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rushlight - Python-based Forward Modelling of Coronal Plasma Models

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Summary

The rushlight Python package provides a framework for creating synthetic images of plasma structures for model-to-data comparisons with coronal events. It handles the projection and alignment of 3D simulated datasets to user-defined locations and orientations relative to the sun. The produced observables are comparable to observations made by instruments such as the Hinode X-Ray Telescope (XRT) and the Solar Dynamics Observatory Atmospheric Imaging Assembly (AIA). rushlight aims to integrate into the growing community of Python-based astrophysics software such as Astropy, SunPy and XRTpy.

Statement of need

rushlight is a Python package which performs forward modelling of simulated 3D plasma datasets in the coronal environment. Its core functionality lies in creating synthetic observables in Soft X-Ray filter bands produced by XRT, and Ultraviolet / Extreme Ultraviolet filter bands produced by AIA.

rushlight adapts some of the core functionality of the FORWARD package, written in the Interactive Data Language (IDL) (Gibson et al., 2016). rushlight is under active development, and aims to be continually improved as to implement more of FORWARD's features.

Part of rushlight's core motivation is to make EUV / SXR forward modelling more accessible to the growing company of astrophysicists who utilize the Python language to develop and share scientific software. To this effect, rushlight has been developed as to be both compatible and scalable with release versions of other astrophysics open-source software, such as Astropy (Astropy Collaboration et al., 2013) (Astropy Collaboration et al., 2018) (Astropy Collaboration et al., 2022), SunPy (Mumford et al., 2020), and XRTpy (Velasquez et al., 2024). By creating a forward-modeling solution built upon newer and actively maintained dependencies, rushlight can be integrated into state-of-the-art solar physics research.

Package Structure

rushlight's modules are organized as to promote the addition of new emission models and instruments to produce synthetic observables with. The package's main functionality comes from the following classes:

- `rushlight.utils.proj_imag_classified.SyntheticImage` - This module is the parent module to all other Synthetic Image classes, regardless of simulated filter type. It is responsible for translating user input into a single object containing both reference and model data. The Python module `yt` is used its ability to orient and project volumetric data from multiple simulation platforms.

- `rushlight.utils.proj_imag_classified.SyntheticFilterImage` - rushlight is intended to be expanded upon by developing other modules similar to SyntheticFilterImage, which overloads the SyntheticImage class to apply the appropriate imaging models specific to UV and SXR observations. Figure **??** showcases multiple simulated filter images produced by rushlight's SyntheticFilterImage class.
- `rushlight.utils.dcube.Dcube` - This module serves to process user provided simulation datasets into a `YTRegion` object. If one is not provided, it can generate a dummy uniform grid dataset.
- `rushlight.utils.rimage.ReferenceImage` - This module processes user provided reference observation maps into `sunpy.map.Map` objects from which coordinate data is later calculated.
- `rushlight.utils.synth_tools.calc_vect` - rushlight accepts user specification of 3 points in 3D space located on the intended projection plane for their simulation data. From these 3 points, it uses the simulated observer's location to calculate the vector that is normal to this plane, and the vector that determines the rotation of the projection relative to the normal axis. These `norm` and `north` vectors, respectively, are used in the yt.off_axis_projection module to calculate projection orientation.
- `rushlight.utils.emission_models.uv.UVModel` - This module is used by `rushlight.utils.proj_imag_classified.SyntheticFilterImage` to interpolate the temperature response function for a specified AIA channel, and then to utilize the density and temperature data from the simulation dataset to estimate the UV intensity of the solar plasma.
- `rushlight.utils.emission_models.xrt.XRTModel` - Similar to `rushlight.utils.emission_mode` this module instead interpolates the temperature response function for a specified combination of XRT filters to estimate the SXR intensity of the simulation dataset.

As referenced above, rushlight is dependent on a number of packages to provide critical functionality.

The yt Python package (Turk et al., 2011) provides the yt.off_axis_projection module, which is used for translating the provided 3D simulation dataset into a 2D image by efficiently integrating temperature and density information along an arbitrary line of sight.

