

Modeling Processes

Case Study of a Dust Storm

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Rajesh Kumar (NCAR)

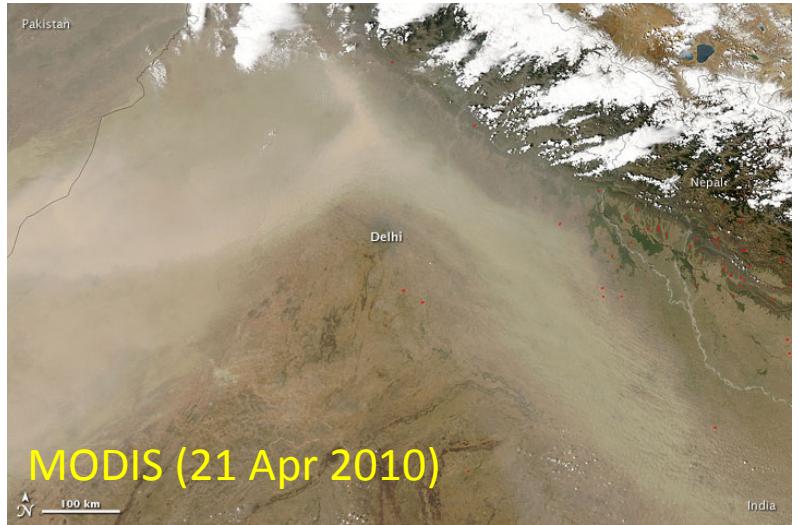
Kumar et al. (2014) *Atmos. Chem. Phys.*, Dust effects on radiation
and aerosol optical properties

Kumar et al. (2014) *Atmos. Chem. Phys.*, Dust effects on chemistry

Exercises will be on these topics:

1. Calculation of photodissociation rate constants
2. Effects of aerosols on photodissociation rate constants

Why Model Dust Storms and Chemistry?

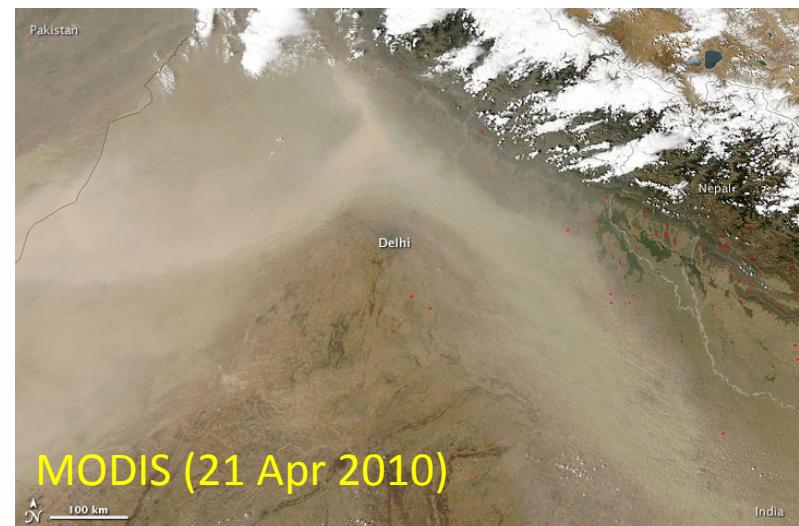


- Dust storms often occur during the pre-monsoon (MAM) season in northern India and affect day to day life.
- **GOAL:** Understand the effect of these dust storms on regional scale aerosol optical properties, radiation budget and tropospheric chemistry.

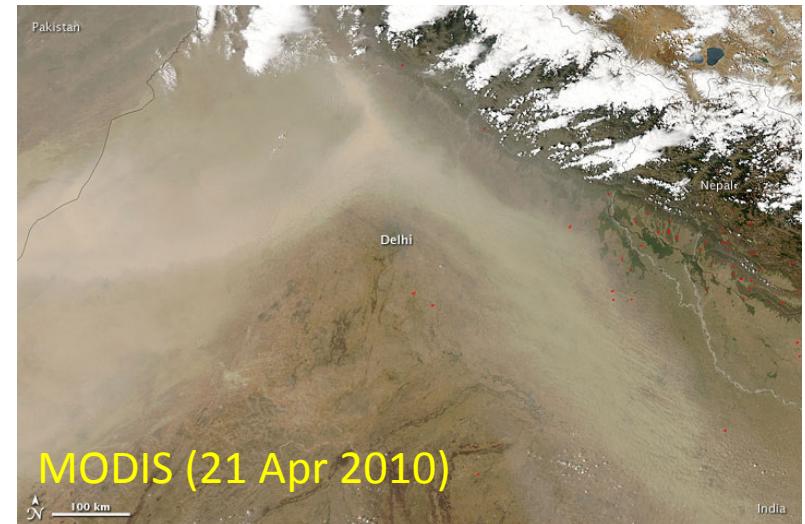
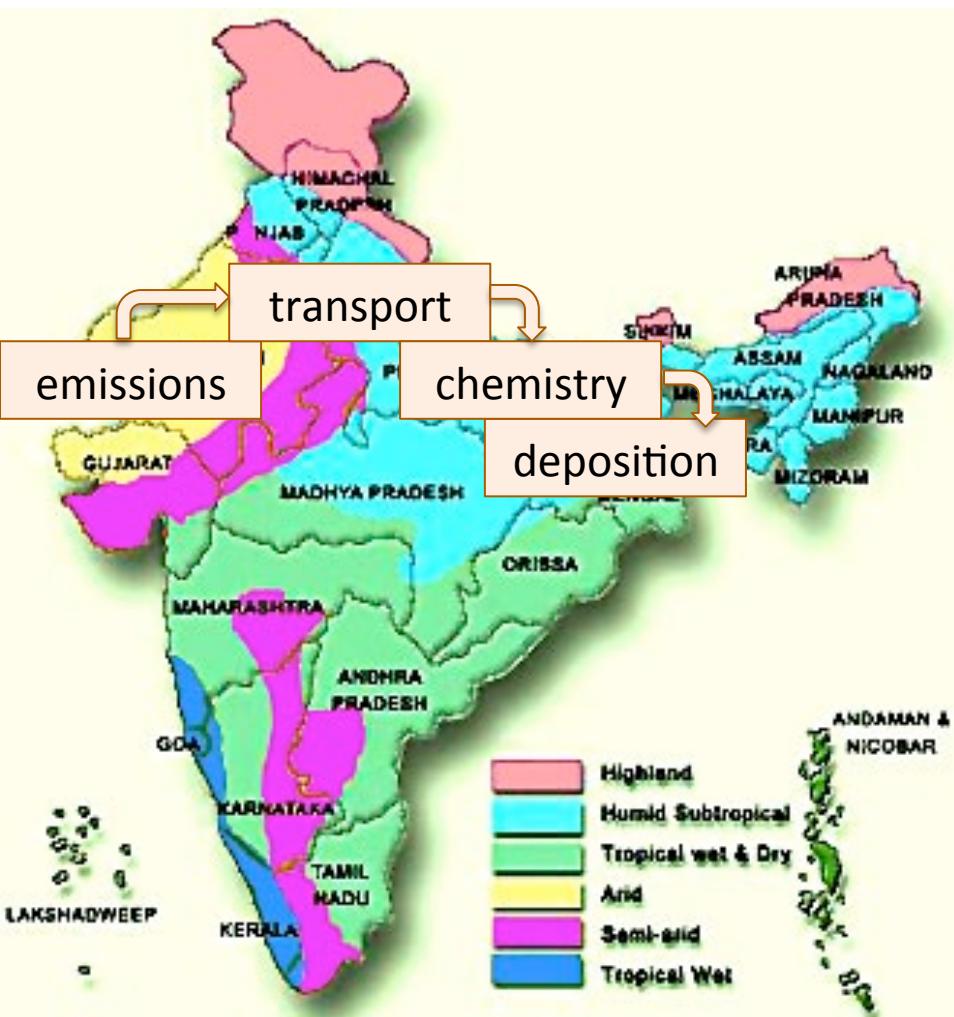
What Causes the Dust to Reside over the IGP?



Dust from the Thar Desert gets channeled by the topography of the Indo-Gangetic Plain



What Processes Happen from the Desert to the IGP?



Modeling dust storm effects on aerosols and trace gases

- Emissions
 - Dust emissions = $f(\text{wind, soil type \& moisture})$
 - Sea salt emissions = $f(\text{wind})$
 - Anthropogenic emissions = prescribed
 - Biomass burning emissions = $f(\text{fire size, vegetation})$
 - Emissions from vegetation = $f(\text{vegetation type, T, PAR})$
- Transport
- Chemistry
- Deposition

Modeling emissions of aerosols and trace gases

- Dust Emissions

$$Dust_{emis} = C (f_{size} \times erod \times area) (wspd_{10m})^2 (wspd_{10m} - u_{thres}) dt$$

C = tuning factor

→ Improving dust emissions for different deserts is important

- Sea Salt Emissions

$$SS_{emis} = 4/3(\pi (r_{dry})^3 \rho_{SS} frh dF_n dr) dt$$

These equations are from the WRF-Chem GOCART emissions modules. They can easily vary among models. References are Ginoux et al. (2001, 2004); Chin et al. (2002).

