## Lecture 16

June 4, 2016

## 1 The power-law lens

The power-law lens has a convergence profile of the kind

$$\kappa(x) = \frac{3-n}{2}x^{1-n}$$

The corresponding mass profile is

$$m(x) = x^{3-n}$$

which implies that the deflection angle is

$$\alpha(x) = x^{2-n}$$

and that the shear profile is

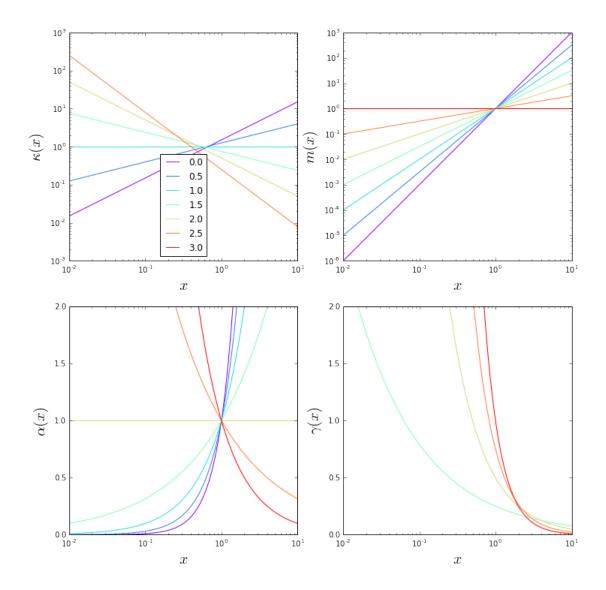
$$\gamma(x) = \frac{m(x)}{x^2} - \kappa(x) = \frac{n-1}{2}x^{1-n}$$
.

Here are some plots of some of the relevant quantities.

```
In [1]: import numpy as np
        import matplotlib.pyplot as plt
        %matplotlib inline
        from matplotlib.pyplot import cm
        fig,ax=plt.subplots(2,2,figsize=(12,12))
        x=np.logspace(-2,1.0,1000)
       n=np.linspace(0,3,7)
       def kappa(x,n):
            return((3-n)/2.0*np.abs(x)**(1-n))
        def gamma(x,n):
            return((n-1)/2.0*np.abs(x)**(1-n))
        def mass(x,n):
            return(np.abs(x)**(3-n))
        def alpha(x,n):
            return(np.sign(x)*np.abs(x)**(2-n))
        color=iter(cm.rainbow(np.linspace(0,1,n.size)))
        ax[0,0].set_xscale('log')
        ax[0,0].set_yscale('log')
        ax[0,1].set_xscale('log')
        ax[0,1].set_yscale('log')
        ax[1,0].set_xscale('log')
```

```
#ax[1,0].set_yscale('log')
ax[1,1].set_xscale('log')
#ax[1,1].set_yscale('log')
ax[1,0].set_ylim([0,2])
ax[1,1].set_ylim([0,2])
for i in range(n.size):
    c=next(color)
   ka=kappa(x,n[i])
    m=mass(x,n[i])
    a=alpha(x,n[i])
    g=gamma(x,n[i])
    ax[0,0].plot(x,ka,color=c,label=str(n[i]))
    ax[0,1].plot(x,m,color=c,label=str(n[i]))
    ax[1,0].plot(x,a,color=c,label=str(n[i]))
    ax[1,1].plot(x,g,color=c,label=str(n[i]))
ax[0,0].legend(loc='best')
ax[0,0].set_xlabel('$x$',fontsize=20)
ax[0,0].set_ylabel('$\kappa(x)$',fontsize=20)
ax[0,1].set_xlabel('$x$',fontsize=20)
ax[0,1].set_ylabel('$m(x)$',fontsize=20)
ax[1,0].set_xlabel('$x$',fontsize=20)
ax[1,0].set_ylabel(r'$\alpha(x)$',fontsize=20)
ax[1,1].set_xlabel('$x$',fontsize=20)
ax[1,1].set_ylabel(r'$\gamma(x)$',fontsize=20)
```

