

# Trajectories with prescribed itineraries in CR3BP

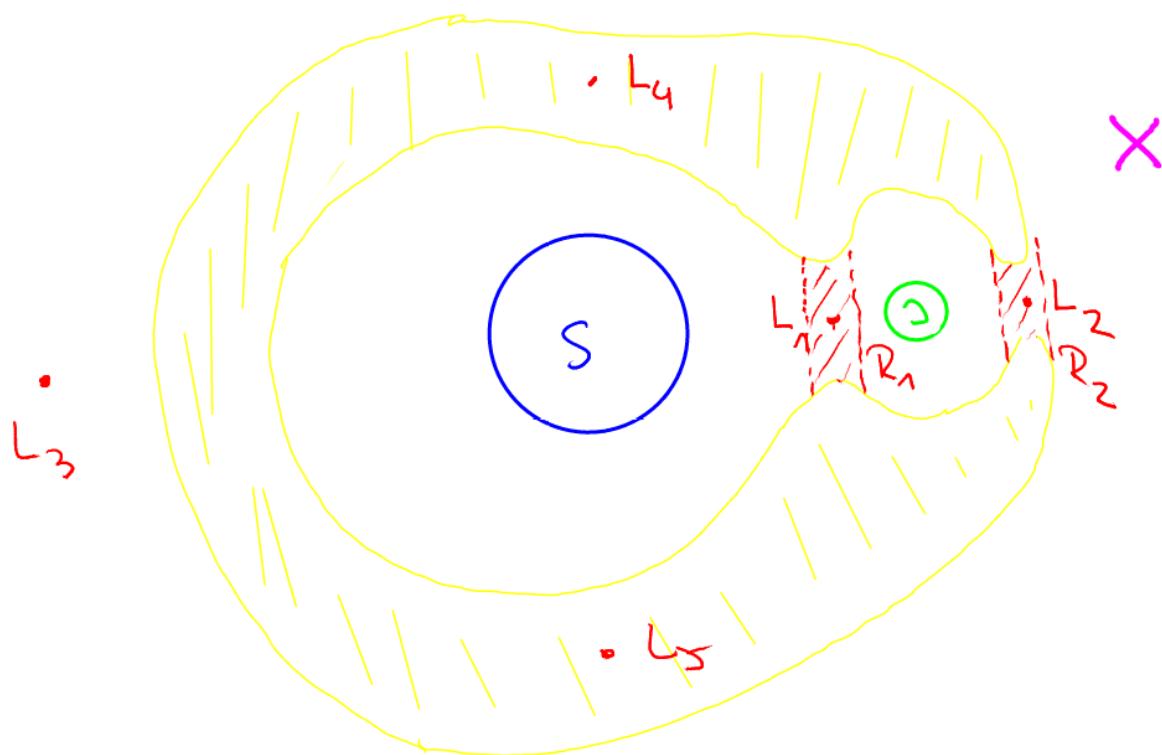
25.1.2017, Univ Augsburg, Pavel

## Outline :

- 1) Motivation: Orbits with prescribed itineraries
- +  
H  
F  
O  
R  
T 2) Find them using tube dynamics of Lagrangian points
- 3) When is their existence granted?
- D  
A  
X.  
n. 4) How to find Lyapunov orbits numerically?
- R  
S 5) How to find stable and unstable manifolds and Poincaré sections numerically?
- 6) Program

①  
o) Motivation: Orbits with prescribed itinerary

CR3BP, rotating frame, fixed E



Hill's Region  $\rightarrow$  two necks  $\Rightarrow$  realms

X, D, S

Goal: Find IC of an orbit which follows

$(\dots, X, D, S, D, X, \dots)$

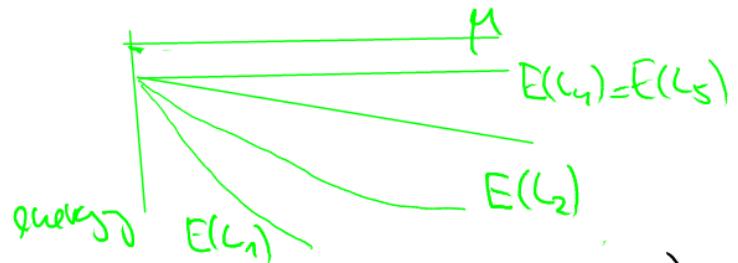
(2)

1) Find 'em using the dynamics of Logr. p.t.s.

a)  $\exists$  hyperbolic equilibrium points  $L_1$  and  $L_2$   
(and  $L_3, L_4, L_5$ )

b) Linearize Ham. system around  $L$  for  $E$

Slightly above  $E(L)$



$\mu \ll 1 \rightarrow E(L_2)$  only slightly above  $E(L_1)$

↑ so  $E$  can be chosen slightly above both

$$S \rightarrow 9.537 \times 10^{-4} \quad (\text{lin. } L \sim 0, E \sim E - E(L))$$

$\Rightarrow \forall E > E(L) \exists!$  p.o. of the linearized

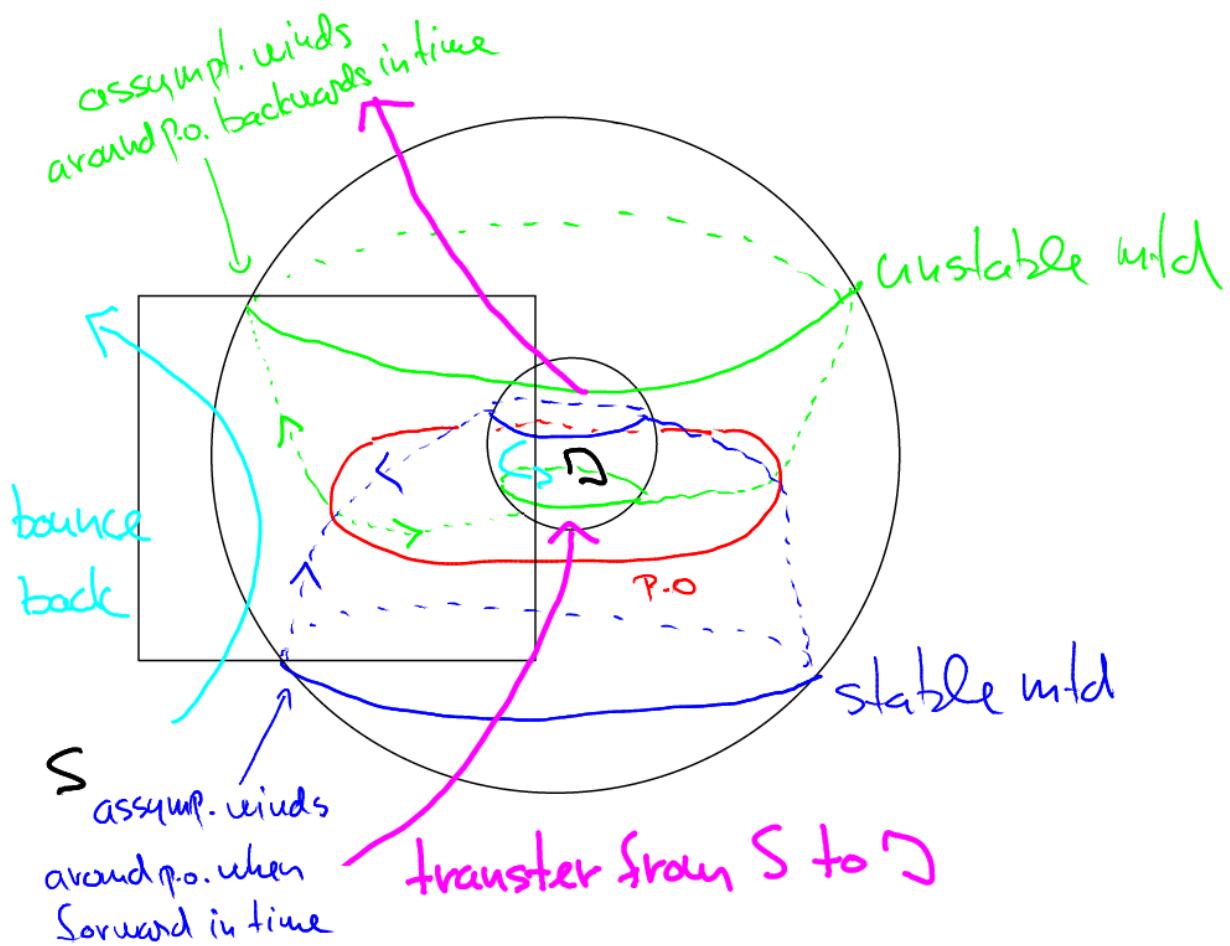
system and its neighborhood in the energy hyper-

surface in  $H(E) \cap \mathbb{R}^2$  looks like  $S^2 \times I$ , where the  
 (energy hypersurface)

dynamics is in the DeGieter representation

picturized as follows:

(3)



Note: 1) For  $E \rightarrow E(L)$ , the circles degenerate to points in the plane, the p.o. to L, in particular.

c) Nosé's theorem about existence of a real analytic change of coordinates  $(x_1, v_x, v_y) \sim (\xi, \eta, \zeta, \epsilon_1, \epsilon_2)$  around L in  $T^*\mathbb{R}^2$ , where the solutions of the original system have manifestly the same form as the solutions of the linearized system  $\rightarrow$

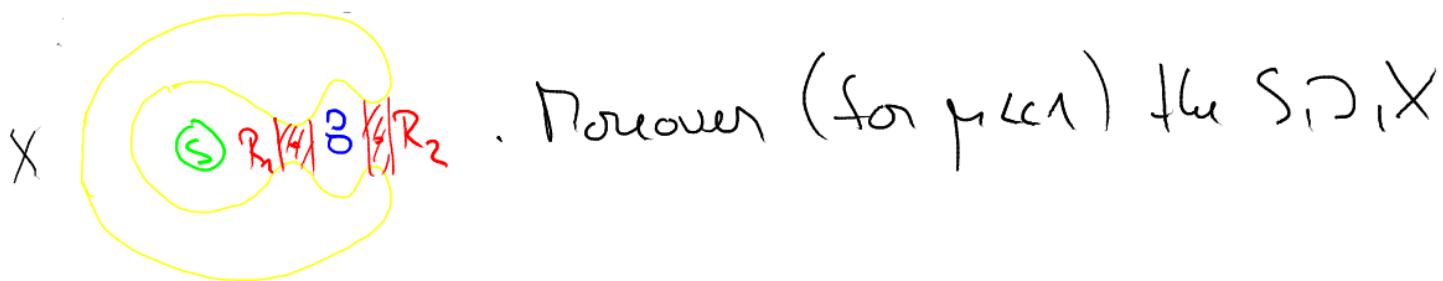
(4)

$\Rightarrow \forall E > E(L)$  slightly above  $\exists!$  periodic orbit  
 $\text{Lyapunov orbit } \gamma_L$

and its nbhd  $R$  in  $M(E)$  such that the dynamics  
in  $R$  is described qualitatively by the McGehee  
spheres.

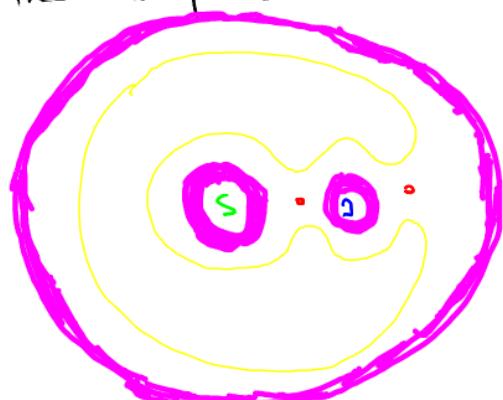
d) Pair of McGehee theorems  $\Rightarrow$

$\exists \sigma$  open nbhd of the graph  $E_L(\mu)$  for  $\mu \ll 1$ ,  
so that for  $(\mu, \varepsilon) \in \sigma$  the region  $R$  as above can  
be chosen, s.t. it projects to  $R \subset \mathbb{R}^2$  in the form



are separated by KAM-tori from the rest of the

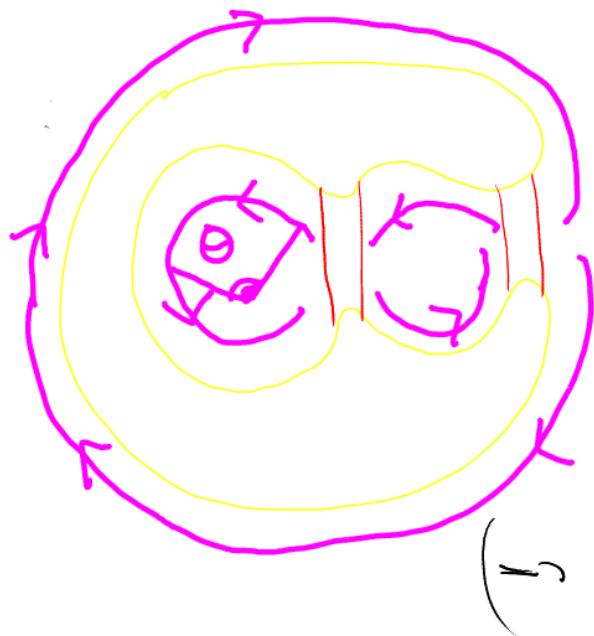
dynamics



(The smaller  $\mu$ , the

(5)

Closer are the KAN-tori to the bdd.). Moreover,



→ meridional angular coordinate  $\theta$  on complements of  $R$  in  $D(p_1, \epsilon)$  which strictly increases along orbits  
 (→ orbits go around  $X, S$  in our direction)

2) To the hyperbolic equilibrium point  $L$

(normal definition), equivalently to  $\gamma_L$  ( $\omega$ -limit set definition) are associated stable and unstable manifolds  $\underline{W}_s(L)$  and  $\underline{W}_u(L)$ .

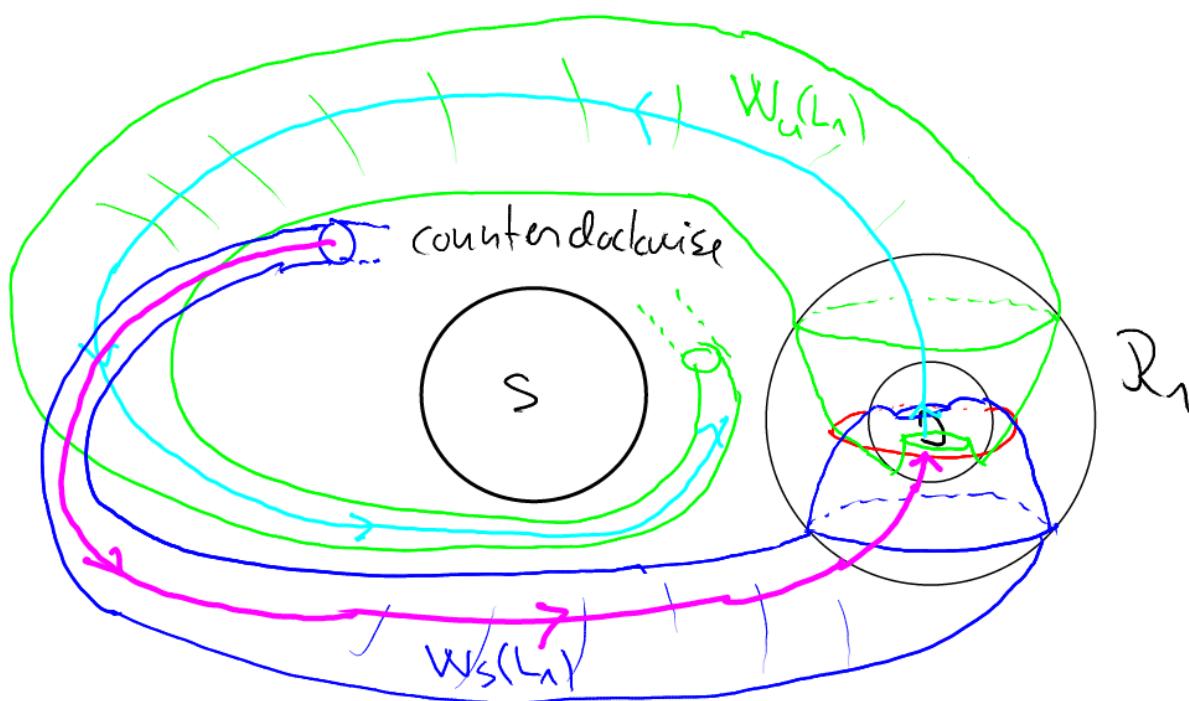
Properties: ① Global 2-D submanifolds of  $D(\epsilon)$ ,  
 (energy hypersurf.)