Modeling emissions of aerosols and trace gases

- Anthropogenic Emissions
 - Several emissions inventories available
 - See ECCAD web site eccad.sedoo.fr/ and Sachin Ghude's lecture
- Biomass Burning Emissions

GFED www.globalfiredata.org/

QFED http://gmao.gsfc.nasa.gov/research/science_snapshots/global_fire_emissions.php

FINN <https://www2.acd.ucar.edu/modeling/finn-fire-inventory-ncar>

Forecast <http://www.acd.ucar.edu/acresp/forecast/fire-emissions.shtml>

- Biogenic Emissions

MEGAN

<https://www2.acd.ucar.edu/modeling/model-emissions-gases-and-aerosols-nature-megan>

BEIS

<http://www.epa.gov/ttn/chief/emch/biogenic/beis/index.html>

Modeling transport of aerosols and trace gases

- Emissions
- Transport
 - Resolved on grid of model
 - Parameterized motions in the boundary layer (i.e. diffusivity to represent large eddy motions)
 - Parameterization of convective transport
- Chemistry
- Deposition

Modeling transport of aerosols and trace gases

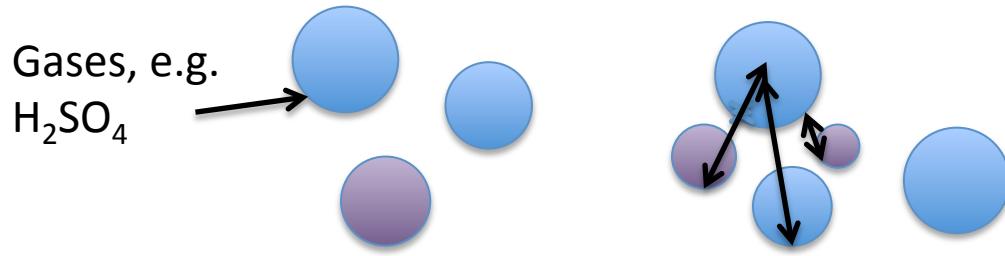
- Transport
 - See lectures by Mark Lawrence and Federico Fierli

Modeling chemistry of aerosols and trace gases

- Emissions
- Transport
- Chemistry
 - Aerosol growth by condensation and coagulation
 - Photodissociation reactions
 - Reactions between trace gases
 - Reactions between gas and aerosol
 - Reactions in cloud and rain drops
- Deposition

Modeling aerosol physics and chemistry

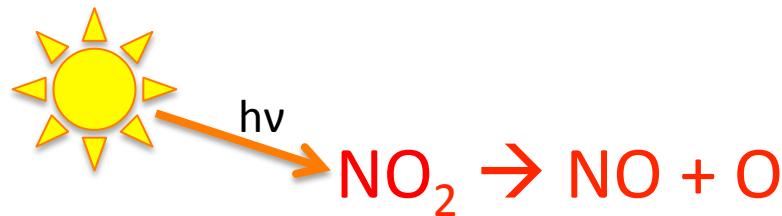
- Aerosol growth by condensation and coagulation



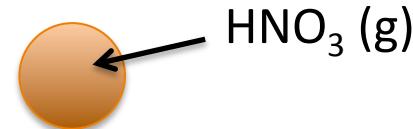
- Photodissociation reactions
- Reactions between trace gases
- Reactions between gas and aerosol
- Reactions in cloud and rain drops

Modeling chemistry of aerosols and trace gases

- Aerosol growth by condensation and coagulation
- Photodissociation reactions



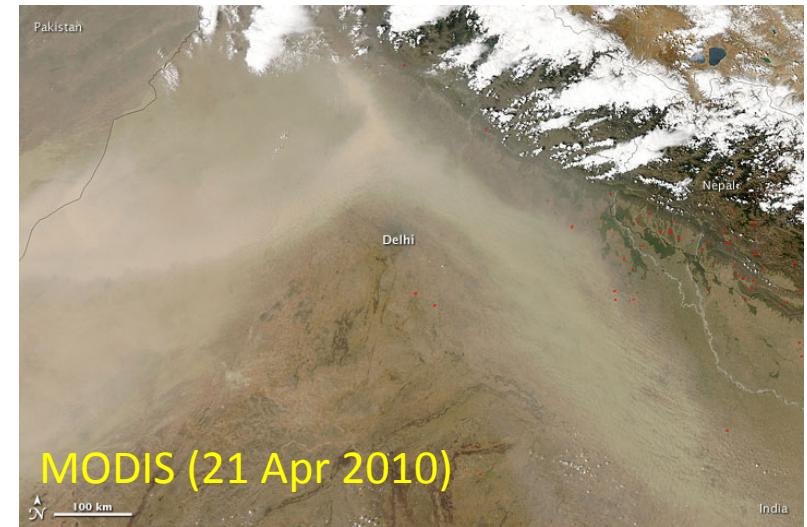
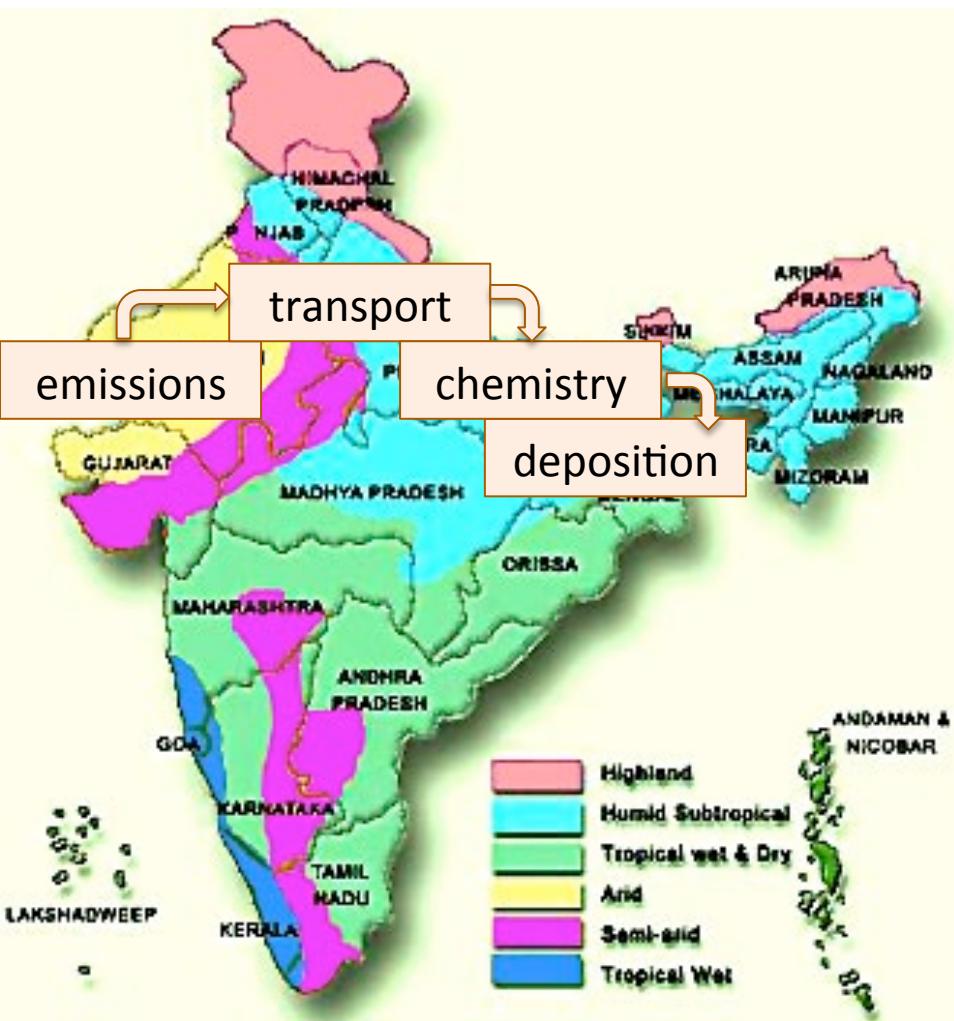
- Reactions between trace gases, e.g. $\text{NO} + \text{O}_3 \rightarrow \text{NO}_2$
- Reactions between gas and aerosol
 - $\text{HNO}_3 \text{ (g)} + \text{dust} \rightarrow 0.5 \text{ NOx}$
- Reactions in cloud and rain drops



Modeling deposition of aerosols and trace gases

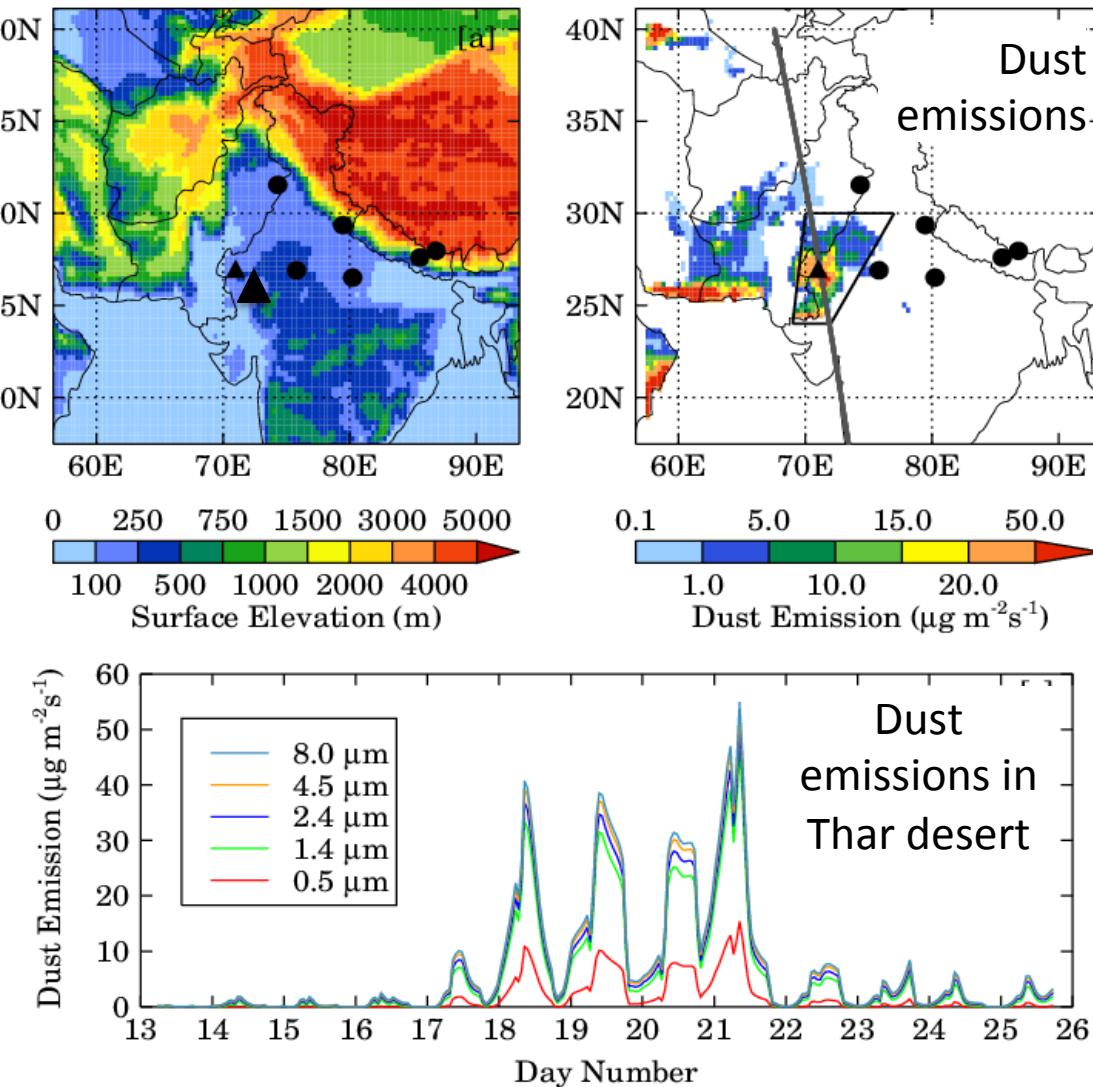
- Dry Deposition
 - Deposition velocity, vegetation (canopy or not), Henry's Law constant
 - Wesely (1989) parameterization often used
- Wet Deposition
 - Amount of cloud water, Henry's Law constant, production of precipitation, evaporation
- Henry's Law (M/atm)
 $[H_2O_2 \text{ (aq)}] = K_H p_{H_2O_2 \text{ (g)}}$
 $K_H = \text{Henry's Law coefficient} = f(\text{temperature})$