Out[1]: <matplotlib.text.Text at 0x10fa3f910>



## 1.1 Magnification and critical lines

The magnification is

$$\det A = (1 - \kappa - \gamma)(1 - \kappa + \gamma)$$

In particular, the inverse tangential and the radial magnifications are

$$\lambda_t = 1 - x^{1-n}$$

and

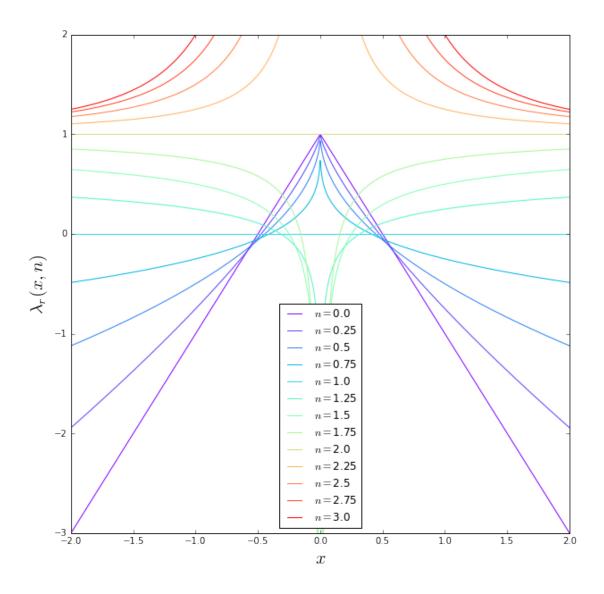
$$\lambda_r = 1 - (2 - n)x^{1 - n}$$

The critical lines are given by

$$x_t = 1$$

and

$$x_r = \left(\frac{1}{2-n}\right)^{\frac{1}{1-n}} = (2-n)^{\frac{1}{n-1}}$$



```
In [3]: n=np.linspace(0.0,3.0,13)

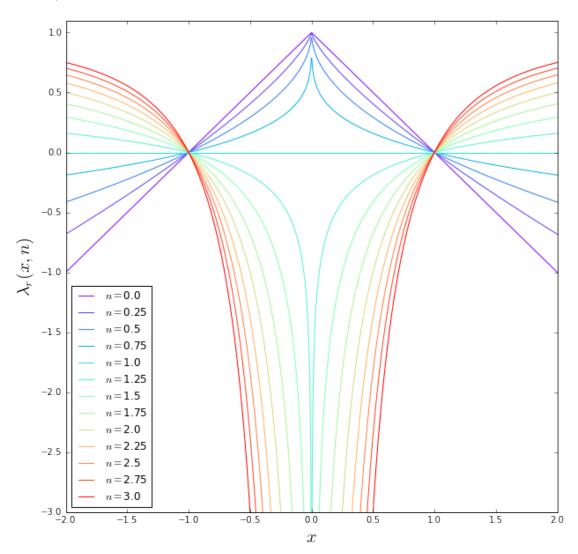
def lambdat_pow(x,n):
    return 1-np.abs(x)**(1.0-n)

fig,ax=plt.subplots(1,1,figsize=(10,10))
    color=iter(cm.rainbow(np.linspace(0,1,n.size)))

for i in range(n.size):
        c=next(color)
        ax.plot(x,lambdat_pow(x,n[i]),color=c,label=r'$n=$'+str(n[i]))

ax.set_xlabel(r'$x$',fontsize=20)
    ax.set_ylabel(r'$\lambda_r(x,n)$',fontsize=20)
    ax.legend(loc='best')
    ax.set_ylim([-3,1.1])
```

