

# Angular dependence of the atmospheric neutrino flux with IceCube data

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The IceCube Collaboration

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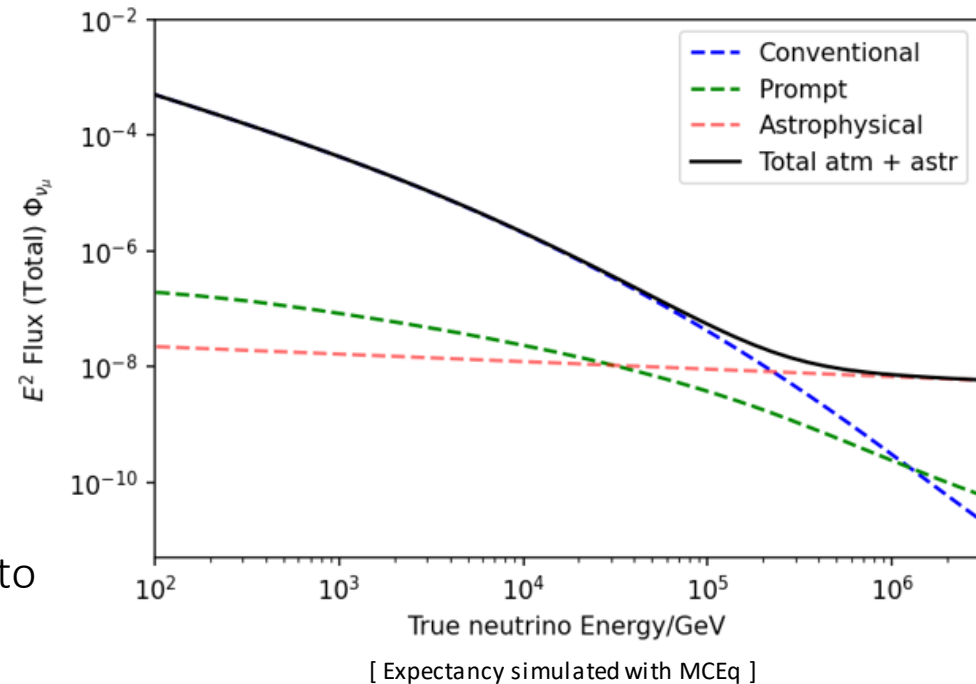
**Karlsruhe24**

# Neutrino flux and its components

The total neutrino flux arriving to Earth is comprised from several sources separated depending on their creation process. We differentiate three components;

- **Conventional:** decay of long-lived particles, majorly  $\pi$  and K
- **Prompt:** decay of short-lived particles usually charmed
- **Astrophysical:** created at astrophysical sources in different mechanisms

Components jointly comprise the total muon neutrino flux aimed to be determined in a model-independent approach in this analysis.