Modeling dust storm effects on aerosols and trace gases



Use the Weather Research and Forecasting model coupled with Chemistry (WRF-Chem) to learn what processes affect aerosols and trace gases

WRF-Chem set-up

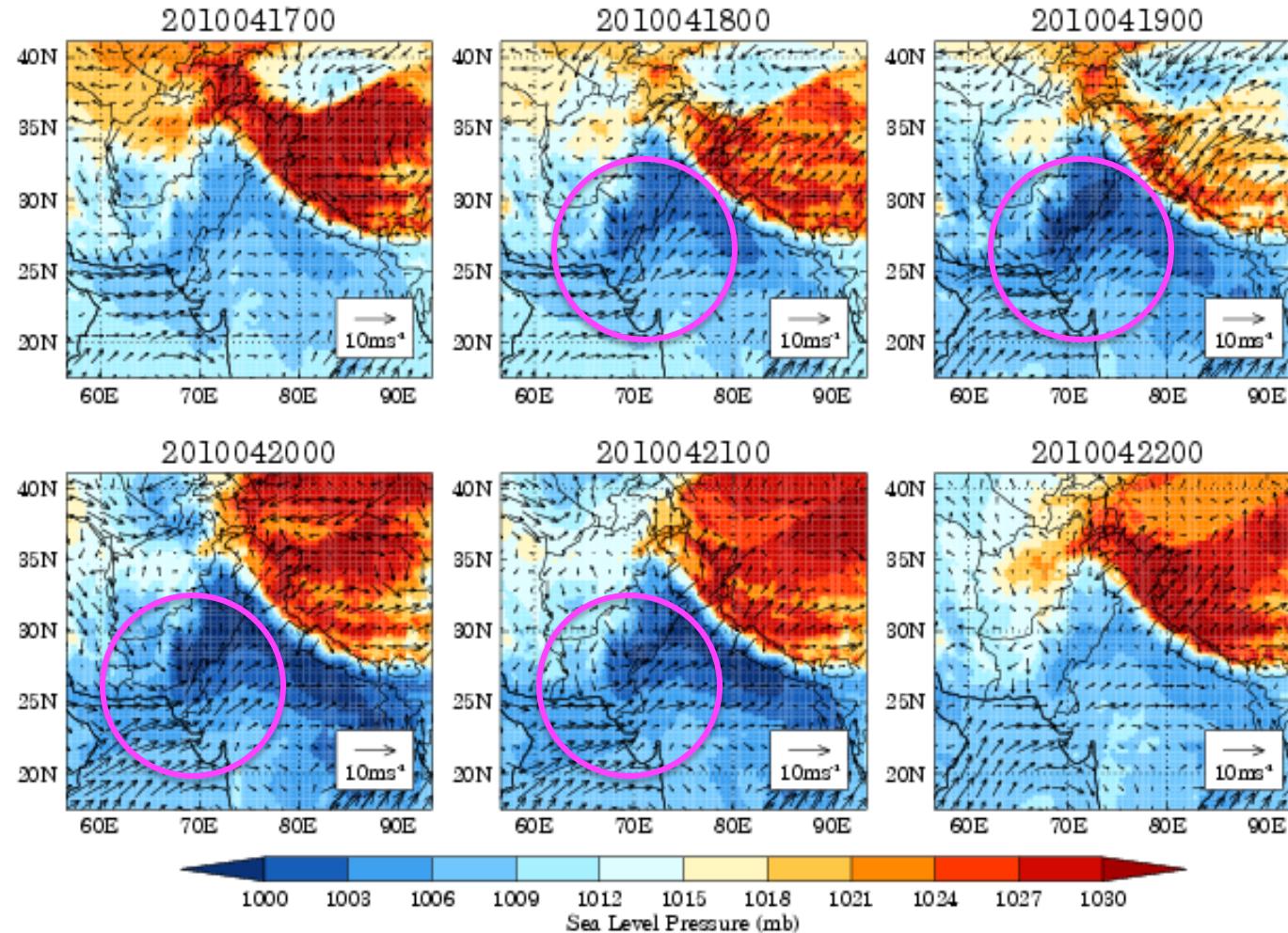


Grid spacing: 30 km
Grid points $(x,y,z) = (120,90,51)$
Simulation period: 10-25 Apr 2010

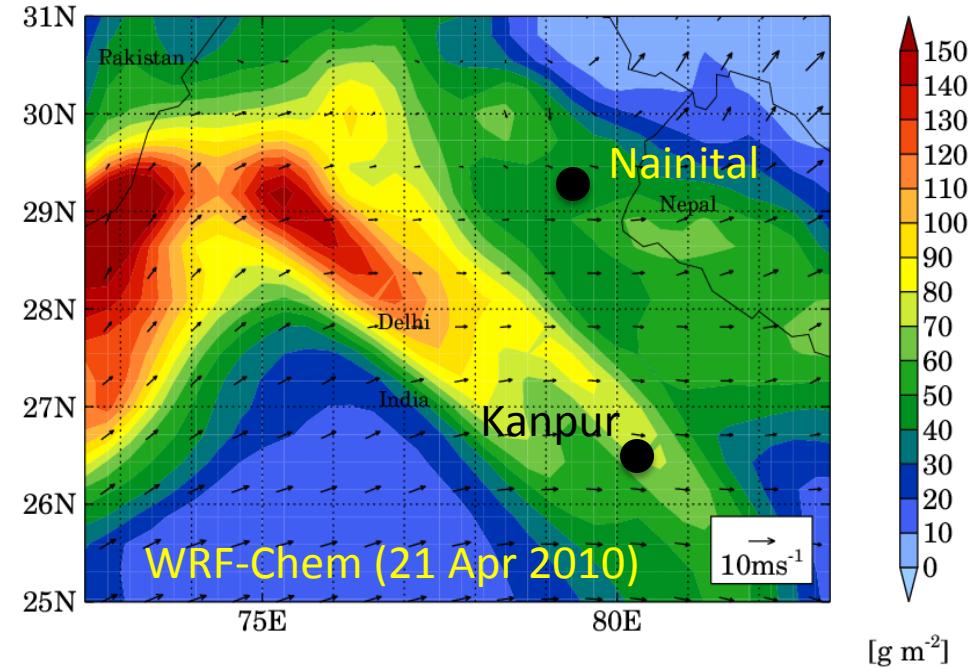
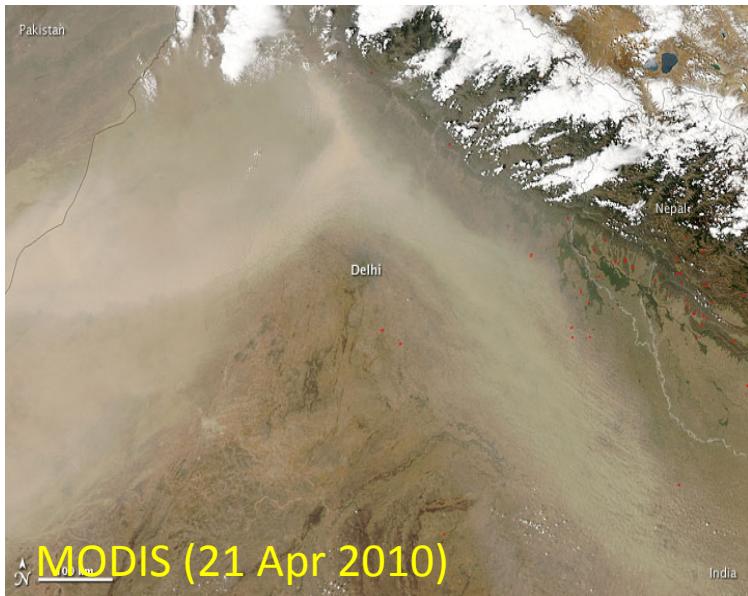
Microphysics: Thompson
Cumulus: Kain-Fritsch
Surface and PBL: MYJ Scheme
Radiation: RRTMG

Chemical Mechanism:
MOZART gas + GOCART aerosols
Photolysis: F-TUV
Anthro Emissions: MACCity
BB Emissions: FINN v1
Biogenic Emissions: MEGAN
Initial and Boundary Conditions:
Meteorology: NCEP FNL
Chemistry: MOZART-4 CTM

