/wiki/Liquid_hydrogen) and [liquid helium](/wiki/Liquid_helium), and finally outside the [Frenkel line](/wiki/Frenkel_line) a gaseous outer layer.[[6]](#cite_note-6) Saturn has a pale yellow hue due to [ammonia](/wiki/Ammonia) crystals in its upper atmosphere. [Electrical current](/wiki/Electrical_current) within the metallic hydrogen layer is thought to give rise to Saturn's planetary [magnetic field](/wiki/Magnetic_field), which is weaker than Earth's, but has a [magnetic moment](/wiki/Magnetic_moment) 580 times that of Earth due to Saturn's larger size. Saturn's magnetic field strength is around one-twentieth of Jupiter's.[[7]](#cite_note-7) The outer [atmosphere](/wiki/Atmosphere) is generally bland and lacking in contrast, although long-lived features can appear. [Wind speeds](/wiki/Wind_speed) on Saturn can reach [Template:Convert](/wiki/Template:Convert), higher than on Jupiter, but not as high as those on [Neptune](/wiki/Neptune).[[8]](#cite_note-8) Saturn has a prominent [ring system](/wiki/Rings_of_Saturn) that consists of nine continuous main rings and three discontinuous arcs and that is composed mostly of ice particles with a smaller amount of rocky debris and [dust](/wiki/Cosmic_dust). Sixty-two[[9]](#cite_note-9) [moons](/wiki/Moons_of_Saturn) are known to orbit Saturn, of which fifty-three are officially named. This does not include the hundreds of [moonlets](/wiki/Moonlet) comprising the rings. [Titan](/wiki/Titan_(moon)), Saturn's largest moon, and the second-largest in the Solar System, is larger than the planet [Mercury](/wiki/Mercury_(planet)), although less massive, and is the only moon in the Solar System to have a substantial atmosphere.[[10]](#cite_note-10)[Template:TOC limit](/wiki/Template:TOC_limit)

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## Physical characteristics[[edit](/index.php?title=(none)&action=edit&section=1)]

[left|thumb|Composite image roughly comparing the sizes of Saturn and Earth|200x200px](/wiki/File:Saturn,_Earth_size_comparison.jpg)

Saturn is a [gas giant](/wiki/Gas_giant) because it is predominantly composed of hydrogen and helium ('gas'). It lacks a definite surface, though it may have a solid core.<ref name=melosh2011/> Saturn's rotation causes it to have the shape of an [oblate spheroid](/wiki/Oblate_spheroid); that is, it is flattened at the poles and bulges at its equator. Its equatorial and polar radii differ by almost 10%: 60,268 km versus 54,364 km, respectively.[[11]](#cite_note-11) Jupiter, [Uranus](/wiki/Uranus), and [Neptune](/wiki/Neptune), the other giant planets in the Solar System, are also oblate but to a lesser extent. Saturn is the only planet of the Solar System that is less dense than water—about 30% less.<ref name=preserve/> Although Saturn's [core](/wiki/Planetary_core) is considerably denser than water, the average [specific density](/wiki/Relative_density) of the planet is 0.69 g/cm3 due to the atmosphere. Jupiter has 318 times the Earth's mass,[[12]](#cite_note-12) while Saturn is 95 times the mass of the [Earth](/wiki/Earth),[[11]](#cite_note-11) Together, Jupiter and Saturn hold 92% of the total planetary mass in the Solar System.<ref name=ssr152\_1\_423/>

On 8 January 2015, NASA reported determining the center of the planet Saturn and its [family of moons](/wiki/Moons_of_Saturn) to within [Template:Convert](/wiki/Template:Convert).[[13]](#cite_note-13)

### Internal structure[[edit](/index.php?title=(none)&action=edit&section=2)]

[thumb|600x600px|Diagram of Saturn, to scale.|left](/wiki/File:Saturn_diagram.svg) Despite consisting mostly of hydrogen and helium, most of Saturn's mass is not in the [gas](/wiki/Gas) [phase](/wiki/Phase_(matter)), because hydrogen becomes a [non-ideal liquid](/wiki/Ideal_solution) when the density is above 0.01 g/cm3, which is reached at a radius containing 99.9% of Saturn's mass. The temperature, pressure, and density inside Saturn all rise steadily toward the core, which causes hydrogen to transition into a metal in the deeper layers.<ref name=ssr152\_1\_423/>

Standard planetary models suggest that the interior of Saturn is similar to that of [Jupiter](/wiki/Jupiter), having a small rocky core surrounded by [hydrogen](/wiki/Hydrogen) and [helium](/wiki/Helium) with trace amounts of various [volatiles](/wiki/Volatiles).<ref name=guillot\_et\_al2009/> This core is similar in composition to the Earth, but more dense. Examination of Saturn's [gravitational moment](/wiki/Gravitational_moment), in combination with physical models of the interior, allowed French astronomers Didier Saumon and Tristan Guillot to place constraints on the mass of Saturn's core. In 2004, they estimated that the core must be 9–22 times the mass of the Earth,<ref name=science305\_5689\_1414/><ref name=apj609\_2\_1170/> which corresponds to a diameter of about 25,000 km.[[14]](#cite_note-14) This is surrounded by a thicker liquid [metallic hydrogen](/wiki/Metallic_hydrogen) layer, followed by a liquid layer of helium-saturated [molecular hydrogen](/wiki/Molecular_hydrogen) that gradually transitions to a gas with increasing altitude. The outermost layer spans 1,000 km and consists of gas.[[15]](#cite_note-15)[[16]](#cite_note-16) Saturn has a hot interior, reaching 11,700 °C at its core, and it radiates 2.5 times more energy into space than it receives from the Sun. Jupiter's [thermal energy](/wiki/Thermal_energy) is generated by the [Kelvin–Helmholtz mechanism](/wiki/Kelvin–Helmholtz_mechanism) of slow [gravitational compression](/wiki/Gravitational_compression), but this alone may not be sufficient to explain heat production for Saturn, because it is less massive. An alternative or additional mechanism may be generation of heat through the "raining out" of droplets of helium deep in Saturn's interior. As the droplets descend through the lower-density hydrogen, the process releases heat by [friction](/wiki/Friction) and leaves Saturn's outer layers depleted of helium.<ref name=de\_pater\_lissauer2010/><ref name=nasa\_saturn/> These descending droplets may have accumulated into a helium shell surrounding the core.<ref name=guillot\_et\_al2009/>

