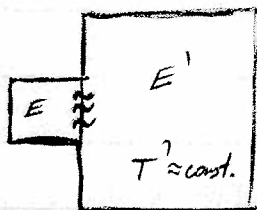


Canonical ensemble

system in contact with heatbath at temperature T



E : energy of system of interest
 E' : energy of heatbath
 composite system: $E^{(0)}$

$$H(p, q) = E$$

$$E + E' = E^{(0)} = \text{const.}$$

$$E \ll E' \lesssim E^{(0)}$$

system of interest (not connected: in touch with heatbath)

system and heatbath are quasi-independent:

$s = (p, q)$, etc

$$p(p, q) = \frac{\Omega'(E')}{\Omega^{(0)}(E^{(0)})}$$

probability of being in a particular state with (p, q) (energy $E = E^{(0)} - E'$)

$$\ln p = \text{const.} + \ln \Omega'(E^{(0)} - E) = \quad \text{expand about } E' \lesssim E^{(0)}$$

$$= \text{const.}' - \left. \frac{\partial \ln \Omega'}{\partial E'} \right|_{E' = E^{(0)}} E + \dots = \text{const.} - \frac{E}{kT}$$

T' : temperature of heatbath

we showed that in thermal equilibrium, $T = T'$ (system and heatbath)

$$p(p, q) = \frac{1}{Z} e^{-\frac{E}{kT}} = \frac{1}{Z} e^{-\beta H(p, q)}$$

$$\beta = \frac{1}{kT}$$

"Boltzmann" factor

elliptical
microcav

$$\langle f(p, q) \rangle = \int \frac{d^3p d^3q}{N! h^{3N}} p(p, q) f(p, q)$$

$$\int \frac{d^3p d^3q}{N! h^{3N}} p(p, q) = 1$$

normalization

$$Z = \int \frac{d^3p d^3q}{N! h^{3N}} e^{-\beta \mathcal{H}(p, q)}$$

partition function
(normalization factor for p as well)

one can discretize systems with formal "function" numbers s

$$p_s = \frac{e^{-\beta E_s}}{Z}$$

$$Z = \sum_s e^{-\beta E_s}$$

discrete

$$d\Gamma \equiv \frac{d^3p d^3q}{N! h^{3N}}$$

$$Z = \int d\Gamma e^{-\beta \mathcal{H}(p, q)}$$

generating function

$$\int p(p, q) d\Gamma = 1$$

normalize!

use this def:

(extensive variable)

$$\langle E \rangle = \langle \mathcal{H}(p, q) \rangle = \int d\Gamma p(p, q) \mathcal{H}(p, q) = \int d\Gamma \mathcal{H}(p, q) \frac{e^{-\beta \mathcal{H}(p, q)}}{Z} =$$

$$\frac{1}{Z} \int d\Gamma \mathcal{H}(p, q) e^{-\beta \mathcal{H}(p, q)} = -\frac{\partial}{\partial \beta} \ln Z \sim \mathcal{O}(N) \text{ (extensive)}$$

$$-\frac{\partial \langle E \rangle}{\partial \beta} = \left[\frac{\int d\Gamma \mathcal{H}^2 e^{-\beta \mathcal{H}}}{Z} - \frac{(\int d\Gamma \mathcal{H} e^{-\beta \mathcal{H}})^2}{Z^2} \right] = \langle E^2 \rangle - \langle E \rangle^2 \sim \mathcal{O}(N)$$

extensive
~~ext~~
ext

$$C_V = \left(\frac{\partial \langle E \rangle}{\partial T} \right)_V = \left(\frac{\partial \langle E \rangle}{\partial \beta} \right)_V \frac{d\beta}{dT} = -\frac{1}{kT^2} \left(\frac{\partial \langle E \rangle}{\partial \beta} \right)_V$$