The SunPy (Mumford et al., 2020) package is the modus operandi by which the 2D observables are made accessible by the package. It provides the sunpy.map.header_helper.make_fitswcs_header module, which allows the synthetic observable's properties (such as time and coordinate data) to be customized and exported as a .fits file along with the projected image. It also provides the sunpy.map class, which allows easy access to the observable's properties and World Coordinate System (WCS), as well as visualization of the final synthetic observable in the observation plane.

The Astropy (Astropy Collaboration et al., 2013) (Astropy Collaboration et al., 2018) (Astropy Collaboration et al., 2022) package is used largely for its astropy.coordinates.SkyCoord module, which provides the flexible framework for calculating the relative positions of observables in physical space, as well as transformation to and from different reference frames (eg. Heliographic Stonyhurst, Heliocentric, Helioprojective). Additionally, the astropy.units module is used extensively to dynamically determine the resulting units of various physical calculations across rushlight's observable alignment procedure.

The CoronalLoopBuilder (Yu, 2023) package provides a convenient way for users to define a projection plane that passes through a particular point in space defined in the Heliographic Stonyhurst coordinate system. CoronalLoopBuilder is heavily used in the Jupyter Notebook example files provided with rushlight's documentation, allowing the user to make dynamic comparisons between their loop object, and their observable.

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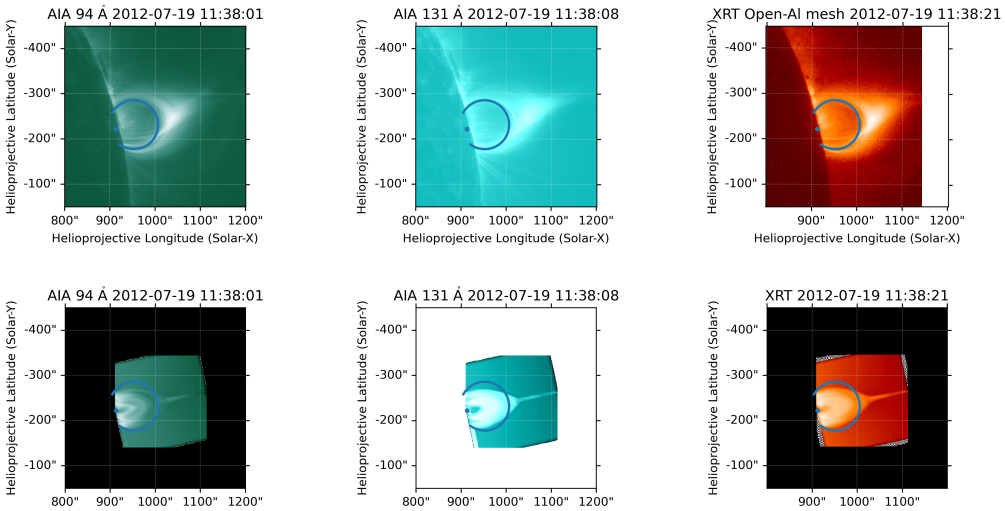

Figure 1: The top row of images are SunPy maps created from observables captured by AIA / XRT spacecraft of an eruptive solar flare event occurring on 2012-07-19. The bottom row of images are synthetic observable SunPy maps created by the alignment and subsequent projection of UV emission along the spacecraft's line of sight. An identical Coronal Loop Builder object is plotted in each case to provide a visual comparison of scaling and alignment across the real and synthetic observables.

Ongoing Projects

rushlight is currently being utilized to support the visualization and calculation of EUV and SXR emissions of a Magnetohydrodynamic simulation of an eruptive solar flare (Oparin et al., 2026). The synthetic observables produced by rushlight serve as the basis for comparison between the flare simulation and documented flare events, providing both a method of model validation, and a projected field of expected plasma properties that are used to further investigate the underlying magnetic field structure of the flare events.

Acknowledgements

Test

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