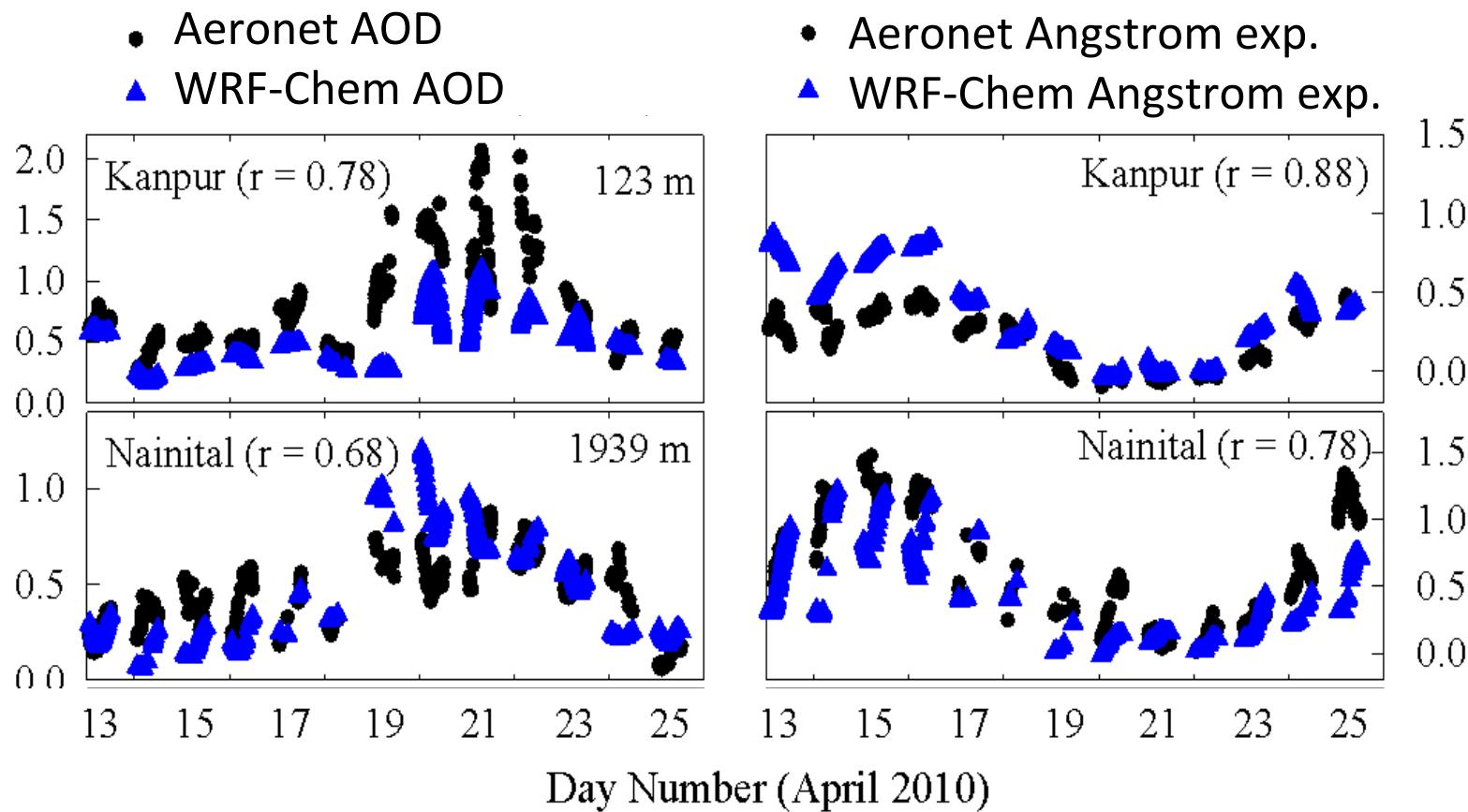
A low pressure region over the Thar Desert generated this dust storm.



WRF-Chem captures spatial distribution of the dust storm



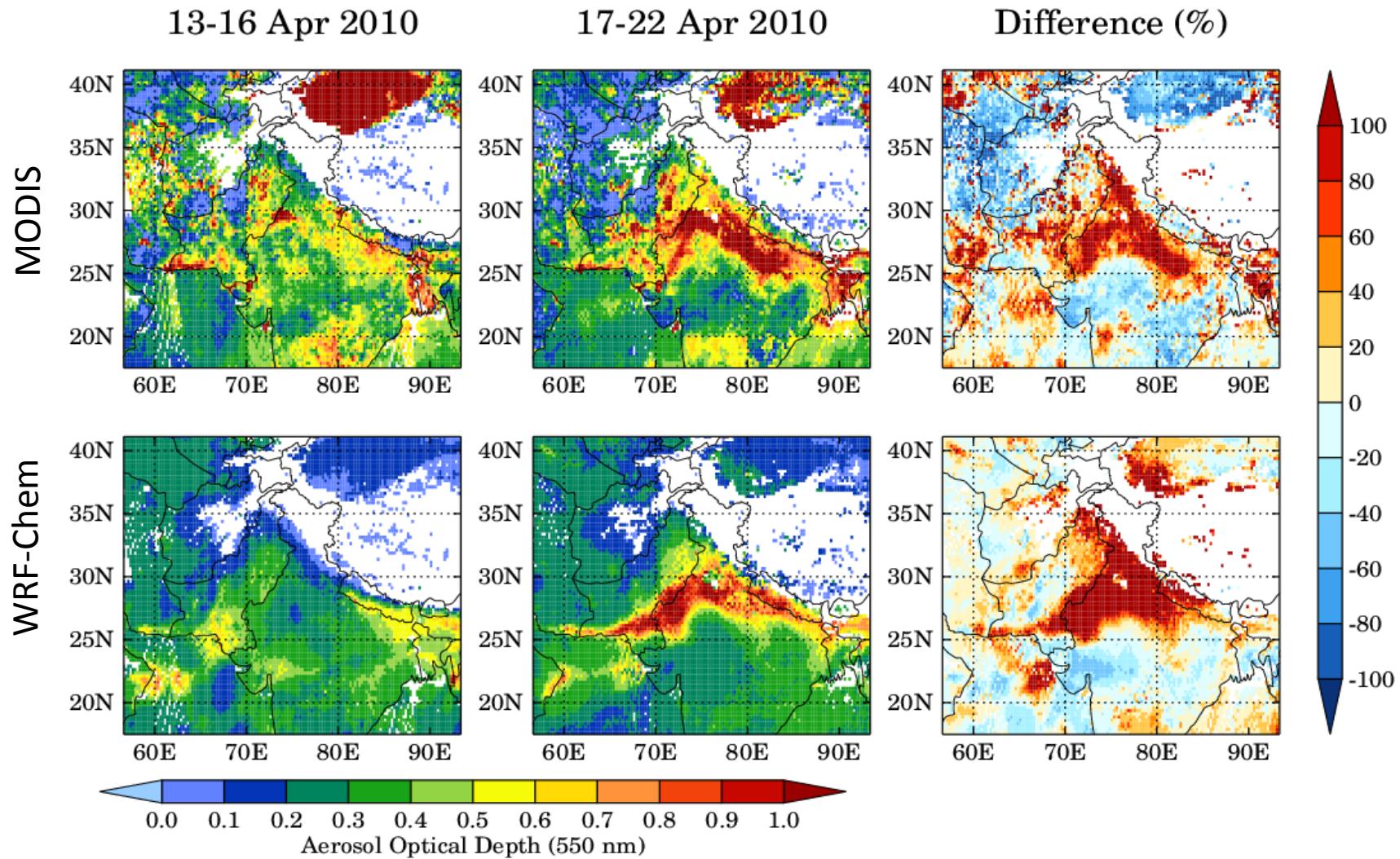
WRF-Chem captures AOD and Angstrom exponent



AOD – integrated extinction coefficient over a vertical column of unit cross section.

Angstrom exponent – inverse relation with aerosol size, smaller for larger aerosols and vice versa.

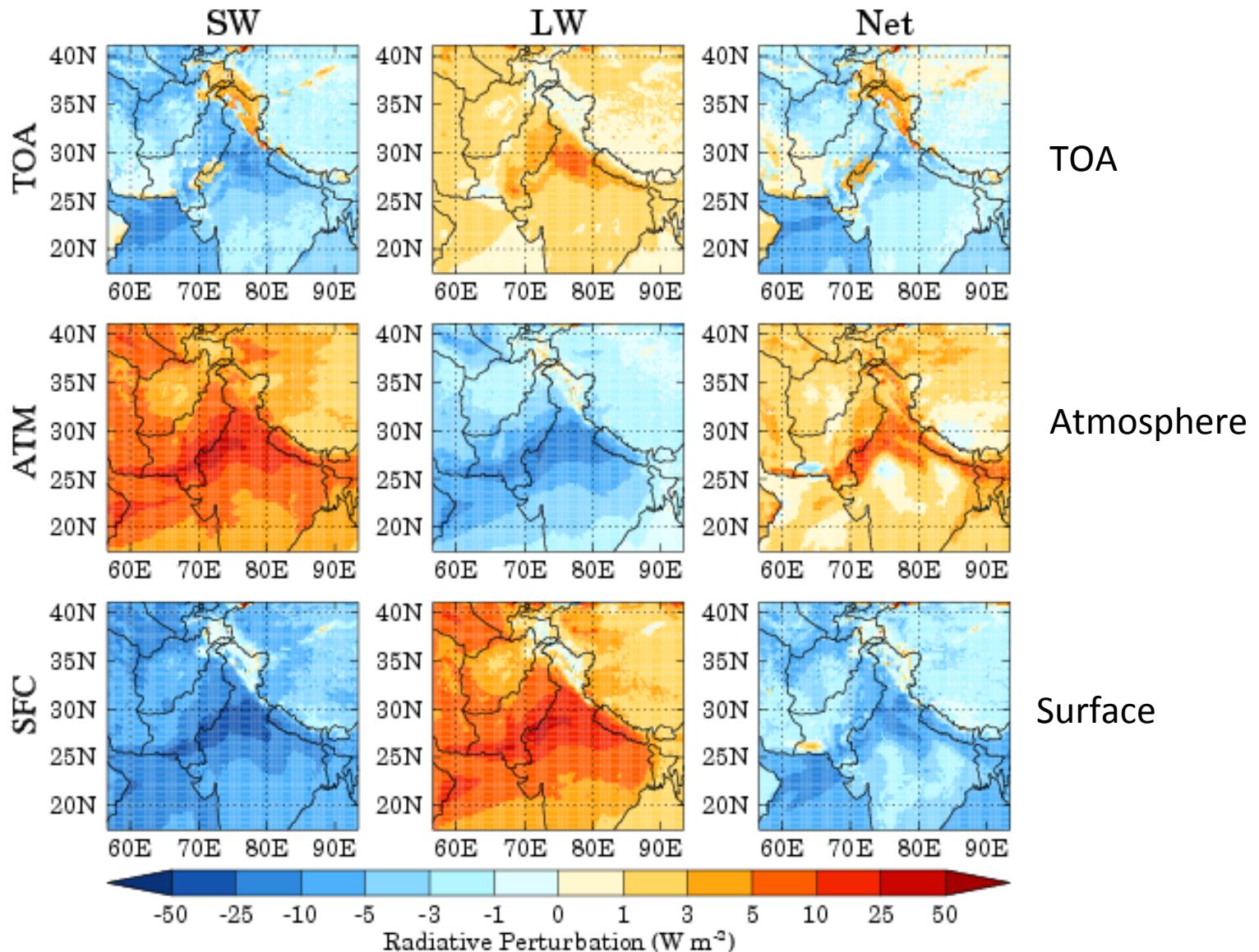
Dust Storm almost doubled the regional aerosol loading