$\ln Z \sim$ cumulant
generating
function

$$C_V = \frac{1}{kT^2} (\langle E^2 \rangle - \langle E \rangle^2)$$

$$\Delta E = E - \langle E \rangle$$

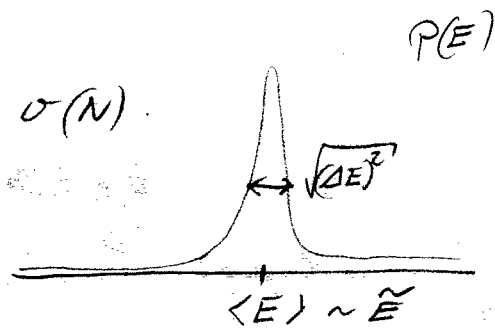
$$\langle E^2 \rangle - \langle E \rangle^2 = kT^2 C_V$$

(linear response)
fluct.-dissip
relations

$$\tau_E \sim o(N)$$

$$\langle (\Delta E)^2 \rangle = \langle E^2 \rangle - \langle E \rangle^2 \sim o(N)$$

$$\frac{\sqrt{\langle (\Delta E)^2 \rangle}}{\langle E \rangle} \sim \frac{1}{\sqrt{N}}$$



The distribution of the energy

$$P(p, q) = \frac{e^{-\beta \mathcal{H}(p, q)}}{Z}$$

$$Z = \int d\Gamma e^{-\beta \mathcal{H}(p, q)}$$

N-particle
density of state
↓

$$P(p, q) d\Gamma = P(E) dE \Rightarrow P(E) = P(p, q) \frac{d\Gamma}{dE} = P(\mathcal{H}(p, q)) g(E)$$

$$P(E) = \frac{e^{-\beta E}}{Z} g(E)$$

probability density that
the macroscopic system has energy
between E and $E+dE$

$$P(E) = \frac{1}{Z} e^{\ln(g(E)) - \beta E} = \frac{1}{Z} e^{f(E)}$$

$$f(E) = \ln(g(E)) - \beta E$$

expand about \tilde{E} , which maximizes f

$$f(E) \approx f(\tilde{E}) + \frac{1}{2} \left. \frac{\partial^2 f}{\partial E^2} \right|_{\tilde{E}} (E - \tilde{E})^2 + \dots$$

$$\left. \frac{\partial f}{\partial E} \right|_{\tilde{E}} = 0$$

$$\frac{\partial \ln(g(E))}{\partial E} \bigg|_{\tilde{E}} - \beta = 0$$

⇓

$$\tilde{E}(\beta) \leftrightarrow \beta(\tilde{E})$$

"equilibrium" energy \tilde{E}
for a given β

$$\left. \frac{\partial^2 f}{\partial E^2} \right|_{\tilde{E}} = \left. \frac{\partial^2 \ln(g(E))}{\partial E^2} \right|_{\tilde{E}} = \left. \frac{\partial \beta(\tilde{E})}{\partial \tilde{E}} \right|_{\tilde{E}}$$

$$= \frac{\partial}{\partial \tilde{E}} \left(\frac{1}{kT} \right) = -\frac{1}{kT^2} \frac{\partial T}{\partial \tilde{E}} = -\frac{1}{kT^2 C_V}$$

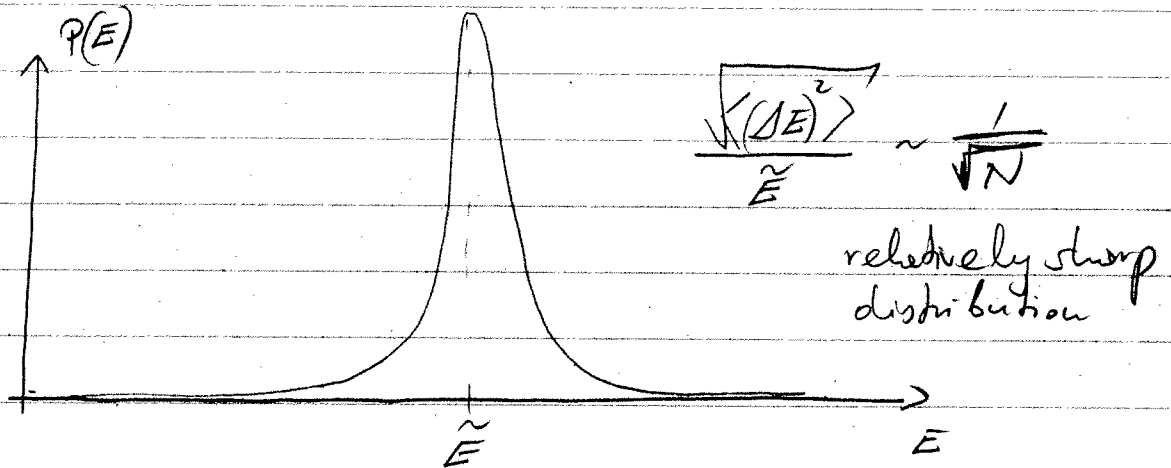
$$P(E) \approx \frac{e^{\ln(g(\tilde{E})) - \beta \tilde{E}}}{Z} e^{-\frac{(E - \tilde{E})^2}{2 k T^2 C_V}}$$