they are annuli (cylinders)

②  $\underline{W}_s(L)$  is the set of orbits which asymptotically wind around  $\gamma_L$  in the forward time.

$W_u(L)$  is the set of orbits which asymptotically wind around  $\gamma_L$  in the backward time. (6)

- (3) They wind around  $S, \gamma, X$  in the direction of increasing  $\Theta$ .
- (4) They start in  $\mathbb{R}$  from  $\gamma_L$ . Eg (picture in  $D(\epsilon)$ ):



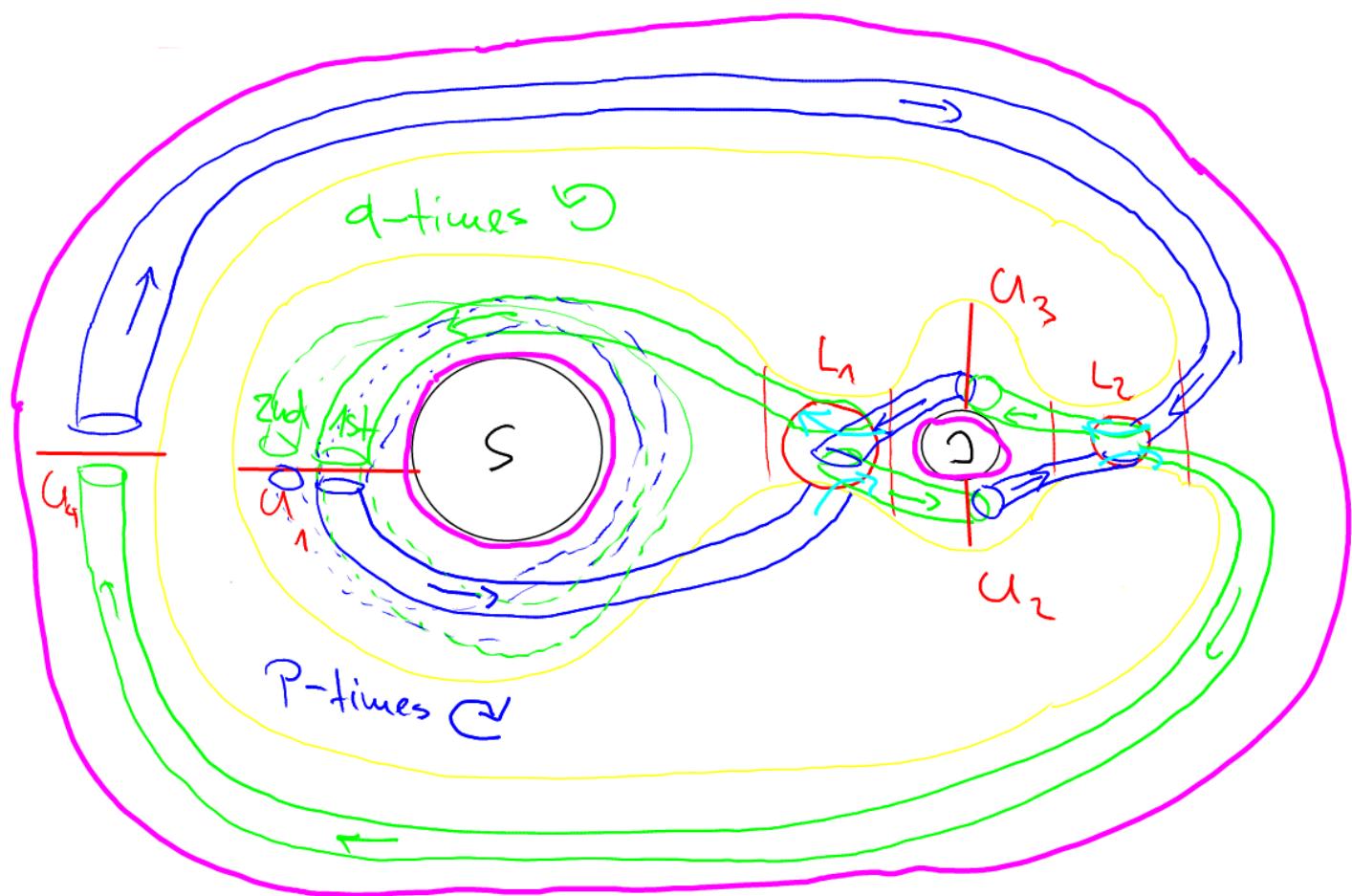
- (5) All trajectories inside of the tubes bounded by  $W_s(L), \text{ resp. } W_u(L)$  are transfer trajectories ( $\dots S \circ \dots$ , resp.  $\dots \circ S \dots$ ). The closer to  $\partial W$ , the more time they spend winding around  $\gamma_L$ . In particular  
 (Note:  $W_s(L) \cap W_u(L)$  homoclinic pts,  $W_s(L_1) \cap W_u(L_2)$  heteroclinic pts)

(7)

if  $IC \in \overset{\circ}{W_u(L_i)} \cap \overset{\circ}{W_s(L_i)}$   $\Rightarrow$  the corresponding orbit has itinerary (...DS...).

f) The concept of the tube dynamics is to search for trajectories with a given itinerary by searching for a "common intersections" of  $\overset{\circ}{W_s(L)}$  and  $\overset{\circ}{W_u(L)}$ .

The whole picture:



(8)

This is feasible by looking at intersections with suitable sections  $U_1, U_2, U_3, U_4$  in  $\Pi(\epsilon)$ :

$$U_1 = \{(x, \dot{x}) \mid \gamma=0, x<0, \dot{x}(x, \dot{x}, \epsilon) < 0\}$$

(i.e. projection of  $\Pi(\epsilon) \cap \{\gamma=0, x<0, \dot{x}<0\}$  to the  $(x, \dot{x})$ -plane)

$$U_2 = \{(\gamma, \dot{\gamma}) \mid x=0, \gamma<0, \dot{x}(\gamma, \dot{\gamma}, \epsilon) > 0\}$$

$$U_3 = \{(\gamma, \dot{\gamma}) \mid x=0, \gamma>0, \dot{x}(\gamma, \dot{\gamma}, \epsilon) < 0\}$$

$$U_4 = \{(x, \dot{x}) \mid \gamma=0, x<-1, \dot{x}(x, \dot{x}, \epsilon) > 0\}$$

