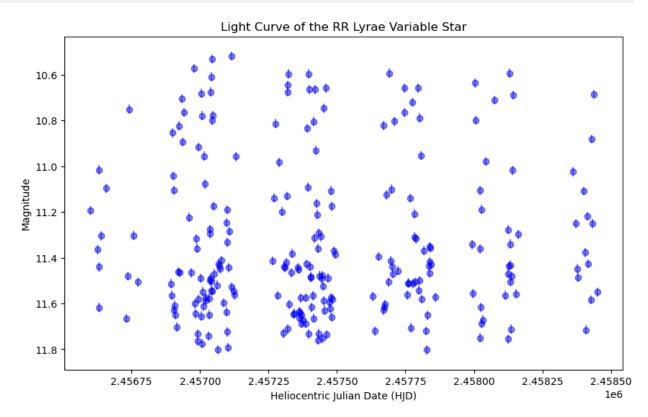
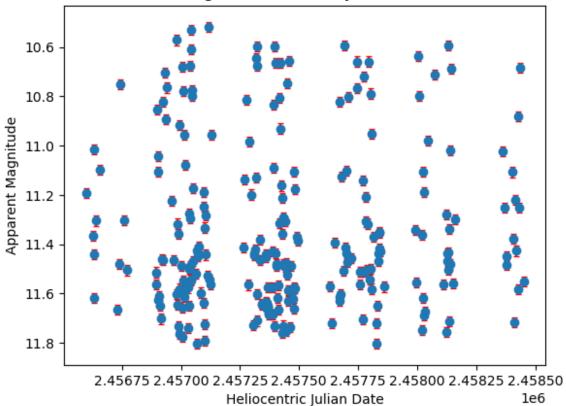
```
import pandas as pd
# Load the dataset
data = pd.read csv('RR GEM TEST DATA.csv')
data
                            mag_err
              hjd
                      mag
0
                   11.130
                               0.02
     2.457317e+06
1
                   11.411
                               0.02
     2.457077e+06
2
                   10.766
                               0.02
     2.457745e+06
3
     2.456657e+06
                   11.097
                               0.02
4
     2.457451e+06
                   10.747
                               0.02
                                . . .
243
    2.457018e+06
                   11.078
                               0.02
244
     2.456960e+06
                   11.224
                               0.02
245
    2.457446e+06
                   11.477
                               0.02
246 2.457265e+06
                   11.414
                               0.02
247 2.457789e+06
                   11.317
                               0.02
[248 rows x 3 columns]
data.tail()
              hjd
                      mag
                            mag err
243
     2.457018e+06
                   11.078
                               0.02
                               0.02
244
    2.456960e+06
                   11.224
                               0.02
245
     2.457446e+06
                   11.477
     2.457265e+06
                   11.414
                               0.02
246
247 2.457789e+06
                   11.317
                               0.02
data.duplicated().sum()
0
# Check for missing values
print(data.isnull().sum())
hjd
           0
mag
           0
mag err
           0
dtype: int64
import matplotlib.pyplot as plt
# Plot the Light Curve
plt.figure(figsize=(10, 6))
plt.errorbar(data['hjd'], data['mag'], yerr=data['mag err'], fmt='o',