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electron  
neutrino

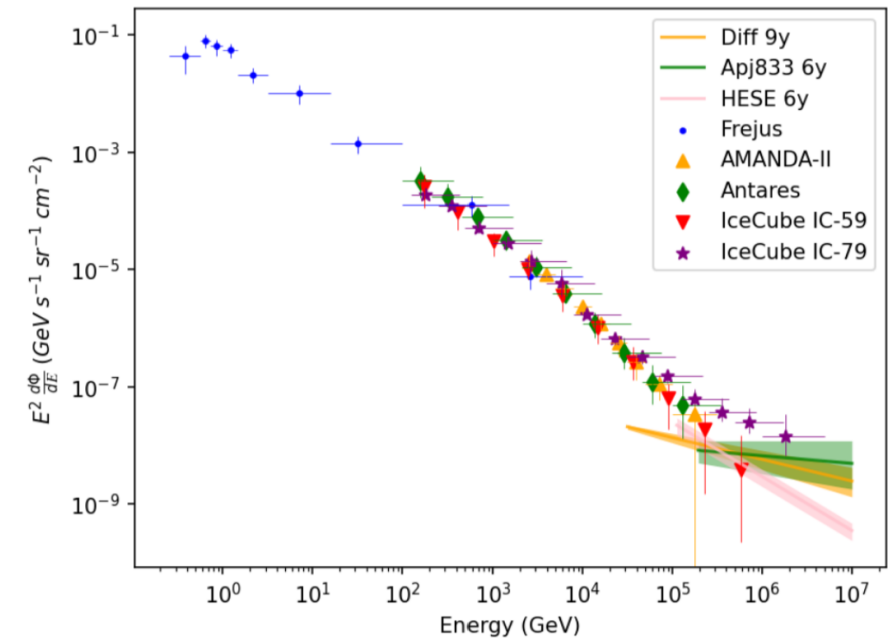


muon  
neutrino



tau  
neutrino

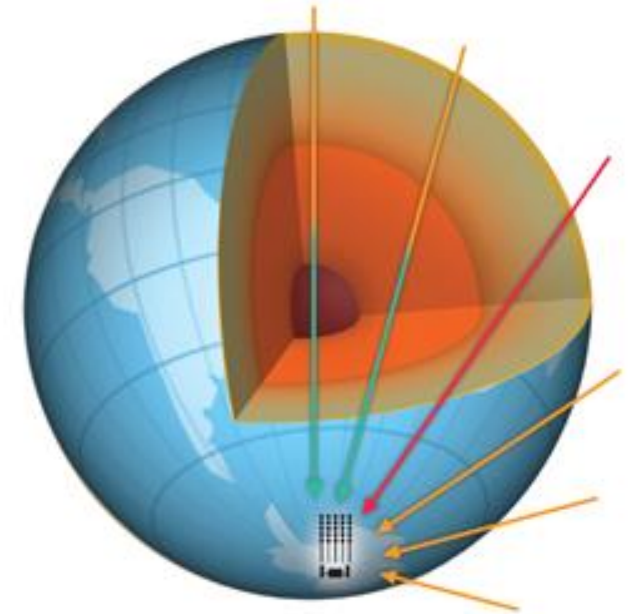
Experimental results of the muon  
neutrino flux up to date.



# Angular dependence

Differences in primary particles of the three components inevitably lead to a zenith dependence arising in Earth's atmosphere.

- **Conventional:** longer mean free path of primaries
- **Prompt:** short mean free path, atmospheric effects irrelevant
- **Astrophysical:** primaries at source, atmospheric effects irrelevant



Credit: IceCube Collaboration

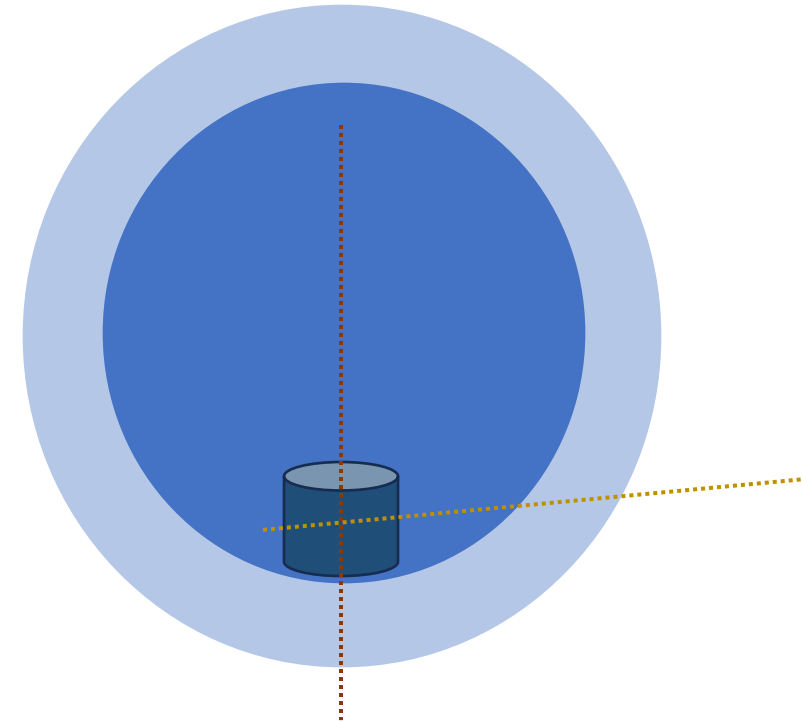
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--> Only conventional flux is anisotropic!

Longer path in low air density --> fewer secondary interactions --  
> lower energy loss --> higher energy neutrinos



