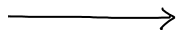


# Lecture 3: Image Geometry

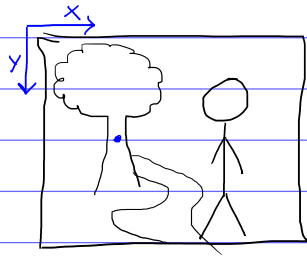
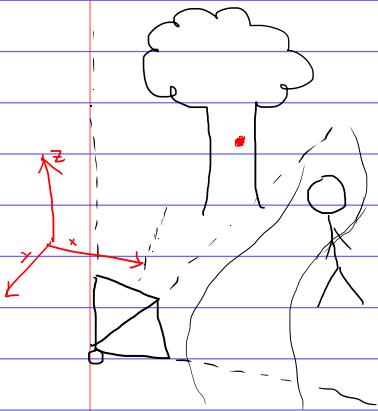
## Topics:

- Frames of reference
- Image planes
- Projection + transformations
- Homogeneous coordinates
- Two-camera geometry
- Epipolar geometry
- Depth from disparity

3D Scene



2D Image

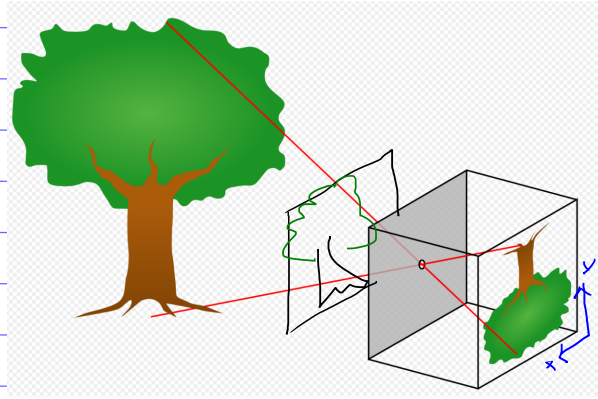


How does  $[x, y, z]$  relate to  $[x, y]$ ?

→ How do cameras work?

# Pinhole camera model:

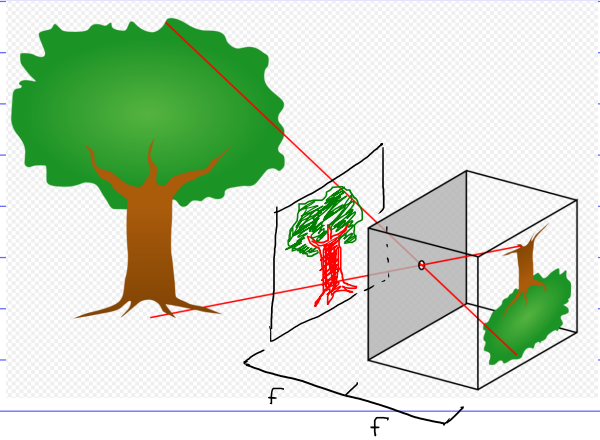
- + Simple and accurate
- Does not account for focus or depth of field