### Atmosphere[[edit](/index.php?title=(none)&action=edit&section=3)]

[thumb|125px|right|](/wiki/File:PIA18354-Saturn-MethaneBands-20150906.jpg)[Methane](/wiki/Methane) bands circle Saturn (rel 1 February 2016) The outer atmosphere of Saturn contains 96.3% molecular hydrogen and 3.25% helium by volume.[[17]](#cite_note-17) The proportion of helium is significantly deficient compared to the abundance of this element in the Sun.<ref name=guillot\_et\_al2009/> The quantity of elements heavier than helium is not known precisely, but the proportions are assumed to match the primordial abundances from the formation of the Solar System. The total mass of these heavier elements is estimated to be 19–31 times the mass of the Earth, with a significant fraction located in Saturn's core region.[[18]](#cite_note-18) Trace amounts of ammonia, [acetylene](/wiki/Acetylene), [ethane](/wiki/Ethane), [propane](/wiki/Propane), [phosphine](/wiki/Phosphine) and [methane](/wiki/Methane) have been detected in Saturn's atmosphere.<ref name=baas15\_831/><ref name=cain2009\_24029/><ref name=pfsaa2008/> The upper clouds are composed of ammonia crystals, while the lower level clouds appear to consist of either [ammonium hydrosulfide](/wiki/Ammonium_hydrosulfide) (NH4SH) or water.<ref name=martinez20050905/> [Ultraviolet](/wiki/Ultraviolet) radiation from the Sun causes methane [photolysis](/wiki/Photolysis) in the upper atmosphere, leading to a series of [hydrocarbon](/wiki/Hydrocarbon) chemical reactions with the resulting products being carried downward by eddies and diffusion. This [photochemical cycle](/wiki/Photochemical_cycle) is modulated by Saturn's annual seasonal cycle.<ref name=pfsaa2008/>

#### Cloud layers[[edit](/index.php?title=(none)&action=edit&section=4)]

[thumb|left|A global storm girdles the planet in 2011. The head of the storm (bright area) passes the tail circling around the left limb.](/wiki/File:Saturn_Storm.jpg)

Saturn's atmosphere exhibits a banded pattern similar to Jupiter's, but Saturn's bands are much fainter and are much wider near the equator. The nomenclature used to describe these bands is the same as on Jupiter. Saturn's finer cloud patterns were not observed until the flybys of the Voyager spacecraft during the 1980s. Since then, Earth-based [telescopy](/wiki/Telescope) has improved to the point where regular observations can be made.<ref name=emp105\_2\_143/>

The composition of the clouds varies with depth and increasing pressure. In the upper cloud layers, with the temperature in the range 100–160 K and pressures extending between 0.5–2 [bar](/wiki/Bar_(unit)), the clouds consist of ammonia ice. Water [ice clouds](/wiki/Ice_cloud) begin at a level where the pressure is about 2.5 bar and extend down to 9.5 bar, where temperatures range from 185–270 K. Intermixed in this layer is a band of ammonium hydrosulfide ice, lying in the pressure range 3–6 bar with temperatures of 290–235 K. Finally, the lower layers, where pressures are between 10–20 bar and temperatures are 270–330 K, contains a region of water droplets with ammonia in aqueous solution.<ref name=dougherty\_esposito2009/>

Saturn's usually bland atmosphere occasionally exhibits long-lived ovals and other features common on Jupiter. In 1990, the [Hubble Space Telescope](/wiki/Hubble_Space_Telescope) imaged an enormous white cloud near Saturn's equator that was not present during the Voyager encounters and in 1994, another, smaller storm was observed. The 1990 storm was an example of a [Great White Spot](/wiki/Great_White_Spot), a unique but short-lived phenomenon that occurs once every Saturnian year, roughly every 30 Earth years, around the time of the northern hemisphere's [summer solstice](/wiki/Summer_solstice).<ref name=icarus176\_1\_155/> Previous Great White Spots were observed in 1876, 1903, 1933 and 1960, with the 1933 storm being the most famous. If the periodicity is maintained, another storm will occur in about 2020.[[19]](#cite_note-19) The winds on Saturn are the second fastest among the Solar System's planets, after Neptune's. [Voyager](/wiki/Voyager_program) data indicate peak easterly winds of 500 m/s (1800 km/h).[[20]](#cite_note-20) In images from the [Cassini](/wiki/Cassini–Huygens) spacecraft during 2007, Saturn's northern hemisphere displayed a bright blue hue, similar to Uranus. The color was most likely caused by [Rayleigh scattering](/wiki/Rayleigh_scattering).[[21]](#cite_note-21) [Thermography](/wiki/Thermography) has shown that Saturn's south pole has a warm [polar vortex](/wiki/Polar_vortex), the only known example of such a phenomenon in the Solar System.<ref name=MCP>[Template:Cite web](/wiki/Template:Cite_web)</ref> Whereas temperatures on Saturn are normally −185 °C, temperatures on the vortex often reach as high as −122 °C, suspected to be the warmest spot on Saturn.<ref name=MCP/>