= const. e

↑ Gaussian
energy fluctuations

$$\tilde{E} \simeq \langle E \rangle \sim N \quad (\text{extensive quantity})$$

variance $\langle (\Delta E)^2 \rangle = \langle (E - \tilde{E})^2 \rangle = kT^2 C_V \sim N$ also extensive
as discussed earlier



In the $N \rightarrow \infty$ limit, $P(E)$ asymptotically becomes Gaussian

Connection with thermodynamics

$$P(E) = \frac{g(E) e^{-\beta E}}{Z}$$

$g(E)$

N -particle density of states

$$\int P(E) dE = 1$$

$$Z = \int g(E) e^{-\beta E} dE$$

"Rough" derivation: for macroscopic systems E is sharply distributed about \tilde{E}

$$Z \approx g(\tilde{E}) e^{-\beta \tilde{E}} \delta E$$



$$\ln Z \approx \ln g(\tilde{E}) - \beta \tilde{E} + \ln \delta E$$

$$|\ln \delta E| \sim -(\ln \sqrt{N})$$

$$-kT \ln Z \approx -kT \ln g(\tilde{E}) + \tilde{E} \approx E - T \cdot k \ln g(\tilde{E})$$

$$N \rightarrow \infty \quad E \approx \langle E \rangle \approx \tilde{E}$$



this is the reason why thermodynamics works!

$$F \equiv -kT \ln Z$$

Helmholtz free energy

$$S = k \ln g(\tilde{E})$$

entropy
(as we had)

$$F = E - TS$$

$$E = E(S, V, N)$$

$$dE = T dS - P dV + \mu dN$$

$$\Rightarrow F = F(T, V, N) \quad dF = -S dT - P dV + \mu dN$$

$$S = -\left(\frac{\partial F}{\partial T}\right)_{V, N}$$

$$P = -\left(\frac{\partial F}{\partial V}\right)_{T, N}$$

$$\mu = \left(\frac{\partial F}{\partial N}\right)_{T, V}$$

More "rigorous" derivation: (essentially the same approx. used for $P(E)$ - "saddle point method")

$$Z = \int d\Gamma e^{-\beta \mathcal{H}(\Gamma)} = \int \frac{d\Gamma}{dE} dE e^{-\beta E} = \int dE g(E) e^{-\beta E}$$

$$= \int dE e^{\ln(g(E)) - \beta E} \approx \int_{-\infty}^{+\infty} dE e^{\ln g(\tilde{E}) - \beta \tilde{E} - \frac{(E - \tilde{E})^2}{2KT C_V}}$$

$$= e^{\ln g(\tilde{E}) - \beta \tilde{E}} \int_{-\infty}^{+\infty} dE e^{-\frac{(E - \tilde{E})^2}{2KT C_V}} = e^{\ln g(\tilde{E}) - \beta \tilde{E}} \sqrt{2\pi KT C_V}$$

$$-KT \ln Z \approx -KT \left(\ln g(\tilde{E}) - \frac{1}{KT} \tilde{E} \right) - KT \ln \sqrt{2\pi KT C_V}$$

