

1. Calculating the matrix elements  $\langle \mathbf{x}_1 \mathbf{x}_2 \dots \mathbf{x}_N | e^{-\beta H} | \mathbf{x}_1 \mathbf{x}_2 \dots \mathbf{x}_N \rangle$ , in class we formally obtained the partition function in the canonical ensemble for an  $N$ -particle non-interacting non-relativistic quantum system

$$Z_N = \text{Tr}(e^{-\beta H}) = \int d^{3N} \mathbf{x} \langle \mathbf{x}_1 \mathbf{x}_2 \dots \mathbf{x}_N | e^{-\beta H} | \mathbf{x}_1 \mathbf{x}_2 \dots \mathbf{x}_N \rangle,$$

where

$$\langle \mathbf{x}_1 \mathbf{x}_2 \dots \mathbf{x}_N | e^{-\beta H} | \mathbf{x}_1 \mathbf{x}_2 \dots \mathbf{x}_N \rangle = \frac{1}{N! \lambda^{3N}} \sum_P \delta_P f(P\mathbf{x}_1 - \mathbf{x}_1) f(P\mathbf{x}_2 - \mathbf{x}_2) \dots f(P\mathbf{x}_N - \mathbf{x}_N)$$

$$f(u) = e^{-\frac{\pi u^2}{\lambda^2}}, \quad \lambda = \left( \frac{h^2}{2\pi m k T} \right)^{1/2}, \quad \delta_P \equiv 1 \text{ for Bosons and } \delta_P = (-1)^{[P]} \text{ for Fermions.}$$

In order to study lowest-order quantum corrections to classical non-interacting systems ( $\lambda^3 \left( \frac{N}{V} \right) \ll 1$ ), it is sufficient to consider only the trivial permutation (“no permutation”) and two-particle permutations in the above expression, i.e.,

$$\langle \mathbf{x}_1 \mathbf{x}_2 \dots \mathbf{x}_N | e^{-\beta H} | \mathbf{x}_1 \mathbf{x}_2 \dots \mathbf{x}_N \rangle = \frac{1}{N! \lambda^{3N}} \left\{ 1 \pm \sum_{i < j} f^2(r_{ij}) + \dots \right\}$$

where  $r_{ij} = |\mathbf{x}_i - \mathbf{x}_j|$ , and the  $\pm$  sign are for Bosons/Fermions).

(a) Find the equation of state in this first order approximation.

(b) Find  $\langle E \rangle$  as a function of  $T, V$ , and  $N$  in the same approximation.

2. We defined in class the Fermi-Dirac (+) and Bose-Einstein (-) integrals

$$f_\nu^\pm(z) = \frac{1}{\Gamma(\nu)} \int_0^\infty \frac{dx x^{\nu-1}}{z^{-1} e^x \pm 1}.$$

Prove that for  $z \ll 1$

$$f_\nu^\pm = \sum_{l=1}^{\infty} (\mp 1)^{l-1} \frac{z^l}{l^\nu} = z \mp \frac{z^2}{2^\nu} + \frac{z^3}{3^\nu} \mp \dots$$

3. Working in the grand-canonical ensemble, we obtained in class for the ideal non-relativistic Fermi (+) and Bose (-) systems with spin  $s$  :

$$N = (2s + 1) \frac{V}{\lambda^3} f_{3/2}^{\pm}(z),$$

$$E = \frac{3}{2} kT (2s + 1) \frac{V}{\lambda^3} f_{5/2}^{\pm}(z),$$

where  $z = e^{\mu/kT}$  and  $\lambda = \left( \frac{h^2}{2\pi m kT} \right)^{1/2}$ .

(a) Obtain the energy  $E$  as a function of  $T, V, N$ , and the equation of state up to *first order* in  $\frac{\lambda^3 N}{V}$ . Note that after using the small  $z$  approximation for  $f_{\nu}^{\pm}(z)$ ,

$z$  must be eliminated from the equations in favor of  $\frac{\lambda^3 N}{V}$ . To this end you should

use the first equation above for  $N$ . How do your final results compare with those of Problem 1.(a) and (b)?

(b) Using  $E(T, V, N)$ , obtain the specific heat of the quantum gas up to the same order in  $\frac{\lambda^3 N}{V}$ .

4. Consider He gas at room temperature and atmospheric pressure, and determine whether the classical approximation for the equation of state ( $PV = NkT$ ) is justified or not. Repeat the above considerations for the electron “gas” at room temperature.