

Tremolo-X Manual¹ (Version 1.6)

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Chapter 1

Tutorial

This tutorial will introduce you to the most common features of Tremolo-X by guiding you through a variety of example simulations. It is assumed that you read the chapter "First Steps" prior to starting with this tutorial.

All files mentioned here already exit in your Tremolo-X home directory in the folder tutorial. The leading number in the folder name corresponds to the section number in this chapter.

1.1 Optimizing an initial particle setup

Very often the setup of a simulation requires the assembly of a data file (with the particle positions) by hand or script. As a result the particle distribution is almost always non-optimal in the sense that the relative positions create local energy spikes which adversely affect the stability and equilibrium property of the simulation.

To counter these effects a simulation is usually preceded by an optimization phase, in which particle positions are slightly modified towards a (static) energy minimum.

We demonstrate the optimization procedure for an argon gas with reduced units.

We start by writing up the file argon.tremolo

```
global: defaultpath = ".";
global: projectname="argon";
global: comment="Reduced argon example";
global: systemofunits=custom;
```

First we set the defaultpath. Even though we do not intend to use an external path to files in this example, this must always be set. In the second line we provide the projectname, all files for this simulation must carry this name (input) or respectively will carry this name (output). The comment can be freely used to add information about the simulation. In the last line we specify that we will use a custom system of units. This means that we

have to specify a few base magnitudes, from which any others are derived. Those magnitudes are specified by line pairs, the first fixing the physical unit to be used, whereas the second specifies the amount:

In the above example, a length unit corresponds to 3.4Å, whereas mass is measured in 39.948 atomic mass units. (Thus a length of 2 in Tremolo-X units is 6.8Å, whereas a Tremolo-X mass of 0.5 corresponds to 19.974u.

Now we turn to the particles and their potentials, those are specified in the argon.potentials file. First we specify the particle types present in the simulation. In this case we use only one particle, which we name Argon, with chemical symbol Ar. The element name appears in some of the output files. Then we fix the sigma, epsilon, sigma14 and epsilon14 values [2]. (Generally you should keep to the convention sigma = sigma14 and epsilon = epsilon14, unless you have parameter sets for complex molecules which tell you otherwise.) Those are used in the computation of various potentials, namely the Lennard Jones potential which we will use here.

Due to our custom choice of units, all those values are set to one. The same holds true for the particle mass. Degrees of freedom are three (this is the norm), the particles hold no charge. In the same file we enter the actual potentials to be used in between particles. Since only one type of particle is present there are not many mixtures to keep track of.

We denote the pair of particle types which are affected by the potential. For the first Lennard Jones potential we specify a cut of radius r_{cut} of 12.0 TODO: Why? (custom units), outside of which the interaction between two particles is set to 0. Since that introduces a discontinuity, we also second potential, which is a splined Lennard Jones Potential. It gets an additional parameter r_l , after which a spline is used to interpolate from the current value to 0 a r_{cut} . Note that both types of potentials are full LJ potentials in their own right. Using both would cause approximately twice the normal potential (and twice the normal forces) to be calculated between particles. Nevertheless, it does not hurt to supply any applicable potentials in this file, as they can be switched on and off at a different location, determining those which contribute to the particle interaction.

This location is the argon.validates file.

```
validate: particle_type=Argon, state=on;
validate: force_type=lennardjones, state=off;
validate: force_type=ljspline, state=on;
```

Here we validate the use of the particle type "Argon" and decide which of the two potentials shall remain in effect, while making sure that the other one is switched off.

The majority of parameter choices are made in the argon.parameters file. Here we will determine the domain, the number of optimization steps and make choices about the output created by the program. First we state that this simulation is to be an optimization, not a dynamic simulation.:

```
integration: type=optimization;
```

Next we determine the domain as cubic shaped with a side length of 81.05. The borders are to be periodic, so a particle leaving the cube over one side will reenter it via the other.

The choice of the cell size (cellrcut) is more algorithmic that geometric, but we have to ensure that cellrcut < size.

Remark: In the parallel case this condition changes to

```
{\tt cellrcut} < \frac{{\tt size}}{\# \ {\tt of \ processors \ per \ dimension}}.
```

In the next block we set the options for parameterization. In this example we want to optimize particle positions only. Optimization is to be done by conjugate gradient (cg) method with 2001 steps (for algorithmic reasons the step number must be divisible by 3, otherwise it would be adjusted upwards).

The CG method is reset periodically to speed up convergence and furthermore we use strong Wolfe conditions on the cg line search. mean_force_eps and mean_force_eps_rel are the cutoff values for the absolute and relative mean force values respectively. The last parameter in line prefactor, is possibly the most important, as it directly controls the size of changes. Depending on the energy surface of the sample, a small value might be imperative to prevent the optimization from failing. On the other hand a small value prevents significant changes and slows the method down.

What is meant by visuals?