# Out[3]: (-3, 1.1)



```
In [21]: n=np.linspace(0.0000001,3,90)

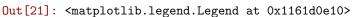
def radcl(n):
    return((2.0-n)**(1./(n-1)))

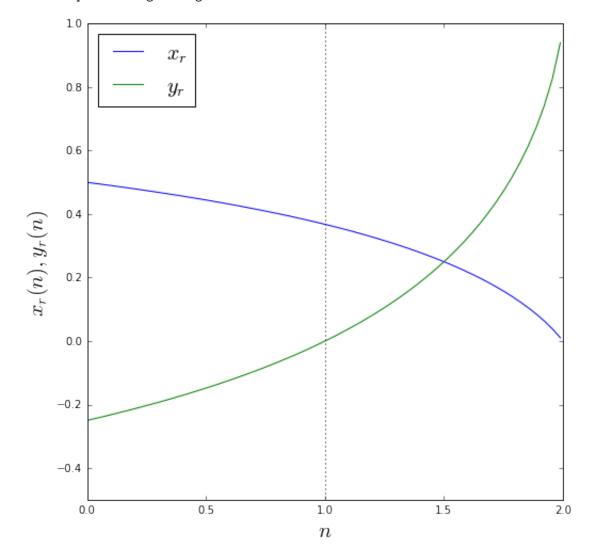
def radcau(n):
    return -(radcl(n)-alpha(radcl(n),n))

rl=radcl(n)
    rc=radcau(n)
    fig,ax=plt.subplots(1,1,figsize=(8,8))
    ax.plot(n,rl,label=r'$x_r$')
    ax.plot(n,rc,label=r'$y_r$')
```

```
ax.set_xlabel('$n$',fontsize=20)
ax.set_ylabel('$x_r(n),y_r(n)$',fontsize=20)
ax.plot([1.0,1.0],[-1,1],':',color='black')
ax.set_ylim([-0.5,1.0])
ax.legend(loc='best',fontsize=20)
```

/Users/massimo/anaconda/envs/python2/lib/python2.7/site-packages/ipykernel/\_main\_.py:4: RuntimeWarning





The radial critical line only exists if n < 2. For n > 2 there is no radial critical line.

# Multiple images

### Lenses with n < 2

Let's discuss the occurrence of multiple images. If n < 2, the deflection angle is zero at the origin. In this case, as it emerges from the image diagram, multiple images can form only if dy/dx < 0 somewhere in the lens. This translates into a condition on the derivative of the deflection angle:

$$\frac{d\alpha}{dx} > 1$$

Being  $\alpha(x) = x^{2-n}$ , we have that

$$\frac{d\alpha}{dx} = (2-n)x^{1-n} ,$$

which means that  $d\alpha/dx > 1$  for

$$x < (2-n)^{\frac{1}{n-1}} = x_r \; ,$$

in case n > 1, and

 $x > x_r$ 

otherwise.

Therefore, for n < 2, there are multiple images. How many? For  $y = y_r$ ,  $d\alpha/dx = 1$  meaning that that the line  $x - y_r$  is tangent to  $\alpha(x)$  in  $x_r$ . For n > 1,  $y < y_r$ , there are three images of the source. Otherwise, only one. In  $y = y_r$ , two images merge.