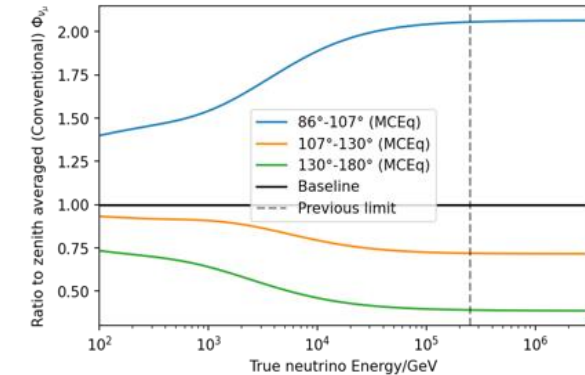
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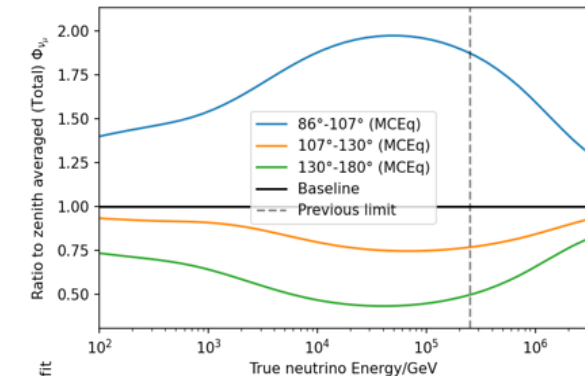
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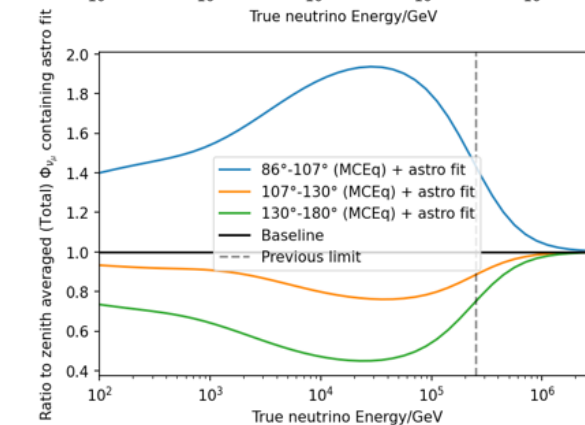
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Conventional



Conventional  
+  
prompt



Conventional  
+  
Prompt  
+  
Astrophysical