In the last block of this file we specify the type and intervals of output created. Visuals shall be created every 5 time units or 10 iteration steps, whereas the particle data shall be written every 500 time units or 10 iteration steps. Any other measured quantities are written every 0.5 time units or after 1 iteration step. In this example we measure only energy.

```
output {
    Outvis: T_Start=0, T_Delta=5.0, Step_Delta=10;
    Outdata: T_Start=0, T_Delta=500, Step_Delta=10;

Outm: T_Start=0, T_Delta=0.5, Step_Delta=1;
    energy: measure=on;
};
```

The last item missing before we can start the simulation is the argon.data file supplying the initial particle positions. This file may not contain any free comments, all lines must match a particular format. Here we display only the first few lines of the file:

```
\# ATOMDATA Id x=3 u=3 type
# INPUTCONV temp 2.7
                           14.25174
                                             47.74411
1
         80.10823
     0.0
              0.0
                       0.0
                                Argon
         70.29545
                           7.599451
                                             58.29292
     0.0
              0.0
                       0.0
                                Argon
```

3	66.10589	2.939586	69.22116	\
\hookrightarrow	0.0 0.0	0.0 Argon		
4	77.18604	63.33052	73.03413	\
\hookrightarrow	0.0 0.0	0.0 Argon		
5	76.02857	54.41654	24.62265	\
\hookrightarrow	0.0 0.0	0.0 Argon		
6	16.96877	11.43903	21.40026	\
\hookrightarrow	0.0 0.0	0.0 Argon		
7	53.09684	7.723396	47.48021	\
\hookrightarrow	0.0 0.0	0.0 Argon		
8	77.50526	70.68508	49.97302	\
\hookrightarrow	0.0 0.0	0.0 Argon		

The first line sets the layout of the particle list. It always begins with # ATOMDATA followed by the attributes set in the file. In this case this is the particle id, the particles position (3 columns), the particles velocity (3 columns) and the particle type. As you can see we already supply spatial coordinates, however we set all velocities to zero. Those will be set by the other # preceded line in the file, which performs some manipulation on the data provided. Such lines always begin with # INPUTCONV. In this instance we set the temperature of the sample. Using a Maxwell-Boltzmann distribution each particle is then assigned a random velocity. The temperature is measured in the unit system we provided in the argon.tremolo file.

Now all pieces are in place to start the optimization on our particles. You can use the commands you learned in the "Quickstartguide" (chapter ??) to run the optimization.

The optimized particle positions are written to argon.data.999. The file looks very much like the original data file, though we find that the INPUTCONV line has been removed and two lines have been added at the top (Their use is explained in section 1.5). Furthermore the velocity columns now show non-zero entries - particle speeds corresponding to the specified temperature.

a '9' is missing in the file name

1.1.1 Exercises

- Increase the prefactor to $1.0e^{-2}$. What happens? The message tells you which particle causes the problem have a look at it and the surrounding particles in the data file.
- Have a look at the potential energy curve, how does it behave?

¹Note that if you plot these distributions, you see the shape of a Maxwell-Boltzmann distribution, but you seldomly find the parameters to be matching exactly. This is due to finite size effects, even very large simulations contain few particles compared to the real world, where you might have particle numbers by the mol.

1.2 Setting up a basic simulation

Now we are ready to start the actual simulation. Fortunately we did most of the work required setting up the optimization, so we now only have to amend very few lines in the argon.parameters file and make sure that we use the optimized data instead of the original. First we need to change the line specifying the integration type from "optimization" to "dynamics":

```
integration: type=dynamics;
```

The **cellrcut** and **domain** blocks remain untouched. We could remove the block with the optimization parameters, but since we changed the integration type to dynamic (removing optimization), we can also leave it in place without harm.

However we need to add a new block, setting the parameters for the dynamics:

```
dynamics {
    ensemble: ensemble=NVE;
    propagator, verlet: delta_T=0.5e-3, endtime=1000;
};
```

Here we set particle number, domain volume and total energy to be constant (NVE ensemble). For the integration of the particle trajectories we choose a standard verlet algorithm with a time step of 0.005 custom time units. The total simulation time will be 1000 custom time units.

At last we make a small addition to the **output** block; we would like to gain some insight into the velocity distribution of the particles:

```
 \begin{array}{c} \vdots \\ \text{energy: measure=on;} \\ \\ \text{analyze } \{ \\ \\ \text{velocity: measure=on, meanmeasure=off, vis=off,} \\ \\ \hookrightarrow \\ \text{min=0.0, max=25.0, n\_bin=50;} \\ \}; \end{array}
```

After the energy line we add a sub-block for analysis (within the output block). We consider velocities between 0 and 25 (custom units) and use bins of 0.5 (custom units) each.

Now we have to make sure that we use the optimized data instead of the original. (As a general rule one should keep copies of the original at all times.) So, after having copied the original argon.data file somewhere safe, we rename the argon.data.9999 file as argon.data. For this example we can still ignore the extra lines.

Now you can start the simulation as we did with the optimization before.

Do you mean 0.0005 or is this a mistake in the listing above?

Which extra lines?

1.2.1 Exercises

- Compare the values of the argon.etot, argon.ekin and argon.epot files (plot them in the same graph). What do you notice?
- Take a look at the velocity distribution in the argon.histogram file.
- Try starting the simulation with the original data instead of the optimized.

Is there something particular one have to look at?

- In order to smooth measurement curves and remove static, one often uses mean measurements over intervals. Switch mean measurements on for the energy measurements and compare the curves of the mean measurement with those of the regular measurement (they are written to separate files.)
- Have a look at the different energy curves. Then, in the argon.data file, change the temperature value to 3.0. (Use the #INPUTCONV temp line from the previous lesson.) The individual velocity values are then overwritten. How does this affect the different energies?