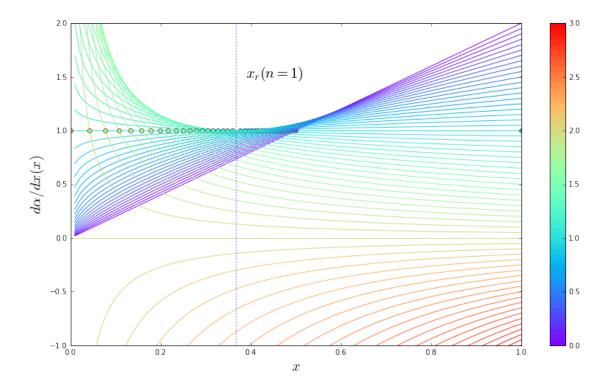
```
In [22]: def dadx(x,n):
             return((2.0-n)*x**(1.0-n))
         x=np.linspace(0.01,1,100)
         n=np.linspace(0,3.0,61)
         \#xm=xmax(n)
         color=iter(cm.rainbow(np.linspace(0,1,n.size)))
         Z = [[0,0],[0,0]]
         levels = n#range(min, max+step, step)
         CS3 = plt.contourf(Z, levels, cmap=cm.rainbow)
         plt.clf()
         fig,ax=plt.subplots(1,1,figsize=(14,8))
         ax.set_ylim([-1,2])
         \#ax.plot(n,xm,'-')
         xr=radcl(n)
         yr=np.zeros(n.size)+1.0
         for i in range(n.size):
             c=next(color)
             dd=dadx(x,n[i])
             ax.plot(x,dd,'-',color=c)
             ax.plot(xr[i],yr[i],'o',markeredgecolor='black',markersize=6,c=c)
         xr=radcl(1.001)
         xx=[xr,xr]
         yy = [-1.0, 2.0]
         #ax.plot(x,dd,'--',color='black')
         ax.plot(xx,yy,':')
         cb=plt.colorbar(CS3)
         cb.set_ticks([0.0,0.5,1.0,1.5,2.0,2.5,3.0])
```

```
ax.set_xlabel('$x$',fontsize=20)
ax.set_ylabel(r'$d\alpha/dx(x)$',fontsize=20)
ax.text(0.39,1.5,'$x_r(n=1)$',fontsize=20)
```

/Users/massimo/anaconda/envs/python2/lib/python2.7/site-packages/ipykernel/\_main\_.py:4: RuntimeWarning /Users/massimo/anaconda/envs/python2/lib/python2.7/site-packages/ipykernel/\_main\_.py:4: RuntimeWarning

Out[22]: <matplotlib.text.Text at 0x118b134d0>

<matplotlib.figure.Figure at 0x114dc85d0>



The following are some image diagrams for lenses with n=0.1,1.0,1.5

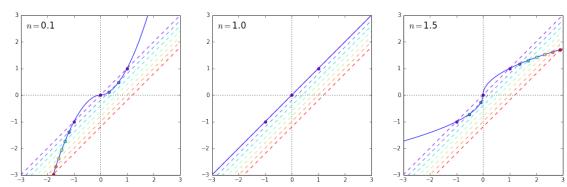
```
for i in range(len(n)):
    a=alpha(x,n[i])
    ax[i].plot(x,a,'-')
    color=iter(cm.rainbow(np.linspace(0,1,len(ys))))
    for j in range(len(ys)):
        c=next(color)
        fu=f(x,ys[j])
        ax[i].plot(x,fu,'--',c=c)
        def func(xx):
            return f(xx,ys[j])-alpha(xx,n[i])
        x0 = fsolve(func, 1.0)
        x1 = fsolve(func, -1.0)
        x3 = fsolve(func, 0.0)
        if (np.abs(func(x0))<1e-8):
            ax[i].plot(x0,alpha(x0,n[i]),'o',markersize=5,c=c)
        if (np.abs(func(x1))<1e-8):
            ax[i].plot(x1,alpha(x1,n[i]),'o',markersize=5,c=c)
        if (np.abs(func(x3))<1e-8):
            ax[i].plot(x3,alpha(x3,n[i]),'o',markersize=5,c=c)
    ax[i].set_xlim([-3,3])
    ax[i].set_ylim([-3,3])
    ax[i].text(-2.8,2.5, '$n=$'+str(n[i]), fontsize=15)
    xa=[0.0,0.0]
    ya=[-10,10]
    ax[i].plot(xa,ya,':',color='black')
    xa=[-10.0,10.0]
    ya=[0,0]
    ax[i].plot(xa,ya,':',color='black')
```

/Users/massimo/anaconda/envs/python2/lib/python2.7/site-packages/scipy/optimize/minpack.py:236: Runtime improvement from the last ten iterations.

warnings.warn(msg, RuntimeWarning)

/Users/massimo/anaconda/envs/python2/lib/python2.7/site-packages/scipy/optimize/minpack.py:236: Runtime improvement from the last five Jacobian evaluations.