WRF-Chem Sensitivity Simulations

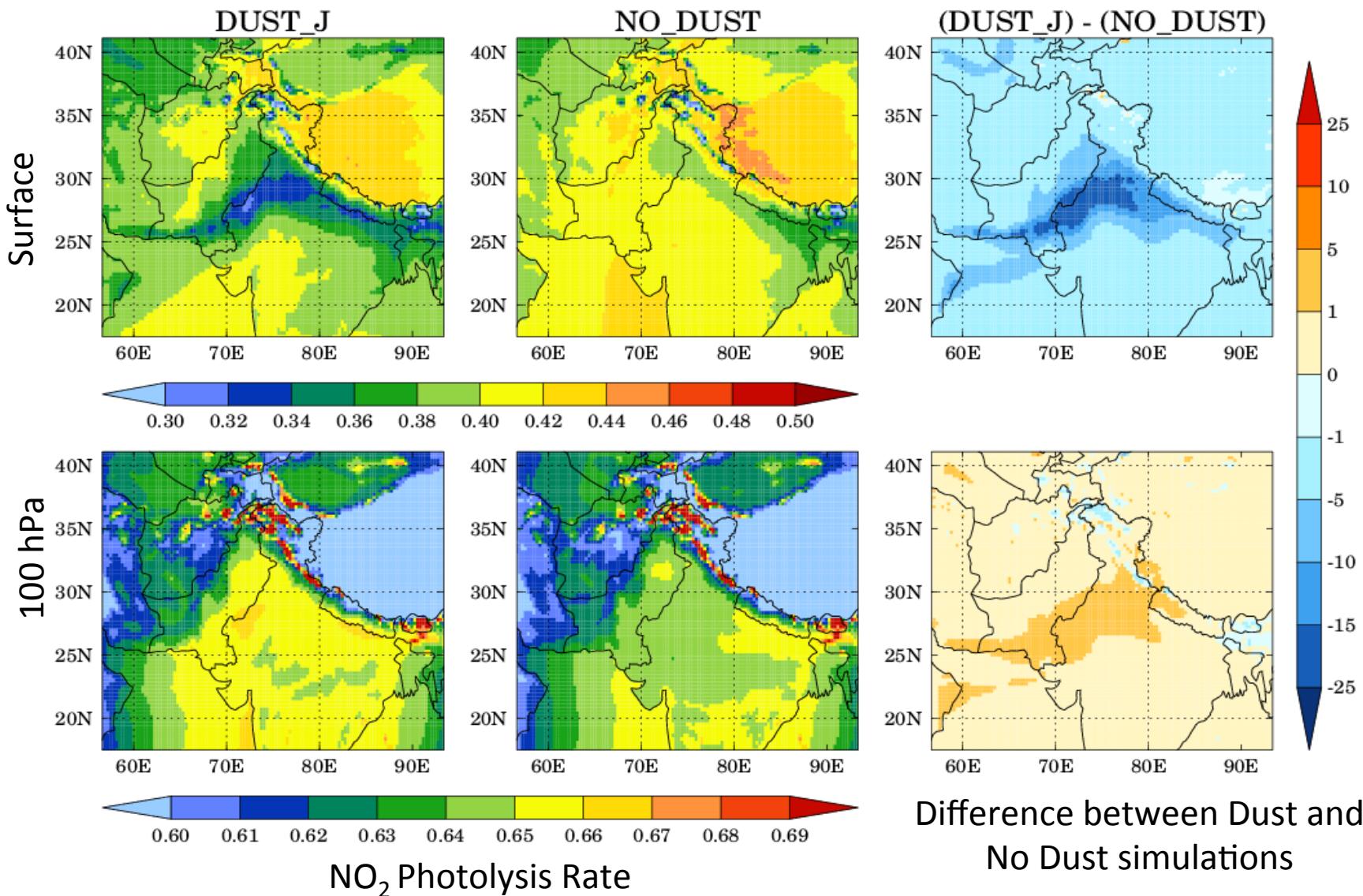
1. Base Case – with Dust emissions and j-values affected = **DUST_J**
2. No Dust emissions Case = **No Dust**

Dust Storm cools the surface and TOA, and warms the atmosphere

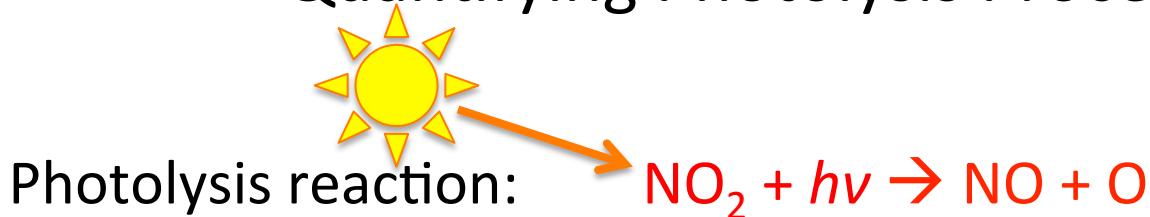


Dust storm decreases photolysis rates at the surface and increases in the upper troposphere

17-22 April 2010



Quantifying Photolysis Processes



Photolysis rates:

$$\frac{d[NO_2]}{dt} \Big|_{hv} = -j[NO_2]$$

$$\frac{d[NO]}{dt} \Big|_{hv} = \frac{d[O]}{dt} \Big|_{hv} = +j[NO_2]$$

Photolysis frequency (s^{-1}) $j = \int_{\lambda} \sigma(\lambda) \phi(\lambda) F(\lambda) d\lambda$

(other names: photo-dissociation rate coefficient, J-value)

Calculation of Photolysis Coefficients

$$J \text{ (s}^{-1}\text{)} = \int \sigma(\lambda) \phi(\lambda) F(\lambda) d\lambda$$