1.3 Using the Berendsen thermostat

In this section we will introduce the first of two different thermostats. Given the previous preparations we require even less changes to the parameter file. All changes are within the dynamics block.

First we have to change the ensemble type:

```
dynamics {
    ensemble: ensemble=NVT;
    propagator, verlet: delta_T=0.5e-3, endtime=1000;
```

Instead of the total energy, we now hold the temperature constant. For this type of thermostat the propagator remains untouched. Then we open a new sub-block for the details of the thermostat:

```
thermostat {
    berendsen: state=on, T_Interval=0.01;
    constanttargettemp: state=on, T_Temp=2.5;
};
};
```

We declare that we use the **berendsen** thermostat and choose to enforce it every second time step by our choice of **T_Interval**. Furthermore we declare that we wish to hold the temperature constant and at which value. Note that we chose a value lower than the originally set temperature (regardless of whether you did the exercise or not.) Those are all the changes required, so go ahead and run the simulation.

which data file has to be used? The one with '9999' after running the last simulation? In difference to the previous example, where the temperature varied around its original value, you will observe a very sharp drop from the original to the designated temperature. Afterwards all temperature values hit the value almost *exactly*. In fact, every second one is exactly at 2.5, whereas every other varies ever so slightly.

1.3.1 Exercises

- Compare the values of the argon.etot, argon.ekin and argon.epot files (plot them in the same graph). What changed compared to the previous lesson?
- Play with the time interval for the thermostat. Make it match the propagator timestep or make it a thousand times as long and observe the effects. In addition to the kinetic energy, observe the potential and total energy as well.

1.4 An alternative: The Nose-Hoover-thermostat

Shouldn't it be a é?

We also introduce a second type of thermostat. Three lines require a change in order to switch to the alternative, all in the dynamics block of the argon.parameter file.

Here we have to make a change to the propagator, in order for the Nose-Hoover thermostat to work, a velocity integrator is required.

What is a velocity integrator?

```
thermostat {
    berendsen: state=off, T_Interval=0.01;
    nosehoover: state=on, F_Mass=1.0;
    constanttargettemp: state=on, T_Temp=2.5;
};
```

In the tutorials-file F_Mass is set to 5.0...

I am not sure what this is supposed to mean...

Instead of deleting the berendsen line we can also switch it off. In the added nosehoover line we do not have to specify an interval for the thermostat, but a virtual mass. This constant determines the strength of the coupling of the particles in the simulation with a virtual heat bath.

Start the simulation as before and observe the temperature behavior.

You will note that the temperature oscillates, first significantly reducing its amplitude. While the amplitude increases again after some time it does not gain the same value as before. Thus, when using the Nose-Hoover thermostat considerations with respect to equilibration are imperative.

I do not understand the last sentence.

1.4.1 Exercises

- Again, compare the values of the argon.etot, argon.ekin and argon.epot files (plot them in the same graph). What changed compared to the previous two lessons?
- Play around with different virtual masses (0.01 100.0) and different starting temperatures. You will note some different behaviors.

1.5 Optimizing the domain

Sometimes it is not possible to determine the optimal size of the domain prior to the simulation. In addition to the use of the barostat which we handle in the next lesson of this tutorial, we will now take a look at the initial optimization

We begin by modifying the the optimization of the sample prior to the actual simulation. In particular we allow that, in addition to the positions of the atoms, also the box may be scaled to minimize the potential energy.

Duplicate 'the'.

In the argon.parameters file from the first lesson we add two lines to the optimization block

where we set the line search parameters for the cell optimization (generally we can use the same parameters as for the particle positions) and may also choose some constraints on which entries of the box matrix are allowed to vary.

Finally, we make one addition to the **common** parameter set by adding an external pressure value:

```
\begin{array}{lll} & common: & algorithm=cg \,, & maxcg=2001, & RT=periodical \,, \\ & \hookrightarrow & maxresetcg=6, & LS=strongwolfe \,, & maxlinesearch=6, \, \\ & \hookrightarrow & mean\_force\_eps=1e-6, & mean\_force\_eps\_rel=1e-10, \, \\ & \hookrightarrow & prefactor=1e-4, & extpressure=0.0024455185; \\ \}; \end{array}
```

After running the optimization, take a look at the file argon.data.9999. In addition to the new coordinates of the particles you find a time stamp and a box matrix entry. This is the domain shape after the optimization has finished. To make use of these values, you need to transfer them to the parameter file, as described in the next lesson.

1.5.1 Exercises

- Change the extpressure value. Check and compare the new box values². (Changes by order of magnitude are advised for clearly visible results.)
- In the constraint matrix, change one of the main axis entries (XX, YY, ZZ) to 0. Check and compare the new box values.
- Change one of the secondary axis entries to 1. Check and compare the new box values. Now rename the constraint to constraint=standard and check again. Repeat the process for constraint=symmetric.

For constraint=isotropic optimization is not possible. Is this on purpose in the manual?

1.6 Introducing barostats

In some cases it is desired to run simulations not only with isothermic, but also isobaric conditions. For this to be possible we allow the volume of the box to be variable and set a barostat similarly to the thermostat.

Again we work in the argon.parameter file and enter the thermostat after (or in place of) the thermostat:

The first 'thermostat' must be replaced by 'barostat'.