# Unfolding

Model-independent approach to reconstructing convolved variables of interest.

Detector response Neutrino energy spectrum  

$$\text{Observables} \leftarrow g(y) = \int A(x, y) f(x) dx + b(y) \rightarrow \text{Background}$$

Discretization

$$\vec{g} = \mathbf{A} \vec{f}.$$

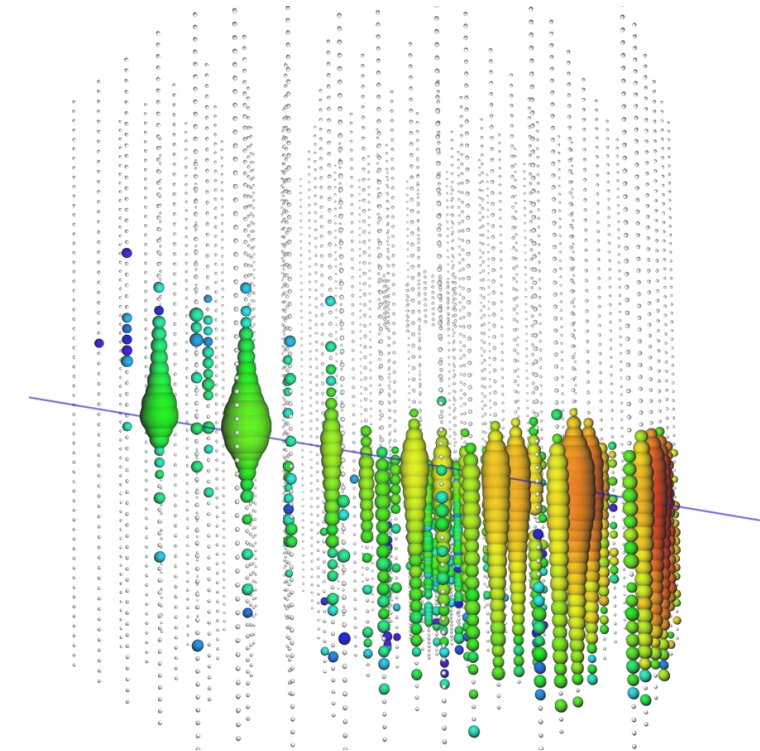
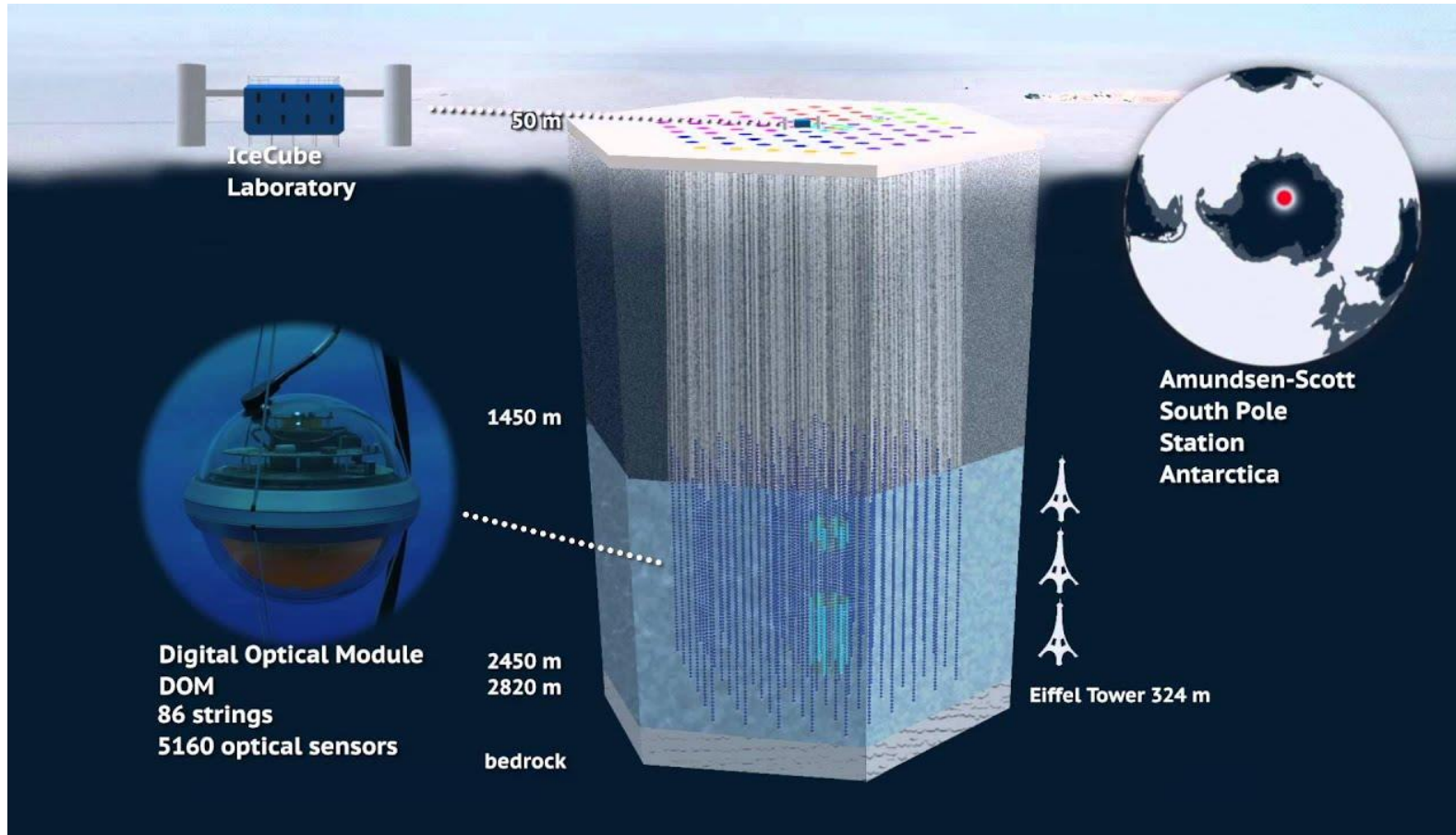
Poisson likelihood  
definition, simplifying