#### North pole hexagonal cloud pattern[[edit](/index.php?title=(none)&action=edit&section=5)]

[thumb|125px|right|Saturn –](/wiki/File:PIA18274-Saturn-NorthPolarHexagon-Cassini-20140402.jpg) [North polar hexagon](/wiki/Saturn's_hexagon) and vortex as well as rings (2 April 2014) [Template:Main](/wiki/Template:Main)

A persisting [hexagonal](/wiki/Hexagon) wave pattern around the north polar vortex in the atmosphere at about 78°N was first noted in the Voyager images.[[22]](#cite_note-22)[[23]](#cite_note-23)[[24]](#cite_note-24) The sides of the hexagon are each about [Template:Convert](/wiki/Template:Convert) long, which is longer than the diameter of the Earth.[[25]](#cite_note-25) The entire structure rotates with a period of [Template:Nowrap](/wiki/Template:Nowrap) (the same period as that of the planet's radio emissions) which is assumed to be equal to the period of rotation of Saturn's interior.<ref name=science247\_4947\_1206/> The hexagonal feature does not shift in longitude like the other clouds in the visible atmosphere.<ref name=pss57\_14\_1671/> The pattern's origin is a matter of much speculation. Most scientists think it is a [standing wave](/wiki/Standing_wave) pattern in the atmosphere. Polygonal shapes have been replicated in the laboratory through differential rotation of fluids.[[26]](#cite_note-26)<ref name=labmodel>[Template:Cite journal](/wiki/Template:Cite_journal) Laboratory experiment of spinning disks in a liquid solution forms vortices around a stable hexagonal pattern similar to that of Saturn's.</ref>

#### South pole vortex[[edit](/index.php?title=(none)&action=edit&section=6)]

[thumb|125px|A storm at Saturn's south pole](/wiki/File:Looking_saturn_in_the_eye.jpg) [HST](/wiki/Hubble_Space_Telescope) imaging of the south polar region indicates the presence of a [jet stream](/wiki/Jet_stream), but no strong polar vortex nor any hexagonal standing wave.[[27]](#cite_note-27) [NASA](/wiki/NASA) reported in November 2006 that Cassini had observed a "[hurricane](/wiki/Hurricane)-like" storm locked to the south pole that had a clearly defined [eyewall](/wiki/Eyewall).[[28]](#cite_note-28)[[29]](#cite_note-29) Eyewall clouds had not previously been seen on any planet other than Earth. For example, images from the [Galileo](/wiki/Galileo_(spacecraft)) spacecraft did not show an eyewall in the [Great Red Spot](/wiki/Great_Red_Spot) of Jupiter.[[30]](#cite_note-30) The south pole storm may have been present for billions of years.[[31]](#cite_note-31) This vortex is comparable to the size of Earth, and it has winds of 550 km/h.[[31]](#cite_note-31)

#### Other features[[edit](/index.php?title=(none)&action=edit&section=7)]

Cassini has observed a series of cloud features nicknamed "String of Pearls" found in northern latitudes. These features are cloud clearings that reside in deeper cloud layers.[[32]](#cite_note-32)

### Magnetosphere[[edit](/index.php?title=(none)&action=edit&section=8)]

[Template:Main](/wiki/Template:Main) [thumb|125px|left|Polar aurorae on Saturn](/wiki/File:Saturn's_double_aurorae_(captured_by_the_Hubble_Space_Telescope).jpg) [thumb|125px|right|Auroral lights at Saturn’s north pole.](/wiki/File:Hubble_sees_a_flickering_light_display_on_Saturn.jpg)[[33]](#cite_note-33) Saturn has an intrinsic [magnetic](/wiki/Magnetic) field that has a simple, symmetric shape – a magnetic [dipole](/wiki/Dipole). Its strength at the equator – 0.2 [gauss](/wiki/Gauss_(unit)) (20 [µT](/wiki/Microtesla)) – is approximately one twentieth of that of the field around Jupiter and slightly weaker than Earth's [magnetic field](/wiki/Magnetic_field).[[7]](#cite_note-7) As a result, Saturn's [magnetosphere](/wiki/Magnetosphere) is much smaller than Jupiter's.[[34]](#cite_note-34) When Voyager 2 entered the magnetosphere, the [solar wind](/wiki/Solar_wind) pressure was high and the magnetosphere extended only 19 Saturn radii, or 1.1 million km (712,000 mi),[[35]](#cite_note-35) although it enlarged within several hours, and remained so for about three days.<ref name=atkinson2010/> Most probably, the magnetic field is generated similarly to that of Jupiter – by currents in the liquid metallic-hydrogen layer called a metallic-hydrogen dynamo.[[34]](#cite_note-34) This magnetosphere is efficient at deflecting the [solar wind](/wiki/Solar_wind) particles from the [Sun](/wiki/Sun). The moon Titan orbits within the outer part of Saturn's magnetosphere and contributes plasma from the [ionized](/wiki/Ion) particles in Titan's outer atmosphere.[[7]](#cite_note-7) Saturn's magnetosphere, like [Earth's](/wiki/Earth's_magnetic_field), produces [aurorae](/wiki/Aurora_(astronomy)).[[36]](#cite_note-36)

## Orbit and rotation[[edit](/index.php?title=(none)&action=edit&section=9)]

The average distance between Saturn and the [Sun](/wiki/Sun) is over 1.4 billion kilometres (9 [AU](/wiki/Astronomical_unit)). With an average orbital speed of 9.69 km/s,[[11]](#cite_note-11) it takes Saturn 10,759 Earth days (or about [Template:Frac](/wiki/Template:Frac) years),<ref name=cain2009/> to finish one revolution around the Sun.[[11]](#cite_note-11) The elliptical orbit of Saturn is inclined 2.48° relative to the orbital plane of the Earth.[[11]](#cite_note-11) The [perihelion and aphelion](/wiki/Apsis) distances are, respectively, 9.022 and 10.053 AU, on average.<ref name=AA>Jean Meeus, *Astronomical Algorithms* (Richmond, VA: Willmann-Bell, 1998). Average of the nine extremes on p 273. All are within 0.02 AU of the averages.</ref> The visible features on Saturn rotate at different rates depending on latitude and multiple rotation periods have been assigned to various regions (as in Jupiter's case).