warnings.warn(msg, RuntimeWarning)



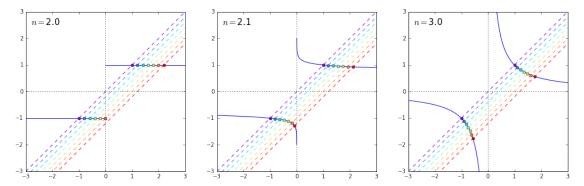
Some noticeable properties: \* these lenses produce three or one image \* three images are formed if the source lays inside the radial critical line, i.e. for  $y < y_r$ , where  $y_r$  corresponds to  $x_r : dy/dx = 0$ 

#### 3 Lenses with n > 2

These lenses always produce multiple images because of the singuarity of their lensing potential. The following are some examples of image diagrams for n = 2.0, 2.1, 3.0.

```
In [7]: n=[2.0,2.1,3.0]
        fig,ax=plt.subplots(1,3,figsize=(17,5))
        x0=np.linspace(-3,-1.e-3,500)
        x1=np.linspace(1.e-3,3,500)
        ys=np.linspace(0,1.2,7)
        for i in range(len(n)):
            a0=alpha(x0,n[i])
            a1=alpha(x1,n[i])
            ax[i].plot(x0,a0,'-',color='blue')
            ax[i].plot(x1,a1,'-',color='blue')
            color=iter(cm.rainbow(np.linspace(0,1,len(ys))))
            for j in range(len(ys)):
                c=next(color)
                fu=f(x,ys[j])
                ax[i].plot(x,fu,'--',c=c)
                def func(xx):
                    return f(xx,ys[j])-alpha(xx,n[i])
                x0_{-} = fsolve(func, 1.0)
                x1_= fsolve(func, -1.0)
                x3_= fsolve(func, 0.0)
                if (np.abs(func(x0_))<1e-8):
                    ax[i].plot(x0_,alpha(x0_,n[i]),'o',markersize=5,c=c)
                if (np.abs(func(x1_))<1e-8):
                    ax[i].plot(x1_,alpha(x1_,n[i]),'o',markersize=5,c=c)
                if (np.abs(func(x3_))<1e-8 and j>0):
                    ax[i].plot(x3_,alpha(x3_,n[i]),'o',markersize=5,c=c)
            ax[i].set_xlim([-3,3])
            ax[i].set_ylim([-3,3])
            ax[i].text(-2.8,2.5,'$n=$'+str(n[i]),fontsize=15)
            xa=[0.0,0.0]
            ya = [-10, 10]
            ax[i].plot(xa,ya,':',color='black')
            xa=[-10.0,10.0]
            ya=[0,0]
            ax[i].plot(xa,ya,':',color='black')
```

/Users/massimo/anaconda/envs/python2/lib/python2.7/site-packages/ipykernel/\_main\_.py:20: RuntimeWarning



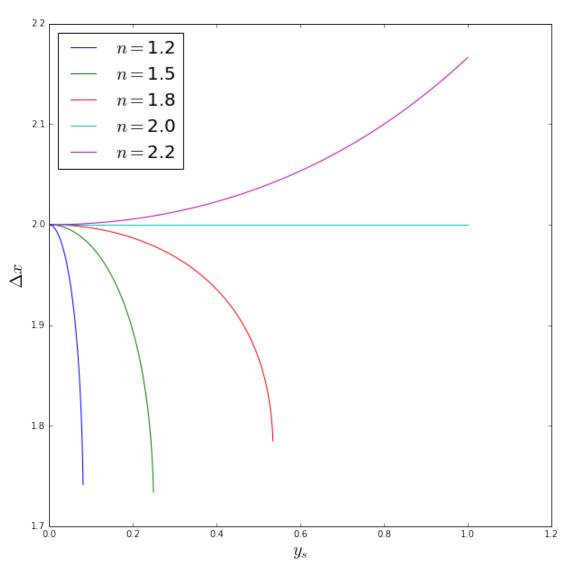
Notice that \* none of these lenses ever has a central image \* while for a lens with n=2 produces multiple images only for  $|y| < y_{cut} = 1$ , all lenses with n>2 always have 2 images \* the  $x_-$  image always moves towards the center of the lens as  $y \to \infty$ 