$\sigma(\lambda)$ =absorption cross section, cm² molec⁻¹

-- probability that photon is absorbed

$\phi(\lambda)$ =photodissociation quantum yield, molec quanta⁻¹

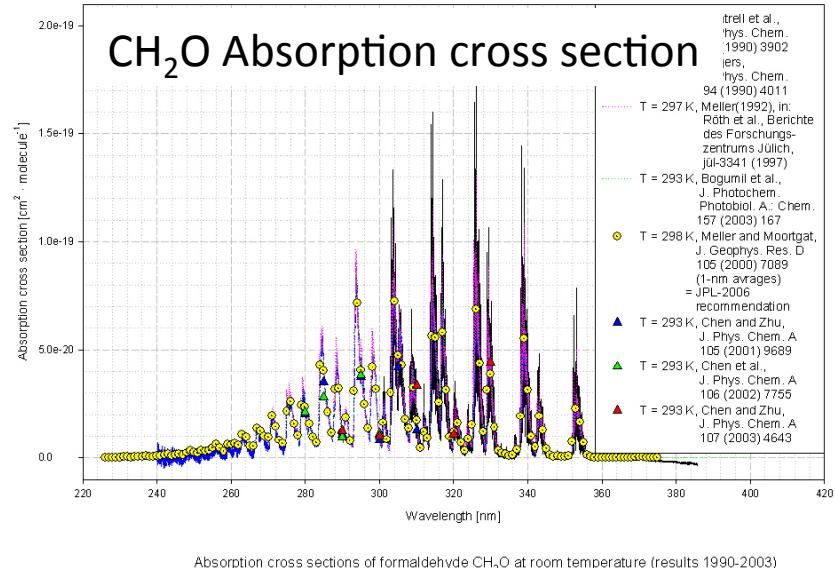
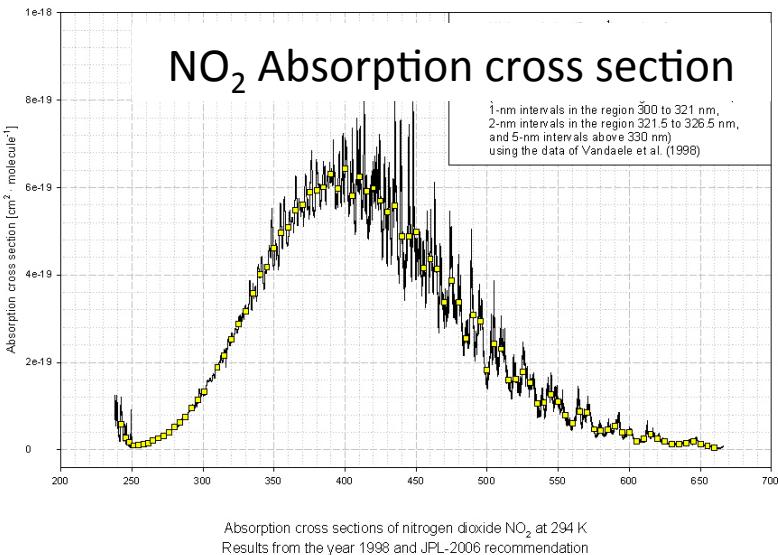
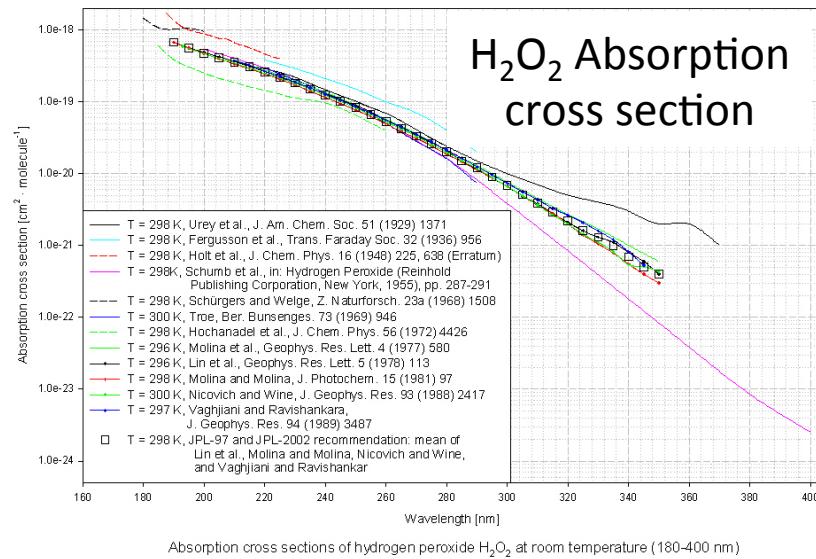
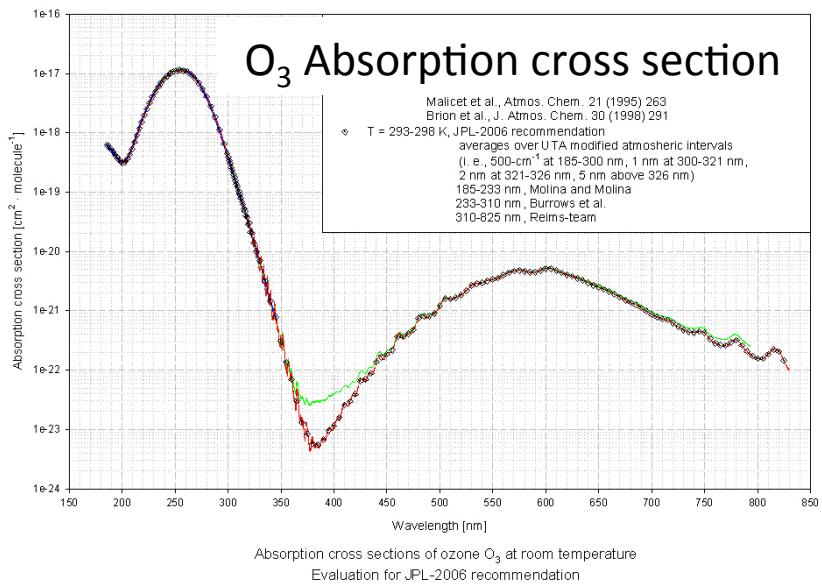
-- probability that absorbed photon causes dissociation

$F(\lambda)$ = spectral actinic flux, quanta cm⁻² s⁻¹ nm⁻¹

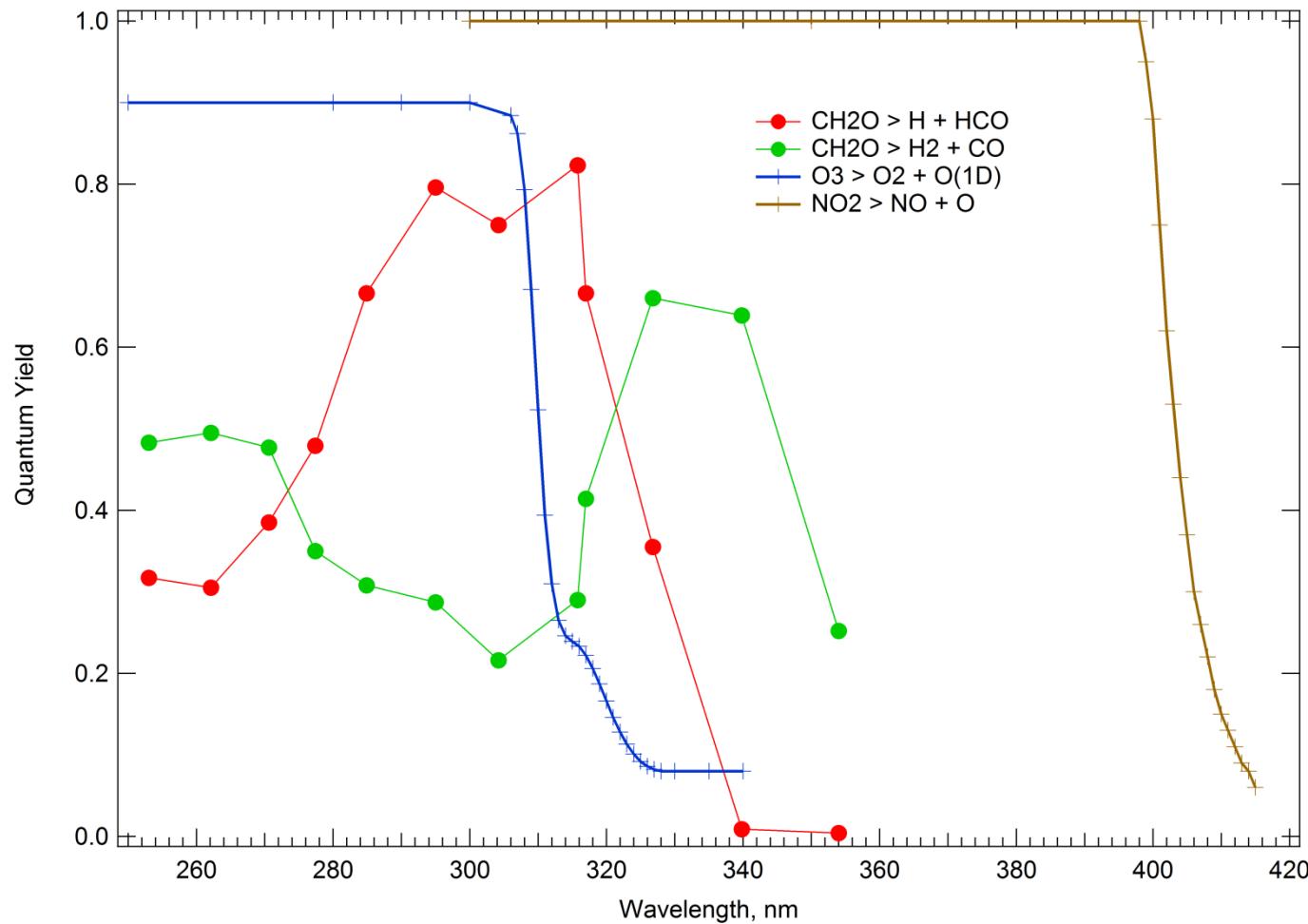
= solar radiation flux onto sphere

-- probability of photon near molecule

Absorption Cross Section Varies with Species and Wavelength



Measured Quantum Yields



Compilations of Cross Sections & Quantum Yields

<http://www.atmosphere.mpg.de/enid/2295>



MPI-Mainz-UV-VIS Spectral Atlas of Gaseous Molecules
A Database of Atmospherically Relevant Species, Including Numerical Data and Graphical Representations
Hannelore Keller-Rudek, Geert K. Moortgat
Max-Planck-Institut für Chemie, Atmospheric Chemistry Division, Mainz, Germany

<http://jpldataeval.jpl.nasa.gov/>



NASA Jet Propulsion Laboratory California Institute of Technology [+ View the NASA Portal](#) Search JPL

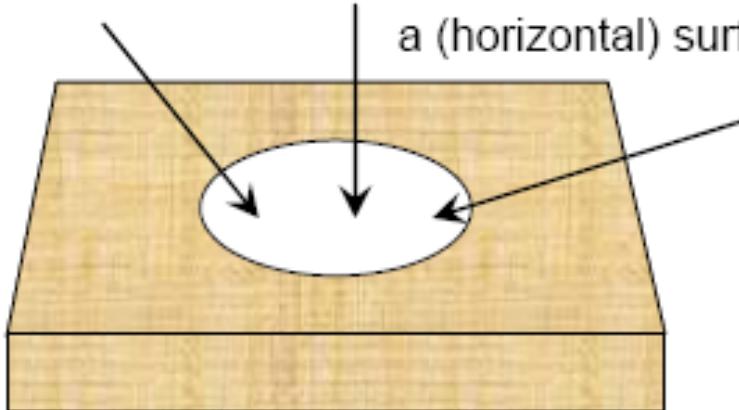
JPL HOME EARTH SOLAR SYSTEM STARS & GALAXIES TECHNOLOGY

NASA/JPL Data Evaluation

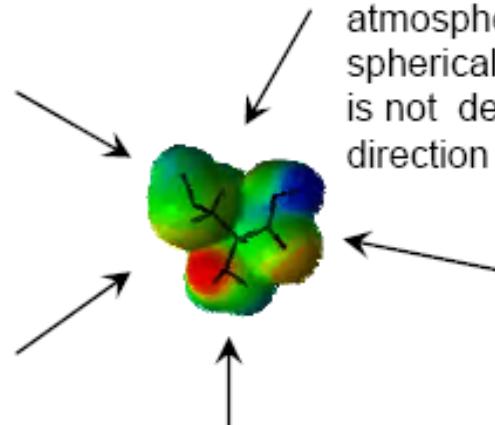
Jet Propulsion Laboratory California Institute of Technology

Integrals Over Angular Incidence

Irradiance vs Actinic Flux



Irradiance: The radiation flux incident on a (horizontal) surface.