The listing and the provided tutorial files differ. Code from the file: ensemble=NPT, nosehoover: state=on. This produces an error telling, that something with the nose hoover thermostat is wrong (NaN)

In the first line we switch the barostat on and specify a virtual mass. In the second line we can choose whether the pressure aimed at shall be constant and if so, at what value. Note that as usually this value is in reduced units. Furthermore, we make our choice of constraints on the allowed changes to the box. First we specify a specific type of constraint, and afterwards we can

²If you run tremolo with increased verbosity (e.g. -v the new box matrix - among other things - will be written to the standard output)

apply additional modification by specifying entries in the box-matrix which may be affected.

Finally we have to deal with the particle data. If we start with a file where we did not use the box-optimization there is nothing else to worry about, we can start the simulation immediately. But lets take the argon.data.9999 file from the previous lesson. (Preferably the one which was optimized with the originally supplied parameters.) Should you attempt to start it right away, you will receive an error message. This is due to the fact that the box-matrix in the data file and the one in the parameter file are different. So go ahead and change the domain in the appropriate line. (Note that the value shown here may be different from your value. If you wish to use a different result (e.g. the one won from symmetric constraints) you need to change the domain type as well. For the appropriate syntax see ??.

Appropriate line of the *.parameters-file... makes it clearer

```
domain {
    size: type=cube, size=7.935811e+01;
    border: bt_xlow=periodic, bt_xhigh=periodic,
    bt_ylow=periodic, bt_yhigh=periodic,
    bt_zlow=periodic, bt_zhigh=periodic;
```

1.6.1 Exercises

- Change the extpressure value. Check and compare the new box values. (changes by order of magnitude are advised for clearly visible results)
- In the constraintmap, change one of the main axis entries (XX, YY, ZZ) to 0. Check and compare the new box values.
- In the constraintmap, change one of the secondary axis entries. Check and compare the new box values.
- Change the f_mass value to 1000. Check and compare the of the box values.
- Go back to the previous lesson and use one of the other constraints types (such as standard). Now transfer this argon.data.9999 file and insert its domain shape in the argon.parameter file. For the appropriate syntax see ??.

1.7 Bonded potentials and measuring bonds

Hitherto only the nonbonded Lennard Jones interaction has been covered in this Tutorial. In most cases the connectivity of atoms is known and it is desired that it stays in its initial configuration. In these cases the indices of the neighbor atoms are set in the appropriate column in the data file and bonded

Do you mean the "Pressure" value instead of "extpressure"?

The type of the constraintmap has to be changed too! This should be mentioned!

Mention that always pairs of coordinates have to be changed?

potentials are specified in the .potentials file. The bond type covered here, named bond is a harmonic potential – computationally inexpensive and with unbreakable bonds. As per definition bonded-type bonds can't be broken, however with harmonic potentials there will be a restoring force proportional to the deflection from the minimal energy distance r_0 . While unphysical for large displacements, this behaviour can also be desired if simulations are carried out under very high kinetic energies to speed processes up, because the general structure of the molecules will persist.

We will set up a butane example and measure the bond distances in the CH3- and CH2-groups. There are three atom types: Methyl-carbon (C in CH3), methylene-carbon (Ci in CH2, i stands for "inner") and Hydrogen (H). The butane data file is build accordingly and the field neighbors=4 is added. The ensemble provided with the tutorial files consists of $5\times5\times5$ butane molecules with a density of $2.71\,\mathrm{kg\,m^{-3}}$. The unit system used is kcalpermole.

Listing 1.1: Header of the .data file

```
# ATOMDATA Id type x=3 u=3 neighbors=4
```

When setting up the butane.potentials file, begin with Lennard Jones interaction, which should additionally work in between molecules. Tremolo-X handles Lennard Jones in bonded molecules in a way, that the potential is not calculated among direct neighbors. However for large molecules, like proteins, intramolecular interaction should be considered: These are controlled via the settings sigma14 and epsilon14, as they take effect from the fourth bond on, hence the naming.

As for the bonded potentials we use bonds, angles and torsions. The angle potential, like the bonds, is harmonic (linear restoring force for angle displacement from an optimal value) and the torsional potential is expressed as cosine series expansion. All parameters are taken from the AMBER94 force field. For the literal definition of the potential terms see the potentials-section in this manual (??).

Listing 1.2: Excerpt from the .potentials file

```
bonds { bond: particle_type1=C, particle_type2=Ci, \hookrightarrow bond_type=harmonic, k_b=310, r_0=1.526;
```

The simulation \$PROJECTNAME.parameters are a simple NVE ensemble with verlet propagator and initial temperature of 0°C. In the analyze section bond distance measurement is set up. It doesn't matter whether the measured pair is bonded in means of the neighbors field: Every pair with specified types undershooting the specified threshold (in this particular case 180 pm for C-C and 140 pm for H-C) are considered bonded and their Ids written to the \$PROJECTNAME.info.bonds (vis) file. The mean value of bond lengths of any pair of types requested in the \$PROJECTNAME.parameters file is written to the \$PROJECTNAME.generalmeas file.