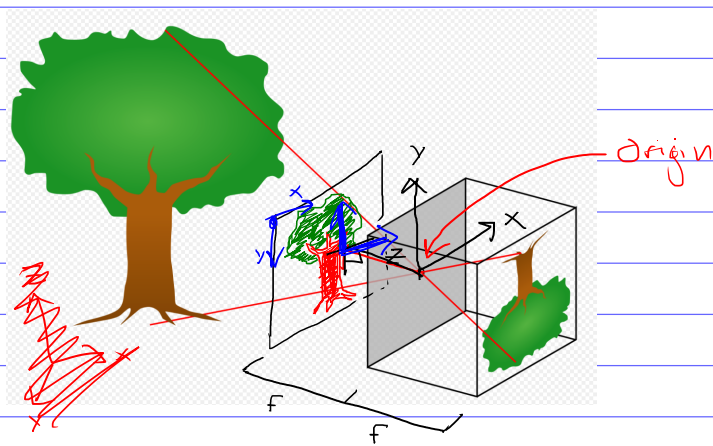


- Produces image flipped  $180^\circ$  - two ways to mitigate this

- Redefine  $x$  and  $y$

→ - Redefine image plane

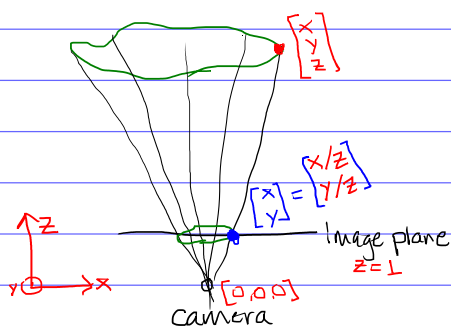




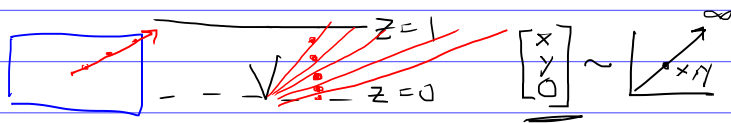
3D  $\longrightarrow$  2D

1. Choose  $[0,0,0] \rightarrow$  pinhole
2. Choose  $\begin{matrix} \uparrow y \\ x \\ \downarrow z \end{matrix} \rightarrow z$ : out of pinhole towards image plane;  $x, y =$  image  $x, y$
3. Choose  $f = 1$
4. Choose  $\begin{matrix} \uparrow y \\ x \\ \downarrow z \end{matrix} \rightarrow$  Same as  $x, y$

Then image formation is linear projection onto  $z=1$

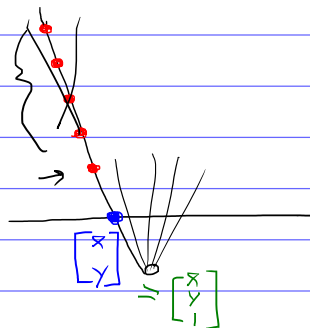


$$\begin{bmatrix} x \\ y \\ z \end{bmatrix} \xrightarrow{\text{proj}} \begin{bmatrix} x/z \\ y/z \\ 1 \end{bmatrix} \xrightarrow{2D} \begin{bmatrix} x/z \\ y/z \end{bmatrix}$$



Notes:

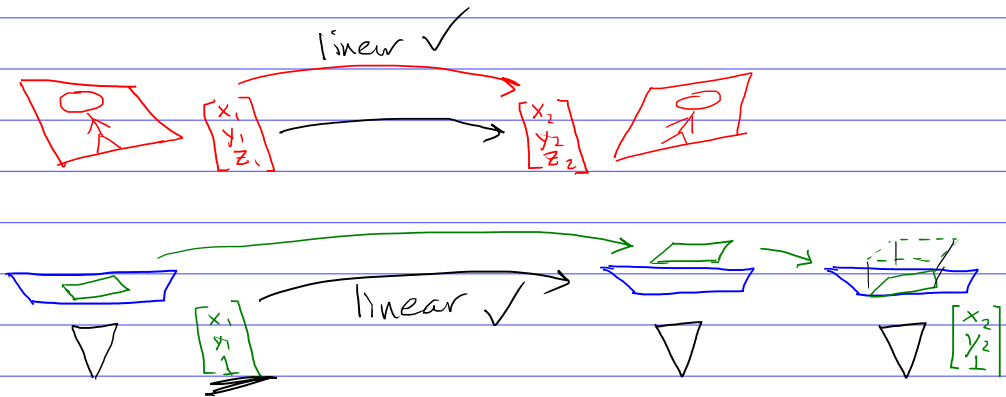
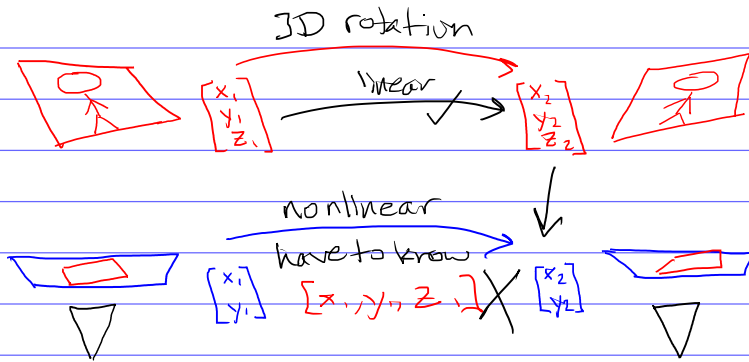
- If  $z=0$ ?  $\rightarrow$  Ideal points (infinity in a direction)
- If  $z < 0$ ?  $\rightarrow$  N/A
- Unique?  $\rightarrow$  Occlusion



# Homogeneous Coordinates

- We almost never know  $\begin{bmatrix} x \\ y \\ z \end{bmatrix}$ , only  $\begin{bmatrix} x \\ y \end{bmatrix}$
- We want to be able to represent 3D transformations in our 2D image space
- Solution: homogeneous coordinates.

$$A \star \begin{bmatrix} x \\ y \\ 1 \end{bmatrix} = \begin{bmatrix} x' \\ y' \\ z' \end{bmatrix} = \begin{bmatrix} x'/z' \\ y'/z' \\ 1 \end{bmatrix}$$



Hom. coordinates are useful for performing 3D transformations in the image plane.