## 4 Image separation vs n

We explore now the sensitivity of image separation on the power-law index n.

```
In [33]: n=[1.2,1.5,1.8,2.0,2.2]
         x0=np.linspace(-3,-1.e-3,500)
         x1=np.linspace(1.e-3,3,500)
         ys=np.linspace(0,1,1000)
         fig,ax=plt.subplots(1,1,figsize=(10,10))
         for i in range(len(n)):
             a0=alpha(x0,n[i])
             a1=alpha(x1,n[i])
             s0=np.zeros(ys.size)
             s1=np.zeros(ys.size)
             s2=np.zeros(ys.size)
             dtheta=[]
             ytheta=[]
             te=[]
             for j in range(len(ys)):
                 def func(xx):
                     return f(xx,ys[j])-alpha(xx,n[i])
                 x0_{-} = fsolve(func, 1.0)
                 x1_= fsolve(func, -1.0)
                 x2_= fsolve(func, 0.0)
                 if (np.abs(func(x0_))<1e-8):
                     s0[j]=x0_[0]
                 if (np.abs(func(x1_))<1e-8 and x1_<0):
                     s1[j]=x1_[0]
                     dtheta.append(np.abs(s1[j]-s0[j]))
```

/Users/massimo/anaconda/envs/python2/lib/python2.7/site-packages/ipykernel/\_main\_.py:10: RuntimeWarning/Users/massimo/anaconda/envs/python2/lib/python2.7/site-packages/ipykernel/\_main\_.py:10: RuntimeWarning



## 5 Time delays

The formula for the light travel time is

$$t(x) = \frac{(1+z_L)}{c} \frac{D_L D_S}{D_{LS}} \frac{\xi_0^2}{D_L^2} \left[ \frac{1}{2} (x-y)^2 - \psi(x) \right] = \frac{(1+z_L)}{c} \frac{D_L D_S}{D_{LS}} \tau(x)$$

For a power-law lens, the lensing potential is

$$\psi(x) = \frac{1}{3-n}x^{3-n}$$

so that

$$\tau(x) = \frac{\xi_0^2}{D_L^2} \left[ \frac{1}{2} (x - y)^2 - \frac{1}{3 - n} x^{3 - n} \right]$$

At the image positions

$$x - y = \alpha(x) = x^{2-n}$$

thus

$$\tau(x_i) = \frac{\xi_0^2}{D_L^2} \left[ \frac{1}{2} x_i^{2(2-n)} - \frac{1}{3-n} x_i^{3-n} \right]$$

We can then compute the time delay between the images:

 $ax.set_ylim([-0.05, 1.5])$ 

$$\Delta t_{ij} \propto \Delta \tau_{ij} = \frac{\xi_0^2}{D_L^2} \left[ \frac{1}{2} \left( x_j^{2(2-n)} - x_i^{2(2-n)} \right) - \frac{1}{3-n} \left( x_j^{3-n} - x_i^{3-n} \right) \right]$$

For n=2, this formula gives:

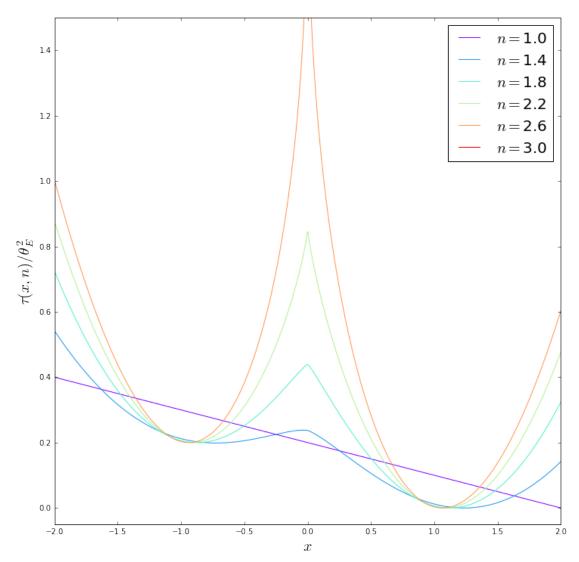
$$\Delta \tau_{ij} = \frac{\xi_0^2}{D_L^2} (x_i - x_j) = \theta_E^2 \left( \frac{\theta_i}{\theta_E} - \frac{\theta_j}{\theta_E} \right) = \frac{1}{2} \left( \theta_i^2 - \theta_j^2 \right) = \Delta \tau_{SIS}$$

```
In [46]: def pot_power(x,n):
             return 1.0/(3.0-n)*np.abs(x)**(3.0-n)
         def travel_time_power(y,x,n):
             return (0.5*(x-y)**2-pot_power(x,n))
         x=np.linspace(-2.0,2.0)
         n=np.linspace(1.0,3.0,6)
         color=iter(cm.rainbow(np.linspace(0,1,n.size)))
         x=np.linspace(-2.0,2.0,1000)
         fig,ax=plt.subplots(1,1,figsize=(12,12))
         for i in range(n.size):
             c=next(color)
             ax.plot(x,travel_time_power(y,x,n[i])-np.amin(travel_time_power(y,x,n[i])),color=c,label=r
             \#ax.plot(x,pot_power(x,n[i]),color=c)
         ax.set_xlabel('$x$',fontsize=20)
         ax.set_ylabel(r'\$\tau(x,n)/\theta_E^2\$',fontsize=20)
         ax.legend(loc='best',fontsize=20)
```

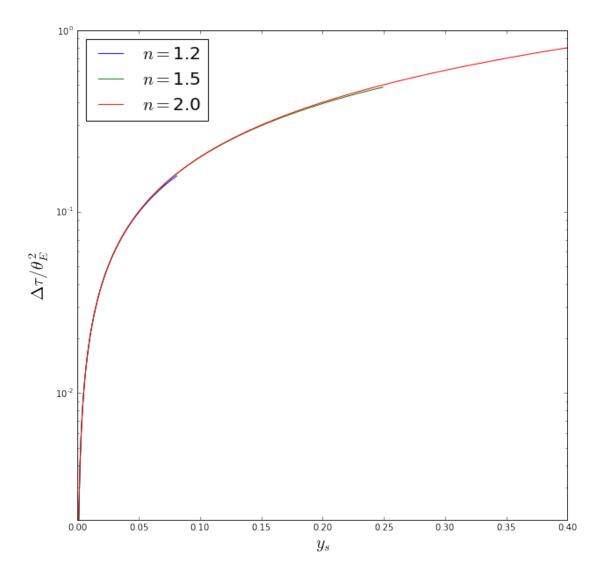
/Users/massimo/anaconda/envs/python2/lib/python2.7/site-packages/ipykernel/\_\_main\_..py:2: RuntimeWarning from ipykernel import kernelapp as app

 $/Users/massimo/anaconda/envs/python2/lib/python2.7/site-packages/ipykernel/\_main\_.py:19: RuntimeWarnings/massimo/anaconda/envs/python2/lib/python2.7/site-packages/ipykernel/\_main\_.py:19: RuntimeWarnings/massimo/anaconda/envs/python2/lib/python2.7/site-packages/ipykernel/\_main\_.py:19: RuntimeWarnings/massimo/anaconda/envs/python2/lib/python2.7/site-packages/ipykernel/\_main\_.py:19: RuntimeWarnings/massimo/anaconda/envs/python2/lib/python2.7/site-packages/ipykernel/\_main\_.py:19: RuntimeWarnings/massimo/anaconda/envs/python2/lib/python2.7/site-packages/ipykernel/\_main\_.py:19: RuntimeWarnings/massimo/anaconda/envs/python2/lib/python2.7/site-packages/ipykernel/\_main\_.py:19: RuntimeWarnings/massimo/anaconda/envs/python2/lib/python2.7/site-packages/ipykernel/\_main\_.py:19: RuntimeWarnings/massimo/anaconda/envs/python2/lib/python2$ 