$$l(\vec{g} | \vec{f}) = \sum_{u=1}^m \left( g_u \ln \left( (\mathbf{A} \vec{f})_u \right) - (\mathbf{A} \vec{f})_u \right) + R(\vec{f})$$

Tikhonov regularization:  
solution assumed to be  
smooth



# The IceCube Neutrino Observatory

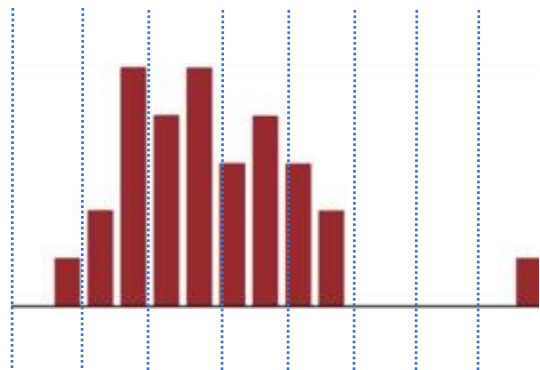


Credit: IceCube Collaboration

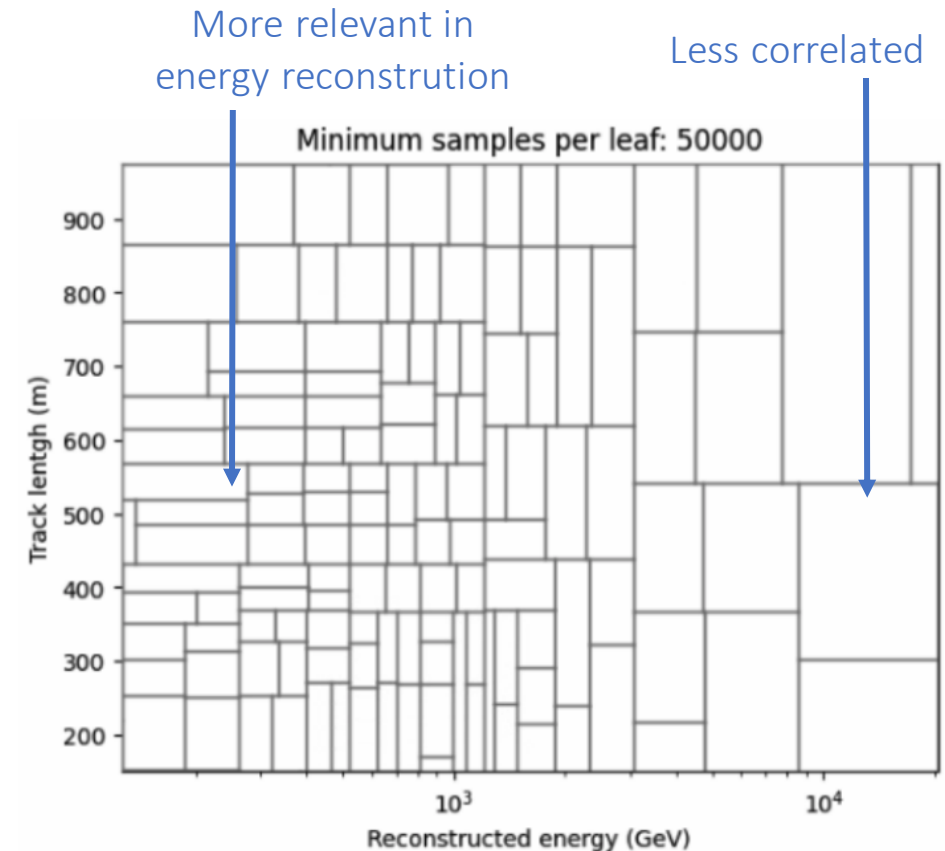
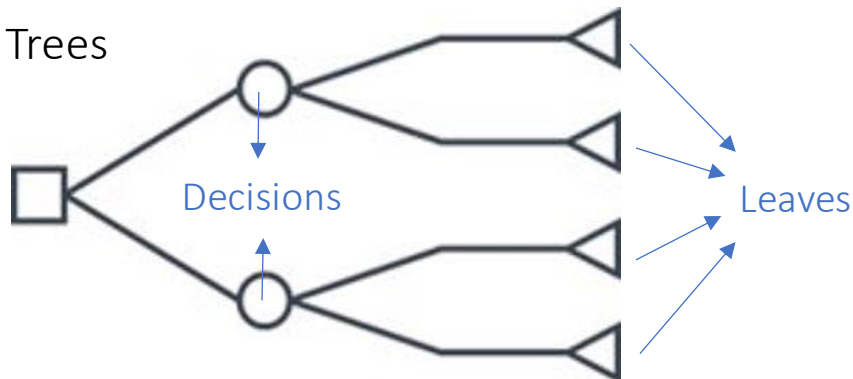


# Rebinning the observable space

Finding optimal binning scheme based on the distribution of observables used in unfolding with the goal of ensuring enough statistics and proper information gain.



- Utilizing Decision Trees



*Example rebinning scheme. The Trees create denser cuts in regions with higher information gain.*