Group	Examples	DoF	Matrix	Invariants
Projective	Homography, perspective warp	8	$\begin{bmatrix} h_{11} & h_{12} & h_{13} \\ h_{21} & h_{22} & h_{23} \\ h_{31} & h_{32} & h_{33} \end{bmatrix}$	Collinearity, intersection, tangency, cross ratio (ratio of ratio of lengths)
Affine	Skew, stretch	6	$\begin{bmatrix} a_{11} & a_{12} & t_x \\ a_{21} & a_{22} & t_y \\ 0 & 0 & 1 \end{bmatrix}$	All of the above, plus parallelism, ratio of areas, ratio of lengths among collinear points (e.g. midpoints), linear combinations of vectors (e.g. centroids)
Similarity	Uniform scaling	4	$\begin{bmatrix} s \cos \theta & -s \sin \theta & t_x \\ s \sin \theta & s \cos \theta & t_y \\ 0 & 0 & 1 \end{bmatrix}$	All of the above, plus ratio of lengths, angles
Euclidean	Translation, rotation	3	$\begin{bmatrix} \cos \theta & -\sin \theta & t_x \\ \sin \theta & \cos \theta & t_y \\ 0 & 0 & 1 \end{bmatrix}$	All of the above, plus length, area

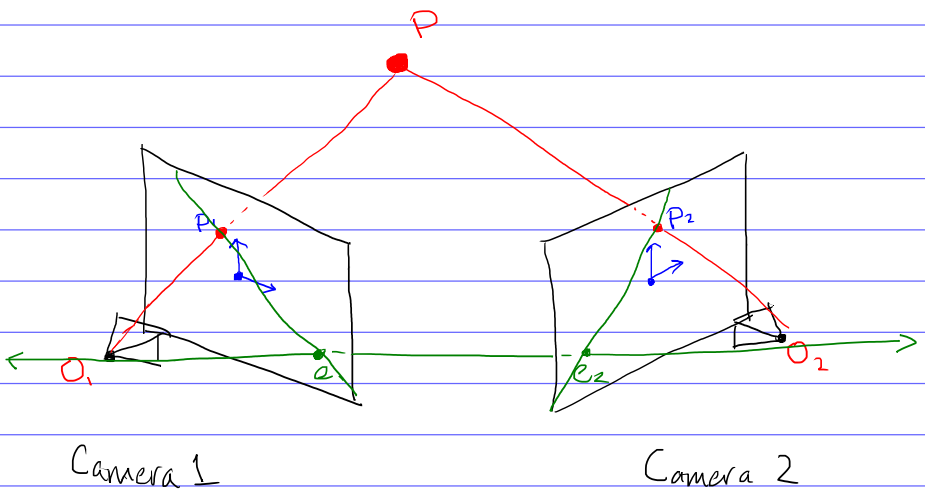


(You do not need to memorize these)

## Two-camera (Epipolar) Geometry

- Also called stereo geometry
- Yields information about scene depth ( $z$ )
- Also useful when we have one moving camera
- Note: We now have three frames of ref. and must be careful with our coordinates

Basic setup:



Terms:

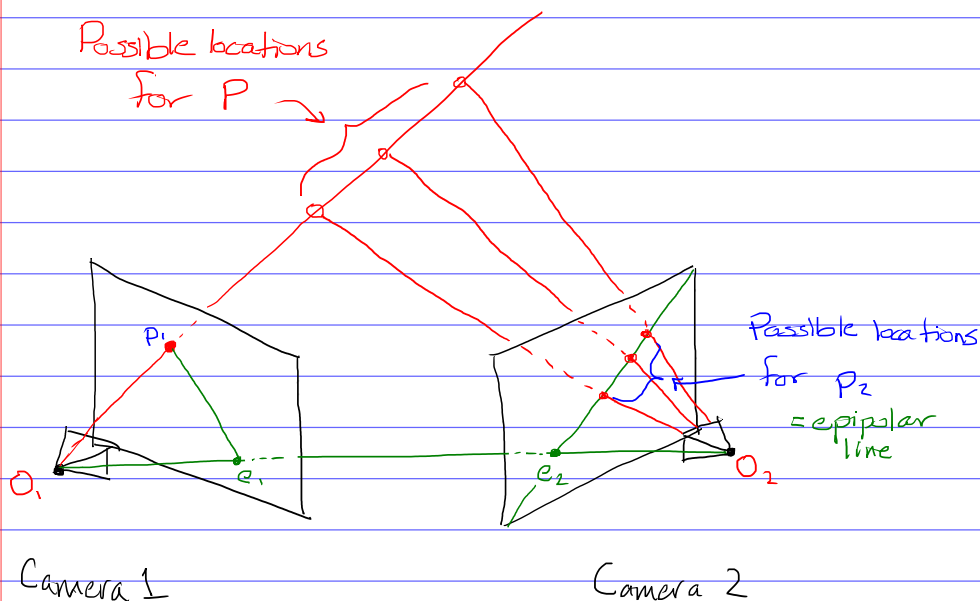
- Baseline - Line  $\overline{O_1 O_2}$
- Epipoles - Intersection of baseline and image plane
- Epipolar plane (containing  $P$ ) - plane  $O_1 O_2 P$
- Epipolar lines - Intersection of ep. plane & image plane

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Q: Suppose we know  $O_1$ ,  $O_2$ , and  $p_1$ .  
We don't know  $P$ . What do we know  
about  $p_2$ ?

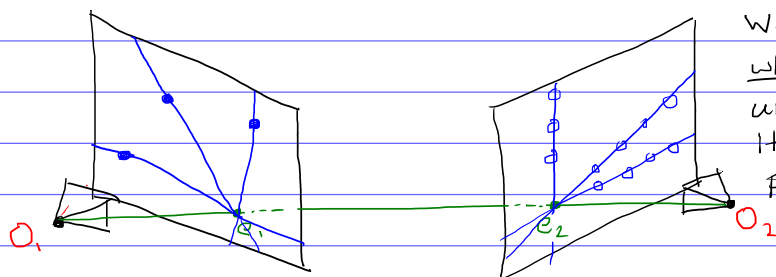
→ Useful in object tracking, scene reconstruction

A: We know  $p_2$  lies on the epipolar line



Intuition: Ep. line  $\overrightarrow{e_1 p_1}$  is  $\overrightarrow{O_2 P}$  projected onto image plane 1, and  $\overrightarrow{e_2 p_2}$  is just  $\overrightarrow{O_1 P}$  projected onto image plane 2.

Takeaway: When we move from one camera position  $O_1$  to another  $O_2$ , every point in our image must appear along an epipolar line.



We don't know where on the line unless we know its real-world position.

# Example: Forward motion

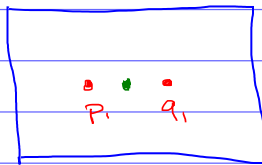
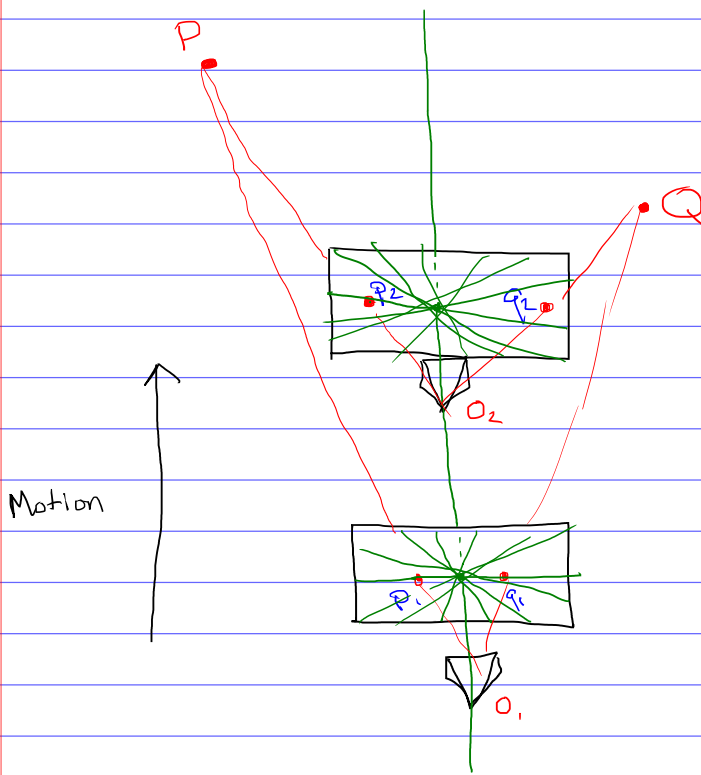