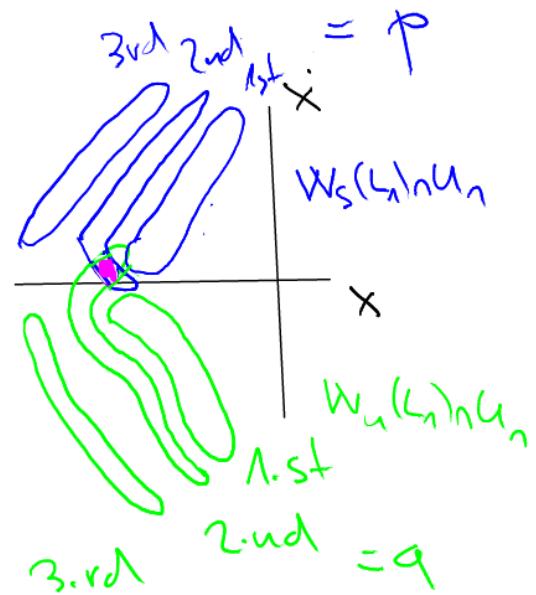
$\Rightarrow$  Get pictures like

this in  $U_1$ :

Note: the symmetry of the EOM

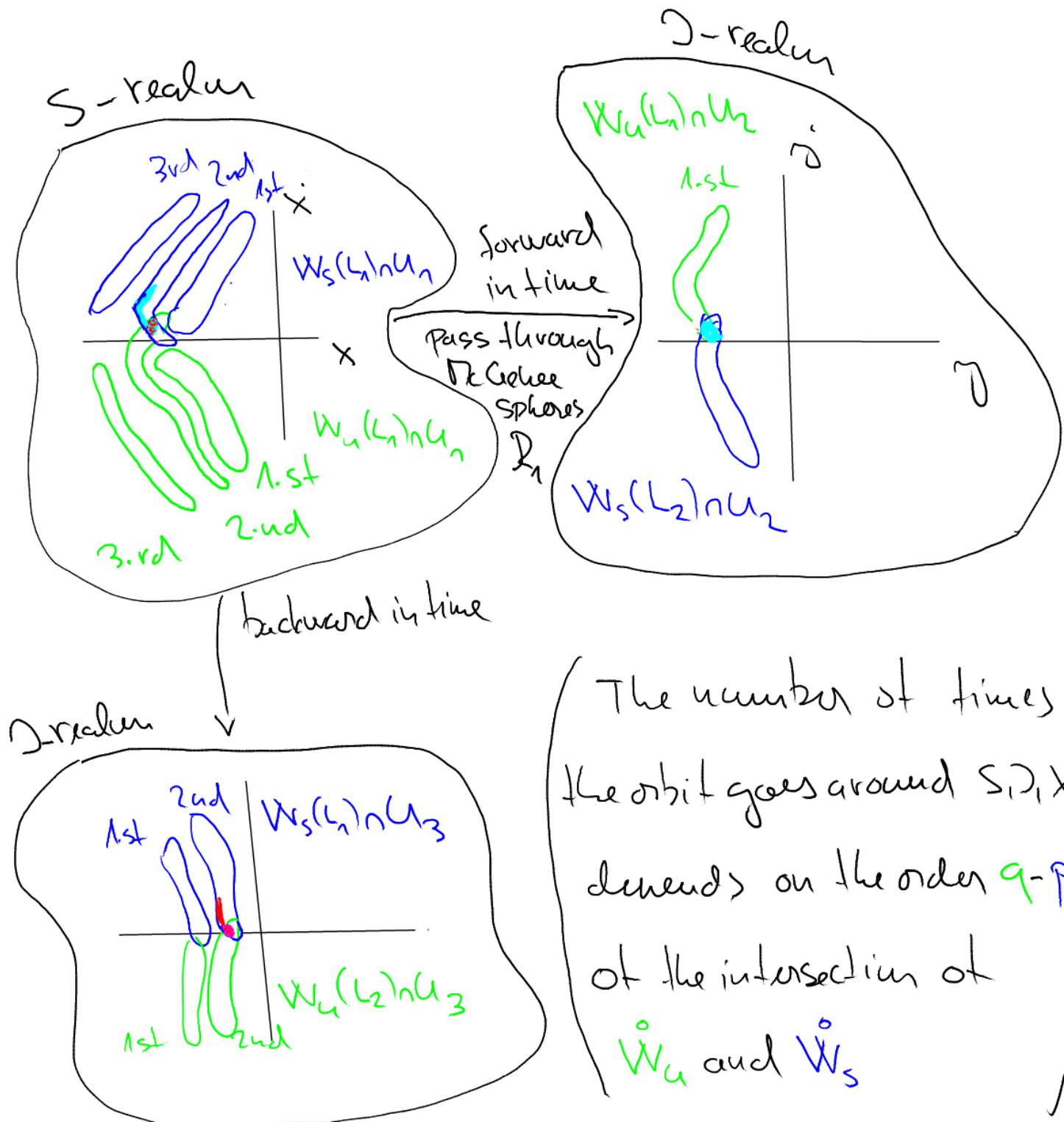
$s: (x, \gamma, v_x, v_\gamma, t) \mapsto (x, -\gamma, -v_x, v_\gamma, t)$

is manifested in  $U_1$



(9)