# Application of unfolding

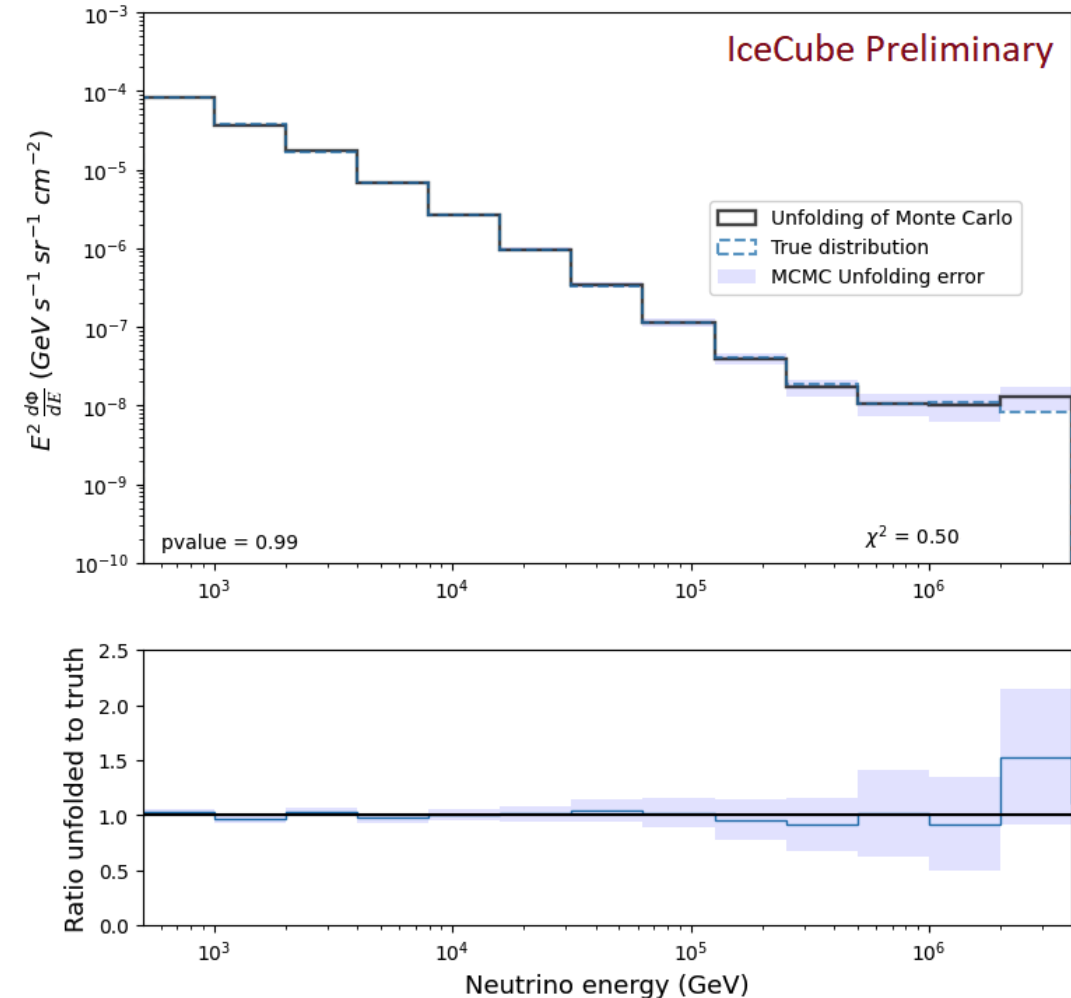
Proof of concept: a pseudosample imitating the expectation for the 11 year dataset is unfolded with Markov Chain Monte Carlo walkers.

- Observables chosen in a three-fold feature selection process and rebinned with the aforementioned process.