Astronomers use three different systems for specifying the rotation rate of Saturn. *System I* has a period of 10 hr 14 min 00 sec (844.3°/d) and encompasses the Equatorial Zone, the South Equatorial Belt and the North Equatorial Belt. The polar regions are considered to have rotation rates similar to *System I*. All other Saturnian latitudes, excluding the north and south polar regions, are indicated as *System II* and have been assigned a rotation period of 10 hr 38 min 25.4 sec (810.76°/d). *System III* refers to Saturn's internal rotation rate. Based on [radio](/wiki/Radio_astronomy) emissions from the planet in the period of the Voyager flybys, it has been assigned a rotation period of 10 hr 39 min 22.4 sec (810.8°/d). Because it is close to System II, it has largely superseded it.<ref name=benton2006/>

A precise value for the rotation period of the interior remains elusive. While approaching Saturn in 2004, Cassini found that the radio rotation period of Saturn had increased appreciably, to approximately 10 hr 45 min 45 sec (± 36 sec).[[37]](#cite_note-37)[[38]](#cite_note-38) The latest estimate of Saturn's rotation (as an indicated rotation rate for Saturn as a whole) based on a compilation of various measurements from the Cassini, Voyager and Pioneer probes was reported in September 2007 is 10 hr 32 min 35 sec.<ref name=Anderson2007/>

In March 2007, it was found that the variation of radio emissions from the planet did not match Saturn's rotation rate. This variance may be caused by geyser activity on Saturn's moon [Enceladus](/wiki/Enceladus_(moon)). The water vapor emitted into Saturn's orbit by this activity becomes charged and creates a drag upon Saturn's magnetic field, slowing its rotation slightly relative to the rotation of the planet.[[39]](#cite_note-39)<ref name=science316\_5823\_442/><ref name=science316\_5823\_380/>

## Natural satellites[[edit](/index.php?title=(none)&action=edit&section=10)]

[Template:Main](/wiki/Template:Main) [thumb|A montage of Saturn and its principal](/wiki/File:Saturn_System_Montage_-_GPN-2000-000439.jpg) [moons](/wiki/Natural_satellite) ([Dione](/wiki/Dione_(moon)), [Tethys](/wiki/Tethys_(moon)), [Mimas](/wiki/Mimas_(moon)), [Enceladus](/wiki/Enceladus_(moon)), [Rhea](/wiki/Rhea_(moon)) and [Titan](/wiki/Titan_(moon)); [Iapetus](/wiki/Iapetus_(moon)) not shown). This famous image was created from photographs taken in November 1980 by the [Voyager 1](/wiki/Voyager_1) spacecraft. Saturn has 62 known [moons](/wiki/Natural_satellite), 53 of which have formal names.[[40]](#cite_note-40)[[41]](#cite_note-41) In addition, there is evidence of dozens to hundreds of moonlets with diameters of 40–500 meters in Saturn's rings,[[42]](#cite_note-42) which are not considered to be true moons. [Titan](/wiki/Titan_(moon)), the largest moon, comprises more than 90% of the mass in orbit around Saturn, including the rings.<ref name=brunier2005/> Saturn's second-largest moon, [Rhea](/wiki/Rhea_(moon)), may have a tenuous [ring system of its own](/wiki/Rings_of_Rhea),[[43]](#cite_note-43) along with a tenuous [atmosphere](/wiki/Atmosphere).<ref name=atkinson20101126/>[[44]](#cite_note-44)[[45]](#cite_note-45)<ref name=ryan20101126/>

[thumb|left|Possible beginning of a new moon (white dot) of Saturn (image taken by Cassini on 15 April 2013).](/wiki/File:PIA18078-PossibleBeginning-NewMoonOfPlanetSaturn-20130415.jpg) Many of the other moons are small: 34 are less than 10 km in diameter and another 14 between 10 and 50 km in diameter.[[46]](#cite_note-46) Traditionally, most of Saturn's moons have been named after [Titans](/wiki/Titan_(mythology)) of Greek mythology. Titan is the only satellite in the [Solar System](/wiki/Solar_System) with a major [atmosphere](/wiki/Atmosphere_of_Titan),[[47]](#cite_note-47)[[48]](#cite_note-48) in which a complex organic chemistry occurs. It is the only satellite with [hydrocarbon lakes](/wiki/Lakes_of_Titan).[[49]](#cite_note-49)[[50]](#cite_note-50) On 6 June 2013, scientists at the [IAA-CSIC](/wiki/Instituto_de_Astrofísica_de_Andalucía) reported the detection of [polycyclic aromatic hydrocarbons](/wiki/Polycyclic_aromatic_hydrocarbons) in the [upper atmosphere](/wiki/Upper_atmosphere) of Titan, a [possible precursor for life](/wiki/PAH_world_hypothesis).[[51]](#cite_note-51) On 23 June 2014, NASA claimed to have strong evidence that [nitrogen](/wiki/Nitrogen) in the atmosphere of [Titan](/wiki/Titan_(moon)) came from materials in the [Oort cloud](/wiki/Oort_cloud), associated with [comets](/wiki/Comets), and not from the materials that formed Saturn in earlier times.[[52]](#cite_note-52) Saturn's moon [Enceladus](/wiki/Enceladus_(moon)), which seems similar in chemical makeup to comets,[[53]](#cite_note-53) has often been regarded as a potential [habitat](/wiki/Planetary_habitability) for [microbial life](/wiki/Microorganism).[[54]](#cite_note-54)<ref name=madrigal20090624/><ref name=spotts20050928/>[[55]](#cite_note-55) Evidence of this possibility includes the satellite's salt-rich particles having an "ocean-like" composition that indicates most of Enceladus's expelled [ice](/wiki/Ice) comes from the evaporation of liquid salt water.[[56]](#cite_note-56)[[57]](#cite_note-57)[[58]](#cite_note-58) In April 2014, NASA scientists reported the possible beginning of a new [moon](/wiki/Natural_satellite), within the [A Ring](/wiki/Rings_of_Saturn#A_Ring), of the [planet](/wiki/Planet) Saturn, which was imaged by Cassini on 15 April 2013.[[59]](#cite_note-59)

