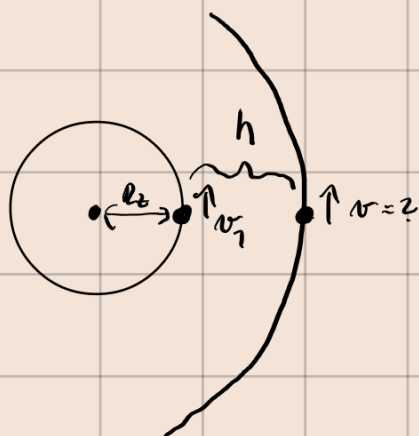


Gravitační pole

9.7 Geostacionární družice



$$\omega = \omega_{\text{země}}$$

$$h = ?$$

$$r = ?$$

$$F_d = F_G$$

$$\cancel{m} \cdot \omega^2 r = \cancel{G} \cdot \frac{\cancel{m} \cdot M_z}{r^2}$$

↑
Gravitační
konstanta

$$\pi^3 \frac{4\pi^2}{T_z^2} = G \cdot M_z$$

↓
 $T_z = 1 \text{ den}$

$$r = R_z + h$$

$$\Rightarrow r = \sqrt[3]{\frac{G \cdot M_z \cdot T_z^2}{4\pi^2}}$$

$$h = \sqrt[3]{\frac{G \cdot M_z \cdot T_z^2}{4\pi^2}} - R_z$$

$$h = 35\,833 \text{ km}$$

$$v = \omega \cdot r$$

$$v = \frac{2\pi}{T_z} \cdot \left(G \cdot M_z \right)^{1/3} \cdot \frac{T_z^{1/3}}{(2\pi)^{2/3}}$$

$$v = \frac{2^{1/3} \pi^{1/3}}{T_z^{1/3}} \left(G M_z \right)^{1/3}$$

$$v = \sqrt[3]{\frac{2\pi G M_z}{T_z}} = 3 \text{ km/s}$$

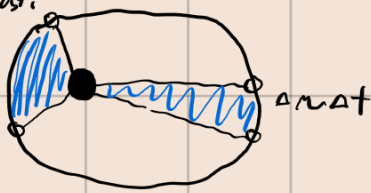
Keplerovy zákony

1. trajektorie planet jsou elipsy, v ohnisku je slunce

2. $\omega = \frac{dS}{dt} = \frac{d\left(\frac{1}{2} |\vec{r} \times d\vec{r}|\right)}{dt} = \frac{1}{2} |\vec{r} \times \frac{d\vec{r}}{dt}| = \frac{1}{2} |\vec{r} \times \underbrace{\left(\frac{\vec{v}}{m}\right)}_{\vec{p}}| = \frac{|\vec{L}|}{2m} = \text{konst.}$

2. zákon zachování \vec{p}

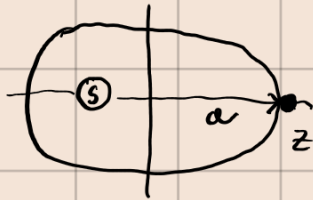
momentu hybnosti



$$\vec{r} \cdot \vec{v} \cdot m = L$$

$$\frac{dL}{dt} = \frac{d(\vec{r} \times m\vec{v})}{dt} = \frac{d\vec{r}}{dt} \times m\vec{v}$$

3.



$$\frac{T^2}{a^3} = \text{konst.} \quad (\text{pro stejným orbitální systém})$$

upgrade 3. kř pro kružnicovou trajektorii

$$m \omega^2 r = \mathcal{L} \frac{m \cdot M}{r^2}$$

$$\frac{4\pi^2 r^3}{T^2} = \mathcal{L} \cdot M$$

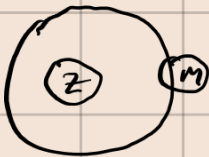
$$\text{konst.} = \frac{r^3}{T^2 M_0} = \frac{\mathcal{L}}{4\pi^3}$$

