

Accepted —. Received —; in original form —

ABSTRACT**Key words:****1 INTRODUCTION****2 THE MAIN SYSTEM OF EQUATIONS**

The crucial part of all the dynamic equations is the non-linear term $(\mathbf{v}\nabla)\mathbf{v}$, containing all the inertial terms. In spherical coordinates:

$$\begin{aligned} (\mathbf{v}\nabla)\mathbf{v} = & \left(v_r \frac{\partial v_r}{\partial r} + \frac{v_\theta}{r} \frac{\partial v_r}{\partial \theta} + \frac{v_\varphi}{r \sin \theta} \frac{\partial v_r}{\partial \varphi} - \frac{v_\theta^2 + v_\varphi^2}{r} \right) \mathbf{e}^r \\ & + \left(v_r \frac{\partial v_\theta}{\partial r} + \frac{v_\theta}{r} \frac{\partial v_\theta}{\partial \theta} + \frac{v_\varphi}{r \sin \theta} \frac{\partial v_\theta}{\partial \varphi} + \frac{v_\theta v_r}{r} - \frac{v_\varphi^2}{r} \cot \theta \right) \mathbf{e}^\theta \\ & + \left(v_r \frac{\partial v_\varphi}{\partial r} + \frac{v_\theta}{r} \frac{\partial v_\varphi}{\partial \theta} + \frac{v_\varphi}{r \sin \theta} \frac{\partial v_\varphi}{\partial \varphi} + \frac{v_\varphi v_r}{r} + \frac{v_\theta v_\varphi}{r} \cot \theta \right) \mathbf{e}^\varphi, \end{aligned} \quad (1)$$

where \mathbf{e} are unit vectors along different coordinate directions, r is radius, θ is polar angle, φ is azimuthal angle. Directly, this expression will be used for the vertical structure equations. On the sphere, the velocities may be converted to vorticity and divergence:

$$\omega = [\nabla \times \mathbf{v}], \quad (2)$$

$$\delta = (\nabla \cdot \mathbf{v}) = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 v_r) + \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta v_\theta) + \frac{1}{r \sin \theta} \frac{\partial v_\varphi}{\partial \varphi}. \quad (3)$$

Taking curl of Euler equation allows to directly derive the equation for vorticity

$$\frac{\partial [\nabla \mathbf{v}]}{\partial t} + (\mathbf{v}\nabla) [\nabla \mathbf{v}] = (\omega \cdot \nabla) \mathbf{v} - \omega (\nabla \cdot \mathbf{v}) + \frac{1}{\rho^2} [\nabla p \times \nabla \rho]. \quad (4)$$

The right-hand-side term is important if motion is baroclinic. Baroclinic term allows to generate vorticity from non-axisymmetric perturbations of density and temperature, hence it is important for any realistic calculations. We will neglect the radial velocities and consider only the radial component of ω that means all the motions are restricted to the surface of the sphere and the vertical relaxation timescale is much smaller than the global dynamical scales. Setting $\omega = \mathbf{e}_r \omega$ and substituting it to equation (4) gives

$$\frac{\partial \omega}{\partial t} + \nabla \cdot (\omega \mathbf{v}) = \frac{1}{\rho^2} [\nabla p \times \nabla \rho]. \quad (5)$$

Additional kinetic terms disappear because of the two-dimensional nature of the flow, not due to incompressibility assumption that also zeros these terms. Current version of the code adopts a fixed equation of state hence $\nabla p \propto \nabla \rho$, and the right-hand side is zero. Note that integration in vertical coordinate is not required in this case.

Another equation comes from taking divergence of Euler equation. It is rather non-trivial to expand the advection term, $(\nabla \cdot ((\mathbf{v}\nabla)\mathbf{v}))$, therefore first let us show that

$$(\nabla \cdot [\mathbf{v}\omega]) = \nabla^2 \frac{v^2}{2} - \nabla \cdot ((\mathbf{v}\nabla)\mathbf{v}). \quad (6)$$

The last term here is identical to the advective left-hand-side term in Euler equation derivative, hence

$$\frac{\partial \delta}{\partial t} = (\nabla \cdot [\mathbf{v}\omega]) - \nabla^2 B. \quad (7)$$

where

$$B = \frac{v^2}{2} + \Phi + h \quad (8)$$

is Bernoulli integral, and $h = \int \frac{1}{\rho} dp$ is enthalpy. For the isothermal case we start with, $p = \rho c_s^2$, where the speed of sound c_s^2 is constant, and $h = c_s^2 \ln \rho$. Integration with height does not change the overall structure of the equation, as the vertical scaleheight is always identical for density and pressure.