### Planetary rings[[edit](/index.php?title=(none)&action=edit&section=11)]

[Template:Main](/wiki/Template:Main) [Template:Multiple image](/wiki/Template:Multiple_image)

Saturn is probably best known for the system of [planetary rings](/wiki/Planetary_ring) that makes it visually unique.[[15]](#cite_note-15) The rings extend from 6,630 km to 120,700 km outward from Saturn's equator, average approximately 20 meters in thickness and are composed of 93% water ice with traces of [tholin](/wiki/Tholin) impurities and 7% amorphous [carbon](/wiki/Carbon).[[60]](#cite_note-60) The particles that make up the rings range in size from specks of dust up to 10 m.[[61]](#cite_note-61) While the other [gas giants](/wiki/Gas_giant) also have ring systems, Saturn's is the largest and most visible.

There are two main hypotheses regarding the origin of the rings. One hypothesis is that the rings are remnants of a destroyed moon of Saturn. The second hypothesis is that the rings are left over from the original [nebular](/wiki/Nebula) material from which Saturn formed. Some ice in the E ring comes from the moon Enceladus's geysers.<ref name=Spahn/> [[62]](#cite_note-62)[[63]](#cite_note-63) In the past, astronomers once thought the rings formed alongside the planet when it formed billions of years ago.[[64]](#cite_note-64) Instead, the age of these planetary rings is probably some hundreds of millions of years.[[65]](#cite_note-65) Beyond the main rings at a distance of 12 million km from the planet is the sparse Phoebe ring, which is tilted at an angle of 27° to the other rings and, like [Phoebe](/wiki/Phoebe_(moon)), orbits in [retrograde](/wiki/Retrograde_motion) fashion.[[66]](#cite_note-66) Some of the moons of Saturn, including [Pandora](/wiki/Pandora_(moon)) and [Prometheus](/wiki/Prometheus_(moon)), act as shepherd moons to confine the rings and prevent them from spreading out.<ref name=russell2004/> [Pan](/wiki/Pan_(moon)) and [Atlas](/wiki/Atlas_(moon)) cause weak, linear density waves in Saturn's rings that have yielded more reliable calculations of their masses.[[67]](#cite_note-67)

## History of observation and exploration[[edit](/index.php?title=(none)&action=edit&section=12)]

[Template:Main](/wiki/Template:Main)

There have been three main phases in the observation and exploration of Saturn. The first era was ancient observations (such as with the [naked eye](/wiki/Naked_eye)), before the invention of the modern telescopes. Starting in the 17th century progressively more advanced telescopic observations from Earth have been made. The other type is visitation by spacecraft, either by orbiting or flyby. In the 21st century observations continue from the Earth (or Earth-orbiting observatories) and from the Cassini orbiter at Saturn.

### Ancient observations[[edit](/index.php?title=(none)&action=edit&section=13)]

[Template:See also](/wiki/Template:See_also)

Saturn has been known since prehistoric times.[[68]](#cite_note-68) In ancient times, it was the most distant of the five known planets in the Solar System (excluding Earth) and thus a major character in various mythologies. [Babylonian astronomers](/wiki/Babylonian_astronomy) systematically observed and recorded the movements of Saturn.<ref name=ptrsl276\_1257\_43/> In [ancient Roman mythology](/wiki/Roman_mythology), the god [Saturnus](/wiki/Saturn_(mythology)), from which the planet takes its name, was the god of agriculture.[[69]](#cite_note-69) The Romans considered Saturnus the equivalent of the [Greek god](/wiki/Greek_mythology) [Cronus](/wiki/Cronus).[[69]](#cite_note-69) The Greeks had made the outermost planet sacred to Cronus,<ref name=evans1998/> and the Romans followed suit. (In modern [Greek](/wiki/Greek_language), the planet retains its ancient name *Cronus*—Κρόνος: *Kronos*.)[[70]](#cite_note-70) The Greek scientist [Ptolemy](/wiki/Ptolemy) based his calculations of Saturn's orbit on observations he made while it was in [opposition](/wiki/Opposition_(planets)).<ref name=ps04\_1893\_862/> In [Hindu astrology](/wiki/Jyotisha), there are nine astrological objects, known as [Navagrahas](/wiki/Navagraha). Saturn, one of them is known as "[Shani](/wiki/Shani)" and judges everyone based on the good and bad deeds performed in life.[[69]](#cite_note-69)<ref name=ps04\_1893\_862/> Ancient [Chinese](/wiki/Chinese_astrology) and Japanese culture designated the planet Saturn as the "earth star" ([Template:Lang](/wiki/Template:Lang)). This was based on [Five Elements](/wiki/Five_elements_(Chinese_philosophy)) which were traditionally used to classify natural elements.[[71]](#cite_note-71)[[72]](#cite_note-72)[[73]](#cite_note-73) In ancient [Hebrew](/wiki/Hebrew_language), Saturn is called 'Shabbathai'.[[74]](#cite_note-74) Its angel is [Cassiel](/wiki/Cassiel). Its intelligence or beneficial spirit is [Agiel](/wiki/Agiel) (layga) and its spirit (darker aspect) is Zazel (lzaz). In [Ottoman Turkish](/wiki/Ottoman_Turkish_language), [Urdu](/wiki/Urdu) and [Malay](/wiki/Malay_language), its name is 'Zuhal', derived from [Arabic](/wiki/Arabic_language) زحل.