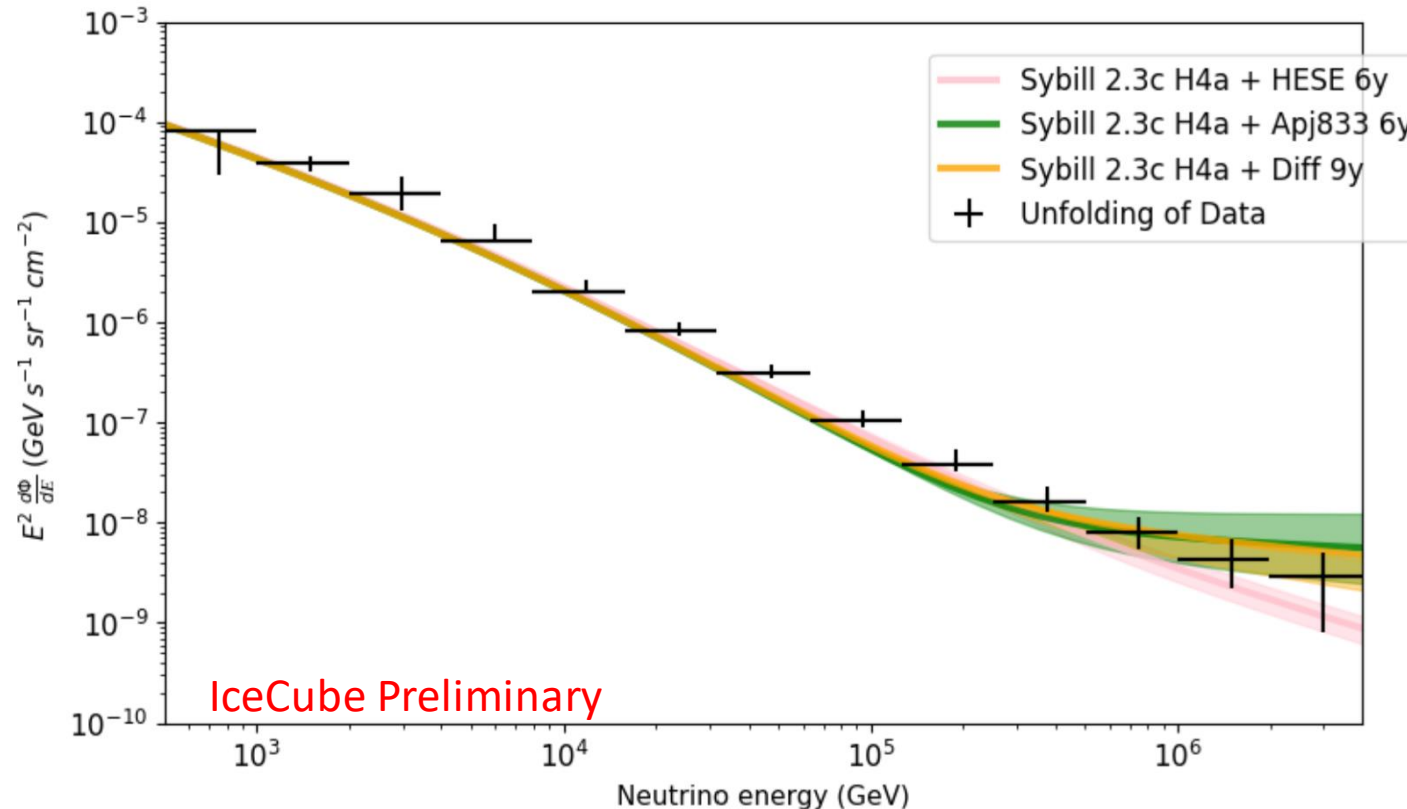
# Application of unfolding

Proof of concept: a pseudosample imitating the expectation for the 11 year dataset is unfolded with Markov Chain Monte Carlo walkers.

- Observables chosen in a three-fold feature selection process and rebinned with the aforementioned process.
- Unfolding repeated for randomized trials, all reconstructions in good agreement with the corresponding true values and with low statistical uncertainty.



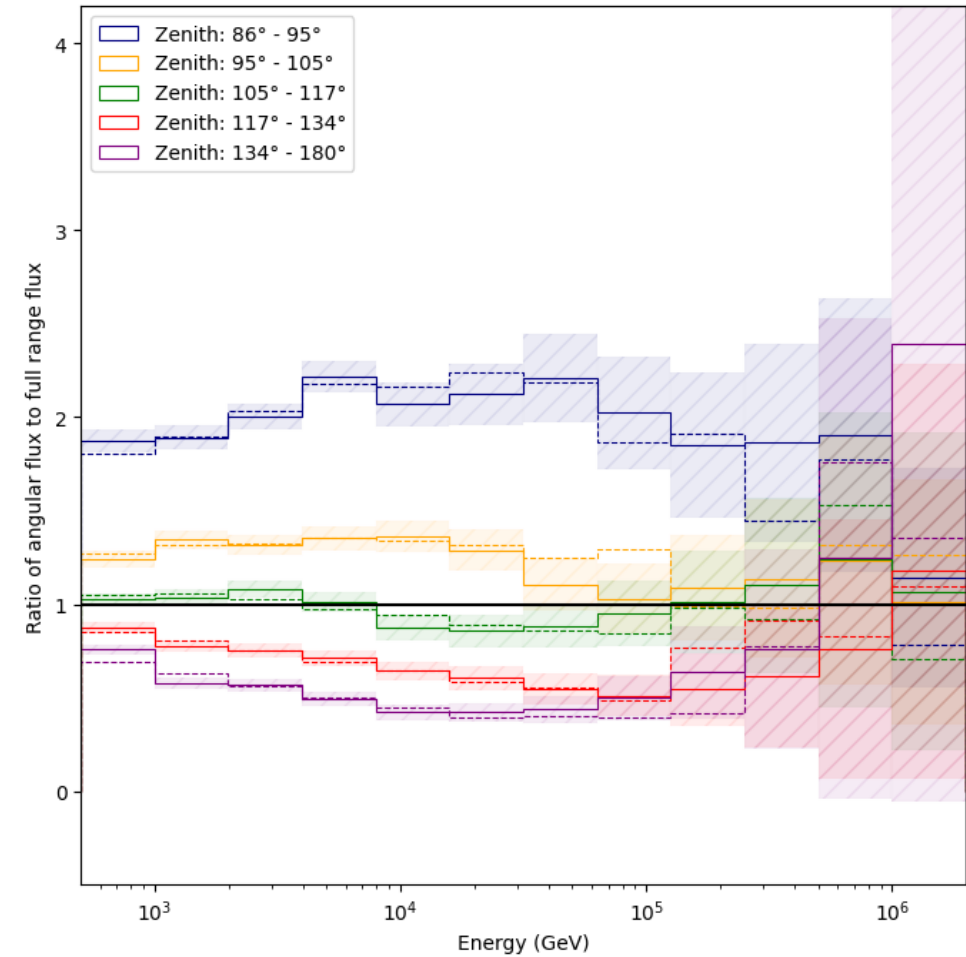
# Application of unfolding



- Results of unfolding 11 years of data
- Atmospheric dominated energy region in good agreement with models.
- Astrophysical dominated region in good agreement with astrophysical fits from other IceCube analyses.
- Crossover region is overestimated in comparison with atmospheric model and astrophysical fit combined prediction.