\tilde{E} is the equilibrium energy, extensive $\sim N$

$k \ln g(\tilde{E})$ is the entropy S , extensive $\sim N$

C_V is the spec. heat, extensive $\sim N$

thus $\ln C_V \sim \ln N$ can be dropped in the $N \rightarrow \infty$ limit compared to N

$$\underbrace{-KT \ln Z}_{\text{Helmholtz free energy}} \approx \underbrace{\tilde{E}}_{E} - T \underbrace{k \ln g(\tilde{E})}_{S \text{ entropy}}$$

Helmholtz free energy

S entropy

$$F = E - TS$$

where used $E = \tilde{E}$ thermodynamic notation

Example:

Ideal Gas

$$\mathcal{H}(p, q) = \sum_{i=1}^{3N} \frac{p_i^2}{2m}$$

3-dim, N particles

$$Z_N = \int d\Gamma e^{-\beta \mathcal{H}(p, q)} = \int \frac{d^3p d^3q}{h^{3N} N!} e^{-\beta \sum_{i=1}^{3N} \frac{p_i^2}{2m}} = \frac{1}{N!} Z_1^N$$

$$\text{where } Z_1 = \int \frac{d^3p d^3q}{h^3} e^{-\beta \frac{p_x^2 + p_y^2 + p_z^2}{2m}} = \frac{V}{h^3} \left[\int_{-\infty}^{+\infty} dp_x e^{-\frac{p_x^2}{2mkT}} \right]^3$$

$$= \frac{V}{h^3} \left(\sqrt{2m\pi kT} \right)^3 = \frac{V}{h^3} (2m\pi kT)^{3/2} = \frac{V}{\lambda^3}$$

$$Z_N = \frac{Z_1^N}{N!} = \frac{V^N}{h^{3N} N!} (2m\pi kT)^{\frac{3N}{2}} = \boxed{\frac{V^N}{N!} \left(\frac{2m\pi kT}{h^2} \right)^{\frac{3N}{2}}}$$

$$\left[\text{remember } \lambda = \sqrt{\frac{h^2}{2\pi m kT}} \right]$$

partition function of
the ideal gas in the
canonical ensemble

Thermodynamics of the ideal gas from the canonical ensemble:

$$F(T, V, N) = -kT \ln Z_N = -NkT \ln \left(V \left(\frac{2\pi m kT}{h^2} \right)^{3/2} \right) + kT \ln N!$$

$$\approx NkT \ln \left[\frac{1}{V} \left(\frac{h^2}{2\pi m kT} \right)^{3/2} \right] + kT (N \ln N - N) =$$

$$- NkT \left[\ln \left(\frac{N}{V} \left(\frac{h^2}{2\pi m kT} \right)^{3/2} \right) - 1 \right]$$

$$\mu = \left(\frac{\partial F}{\partial N} \right)_{T, V} = kT \ln \left[\frac{N}{V} \left(\frac{h^2}{2\pi m kT} \right)^{3/2} \right] = \boxed{kT \ln \left(\frac{\lambda^3}{V/N} \right)}$$

$$\lambda = \sqrt{\frac{h^2}{2\pi m kT}} \quad \text{Thermal wavelength}$$

$$P = - \left(\frac{\partial F}{\partial V} \right)_{T, N} = \frac{NkT}{V} \quad \Rightarrow \quad \boxed{PV = NkT}$$

$$S = - \left(\frac{\partial F}{\partial T} \right)_{V, N} = -Nk \left[\ln \frac{N}{V} \left(\frac{h^2}{2\pi m kT} \right)^{3/2} - 1 \right] + \frac{3}{2} Nk =$$

$$= Nk \left[\ln \left\{ \frac{V}{N} \left(\frac{2\pi m kT}{h^2} \right)^{3/2} \right\} + \frac{5}{2} \right]$$

$$F = E - TS \quad \Rightarrow$$

$$E = F + TS = \frac{3}{2} NkT$$

$$kT = \frac{2E}{3N}$$

$$\boxed{S(E, V, N) = Nk \ln \left(\frac{V}{N} \right) + \frac{3}{2} Nk \ln \left(\frac{E}{N} \right) + \frac{5}{2} Nk + \frac{3}{2} Nk \ln \left(\frac{4\pi m T}{3h^2} \right)}$$