# Process of finding a trajectory by looking at $U \cap W$ :

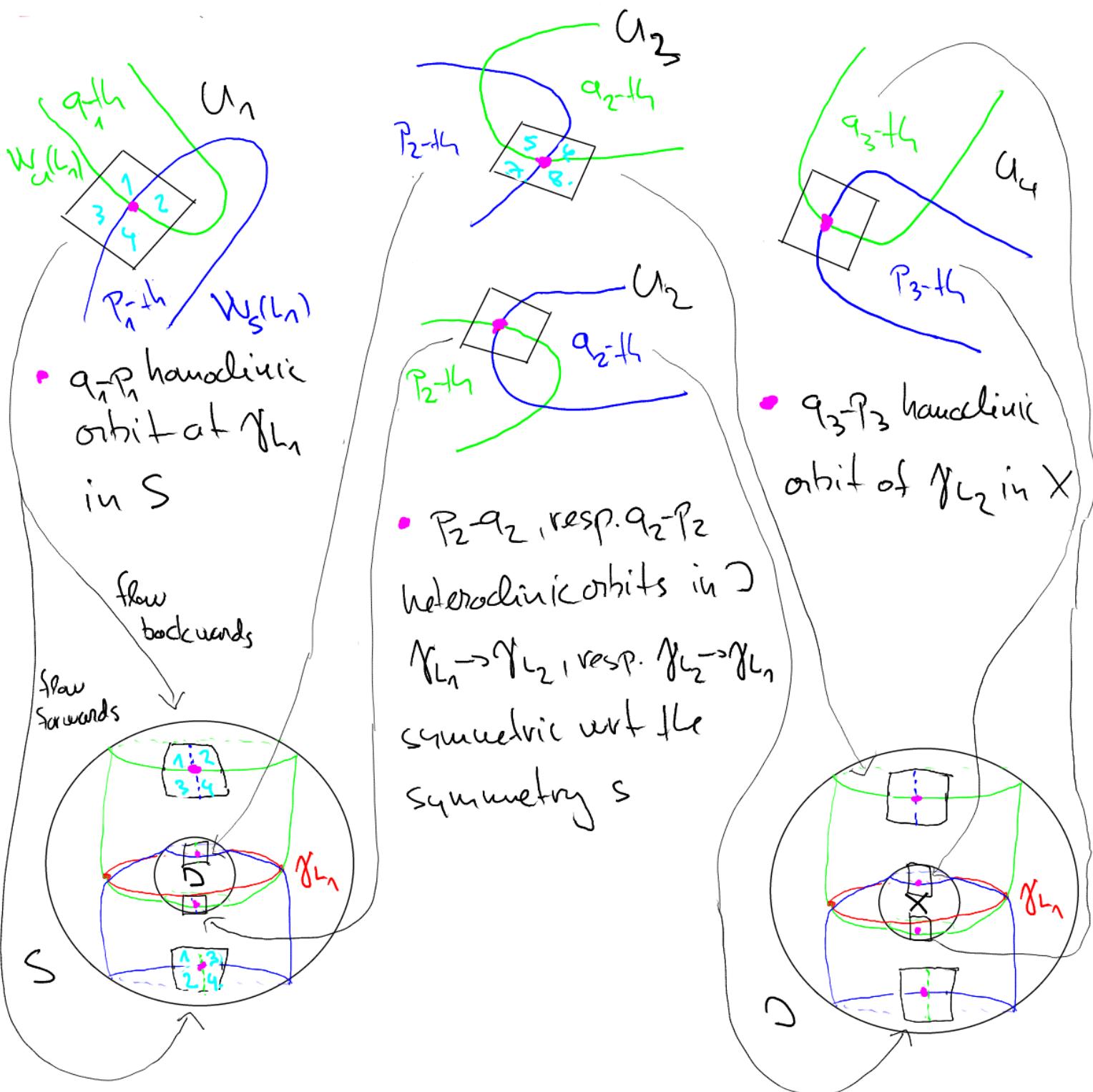


The trajectories in • have itinerary  
 $(\dots X \circ S \circ X \dots)$

2) When is the existence of the trajectories  
guaranteed?

Assume  $\Rightarrow$  transverse homoclinic-heteroclinic

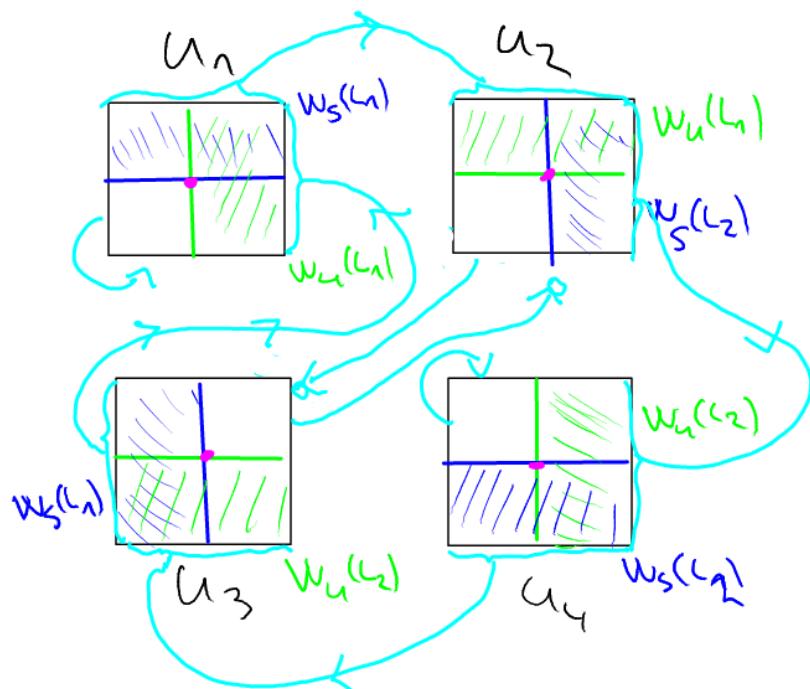
chain, i.e. for given  $E \Rightarrow$  transverse intersections



→ Construct Poincaré map

$$P: U := U_1 \cup U_2 \cup U_3 \cup U_4 \rightarrow U$$

$q \mapsto$  the 1st intersection  
in the forward time

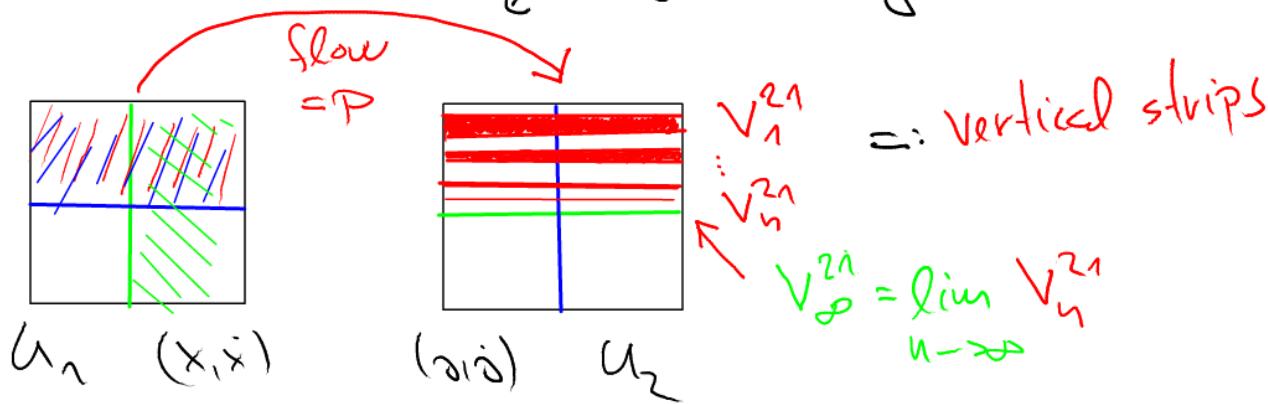


in the forward time  
of  $\phi_t(q)$  with  $U$   
if exists

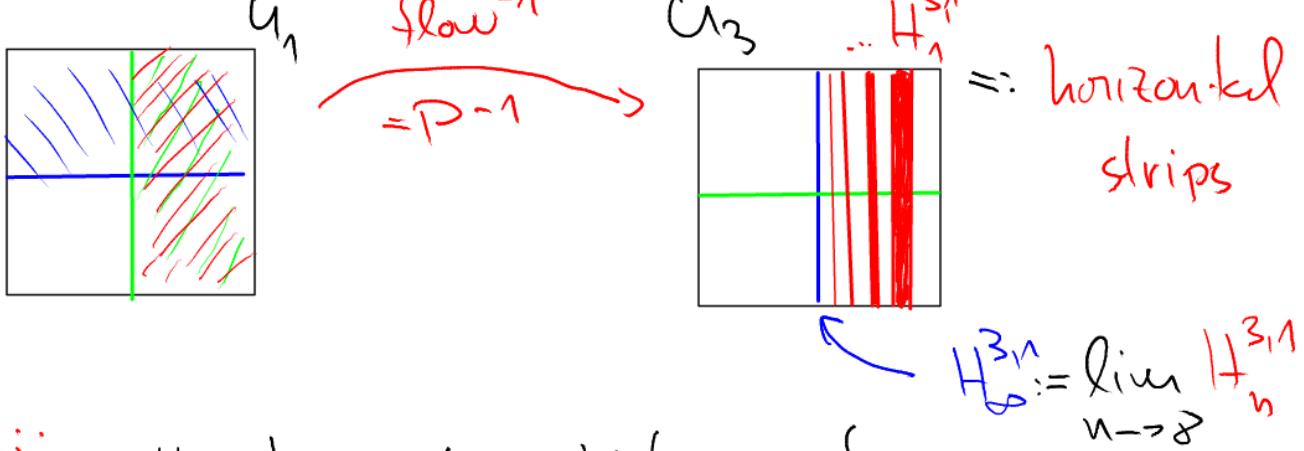
$P$  not defined on  
+ since the orbits  
there remain winding  
around  $\gamma_L$  for  $\infty$  time

Close up:

One gets this spiraling behaviour  
by studying the lin. system in  $R_1$



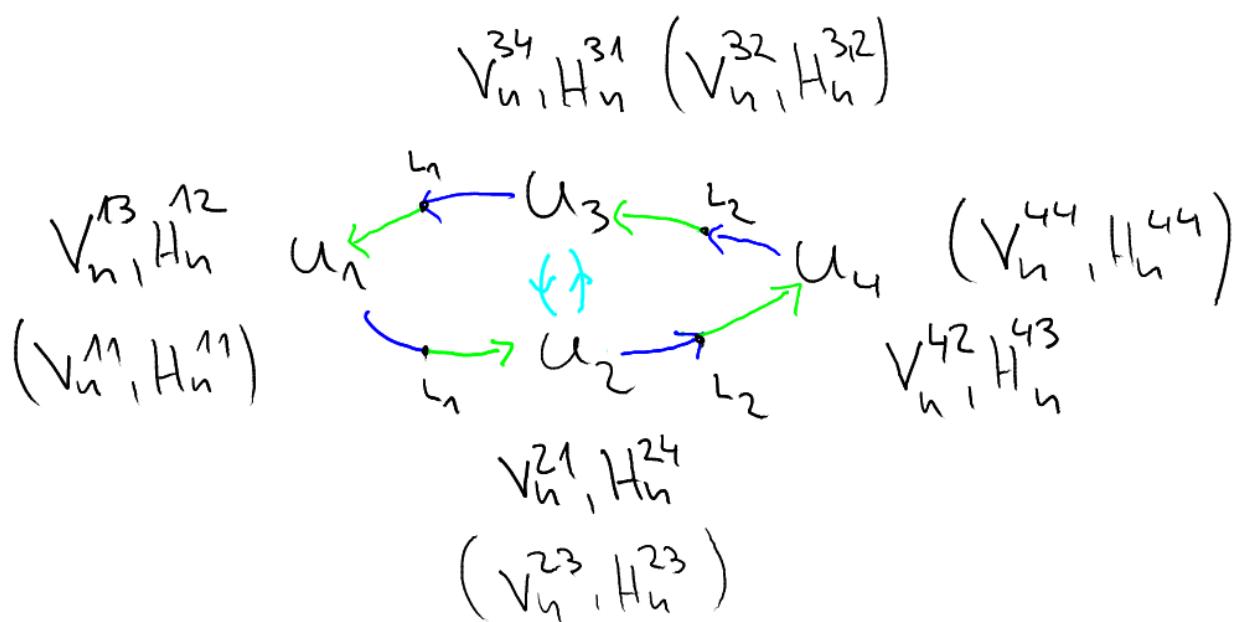
$V_n^{21}$  n-th strip in  $U_2$  which came from  $U_1$



$H_n^{3i}$   $n$ -th strip in  $U_3$  which came from

$U_1$  by the negative flow, in other words, will go to  
 $U_1$  in forward time

→ 8 families of pairwise disjoint vertical and horizontal strips :



Can be proven, the strips satisfy generalized

(B)

Couley-Rosser conditions :

1)  $P$  maps  $(H, \mathcal{H}) \hookrightarrow (V, \mathcal{V})$  (strips condition)

2)  $P$  uniform contraction/expansion in the horizontal/vertical direction

$\Rightarrow$  horseshoe-like chaotic dynamics

"

In particular, the invariant set

$$\Lambda := \bigcap_{n=1}^{\infty} P^n(U) \cap U \cap P^{-n}(U) \quad (P(\Lambda) \subset \Lambda \text{ by def.})$$

of orbits remaining in  $U$  forever can be identified with

biinfinite sequences

address of a point in  $\Lambda$

$j_i$  (its history) future under

$P^n|_P$ )

$$(\dots (u_{i-1}, u_i), (u_{i-2}, u_i), (u_{i-3}, u_i), \dots)$$

•  $u_i \in \{u_1, u_2, u_3, u_4\}$  ... where the point is

•  $n_i$  ... corresponds to index  $n$  of  $|^{ij}u_i|^{ij}v_n$

= number of revolutions around  $\gamma_L$  on the way from  $j$  to  $i$

By construction, every address is traced by an orbit  $\gamma$  (14) going through  $a_i$ , winding around  $\gamma_L$   $n_i$ -times in forward time as  $i \mapsto i+1$ .

→ In particular,  $\exists$  of an orbit with any given itinerary  $(\dots s_{-n} s_0 s_1 \dots)$ ,  $s_i \in \{x, \gamma, S\}$

Note: The number of revolutions around  $S, \gamma, X$  is the same for all orbits and is given by the type  $q_i - p_i$  of the transversal homoclinic/heteroclinic points. Eg.  $q_1 + p_1 - 1, q_3 + p_3 - 1$  for  $S$  and  $X$ , and  $\frac{q_2 + p_2 - 1}{2}$  for  $\gamma$ .

Authors prove numerically existence of transverse homo- and hetero-clinic chains for energies slightly

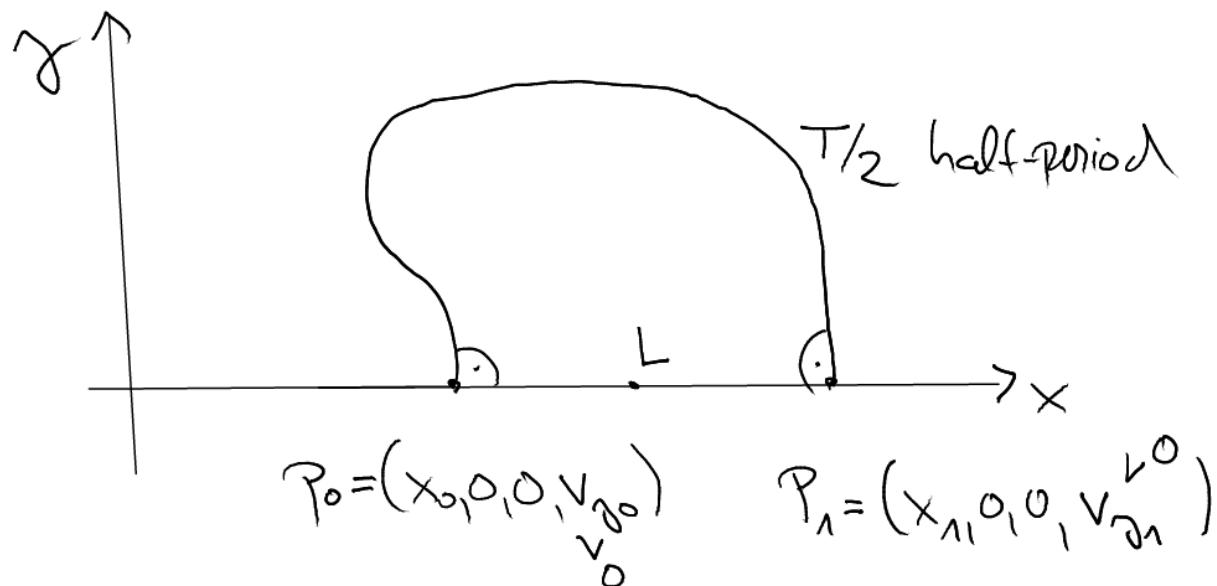
above  $E(L_2)$  and  $\mu = 9.537 \times 10^{-4}$  (Sun-Dupiter)