### European observations (17th–19th centuries)[[edit](/index.php?title=(none)&action=edit&section=14)]

[thumb|](/wiki/File:Saturn_Robert_Hooke_1666.jpg)[Robert Hooke](/wiki/Robert_Hooke) noted the shadows (a and b) cast by both the globe and the rings on each other in this drawing of Saturn in 1666.

Saturn's rings require at least a 15-mm-diameter [telescope](/wiki/Telescope)<ref name=binoculars/> to resolve and thus were not known to exist until [Galileo](/wiki/Galileo) first saw them in 1610.<ref name=chan2000/><ref name=cain2008\_15390/> He thought of them as two moons on Saturn's sides.<ref name=cain2008\_15418/><ref name=cain2008\_46237/> It was not until [Christiaan Huygens](/wiki/Christiaan_Huygens) used greater telescopic magnification that this notion was refuted. Huygens discovered Saturn's moon Titan; [Giovanni Domenico Cassini](/wiki/Giovanni_Domenico_Cassini) later discovered four other moons: [Iapetus](/wiki/Iapetus_(moon)), [Rhea](/wiki/Rhea_(moon)), [Tethys](/wiki/Tethys_(moon)) and [Dione](/wiki/Dione_(moon)). In 1675, Cassini discovered the gap now known as the Cassini Division.<ref name=micek2007/>

No further discoveries of significance were made until 1789 when [William Herschel](/wiki/William_Herschel) discovered two further moons, [Mimas](/wiki/Mimas_(moon)) and [Enceladus](/wiki/Enceladus_(moon)). The irregularly shaped satellite [Hyperion](/wiki/Hyperion_(moon)), which has a [resonance](/wiki/Orbital_resonance) with Titan, was discovered in 1848 by a British team.<ref name=pa54\_122/>

In 1899 [William Henry Pickering](/wiki/William_Henry_Pickering) discovered [Phoebe](/wiki/Phoebe_(moon)), a highly [irregular satellite](/wiki/Irregular_satellite) that does not rotate synchronously with Saturn as the larger moons do.<ref name=pa54\_122/> Phoebe was the first such satellite found and it takes more than a year to orbit Saturn in a [retrograde orbit](/wiki/Retrograde_orbit). During the early 20th century, research on Titan led to the confirmation in 1944 that it had a thick atmosphere – a feature unique among the Solar System's moons.<ref name=apj100\_378/>

### Modern NASA and ESA probes[[edit](/index.php?title=(none)&action=edit&section=15)]

#### Pioneer 11 flyby[[edit](/index.php?title=(none)&action=edit&section=16)]

[thumb|Pioneer 11 image of Saturn](/wiki/File:P11saturnb.jpg)

[Pioneer 11](/wiki/Pioneer_11) carried out the first flyby of Saturn in September 1979, when it passed within 20,000 km of the planet's cloud tops. Images were taken of the planet and a few of its moons, although their resolution was too low to discern surface detail. The spacecraft also studied Saturn's rings, revealing the thin F-ring and the fact that dark gaps in the rings are bright when viewed at high [phase angle](/wiki/Phase_angle_(astronomy)) (towards the Sun), meaning that they contain fine light-scattering material. In addition, Pioneer 11 measured the temperature of Titan.[[75]](#cite_note-75)

#### Voyager flybys[[edit](/index.php?title=(none)&action=edit&section=17)]

In November 1980, the [Voyager 1](/wiki/Voyager_1) probe visited the Saturn system. It sent back the first high-resolution images of the planet, its rings and satellites. Surface features of various moons were seen for the first time. Voyager 1 performed a close flyby of Titan, increasing knowledge of the atmosphere of the moon. It proved that Titan's atmosphere is impenetrable in visible wavelengths, therefore no surface details were seen. The flyby changed the spacecraft's trajectory out from the plane of the Solar System.[[76]](#cite_note-76) Almost a year later, in August 1981, [Voyager 2](/wiki/Voyager_2) continued the study of the Saturn system. More close-up images of Saturn's moons were acquired, as well as evidence of changes in the atmosphere and the rings. Unfortunately, during the flyby, the probe's turnable camera platform stuck for a couple of days and some planned imaging was lost. Saturn's gravity was used to direct the spacecraft's trajectory towards Uranus.[[76]](#cite_note-76) The probes discovered and confirmed several new satellites orbiting near or within the planet's rings, as well as the small [Maxwell Gap](/wiki/Maxwell_Gap) (a gap within the [C Ring](/wiki/Rings_of_Saturn#C_Ring)) and [Keeler gap](/wiki/Keeler_gap) (a 42 km wide gap in the [A Ring](/wiki/A_Ring)).