Listing 1.3: Excerpt from the .parameters file

As we can see the mean values oscillate around a constant, so there is no time dependant development. We assume the ensemble is equilibrated after 20 time units and compare the mean of the mean values for any available pair and get following results:

Bond	$\mathrm{Distance/pm}$	$r_0/{ m \AA}$
C-Ci	152.7(1)	1.526
Ci-Ci	152.7(2)	1.526
C-H	109.1(1)	1.09
$\mathtt{Ci-\!H}$	109.08(9)	1.09

There are no differences in the bond length within statistical error and the values match the r_0 -Parameter because the same parameters were used for both, inner and outer, carbon atom types. The nearest neighbors of the atoms (spatial configuration) are not taken into account thus leading to an identical bond distance.

Do not understand the last sentence.

1.7.1 Exercises

- Increase the temperature or lower the bond strength and observe the magnitude of the oscillation.
- Alter the equilibrium distance r_0 .

1.8 Tersoff potential and stress

After we successfully introduced basic bonded and non-bonded force fields, this chapter shows how to use a bond order potential to determine Young's Modulus of a single graphene sheet. Instead of defining fixed individual neighbors, the potential function will determine the spatial configuration of surrounding carbon atoms by itself. This way the graphene.data file looks rather trivial, the force field parameters for graphene.potentials are taken directly from [4]:

Just like in 1.6 an NPT-ensemble is used, but this time additionally to the external pressure we also support a custom stress tensor, which stretches the domain in xx-direction with linearly increasing strength, starting from 0 in the beginning up to $1e5 \frac{[F][V]}{[A]}$ (kcalpermole units). At this time note that the stress value is not given in units of pressure, as one would expect, but contains a surplus V-Term. The volume can usually be derived from the domain dimensions, but particularly in this tutorial we have to decide how much volume a single graphene sheet has. Also the box vectors need to

Verstehe Satz mit 'box vectors' nicht!

 $^{{}^{3}\}mathbf{F}$ orce, **V**olume and **A**rea. For a detailed explanation of the volume term see chapter ??

be changed individually, so we will choose standard constraints (isotropic constraints would be infeasible due to the coupling of the constraints).

Stress and strain are measured automatically if outm is set. However if you are interested in the stress distribution along individual particles you need to use the local_stress-feature:

```
analyze {
    local_stress: localstress=on;
};
```

If activated, the beta column in the visual .pdb output contains per particle stress, which can be visualized with an external tool like VMD-Viewer (not covered by this tutorial). Due to limitations of the PDB-format the numerical value is clipped at $99.99 \frac{[F][V]}{[A]}$.

After the simulation has been finished, the output is analyzed by plotting a stress-strain diagram. The strain is defined as the length change relative to the initial box xx-length L_0 : $\epsilon(L) = \frac{L-L_0}{L0}$. The argument will be the 43^{rd} column in the file graphene.mbox, which is the domain length in xx direction. The stress is read from column 31 in the same file and divided by the volume of the graphene sheet, which is the product of xx length (column 43), yy length (column 44) and height, which has to be chosen by the user. We set $height = 3.7 \mathring{A}$, the inter layer distance in graphite. Note that since we pull, the stress value from the file is negative. With this information we can now use gnuplot to create the diagram and fit a slope against the initial section, representing the elastic zone where Hook's Law applies. This slope is the Young's Modulus.

```
\begin{array}{lll} & \text{gnuplot} > \text{L0} = 48.19683 \\ & \text{gnuplot} > \text{epsilon}(\text{L}) = (\text{L-L0})/\text{L0} \\ & \text{gnuplot} > \text{plot} \text{ 'graphene.mbox' using } \\ & \hookrightarrow & (\text{epsilon}(\$43)) {:} (-\$31/(\$43 {*\$44 {*3.7}})) \\ & \text{gnuplot} > \text{f(x)} = \text{E*x} \\ & \text{gnuplot} > \text{fit} \quad [0 {:} 0 {.} 04] \quad \text{f(x)} \quad \text{'graphene.mbox' using} \end{array}
```

(epsilon(\$43)):(-\$31/(\$43*\$44*3.7)) via E

Was ist die beta column im PDB-Format?

Wo kann man die Hoehe ablesen?

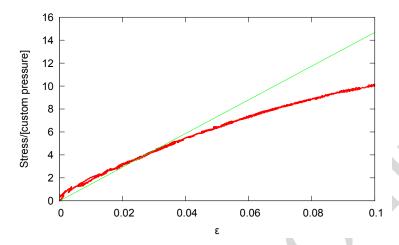


Figure 1.1: Stress-strain diagram of monolayer graphene (red) with green slope fit.

We obtain $E = 146.7(5) \frac{[F]}{[A]}$, in SI-units 1019(3) GPa, which sufficiently reproduces the measurement of 1 TPa in [3].

1.8.1 Exercises

- Change the pull direction to yy. Are there differences?
- Alter the force progression:
 - Increase the simulation length by factor ten.
 - Add a 100 [t] relaxation time at the beginning ($\sigma = 0$).
 - Increase the maximum stress at the end of the simulation by factor five.

Create a new stain-stress plot – is there a difference in the fit quality $(\chi^2_{\text{d.o.f.}})$ and parameter error as supplied by gnuplot)?

Was ist σ ? Bin mir nicht sicher, wie man das machen soll...

1.9 Long ranged potentials 1 - Halley's Comet with N^2

This tutorial covers how to set up simulation to use long ranged potentials like gravity or the coulomb potential. In either case the potential's distance characteristic is $\propto 1/x$ and any kind of cutoff introduces significant errors on the forces and with them error in macroscopical observables.