## 4) How to find Lyapunov orbits numerically? (15)

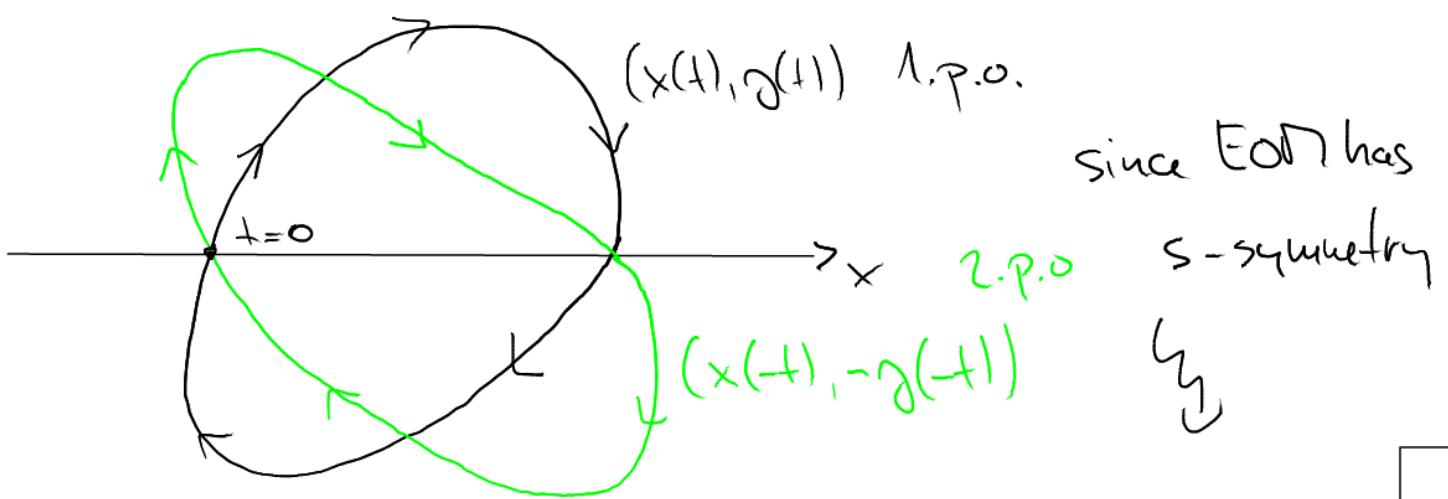
We know from the linearization:

$\forall E > E(L)$  slightly  $\exists!$  p.o.  $\gamma_L$  around  $L$

Claim: Has to be symmetric on reflection about the x-axis. It looks like:



Proof: By contradiction with uniqueness:

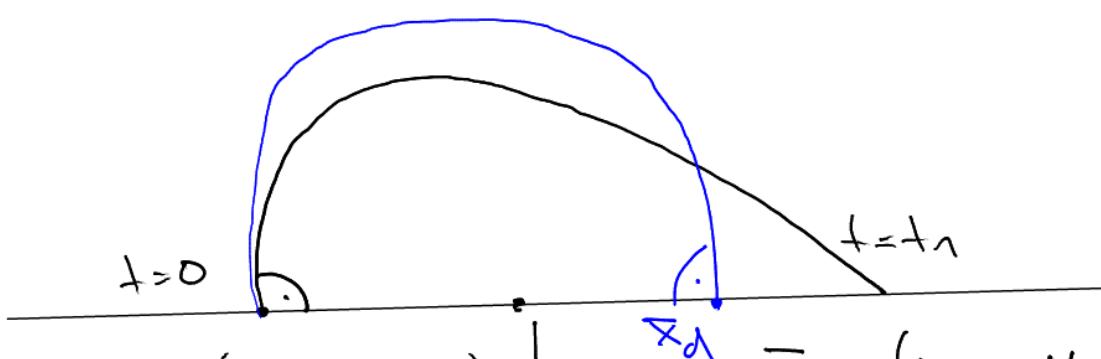


since EOM has  
s-symmetry



16

→ Use shooting ":



$$\bar{x}_0 = (x_0, 0, 0, v_{y0}) \quad L \quad \bar{x}_1 = (x_1, 0, v_{x1}, v_{y1})$$

$$\bar{x}_d = (x_1, 0, 0, v_{y1}) \rightarrow \underline{\delta \bar{x}_1} = \bar{x}_1 - \bar{x}_d = (0, 0, v_{x1}, 0)$$

Noividly adjust  $\delta \bar{x}_0 = \underline{\Phi}(\bar{x}_0, t_1)^{-1} (\bar{x}_1 - \bar{x}_d)$

↑  
state transition matrix along  $\gamma$

$\underline{\Phi}$  is a solution of  $\dot{\underline{\Phi}}(t) = Df(\gamma(t)) \underline{\Phi}(t)$ ,  $\underline{\Phi}(0) = 1$

$$\text{i.e. } \underline{\Phi}(t) = \text{Top} \left( \int_0^t Df(\gamma(u)) du \right) =$$

$$= 1 + \left\{ \int_0^t Df(\gamma(u)) du + \frac{1}{2!} \left\{ \int_0^t \int_0^t Df(\gamma(u_1)) Df(\gamma(u_2)) du_1 du_2 + \dots \right\} \right\}$$

$$= 1 + \left\{ \int_0^t Df(\gamma(u)) du + \int_0^t \int_0^{t_1} Df(\gamma(u_1)) Df(\gamma(u_2)) dt_2 dt_1 + \dots \right\}$$

But we do the following:

1) Fix  $\bar{x}_0$  and adjust  $v_{\bar{x}_0}$  only

$$\rightarrow \delta \bar{x}_0 = (0, 0, 0, \delta v_{\bar{x}_0})$$

2) Take into account sensitivity on the final time  $t_1$

$$\delta \bar{x}(t + \delta t) = \Phi(t + \delta t, \bar{x}_0 + \delta \bar{x}_0) - \Phi(t, \bar{x}_0)$$

$$\begin{aligned} \dot{x} &= \underbrace{\frac{\partial \Phi}{\partial t}(t, \bar{x}_0) \delta t}_{\text{Taylor}} + \underbrace{\frac{\partial \Phi}{\partial x}(t, \bar{x}_0) \delta \bar{x}_0}_{\text{flow}} + \text{h.o.t} \\ &= \bar{\Phi}(t, \bar{x}_0) \end{aligned}$$

$$\rightarrow \delta \bar{x}_1 = \bar{\Phi}(t, \bar{x}_0) \delta \bar{x}_0 + \dot{\bar{x}}(t) \delta t_1 \quad \text{to the 1st. order}$$

$$\begin{pmatrix} 0 \\ 0 \\ v_{x_1} \\ 0 \end{pmatrix} = \bar{\Phi} \begin{pmatrix} 0 \\ 0 \\ 0 \\ \delta v_{\bar{x}_0} \end{pmatrix} + \begin{pmatrix} v_{x_1} \\ v_{\bar{x}_0} \\ \dot{v}_{x_1} \\ \dot{v}_{\bar{x}_0} \end{pmatrix} \delta t_1$$

$$\left. \begin{aligned} v_{x_1} &= \bar{\Phi}_{34} \delta v_{\bar{x}_0} + \dot{v}_{x_1} \delta t_1 \\ 0 &= \bar{\Phi}_{24} \delta v_{\bar{x}_0} + v_{\bar{x}_0} \delta t_1 \end{aligned} \right\} \text{solve for } \delta v_{\bar{x}_0}$$

(18)

$$v_{x_1} v_{y_1} = \Phi_{34} \delta v_{y_0} v_{y_1} + \dot{v}_{x_1} \delta t_1 v_{y_1}$$

Subtract

$$0 = \Phi_{24} \delta v_{y_0} \dot{v}_{x_1} + v_{y_1} \delta t_1 \dot{v}_{x_1}$$

$$\Rightarrow v_x v_{y_1} = (\Phi_{34} v_{y_1} - \Phi_{24} \dot{v}_{x_1}) \delta v_{y_0}$$

$$\Rightarrow \delta v_{y_0} = \left( \Phi_{34} - \frac{1}{v_{y_1}} \Phi_{24} \dot{v}_{x_1} \right)^{-1} v_{x_1}$$

Why do they  
forget this?

Practical problem: How to pick the initial  $(x_0, 0, 0, v_{y_0})$

so that the method converges to the Lyapunov orbit?