Image 1

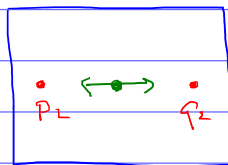
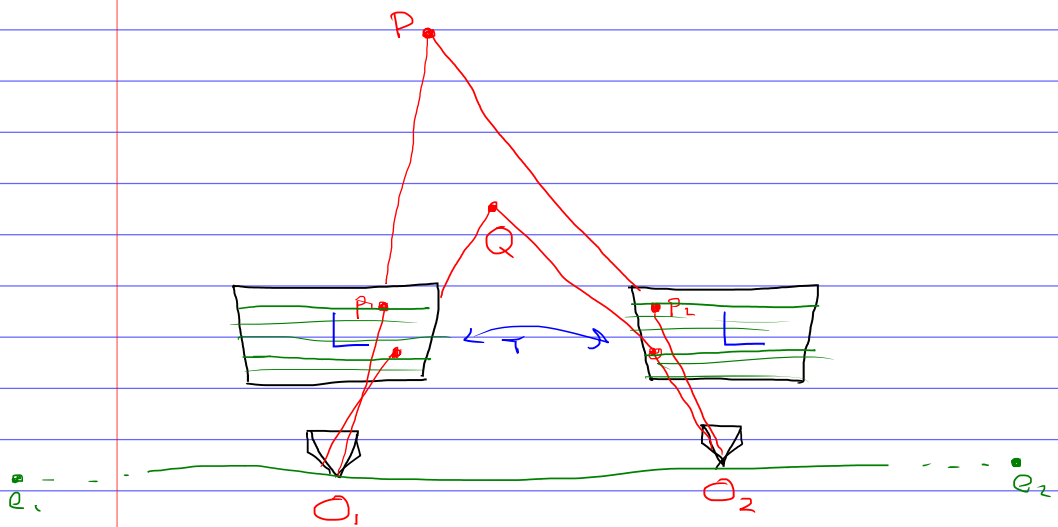


Image 2

# Example: Lateral Movement



→  
Motion

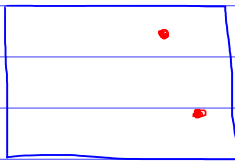


Image 1

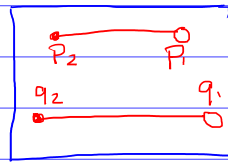


Image 2

Parallax

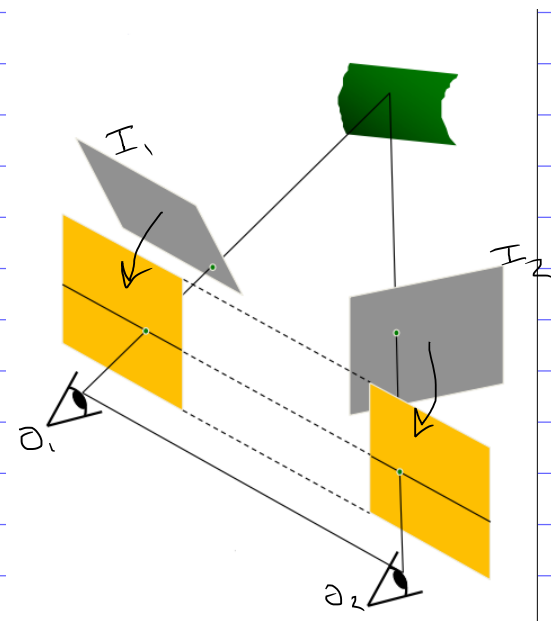


Q: How do we work in stereo algebraically?

A: With difficulty.  $\hat{=}$

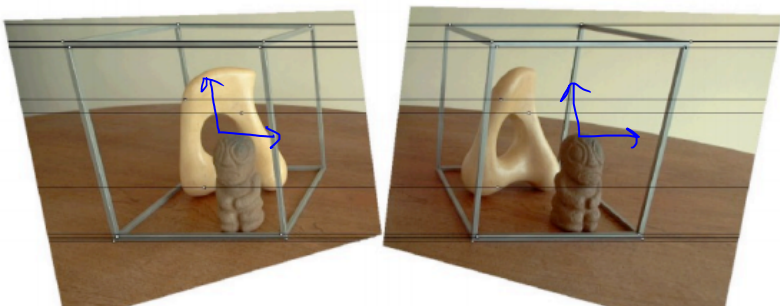
Real answer: Whenever possible, we work with purely lateral motion.