The simplest way to deal with these kind of potentials is using an ordinary pair potential. As stated earlier the potential should not be cut off but has to to fit the linked cell structure of the domain. With only a few particles it is adequate to run the whole simulation in a single cell, the performance drops quadratically with the particle number. In this example we simulate our solar system and calculate when Halley's Comet runs through its perihelion point next time (as of time of writing, year 2013).

Fehlendes Wort nach 'to'?

To obtain handy numbers we set up following unit system in the . tremolo file:

Quantity	Unit	Equals
Length	$1.496\times10^{11}\mathrm{m}$	ua
Time	$86400\mathrm{s}$	d
Mass	$1 \times 10^{23} \mathrm{kg}$	$\approx m_{\rm Earth}/100$

For the .data file we use positions (relative to the Barycenter) and velocities obtained from JPL HORIZONS for sixty astronomical objects of our solar system including the most massive ones. The data represents the state of the solar system on December 19th, 2012.

Tremolo-X doesn't support "gravity" literally, but since the potential is identic to the coulomb potential disregarding the force constant we set up a coulomb simulation with adapted epsilon0inv. It can easily be calculated with Gnu Units:

```
> units -t G "(m*1.496e11)**3 / (kg*1e23 * (s*86400)**2)" 1.4881216e-11
```

Note that the value needs to be multiplied by -1 so equally named "charges" attract each other.

We use an NVE ensemble with verlet propagator, 0.05 d timestep and 100 a endtime, "temperature" and "pressure" can't be applied on the experimental conditions. Also it would not be accurate to wrap around the gravitational forces at the borders of the solar system as it is preferred when homogenous systems are observed, so we choose the domain to three times larger than the solar system, place it in the center (done with a INPUTCONV SHIFT directive in the data file) and set leaving boundary conditions (particles passing the border are removed):

A single linked cell will be large enough to contain the whole ensemble. This is not practical for large scale molecular dynamics because it can't be paralellized, however unless a more sophisticated method like SPME (covered in the next section) is used, it is the only way to obtain accurate results using long ranged potentials.

The coulomb section is set up thusly:

```
coulomb {
    permittivity: epsilon0inv=-1.4881216e-11;
    n2spline: state=on, r_cut=80, r_l=70, i_degree=5;
};
```

The n2spline solver calculates the force for every pair and cuts off beyond 70 ua with a spline taper (like ljspline). The cutoff is chosen large enough to contain the whole solar system so the forces will be exact. If n2spline is used in molecular dynamics with small cutoff the spline interpolation guarantees conservation of energy but the forces will be off their exact result due to the characteristic of the potential.

The bond distance measurement covered earlier is used to measure the distance between Halley's Comet/Earth and the Barycenter:

Finally we set up the objects in the .potentials file. To calculate the gravity potential with the coulomb solver we need to set the "charge" of the particle to its mass. The Barycenter is included as pseudo particle with mass 1 and no charge – this way it won't be affected by any force and stay central.

After running the simulation we extract the time to the minimum of $|\vec{r}_{\text{Barycenter}} - \vec{r}_{\text{Halley}}|$ by plotting the first and fourth column of the .generalmeas file. It's 17750 time units (days), so the next perihelion is 25th July 2061, slightly deviant to the value computed with HORIZONS: 28th July. This relative difference of 0.02% is caused by the simplification of the solar system to sixty objects and the resulting difference in the local density. The minima of the Earth–Barycenter distance are 365.24d apart which matches the definition of the tropical year.

1.9.1 Excercises

Remove any particle but Sun, Jupiter, Barycenter and Halley and compare the relative error of the perihelion.

Simulation hat bei mir nicht funktioniert: Viele Partikel verlassen das Gebiet und ich erhalte nur Nullen als Distanzen... • Only remove Jupiter from the original setup and compare the relative error, observe planetary trajectories and the length of an earth year.

1.10 Long ranged potentials 2 - Sodium chloride with SPME

This part covers a typical usage scenario of coulomb forces in molecular dynamics with more than just a few particles. To maintain a good performance with a large N the potential is seperated into a short ranged part, which is calculated in a linked cell fashion as before, and a long ranged part, which is calculated by Ewald summation in fourier space, to take into account farther particles. In comparison to the Fast Multipole Method, which abstracts groups of far particles into a single one, this way is espacially suitable for periodic systems, like an ionic crystal: In this example we are going to simulate solid NaCl and measure its radial distribution functions. The system of units used is kcalpermole.

In the .potentials-file we set up the short ranged interactions using the Tosi Fumi[1] Potential, which has shown to produce accurate results with this kind of system. The starting configuration in the data file is an NaCl-structure with small random offset for each atom at 20 °C.

We set up an NPT-ensemble in the .parameters file with 1000 hPa pressure maintained by the Parrinello-Barostat with isotropic constraint (the crystal is cubic) and Nose-Hoover-Thermostat for fixed temperature. In the coulomb-section we specify the parameters for the spme-method and a force constant using the vacuum permittivity:

```
\begin{array}{c} \text{coulomb } \{ \\ \text{permittivity: epsilon0inv=} 332; \\ \text{spme: state=} \text{on, r_cut=} 9.0, G=0.32, i\_degree=} 5, \\ \hookrightarrow \text{cellratio=} 4; \\ \}; \end{array}
```

Up to r_cut , which is the short ranged part of the potential, the force is evaluated locally (restricted to neighbored cells) and pair wise, as if n2spline was used. From there it is approximated by bell curves with splitting coefficient G, which is inverse to the standard deviation σ , and applied on a mesh, which is created by splitting the linked cell cellratio times (rounded upward to the next power of two). G should be kept at $0.24 \, \text{Å}^{-1}$ to $0.35 \, \text{Å}^{-1}$.