Actinic flux: The photochemically active radiation flux in the earth's atmosphere. This flux is spherically integrated and is not dependent the direction of the radiation.

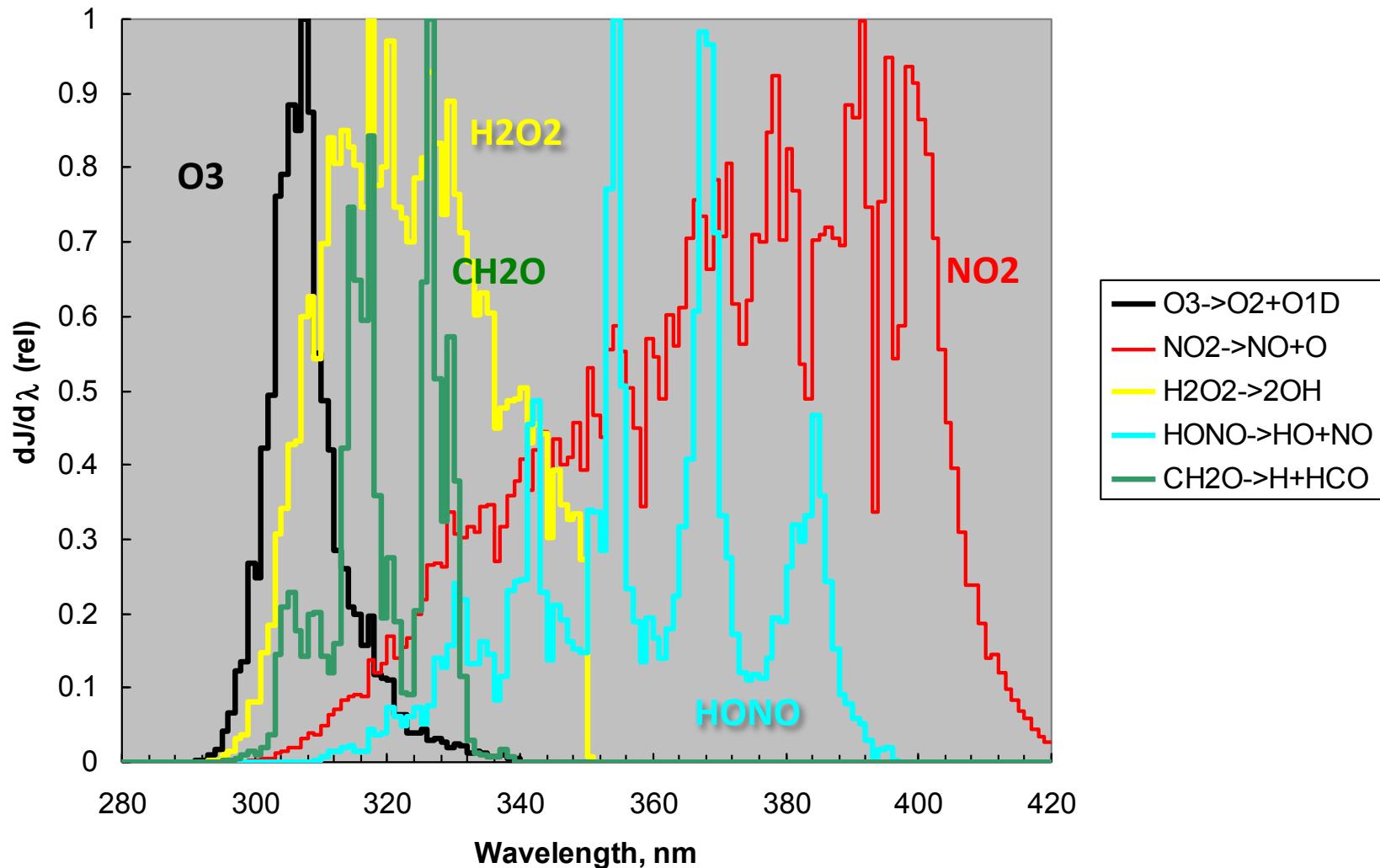
$$E = \int_0^{\frac{\pi}{2}} \int_0^{2\pi} I(\theta, \varphi) \cos \theta \sin \theta d\theta d\varphi$$

Watts m⁻²

$$F = \int_0^{\pi} \int_0^{2\pi} I(\theta, \varphi) \sin \theta d\varphi d\theta$$

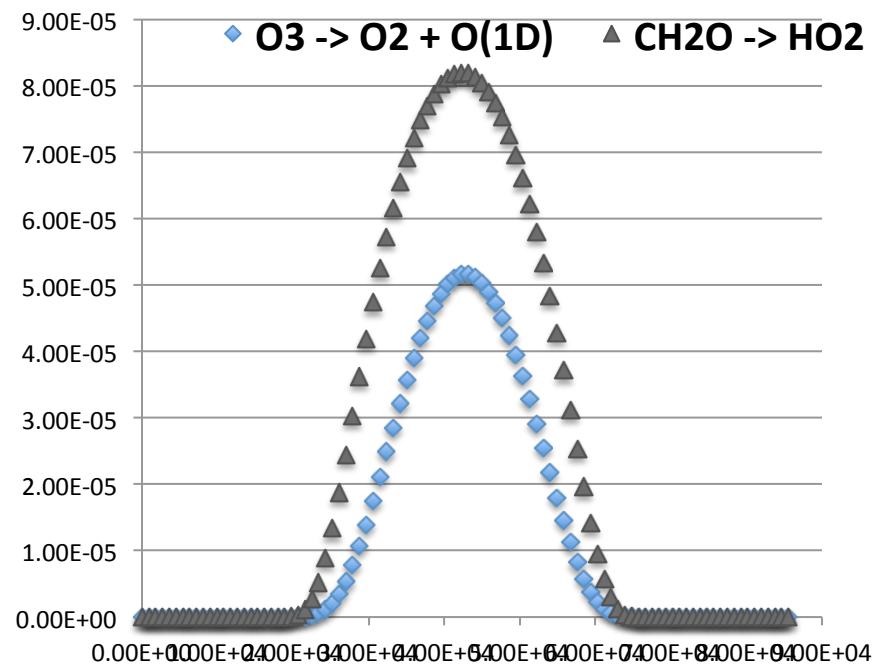
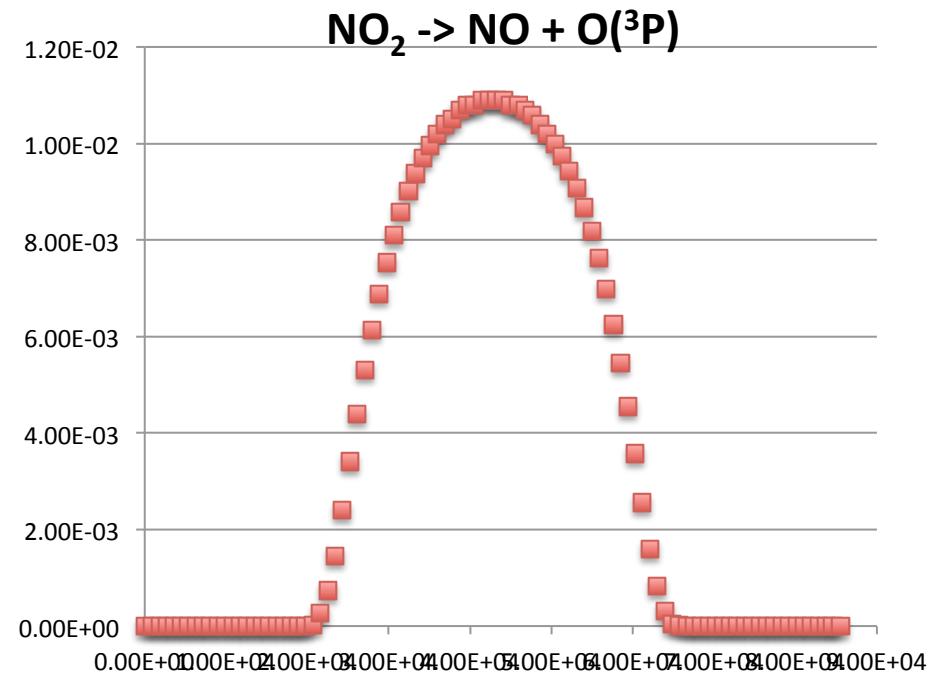
Watts m⁻² or quanta s⁻¹ cm⁻²