→ Rectification: Standardize our image plane w/ homographies



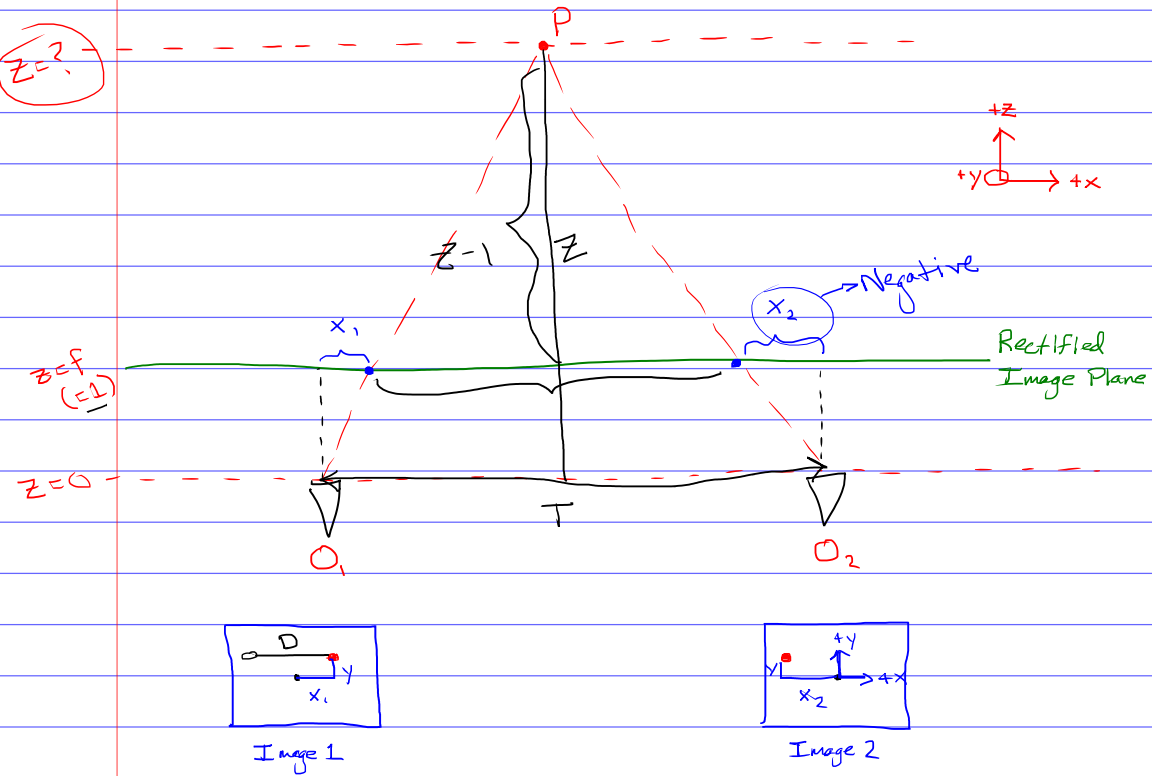
1. Find a new, common image plane
2. Apply a homography on each image to project the scene onto the new plane
3. Use our horizontal epipolar lines to learn about the scene.

Note: We do not move  $O_1$  &  $O_2$ . But we do change the direction they are facing



# Application: Depth from Disparity

Consider this scene: (viewed from above)



$$\frac{T}{Z} = \frac{T - x_1 + x_2}{Z - f} \rightarrow T(Z - f) = Z(T - x_1 + x_2)$$

$$+ fT = Z(+x_1 - x_2)$$

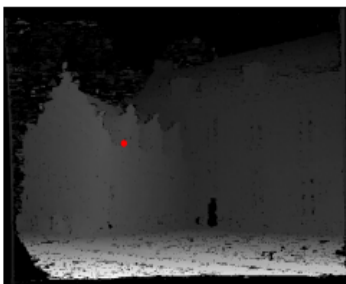
$$Z = \frac{fT}{x_1 - x_2}$$

Depth  $\leftarrow$   $\underbrace{\hspace{1cm}}$   $\rightarrow$  Disparity

Image 1

Disparity

Image 2



$\approx$  Depth