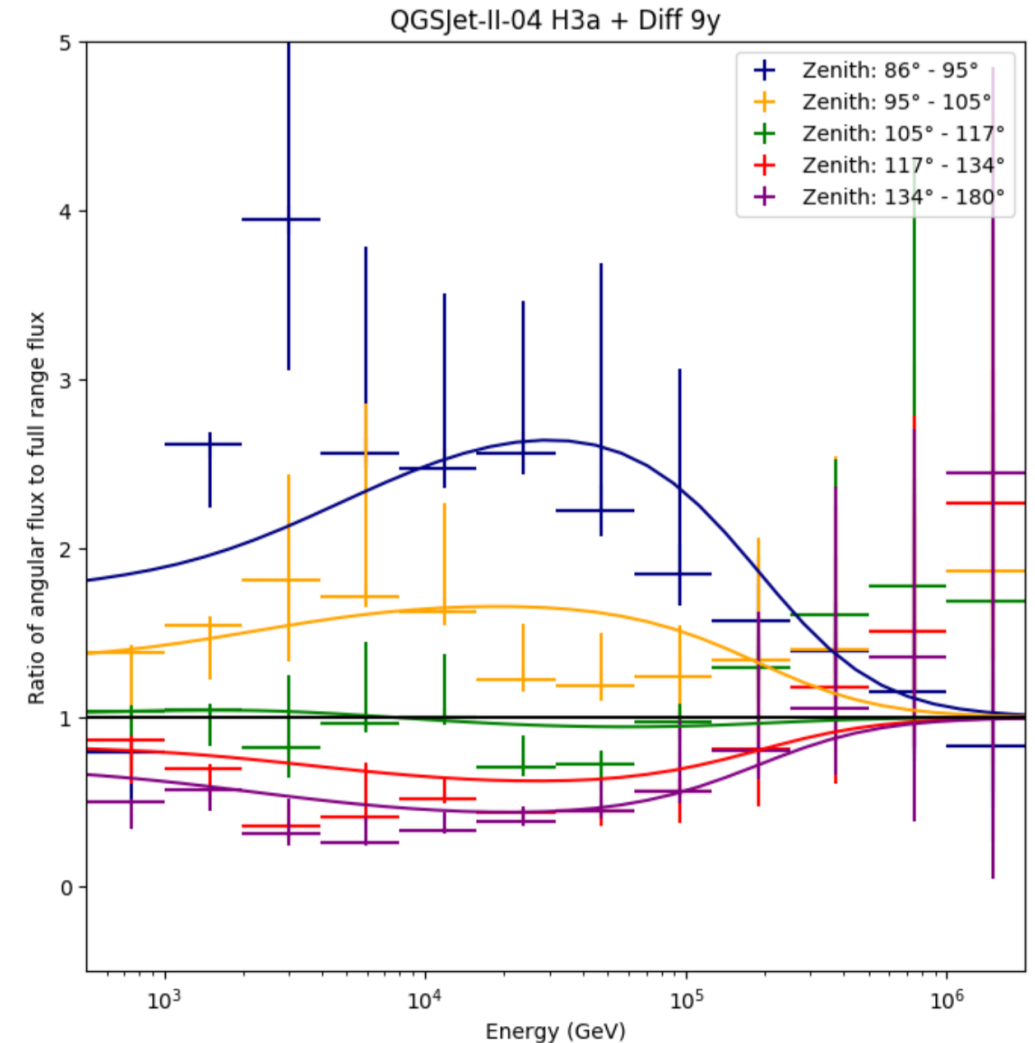
# Angular unfolding

- Five angular bins determined to approximately contain the same number of events, compared to the *baseline* spectrum that covers the full angular range.
- Slightly higher statistical uncertainty due to lower number of events in zenith bins In comparison to expected uncertainty of the full dataset.
- Discrimination of zenith fluxes possible up to 1 PeV with the assumed model and lifetime.



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- Discrimination of zenith fluxes possible up to 1 PeV with the assumed model and lifetime.
- Results show support for the assumed anisotropic behaviour coming from the conventional contribution.





# Conclusion and Outlook

- Presented a rebinning approach applicable to both unfolding and any other areas with the goal of equalizing statistical errors across the entire sample.
- Presented a bin-wise model-independent unfolding algorithm with promising results in flux measurement.
- Flux unfolding in good agreement with predictions, lowest statistical error in measurements up to date.
- Angular unfolding proven as concept, the expected anisotropic behaviour confirmed on the case of five zenith band unfolding in this work. Unfolding possible up to PeV energies.



*Thank you for  
your attention!*