In the analyze-section we set up the measurement of the radial distributions of all atom types:

```
 \begin{array}{lll} \text{output } \{ \\ [\dots] \\ & \text{analyze } \{ \\ & \text{radial: measure=on, meanmeasure=off, vis=off,} \\ & \hookrightarrow & \text{r\_cut=9.0, n\_bin=50;} \end{array}
```

The value of r_cut has to be within the lcs: cellrcut limits, just like the r_cut of the potentials.

Since we now use the SPME method, we have to use a parellel version of Tremolo-X, since the SPME method is not implemented sequentially. Nevertheless you can unse the SPME method and start the parallel version with a single process.

If you do not know how to run the parallel version of Tremolo-X, please check chapter ?? and ??.

1.10.1 Excercises

- Compare the radial distribution histograms from first (ideal NaCl structure) and last timestep.
- Decrease/Increase the temperature and observe the differences in the radial distribution.

1.11 Melting point of Sodium Chloride

In the previous Tutorial the basics for simulating an ionic crystal using the Coulomb- and Tosi-Fumi-Potential have been covered while this one shows how to determine the melting point of NaCl, which is a common application of molecular dynamics. From the different methods described in [5] we are going to use the Voids method: A series of NaCl lattices with increasing defect concentration (removed atoms) is simulated using an NPT-ensemble with temperature timeline. Starting with an ideal lattice, which is a hypothetical, defect free state at 0 K, the temperature at which the lattice breaks ("melting point") is significantly overestimated because the activation energy for this transition is very high. By removing atoms from the crystal it becomes labile and the lattice breaks easily if the temperature is high enough, lowering the activation energy to start the melting process. If the defect concentration is too high the crystal rearranges back into a more stable configuration, increasing the observed melting point. This process leads to a reduced volume which has to be compensated for, using a barostat. Therefore if the measured melting point is plotted against defect concentration one

can see that the measured melting point decreases quickly with increasing defect concentration at first and then oscillates around the actual melting point, which can be obtained by calculating the mean value.

A tricky aspect is how to observe the melting point: Liquid and solid state can be discriminated using the MSD-measurement, potential energy, density, radial distribution or bond length, usually indicated by a rapid slope change which can be seen if these measurements are plotted. Preliminary studies have shown that observing the bond length is the easiest to interpret while very accurate option in this particular example: As the crystal heats up the bond length increases linearily until the crystal breaks and the bond distances relaxes rapidly. The obsverved melting point T_m is the temperature at maximum bond length.

The simulation setup is similar to the previous tutorial apart from the thermostat settings and more measurement options in the nacl.parameters file:

```
thermostat {
    timeline: state=on, [time, temperature, interpolation =
        (0, 0.5822, linear), # 20°C
        (100, 0.5822, linear),
        (1000, 2.9255, linear)]; # 1200°C

    nosehoover: state=on, F_Mass=800;
};
```

TODO: Buchstabendreher beim Gradzeichen im PDF/Text fixen (Encoding-problem mit lstlisting).

With these settings the temperature is held constant at 20 °C for 100 time units and then lineary increased to 1200 °C at the end of the simulation (1000 time units).

Welcher Buchstabendreher?

Every Na-Cl-pair with a distance less than or equal to 4 Å is considered bonded and contributes to the mean distance written to nacl.generalmeas.

When carrying out a series of simulations it is handy to make use of the defaultpath-option in nacl.tremolo. The simulation is organized into a root directory which contains any file but .data and .tremolo and subdirectories containing only these, whith individual nacl.data containing an increasingly more defective crystal. The individual nacl.tremolo files look like this

```
global: defaultpath = "../nacl";
global: projectname="nacl";
global: comment="NaCl";
```

global: systemofunits=kcalpermole;

with the defaultpath set to the parent directory and nacl as basename, so tremolo looks for *nacl*.potentials and so forth. The subdirectories are named after the relative count of cells containing a (single) pair defect. A single cell contains 8 atoms, so the resulting defect concentration is *dirname*/4.

After the simulations are done the times at which the bond lenght peaks are read from nacl.generalmeas and the temperature at this time is looked up in nacl.ekin, fourth column.

Plotting T_m against defect concentration results in following graph:

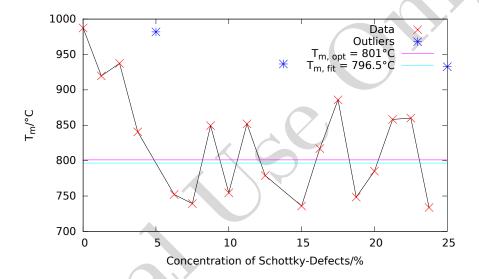


Figure 1.2: Melting points at different defect concentrations. $T_{\rm m,\ opt}$ is the melting point according to literature, $T_{\rm m,\ fit}$ is the mean value of the oscillating region. The black dot-connecting line and outliers are for visualization purposes only.

