

¹ rushlight - Python-based Forward Modelling of ² Coronal Plasma Models

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⁶ Summary

⁷ The rushlight Python package provides a framework for creating synthetic images of plasma
⁸ structures for model-to-data comparisons with coronal events. It handles the projection and
⁹ alignment of 3D simulated datasets to user-defined locations and orientations relative to the
¹⁰ sun. The produced observables are comparable to observations made by instruments such as
¹¹ the Hinode X-Ray Telescope (XRT) and the Solar Dynamics Observatory Atmospheric Imaging
¹² Assembly (AIA). rushlight aims to integrate into the growing community of Python-based
¹³ astrophysics software such as Astropy, SunPy and XRTpy.

¹⁴ Statement of need

rushlight is a Python package which performs forward modelling of simulated 3D plasma datasets in the coronal environment. Its core functionality lies in creating synthetic observables in Soft X-Ray filter bands produced by XRT, and Ultraviolet / Extreme Ultraviolet filter bands produced by AIA.

¹⁹ rushlight adapts some of the core functionality of the FORWARD package, written in the
²⁰ Interactive Data Language (IDL) ([Gibson et al., 2016](#)). rushlight is under active development,
²¹ and aims to be continually improved as to implement more of FORWARD's features.

²² Part of rushlight's core motivation is to make EUV / SXR forward modelling more accessible
²³ to the growing company of astrophysicists who utilize the Python language to develop and
²⁴ share scientific software. To this effect, rushlight has been developed as to be both compatible
²⁵ and scalable with release versions of other astrophysics open-source software, such as Astropy
²⁶ ([Astropy Collaboration et al., 2013](#)) ([Astropy Collaboration et al., 2018](#)) ([Astropy Collaboration](#)
²⁷ et al., 2022), SunPy ([Mumford et al., 2020](#)), and XRTpy ([Velasquez et al., 2024](#)). By creating
²⁸ a forward-modeling solution built upon newer and actively maintained dependencies, rushlight
²⁹ can be integrated into state-of-the-art solar physics research.

³⁰ Package Structure

³¹ rushlight's modules are organized as to promote the addition of new emission models and
³² instruments to produce synthetic observables with. The package's main functionality comes
³³ from the following classes:

- ³⁴ ▪ `rushlight.utils.proj_imag_classified.SyntheticImage` - This module is the parent
³⁵ module to all other Synthetic Image classes, regardless of simulated filter type. It is
³⁶ responsible for translating user input into a single object containing both reference and
³⁷ model data. The Python module `yt` is used its ability to orient and project volumetric
³⁸ data from multiple simulation platforms.

- 39 ■ `rushlight.utils.proj_imag_classified.SyntheticFilterImage` - `rushlight` is in-
40 tended to be expanded upon by developing other modules similar to `SyntheticFilterImage`,
41 which overloads the `SyntheticImage` class to apply the appropriate imaging models
42 specific to UV and SXR observations.
- 43 ■ `rushlight.utils.dcube.Dcube` - This module serves to process user provided simulation
44 datasets into a `YTRegion` object. If one is not provided, it can generate a dummy uniform
45 grid dataset.
- 46 ■ `rushlight.utils.rimage.ReferenceImage` - This module processes user provided refer-
47 ence observation maps into `sunpy.map.Map` objects from which coordinate data is later
48 calculated.
- 49 ■ `rushlight.utils.synth_tools.calc_vect` - `rushlight` accepts user specification of 3
50 points in 3D space located on the intended projection plane for their simulation data.
51 From these 3 points, it uses the simulated observer's location to calculate the vector that
52 is normal to this plane, and the vector that determines the rotation of the projection
53 relative to the normal axis. These norm and north vectors, respectively, are used in the
54 `yt.off_axis_projection` module to calculate projection orientation.
- 55 ■ `rushlight.utils.emission_models.uv.UVModel` - This module is used by
56 `rushlight.utils.proj_imag_classified.SyntheticFilterImage` to interpolate
57 the temperature response function for a specified AIA channel, and then to utilize the
58 density and temperature data from the simulation dataset to estimate the UV intensity
59 of the solar plasma.
- 60 ■ `rushlight.utils.emission_models.xrt.XRTModel` - Similar to `rushlight.utils.emission_mode`
61 this module instead interpolates the temperature response function for a specified
62 combination of XRT filters to estimate the SXR intensity of the simulation dataset.

63 Acknowledgements

64 Test

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