color='blue', alpha=0.5)
plt.xlabel('Heliocentric Julian Date (HJD)')
plt.ylabel('Magnitude')
plt.title('Light Curve of the RR Lyrae Variable Star')
```

```
plt.gca().invert_yaxis()
plt.show()
```



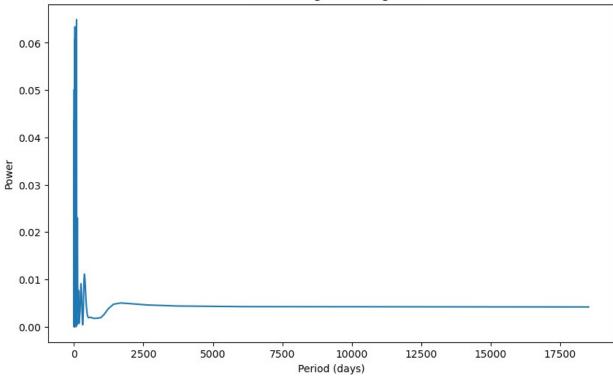
```
# Plot the light curve
plt.errorbar(data['hjd'], data['mag'], yerr=data['mag_err'], fmt='o',
ecolor='r', capsize=2)
plt.gca().invert_yaxis() # Invert y-axis to have brighter magnitudes
on top
plt.xlabel('Heliocentric Julian Date')
plt.ylabel('Apparent Magnitude')
plt.title('Light Curve of RR Lyrae Star')
plt.show()
```





```
from astropy.timeseries import LombScargle
import numpy as np
# Calculate the Lomb-Scargle periodogram
frequency, power = LombScargle(data['hjd'], data['mag']).autopower()
# Find the best frequency
best frequency = frequency[np.argmax(power)]
best period = 1 / best frequency
print(f"Best period: {best period} days")
Best period: 94.06435228428013 days
# Calculate the Lomb-Scargle periodogram
frequency, power = LombScargle(data['hjd'], data['mag']).autopower()
# Plot the periodogram
plt.figure(figsize=(10, 6))
plt.plot(1/frequency, power)
plt.xlabel('Period (days)')
plt.ylabel('Power')
plt.title('Lomb-Scargle Periodogram')
plt.show()
```

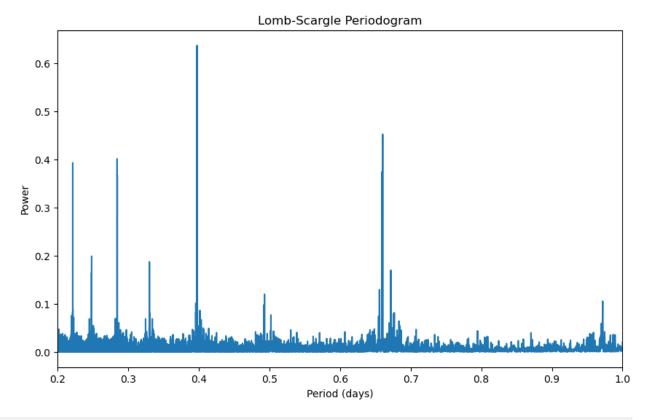




```
from astropy.timeseries import LombScargle
import numpy as np
# Define the period range for RR Lyrae stars
min period = 0.2 # days
max period = 1.0 # days
min frequency = 1 / max period
max frequency = 1 / min period
# Calculate the Lomb-Scargle periodogram within the restricted
frequency range
frequency, power = LombScargle(data['hjd'],
data['mag']).autopower(minimum frequency=min frequency,
maximum frequency=max frequency)
frequency, power
                  , 1.00010793, 1.00021586, ..., 4.99974585,
(array([1.
4.99985378,
        4.99996171]),
array([0.01971643, 0.01060333, 0.00524978, ..., 0.02298379,
0.02124143.