#### Cassini–Huygens spacecraft[[edit](/index.php?title=(none)&action=edit&section=18)]

On 1 July 2004, the [Cassini–Huygens](/wiki/Cassini–Huygens) [space probe](/wiki/Space_probe) performed the SOI (Saturn Orbit Insertion) maneuver and entered orbit around Saturn. Before the SOI, Cassini had already studied the system extensively. In June 2004, it had conducted a close flyby of [Phoebe](/wiki/Phoebe_(moon)), sending back high-resolution images and data. [thumb|left|](/wiki/File:NASA-Cassini-Saturn-TitanFlybyTests-20140617.jpg)[Cassini's](/wiki/Cassini_(spacecraft)) [Titan](/wiki/Titan_(moon)) flyby radio-signal studies (artist's concept) Cassini's flyby of Saturn's largest moon, Titan, has captured radar images of large lakes and their coastlines with numerous islands and mountains. The orbiter completed two Titan flybys before releasing the [Huygens probe](/wiki/Huygens_probe) on 25 December 2004. Huygens descended onto the surface of Titan on 14 January 2005, sending a flood of data during the atmospheric descent and after the landing.<ref name=nature438\_7069\_758/> Cassini has since conducted multiple flybys of Titan and other icy satellites. [thumb|right|Photograph of](/wiki/File:Earth-Moon_system_as_seen_from_Saturn_(PIA17171).jpg) [Earth](/wiki/Earth) and the [Moon](/wiki/Moon) by [Cassini](/wiki/Cassini_spacecraft), visible in the bottom-right. [thumb|Animation of Saturn's north polar vortex seen in infrared light](/wiki/File:Rotatingsaturnhexagon.gif) Since early 2005, scientists have been tracking lightning on Saturn. The power of the lightning is approximately 1,000 times that of lightning on Earth.[[77]](#cite_note-77) In 2006, NASA reported that Cassini had found evidence of liquid water reservoirs that erupt in [geysers](/wiki/Geyser) on Saturn's moon [Enceladus](/wiki/Enceladus_(moon)). Images had shown jets of icy particles being emitted into orbit around Saturn from vents in the moon's south polar region. According to Andrew Ingersoll, California Institute of Technology, "Other moons in the Solar System have liquid-water oceans covered by kilometers of icy crust. What's different here is that pockets of liquid water may be no more than tens of meters below the surface."[[78]](#cite_note-78) Over 100 geysers have been identified on Enceladus.[[79]](#cite_note-79) In May 2011, NASA scientists at an Enceladus Focus Group Conference reported that Enceladus "is emerging as the most habitable spot beyond Earth in the Solar System for life as we know it".[[80]](#cite_note-80)[[81]](#cite_note-81) Cassini photographs have led to other significant discoveries. They have revealed a previously undiscovered planetary ring, outside the brighter main rings of Saturn and inside the G and E rings. The source of this ring is hypothesized to be the crashing of a meteoroid off [Janus](/wiki/Janus_(moon)) and [Epimetheus](/wiki/Epimetheus_(moon)).[[82]](#cite_note-82) In July 2006, Cassini images provided evidence of hydrocarbon lakes near Titan's north pole, the presence of which were confirmed in January 2007. In March 2007, additional images near Titan's north pole revealed hydrocarbon seas, the largest of which is almost the size of the [Caspian Sea](/wiki/Caspian_Sea).[[83]](#cite_note-83) In October 2006, the probe detected an 8,000 km diameter cyclone-like storm with an eyewall at Saturn's south pole.[[84]](#cite_note-84)[thumb|left|At Enceladus's south pole geysers spray water from many locations along the](/wiki/File:Enceladus_geysers.jpg) [tiger stripes](/wiki/Tiger_stripes_(Enceladus)).[[79]](#cite_note-79) From 2004 to 2 November 2009, the probe discovered and confirmed eight new satellites. Its primary mission ended in 2008 when the spacecraft had completed 74 orbits around the planet. The probe's mission was extended to September 2010 and then extended again to 2017, to study a full period of Saturn's seasons.[[85]](#cite_note-85) In April 2013 Cassini sent back images of a hurricane at the planet's north pole 20 times larger than those found on Earth, with winds faster than 530 km/h.[[86]](#cite_note-86) On 19 July 2013, "[The Day the Earth Smiled](/wiki/The_Day_the_Earth_Smiled)", Cassini was pointed towards Earth to capture an image of the [Earth](/wiki/Earth) and the [Moon](/wiki/Moon) (and, as well, [Venus](/wiki/Venus) and [Mars](/wiki/Mars)) as part of a natural light, multi-image portrait of the entire Saturn system. It was the first time NASA informed the people of Earth that a long-distance photo was being taken in advance.[[87]](#cite_note-87)

## Observation[[edit](/index.php?title=(none)&action=edit&section=19)]

[thumb|Amateur telescopic view of Saturn](/wiki/File:Saturn-27-03-04.jpeg) [thumb|Simulated appearance of Saturn as seen from Earth at](/wiki/File:Saturnoppositions-animated.gif) [opposition](/wiki/Opposition_(astronomy)), 2001-2029 Saturn is the most distant of the five planets easily visible to the naked eye from Earth, the other four being [Mercury](/wiki/Mercury_(planet)), [Venus](/wiki/Venus), [Mars](/wiki/Mars) and Jupiter. (Uranus and occasionally [4 Vesta](/wiki/4_Vesta) are visible to the naked eye in dark skies.) Saturn appears to the naked eye in the night sky as a bright, yellowish point of light with an [apparent magnitude](/wiki/Apparent_magnitude) of usually between +1 and 0. It takes approximately 29.5 years for the planet to complete an entire circuit of the [ecliptic](/wiki/Ecliptic) against the background constellations of the [zodiac](/wiki/Zodiac). Most people will require an optical aid (very large binoculars or a small telescope) that magnifies at least 30 times to achieve an image of Saturn's rings, in which clear resolution is present.[[15]](#cite_note-15)<ref name=binoculars/> Twice every Saturnian year (roughly every 15 Earth years), the rings briefly disappear from view, due to the way in which they are angled and because they are so thin.[[88]](#cite_note-88) Such a "disappearance" will next occur in 2025, but Saturn will be too close to the Sun for any ring-crossing observation to be possible.[[89]](#cite_note-89) Saturn and its rings are best seen when the planet is at, or near, [opposition](/wiki/Opposition_(astronomy)), the configuration of a planet when it is at an [elongation](/wiki/Elongation_(astronomy)) of 180°, and thus appears opposite the Sun in the sky. A Saturnian opposition occurs every year—approximately every 378 days—and results in the planet appearing at its brightest. However, both the Earth and Saturn orbit the Sun on eccentric orbits, which means their distances from the Sun vary over time, and therefore so do their distances from each other, hence varying the brightness of Saturn from one opposition to the other. Also, Saturn appears brighter when the rings are angled such that they are more visible. For example, during the opposition of 17 December 2002, Saturn appeared at its brightest due to a favorable [orientation of its rings](/wiki/File:Saturnoppositions.jpg) relative to the Earth,[[90]](#cite_note-90) even though Saturn was closer to the Earth and Sun in late 2003.[[90]](#cite_note-90)[thumb|Saturn eclipses the Sun, as seen from](/wiki/File:Saturn_eclipse.jpg) [Cassini](/wiki/Cassini–Huygens). The rings are clearly visible, including the F Ring. Also, from time to time Saturn is [occulted](/wiki/Occultation) by the Moon (that is, the Moon covers up Saturn in the sky). As with all the planets in the Solar System, occultations of Saturn occur in “seasons”. Saturnian occultations will take place 12 or more times over a 12-month period, followed by about a five-year period in which no such activity is registered.[[91]](#cite_note-91) Australian astronomy experts Hill and Horner explain the seasonal nature of Saturnian occultations:

This is the result of the fact that the moon’s orbit around the Earth is tilted to the orbit of the Earth around the Sun – and so most of the time, the moon will pass above or below Saturn in the sky, and no occultation will occur. It is only when Saturn lies near the point that the moon’s orbit crosses the “plane of the ecliptic” that occultations can happen – and then they occur every time the moon swings by, until Saturn moves away from the crossing point.[[91]](#cite_note-91)

## In culture[[edit](/index.php?title=(none)&action=edit&section=20)]

[thumb|Saturn, from a 1550 edition of](/wiki/File:Saturn-bonatti.PNG) [Guido Bonatti's](/wiki/Guido_Bonatti) *Liber astronomiae* [Template:Further](/wiki/Template:Further)

* [Saturn in astrology](/wiki/Saturn_(astrology)) ([22px](/wiki/File:Saturn_symbol.svg)) is the ruling planet of [Capricorn](/wiki/Capricorn_(astrology)) and, traditionally, [Aquarius](/wiki/Aquarius_(astrology)).
* *Saturn,* [*the Bringer of Old Age*](/wiki/Saturn_(mythology)) is a movement in [Gustav Holst's](/wiki/Gustav_Holst) [*The Planets*](/wiki/The_Planets).
* The [Saturn](/wiki/Saturn_(rocket_family)) family of [rockets](/wiki/Rocket) were developed by a team of mostly German rocket scientists led by [Wernher von Braun](/wiki/Wernher_von_Braun) to launch heavy payloads to Earth orbit and beyond.<ref name=bilstein1999/> Originally proposed as a military satellite launcher, they were adopted as the launch vehicles for the [Apollo program](/wiki/Apollo_program).
* The day Saturday is, like the planet Saturn, named after the Roman god of agriculture, [Saturn](/wiki/Saturn_(mythology)) (linked to the Greek god [Cronus](/wiki/Cronus)).<ref name=plotner20080222/><ref name=reis\_jones2009/>

## See also[[edit](/index.php?title=(none)&action=edit&section=21)]

[Template:Portal](/wiki/Template:Portal) [Template:Wikipedia books](/wiki/Template:Wikipedia_books)

* [Dragon Storm (astronomy)](/wiki/Dragon_Storm_(astronomy))
* [Sixth planet (disambiguation)](/wiki/Sixth_planet_(disambiguation))
* [Space exploration](/wiki/Space_exploration)
* [Timeline of the far future](/wiki/Timeline_of_the_far_future)

[Template:Clear](/wiki/Template:Clear)

## Notes[[edit](/index.php?title=(none)&action=edit&section=22)]

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## References[[edit](/index.php?title=(none)&action=edit&section=23)]

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## Further reading[[edit](/index.php?title=(none)&action=edit&section=24)]

* [Template:Cite magazine](/wiki/Template:Cite_magazine)
* [Template:Cite book](/wiki/Template:Cite_book)
* [Template:Cite book](/wiki/Template:Cite_book)

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[Template:Spoken Wikipedia](/wiki/Template:Spoken_Wikipedia) [Template:Sister project links](/wiki/Template:Sister_project_links)

* [Saturn profile](http://solarsystem.nasa.gov/planets/profile.cfm?Object=Saturn) at NASA's Solar System Exploration site
* [Saturn Fact Sheet](http://nssdc.gsfc.nasa.gov/planetary/factsheet/saturnfact.html), by NASA
* [Gazeteer of Planetary Nomenclature – Saturn (USGS)](http://planetarynames.wr.usgs.gov/?System=Saturn)
* [Cassini–Huygens mission](http://saturn.jpl.nasa.gov/) to Saturn, by NASA
* [Research News](http://www.sciencedaily.com/news/space_time/saturn/) about Saturn
* [General information](http://www.solarviews.com/eng/saturn.htm) about Saturn
* [Studies on the Rings](http://www.affs.org/html/studies_on_the_rings_of_saturn.html) of Saturn
* [Astronomy Cast: Saturn](http://www.astronomycast.com/2007/10/episode-59-saturn/)
* [Saturn in Daytime (12 inch telescope)](https://secure.flickr.com/photos/vontom/5362248377/)
* [Saturn 'Rev 175' Raw Preview](http://www.ciclops.org/view_event/185/Saturn_Rev_175_Raw_Preview?js=1)
* [Template:Cite web](/wiki/Template:Cite_web)
* [BBC In Our Time radio program](http://www.bbc.co.uk/programmes/b06vmr1m)

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[Template:Authority control](/wiki/Template:Authority_control)

[\*](/wiki/Category:Saturn) [Category:Gas giants](/wiki/Category:Gas_giants) [Category:Astronomical objects known since antiquity](/wiki/Category:Astronomical_objects_known_since_antiquity)