Solution: As  $\text{LinToRot}((0, 0, -\alpha_x, 0)) =$

$$= \underbrace{(x_1 - \alpha_x, 0, 0, \underbrace{-\alpha_x v_T})}_{x_0} = \bar{x}_0 \quad \text{for an } \alpha_x$$

so that  $E(\bar{x}_0) = \epsilon$ . Here Lin is the

(19)

coordinate system  $(\xi, \eta, \epsilon_1, \epsilon_2)$  of  $T^* \mathbb{R}^2$  with respect to the basis  $u_1, u_2, R\omega_1, I\omega_1$ , where

$u_1, u_2 \in \mathbb{R}^4$ ,  $\omega_1 = \bar{\omega}_2 \in \mathbb{C}^4$  are eigenvectors of the

matrix  $A$  of the system linearized at  $L \ddot{x} = Ax$ ,

so that the linearized EoN looks like

$$\begin{cases} \dot{\xi} = \lambda \xi & \text{where } \pm \lambda \text{ are eigenvalues for } u_1, u_2 \\ \dot{\eta} = -\lambda \eta & \text{and } \pm i\gamma \text{ eigenvalues for } \omega_1, \omega_2 \\ \dot{\epsilon}_1 = \gamma \epsilon_2 \\ \dot{\epsilon}_2 = -\gamma \epsilon_1 \end{cases}$$

$$\xi(t) = \xi_0 e^{\lambda t}$$

$$\eta(t) = \eta_0 e^{-\lambda t}$$

$$\epsilon(t) = \epsilon_1(t) + i\epsilon_2(t) =$$

$$= \epsilon_0 e^{-i\gamma t}$$

$\rightarrow$  the periodic orbit  $(\eta_0 = 0 = \xi_0)$  ④

goes through  $(0, 0, -\alpha_x, 0)$  for some  $\alpha_x$

Problem 15: This initial seed works only for energies

slightly above  $E(L)$  because of the validity of the linear approximation. When looking for  $\gamma_{L_1}$ , the neck at  $L_2$  is often closed as  $E(L_1) < E < E(L_2)$ .

Sol1: But for very small  $\mu$ ,  $E(L_2)$  is only slightly above  $E(L_1)$  and so  $E$  can be chosen to be slightly above both! E.g.  $\mu = 9.537 \times 10^{-4}$

Sol2: a) Find the Lyapunov orbit  $\gamma_L^{E_0}$  for an energy  $E_0$  slightly above  $E(L)$  using differential correction with the seed from the linearization

b) Calculate  $\Phi_{T_1}^{E_0}$  (monodromy matrix =  $\Phi(T_1)$  after one period)  
along  $\gamma_L^{E_0}$ .

(21)

c)  $\Phi_{\eta}$  has eigenvalues  $1, 1, \lambda, \frac{1}{\lambda}$  where  $\lambda > 1$

(without multiplicities) using the fact it is symplectic.

Pick s.v.  $w$  s.t.  $\Phi_{\eta}w = w$  and  $w \notin \mathcal{F}_L^{\infty}(o)$

d) Shift initial condition by  $\bar{x}_0' = \bar{x}_0 + \varepsilon w$ , where

$\varepsilon < 0$  or  $\varepsilon > 0$  small, so that  $E(\bar{x}_0 + \varepsilon w) > E_0$ .

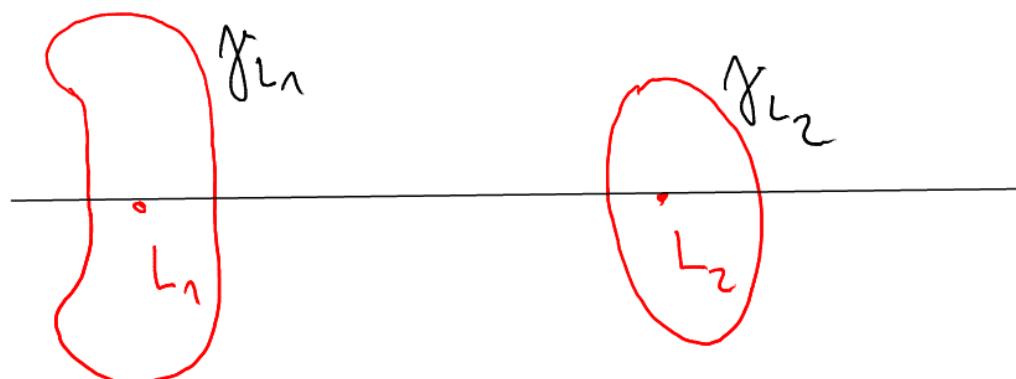
e) Use differential correction with the seed  $\bar{x}_0'$

to find the Lyapunov orbit with  $E(\bar{x}_0') > E_0$ .

$\Rightarrow$  By this process with convenient  $E$  get

Lyapunov orbits for all energies, so that both

necks are open. This is what we need.



(22)

5) How to find stable and unstable manifolds  
and Poincaré sections numerically?

a) Calculate the monodromy matrix  $\Phi_n$  along  $\gamma_L$ ,

$$\Phi_n = \phi(\bar{x}_0, T) \quad (\text{at } \bar{x}_0)$$

b)  $\Phi_n$  has e.v.  $1, 1, \lambda, \frac{1}{\lambda}$  where  $\lambda > 1$

→ eigenvectors  $\psi_0^u$  of  $\lambda > 1$

$\psi_0^s$  of  $1/\lambda < 1$

$\left( \begin{array}{l} \text{if } \lambda > 1, \psi_0^u \text{ of } 1 \text{ in the direction of} \\ \text{the family of periodic orbits} \end{array} \right)$

→  $\psi(t) := \phi(t, 0) \psi_0^s$  eigenvectors of  $\Phi_n(t)$

$\psi_u(t) := \phi(t, 0) \psi_0^u$  (at  $y_L^{(t)}$ )

along  $\gamma$

$\left( \text{This holds because } \Phi_n(t) = \phi(t, 0) \phi_n \phi(t, 0)^{-1} \right)$

(23)

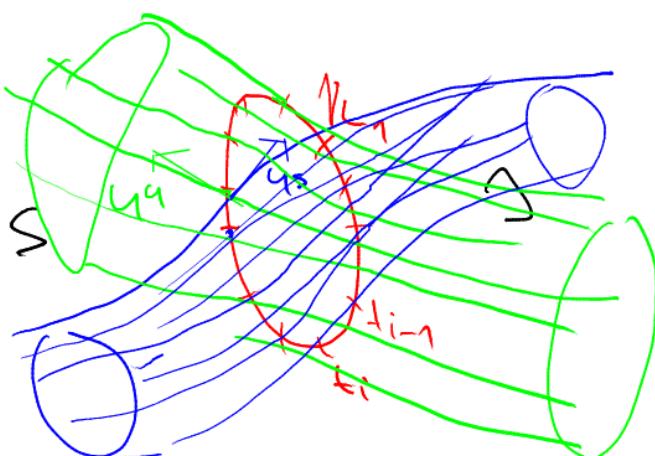
c) Integrate the initial condition  $\underline{\gamma_L(t) + \varepsilon \gamma_s(t)}$ , resp.

$$\underline{\gamma_L(t) + \varepsilon \gamma_u(t)}$$

backwards, resp. forwards in time with  $\varepsilon > 0, \varepsilon \ll 0$  small

(but not too small so that it doesn't wind  $\gamma_L$  for too long time) depending on to which realm we want to go.

Here  $0 \leq t_1 < \dots < t_n \leq T$  are chosen points on  $\gamma_L$ .



## 6) Program

Mathematica library CR3BP.wl and

a) The "explorer" - Demonstrate basic concepts for  $\mu = 0.3$

b) The "new version" - Can be used for more sophisticated calculations

Ex.:  $\mu = \frac{9.537 \times 10^{-4}}{\text{SunJupiter}, L_2, \Delta E = 0.0015}$

→ transverse 1:1 homoclinic intersection point in  $U_4$

Problem for EarthMoon :  $\mu = 1.215 \times 10^{-2}$

Get Lyapunov orbit at  $L_1$  for energy  $E > E(L_2)$

⇒ TODO: Implement Lyapunov orbits to higher energies