Out[46]: (-0.05, 1.5)



```
for i in range(len(n)):
   a0=alpha(x0,n[i])
    a1=alpha(x1,n[i])
    s0=np.zeros(ys.size)
    s1=np.zeros(ys.size)
    ytheta=[]
    td=[]
    dtheta=[]
    for j in range(len(ys)):
        def func(xx):
            return f(xx,ys[j])-alpha(xx,n[i])
        x0_{-} = fsolve(func, 1.0)
        x1_= fsolve(func, -1.0)
        if (np.abs(func(x0_))<1e-8):
            s0[j]=x0_[0]
        if (np.abs(func(x1_))<1e-8 and x1_<0):
            s1[j]=x1_[0]
            dtheta.append(s1[j]+s0[j])
            ytheta.append(ys[j])
            t=time_delay(s0[j],-s1[j],n[i])
            td.append(t)
    ax.set_xlim([0,0.4])
    ax.set_ylim([0,1.0])
    ax.set_yscale('log')
    ax.plot(ytheta,td,'-',label='$n=$'+str(n[i]))
    ax.legend(loc='best',fontsize=20)
    ax.set_xlabel('$y_s$',fontsize=20)
    ax.set_ylabel(r'$\Delta \tau/\theta_E^2$',fontsize=20)
```



In the case of small image separations with respect to the size of the Einstein radius (i.e.  $y \ll 1$ ), Kochanek (2002) derived an approximated formula stating that

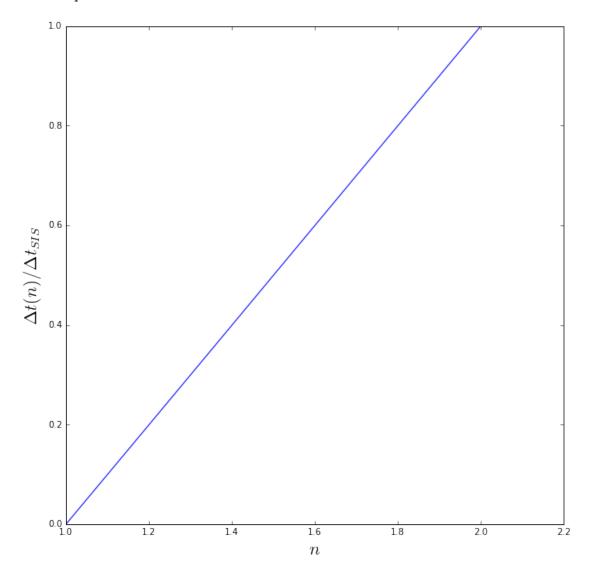
$$\Delta t(n) = (n-1)\Delta t_{SIS} \left[ 1 - \frac{(2-n)^2}{12} \left( \frac{\delta \theta}{\langle \theta \rangle} \right)^2 + \dots \right]$$

The formula however works also in the case of  $\delta\theta/\langle\theta\rangle$  of order unity.

Let us assume to observe an image separation of  $\delta\theta=1$ " and  $\langle\theta\rangle=5$ ". In this case  $\Delta t(n)/\Delta t_{SIS}$  looks like this:

```
ax.plot(n,dt,'-')
ax.set_xlabel('$n$',fontsize=20)
ax.set_ylabel(r'$\Delta t(n)/\Delta t_{SIS}$',fontsize=20)
```

Out[10]: <matplotlib.text.Text at 0x11276f450>



This shows that more centrally concentrated lenses (larger n) produce larger time delays for a fixed image separation.