9.9



$$a_G = 1070000 \text{ km}$$

$$T_G = 7,15 \text{ dne} = 846100 \text{ s}$$



$$a_M = 384400 \text{ km}$$

$$T_M = 27,32 \text{ dne}$$

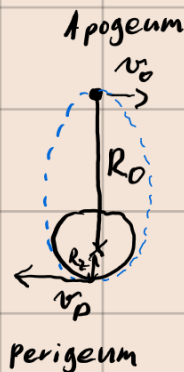
$$\frac{M_j}{M_z} = ?$$

$$\frac{T_G^2 \cdot M_j}{a_G^3} = \frac{T_M^2}{a_M^3} \cdot M_z$$

$$\frac{M_j}{M_z} = \frac{T_M^2}{T_G^2} = \frac{a_G^3}{a_M^3} = 374$$

9.4 - úkol

9.6



1) $v_0 = ? \rightarrow$ traj. kružnice

2) $v_0 = ? \rightarrow$ těsně minul zemi

$$F_d = F_G$$

$$m \cdot \frac{v^2}{R_0} = \mathcal{L} \frac{m \cdot M_Z}{R_0^2}$$

$$v = \sqrt{\frac{\mathcal{L} M_Z}{R_0}}$$

$$\downarrow$$

$$R_0 = R_Z$$

1. kosmická rychlost

$$v_I = \sqrt{\frac{\mathcal{L} M_Z}{R_Z}}$$

1) $L = \text{konst.}$

$$m \cdot v = \text{konst.}$$

$$R_0 \cdot m \cdot v_0 = R_Z \cdot m \cdot v_p$$

$$\boxed{R_0 \cdot v_0 = R_Z \cdot v_p}$$

vice

$$E_{\text{pot}} = - \int \vec{F}_G \cdot d\vec{r} = - \int F_G dr$$

$$= + \int \mathcal{L} \frac{m M_Z}{r^2} dr = - \mathcal{L} \frac{m \cdot M_Z}{r}$$

$$E = E_{\text{kin}} + E_{\text{pot}}$$

$$= \frac{1}{2} m v^2 - \mathcal{L} \cdot \frac{m \cdot M_Z}{r}$$

$$v_{II} = \sqrt{\frac{2 \mathcal{L} M_Z}{R_Z}} =$$

2. kosmická (úniková) rychlost

získan zachování mech. energie

$$\left| \frac{1}{2} m v_0^2 - \frac{\mathcal{L} m M_Z}{R_0} = \frac{1}{2} m v_p^2 - \frac{\mathcal{L} m M_Z}{R} \right|$$

2. vce

$$v_p^2 - v_0^2 = \frac{2 \cdot \mathcal{E} \cdot m_z}{R_z} - \frac{2 \cdot \mathcal{E} \cdot m_z}{R_0} = 2 \cdot \mathcal{E} \cdot m_z \cdot \left(\frac{1}{R_z} - \frac{1}{R_0} \right) = 2 \cdot \mathcal{E} \cdot m_z \cdot \left(\frac{R_0 - R_z}{R_z \cdot R_0} \right)$$

$$\left(\frac{R_0^2}{R_z^2} - 1 \right) v_0^2 = \left(\frac{R_0^2 - R_z^2}{R_z^2} \right) \cdot v_0^2 = 2 \cdot \mathcal{E} \cdot m_z \cdot \left(\frac{R_0 - R_z}{R_z \cdot R_0} \right)$$

$$v_0^2 = 2 \cdot \mathcal{E} \cdot m_z \cdot \left(\frac{R_0 - R_z}{R_z \cdot R_0} \right) \cdot \left(\frac{R_z^2}{R_0^2 - R_z^2} \right)$$

$$v_0 = \sqrt{2 \cdot \mathcal{E} \cdot m_z \cdot \frac{R_z}{R_0(R_0 + R_z)}}$$

Soustava hmotných bodů:

IMPULZOVÉ VĚTKY

$$1. IV: \vec{I} = \Delta \vec{p} = \int \vec{F} dt$$

$$\frac{d \Sigma \vec{p}}{dt} = \Sigma \vec{F} = \underbrace{\Sigma \vec{F}^{(i)}}_{=0 \text{ díky 3. N. Z.}} + \Sigma \vec{F}^{(e)}$$

$$\frac{d \Sigma \vec{p}}{dt} = \left(\Sigma \vec{F}^{(e)} \right) = 0$$

izolovaná soustava

$$\frac{d \Sigma \vec{p}}{dt} = 0 \quad \text{zákon zachování hybnosti}$$

$$\Sigma \vec{p} = \text{konst.}$$

2. I.V.

$$\vec{M} = \vec{r} \times \vec{F}$$

$$\frac{d \Sigma \vec{L}}{dt} = \Sigma \vec{M}^{(e)}$$

moment síly

$$\vec{L} = \vec{r} \times \vec{p}$$

moment hybnosti

$2 \rightarrow M_H$

$$\frac{d\vec{L}}{dt} = 0 \rightarrow \vec{L} = \text{konst.}$$