The oscillating region starts after $7.5\,\%$, the resulting mean value differs by $4.5\,\mathrm{K}$ from the exact value of $801\,^\circ\mathrm{C}$, which is a relative error as low as $0.6\,\%$.

1.12 The EAM potential - Observing phase transition in Metall

The "embedded atom method" (EAM) is a standard potential used in the analysis of metalls and alloys. They are used quite successfully in the investigation of fractures, surface reactions martensite-austenite transitions and phasechanges in condensed matter, nano particles and thin films.

In this tutorial lesson we will demonstrate the use of the EAM potential and how a phase transition can be analyzed with Tremolo-X. We will heat a Fe-Ni nanoparticle from 100 K to 800 K and observe it changing its lattice structure from bcc to fcc/hcp.

In order to use EAM potentials, the user must have a file with EAM parameters in the format "eam/fs" (a generalized EAM type by Finnis-Sinclaire) or in the "eam/alloy" format, which is slightly modified from DY-NAMO setfl file formats. For a detailled description of the EAM formats supported, please check the appropriate section ??

The unit system of the eam parameters file determines the units which need to be used throughout the simulation. As a result in this example SI units will be used. The respective entries for iron and nickel particle values in the potential file are consequently:

```
particles
         particle:
                            particle type=Fe,
              element name=Fe,
                            sigma = 1.0,
                                                        epsilon = 0.0,
                            sigma14=1.0,
                                 epsilon 14 = 0.0,
                            mass = 9.2732785e - 26,
                            free = 3,
                            charge=0;
         particle:
                            particle type=Ni,
              element name=Ni,
                            sigma = 1.0,
                                                        epsilon = 0.0,
                            sigma14 = 1.0,
                                 epsilon14 = 0.0,
                            mass = 9.7462664e - 26,
                            free = 3,
                            charge=0;
         };
```

and subsequently we specify the format ("alloy" format) and the filename in the following way:

```
eam {
    setfl: file="Fe-Ni-MeyerEntel-1995.eam.alloy";
};
```

In the parameter file we need to specify the following:

```
########### Section Ensemble and Propagator:
dynamics {
                          ensemble=NVT;
        ensemble:
                          verlet: delta T=1.0e-15,
        propagator,
              endtime = 8.0e - 11,
              timeinteps=1e-07,
                                        maxiteration=100;
        thermostat
                 berendsen:
                              state=on,
                              T_Interval = 1.0e - 15;
                              state=on,
                 timeline:
                        [time, temperature, interpolation=
                                    1.3806503e-21, linear), #
                             100 \mathrm{K} \ / \ 7.2429638 \, \mathrm{e}{+22} \ \mathrm{K}
                                   1.3806503e-21, linear),
                        (1e-14,
                        (7.56e-11, 1.1045202e-20, linear), #
                             800K
                        (8.0e-11, 1.1045202e-20, linear);
                          };
############ Section Solver and Parallelization:
        cellrcut = 5.6001e - 10;
############ Section Output Measurement:
output
         Outvis: T Start=0, T Delta=5.0e-13, Step Delta=100;
                 T_Start=0, T_Delta=1.0e-15, Step_Delta=100;
        Outm:
        Outmm:
                 T_Start=0, T_Delta=5.0e-14,
              T Deltam=4.0e-14, Step Delta=20,
              Step Deltam=10;
        Outdata: T Start=0, T Delta=10.0e-13, Step Delta=50;
        energy: measure=on,
                                 meanmeasure=off;
         analyze {
           radial: measure=on, meanmeasure=off, vis=on,
                r cut = 5.6001e - 10, n bin = 56;
                 radialdistribution
                      radialdist: particle type1=Ni, \
                           particle type2=Ni;
                      radialdist: particle type1=Ni, \
                           particle type2=Fe;
                      radialdist: particle type1=Fe, \
                           particle type2=Fe;
                       };
                     };
         };
```

Warum? Verstehe ich nicht...

Note that while we use SI units, the temperature has a scaling prefactor, which needs to be accounted for when setting the thermostat. While we use periodic boundaries, those will not be relevant to our simulation, as we set the simulation domain to a cube of 60Å, whereas the nanoparticle has a

1.12. THE EAM POTENTIAL - OBSERVING PHASE TRANSITION IN METALL29

radius of 20 Å. Thus no interaction with or across the domain boundaries is present.

We now can run the simulation .After having run the simulation, we can take a look at the potential energy curve and notice that its slope changes around 4.64e-11s into the simulation. It is at this time that the transformation takes place.

Ich erhalte zu Beginn der Simulation zwei Fehlermeldungen...

We will now analyze the radial distribution of the sample at the beginning of the simulation, during the melting phase and after the transition phase. For this we will use the tool "CalcRadialHist" delivered with Tremolo-X.

By calling

CalcRadialHisto eam.histogram 5e-15 15.3e-15 0.0 5.6e-10 \hookrightarrow 3.0e-10 1 56 545 2196 216000 3.0e-10 > RadialHist

we average the radial distribution over the time interval from 5 to 15 fs and write the results to the file RadialHist. The same can be done for the time intervals 4e-11s to 4.3e-11s and 5e-11s to 5.3-11s. When those three distributions are compared, we observe, that around 4e-11s the original configuration has been severly melted, but that around 5e-11s the atoms have rearanged themselves, which can be seen from the clearly shifted peaks in the radial distribution function.



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