        0.0178524 ]))
# Plot the periodogram
plt.figure(figsize=(10, 6))
plt.plot(1/frequency, power)
```

```
plt.xlabel('Period (days)')
plt.ylabel('Power')
plt.title('Lomb-Scargle Periodogram')
plt.xlim(min_period, max_period) # Restrict x-axis to the period
range of interest
plt.show()

# Find the best frequency
best_frequency = frequency[np.argmax(power)]
best_period = 1 / best_frequency
print(f"Best period: {best_period} days")
```



```
Best period: 0.397273138821235 days

# Convert HJD to phase using the best period
data['phase'] = (data['hjd'] % best_period) / best_period

# Sort by phase for a cleaner plot
data = data.sort_values('phase')

# Plot the phased light curve
plt.figure(figsize=(10, 6))
plt.scatter(data['phase'], data['mag'], color='blue')
plt.xlabel('Phase')
```

```
plt.ylabel('Apparent Magnitude')
plt.gca().invert_yaxis() # Invert y-axis to match the astronomical
convention
plt.title('Phased Light Curve of RR Lyrae Star')
plt.show()
```

Phased Light Curve of RR Lyrae Star 10.6 10.8 Apparent Magnitude 11.0 11.2 11.4 11.6 11.8 0.0 0.2 0.4 0.6 0.8 1.0 Phase

```
# Find maxima and minima
max_magnitude = data['mag'].max()
min_magnitude = data['mag'].min()

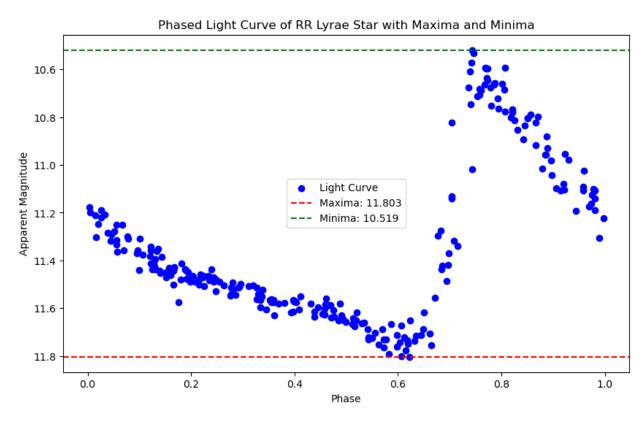
# Calculate amplitude
amplitude = max_magnitude - min_magnitude

print(f"Maxima (Magnitude): {max_magnitude}")
print(f"Minima (Magnitude): {min_magnitude}")
print(f"Amplitude: {amplitude}")

Maxima (Magnitude): 11.803
Minima (Magnitude): 10.519
Amplitude: 1.2840000000000007

# Plot the phased light curve with maxima and minima
plt.figure(figsize=(10, 6))
plt.scatter(data['phase'], data['mag'], color='blue', label='Light
Curve')
```

```
plt.axhline(y=max_magnitude, color='red', linestyle='--',
label=f'Maxima: {max_magnitude}')
plt.axhline(y=min_magnitude, color='green', linestyle='--',
label=f'Minima: {min_magnitude}')
plt.xlabel('Phase')
plt.ylabel('Apparent Magnitude')
plt.gca().invert_yaxis() # Invert y-axis to match the astronomical
convention
plt.title('Phased Light Curve of RR Lyrae Star with Maxima and
Minima')
plt.legend()
plt.show()
```



```
# Summary of Analysis for RR Lyrae Variable Star Dataset
# 1. Load and Explore the Dataset
# - Loaded the RR Lyrae Variable Star dataset
# - Checked for duplicate entries and missing values in the dataset
# 2. Plot the Light Curve
# - Plotted the light curve showing the variation in magnitude over time
# - Inverted the y-axis to display brighter magnitudes at the top
# 3. Calculate the Lomb-Scargle Periodogram
```

- # Used the Lomb-Scargle method to analyze periodicity in the star's brightness variations
- # Determined the best period of the star's brightness variations
- # 4. Plot the Lomb-Scargle Periodogram
- # Visualized the Lomb-Scargle periodogram to show the power spectrum of the star's brightness variations
- # 5. Phased Light Curve Analysis
- # Converted the Heliocentric Julian Date (HJD) to phase using the best period
- # Sorted the data by phase for a clearer plot
- # Plotted the phased light curve to show the periodic behavior of the star
- # 6. Identify Maxima, Minima, and Amplitude
- # Found the maximum and minimum magnitudes in the dataset
- # Calculated the amplitude of the star's brightness variations
- # 7. Plot Phased Light Curve with Maxima and Minima
- # Visualized the phased light curve with markers for maxima and minima
- # Highlighted the maximum and minimum magnitudes on the plot
- # This code provides a comprehensive analysis of the RR Lyrae Variable Star dataset, including period determination, phased light curve visualization, and identification of key features in the star's brightness variations.