2.1 Vertical balance

3 SPECTRAL CODE

4 DISSIPATION

To stabilize the algorithm, we use a hyperdiffusion method selective for the highest harmonics. Dissipation operator for both quantities, vorticity and divergence, equals

$$D = \frac{1}{t_D} \left(\nabla^4 - \frac{4}{R^4} \right), \quad (9)$$

where the second, correction, term serves to ensure conservation of the total angular momentum and mass (see ?). In the spectral domain, this corresponds to

$$\tilde{D} = \frac{1}{t_D} \left(ik^4 - \frac{4}{R^4} \right) \simeq e \quad (10)$$

5 INERTIAL MODES AND SONIC HORIZONS

5.1 Derivation of the dispersion equation

Using WKB method (?), one can linearize the set of dynamic equations and come up with a dispersion relation for short-wavelength waves on a sphere in presence of differential rotation. Density variations $\rho = \rho_0 + \delta\rho(\theta, \varphi, t)$, azimuthal velocity $v_\varphi = \Omega(\theta) \sin \theta + \delta v_\varphi(\theta, \varphi, t)$, and latitudinal velocity $v_\theta = \delta v_\theta(\theta, \varphi, t)$. Latitudinal velocity has the same (first) order as the azimuthal velocity perturbation. All the perturbations will be expressed in exponential form $\propto \exp(i\omega t - k_\theta \theta - k_\varphi \varphi)$. **Will spherical harmonics be better?** Wavenumbers would be expressed in the units of inverse sphere radius.

Continuity equation perturbation, written in WKB assumptions, becomes

$$(\omega - k_\varphi \Omega) \frac{\delta\rho}{\rho} = k_\theta v_\theta + \frac{1}{\sin \theta} k_\varphi \delta v_\varphi. \quad (11)$$

The two tangential Euler equations may be in general form written, ignoring second-order terms, as

$$\frac{\partial v_\theta}{\partial t} + \frac{v_\varphi}{\sin \theta} \frac{\partial v_\theta}{\partial \varphi} - v_\varphi^2 \cot \theta = -c_s^2 \frac{\partial \ln \rho}{\partial \theta} \quad (12)$$

and

$$\frac{\partial v_\varphi}{\partial t} + v_\theta \frac{\partial v_\varphi}{\partial \theta} + \frac{v_\varphi}{\sin \theta} \frac{\partial v_\varphi}{\partial \varphi} + v_\varphi v_\theta \cot \theta = -\frac{c_s^2}{\sin \theta} \frac{\partial \ln \rho}{\partial \varphi}. \quad (13)$$

In WKB approach, and after substitution the expression for the variations of ρ from (11), the equations become, respectively,

$$(\tilde{\omega}^2 - c_s^2 k_\theta^2) v_\theta = \left(c_s^2 \frac{k_\theta k_\varphi}{\sin \theta} - 2i\Omega \tilde{\omega} \cos \theta \right) \delta v_\varphi, \quad (14)$$

and

$$\left(\tilde{\omega} (\tilde{\omega} - k_\varphi \Omega) - \frac{c_s^2 k_\varphi^2}{\sin^2 \theta} \right) \delta v_\varphi = \left(c_s^2 k_\theta k_\varphi + i\tilde{\omega} \frac{\partial \Omega \sin^2 \theta}{\partial \theta} \right) \frac{v_\theta}{\sin \theta}, \quad (15)$$

where $\tilde{\omega} = \omega - k_\varphi \Omega$. Excluding the velocity components yields a dispersion equation

$$(\tilde{\omega}^2 - c_s^2 k_\theta^2) (\tilde{\omega} - k_\varphi \Omega) - \frac{c_s^2 k_\varphi^2}{\sin^2 \theta} \tilde{\omega} = i c_s^2 k_\varphi k_\theta \sin \theta \frac{\partial \Omega}{\partial \theta} + 2\Omega \tilde{\omega} \cot \theta \frac{\partial}{\partial \theta} (\Omega \sin^2 \theta). \quad (16)$$

Two important specific cases may be reproduced when $\Omega \rightarrow 0$ (sonic waves) and $c_s \rightarrow 0$ (inertial waves):

$$\omega_{\text{sonic}} = c_s^2 \left(k_\theta^2 + \frac{k_\varphi^2}{\sin^2 \theta} \right), \quad (17)$$

$$\omega_{\text{inertial}} = \frac{3}{2} k_\varphi \Omega \pm \frac{1}{2} \sqrt{(k_\varphi \Omega)^2 + 4\kappa^2}, \quad (18)$$

where

$$\kappa^2 = 2\Omega \cot \theta \frac{\partial}{\partial \theta} (\Omega \sin^2 \theta) \quad (19)$$

is the latitudinal epicyclic frequency. In the case of k_θ and $\Omega = \text{const}$, $\omega = \kappa = 2\Omega \cos \theta$ reproduces the Coriolis oscillation regime. Isomomentum rotation, on the other hand, reproduces $\kappa = 0$ and does not have oscillating purely latitudinal modes.