Photolysis Frequencies as Function of Wavelength



surface, overhead sun

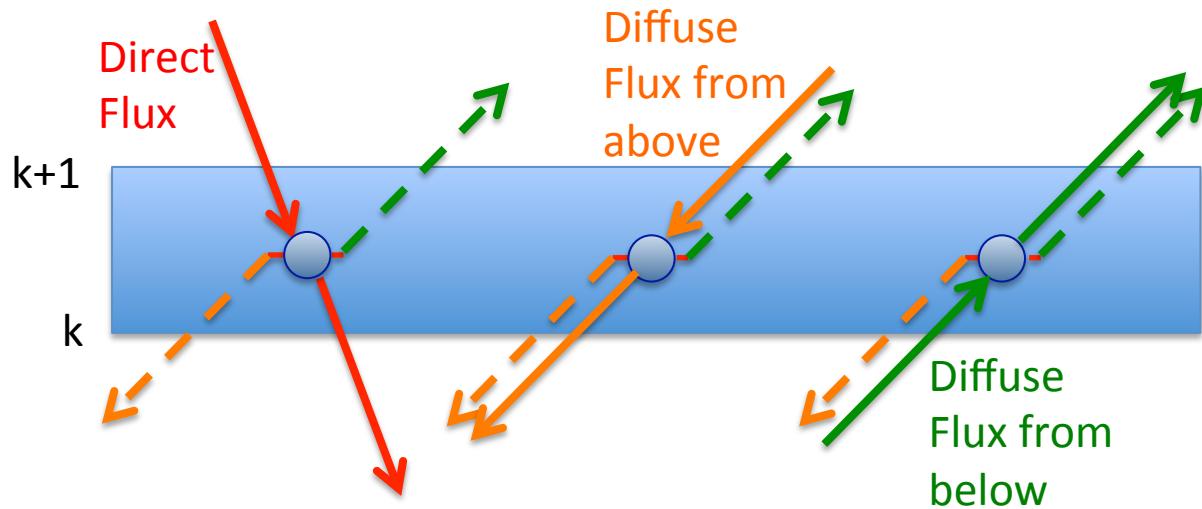
Photolysis Frequencies and Time of Day



NO_2 photolysis has broader parabola
-- result of diffuse vs direct radiation

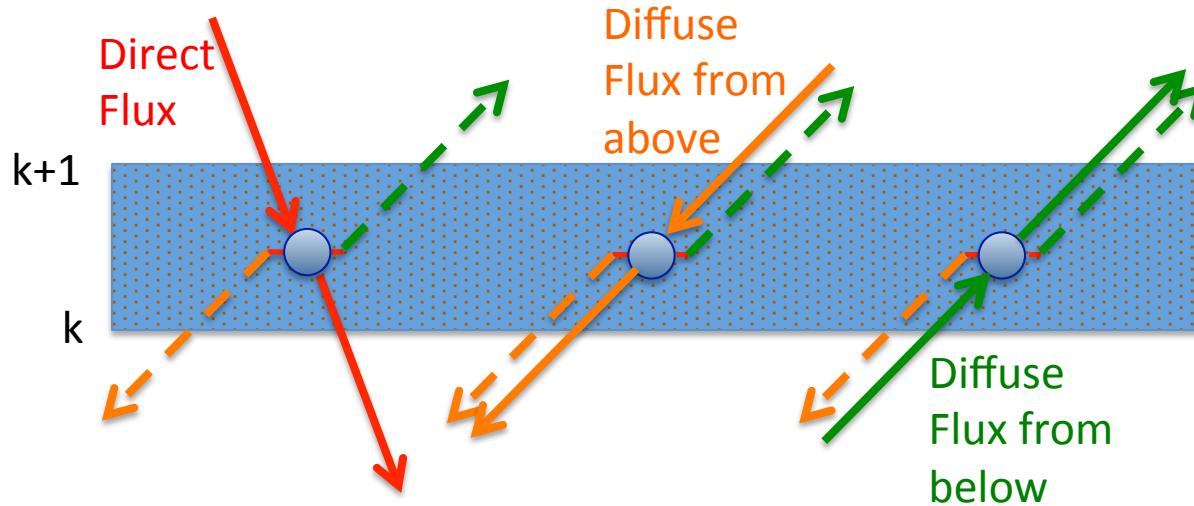
Calculating Photolysis Frequencies

Need to do Radiative Transfer Calculations



Aerosols and Clouds Affect Photolysis Rates

characteristics of the aerosol or cloud layer provides information to estimate their effect



Must specify three optical properties:

Optical depth, $\Delta\tau$

Single scattering albedo, $w_o = \text{scatt.}/(\text{scatt.}+\text{abs.})$

Asymmetry factor, g : forward fraction, $f \sim (1+g)/2$

$$\text{Vertical optical depth, } \Delta\tau(\lambda, z) = \sigma(\lambda, z) n(z) \Delta z$$

for molecules: $\Delta\tau(\lambda, z) \sim 0 - 30$

Rayleigh scatt. $\sim 0.1 - 1.0$

O_3 absorption $\sim 0 - 30$

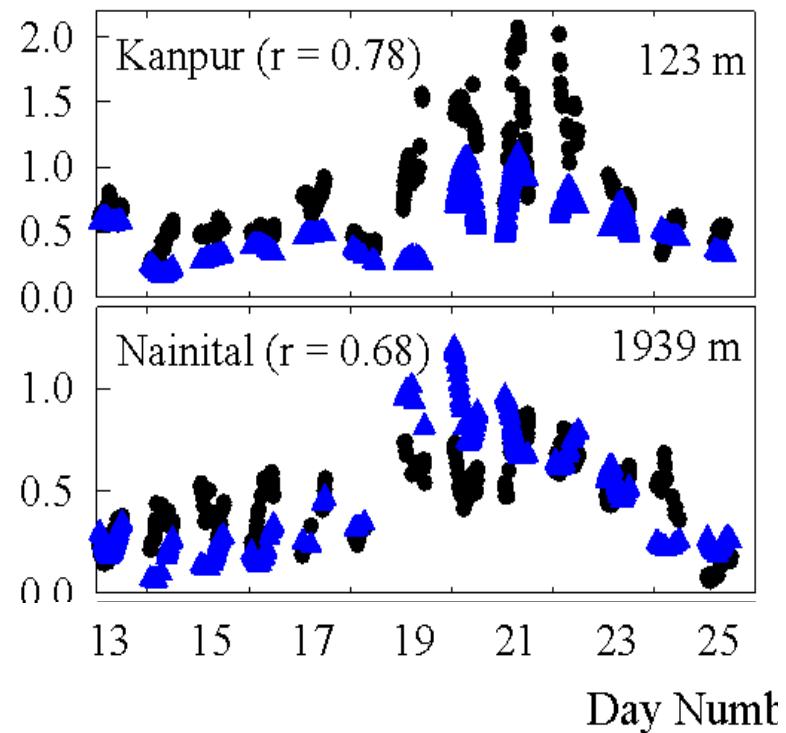
for aerosols: 0.01 - 5.0

for clouds: 1-1000

cirrus $\sim 1-5$

cumulonimbus $\sim > 100$

- Aeronet AOD (500 nm)
- ▲ WRF-Chem AOD (500 nm)



Single Scattering Albedo

$$w_o(\lambda, z) = \text{scatt.}/(\text{scatt.+abs.})$$

SSA range: 0 - 1

limits: pure scattering = 1.0
pure absorption = 0.0

for molecules, strongly λ -dependent, depending on absorber amount, esp. O₃

for aerosols:

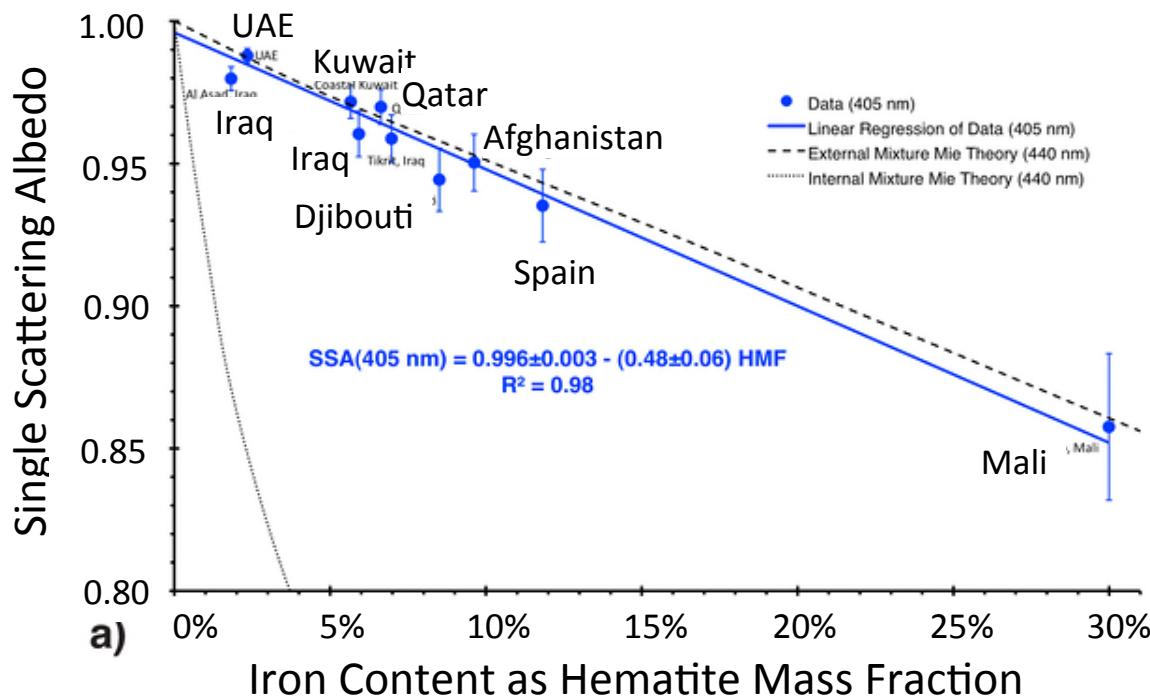
sulfate ~ 0.99

soot, organics ~ 0.8 or less,

not well known but probably higher at shorter λ , esp. in UV

for clouds: typically 0.9999 or larger (vis and UV)

Single scattering albedo of fine mineral dust aerosols controlled by iron concentration



Moosmuller et al. (2012)

Journal of Geophysical Research: Atmospheres

Volume 117, Issue D11, D11210, 8 JUN 2012 DOI: 10.1029/2011JD016909

<http://onlinelibrary.wiley.com/doi/10.1029/2011JD016909/full#jgrd17711-fig-0003>

Asymmetry factor, $g(\lambda, z)$

range -1 to +1

pure back-scattering = -1

isotropic or Rayleigh = 0

pure forward scattering = +1

$$g = \frac{1}{2} \int_{-1}^{+1} P(\Theta) \cos \Theta d(\cos \Theta)$$

strongly dependent on particle size

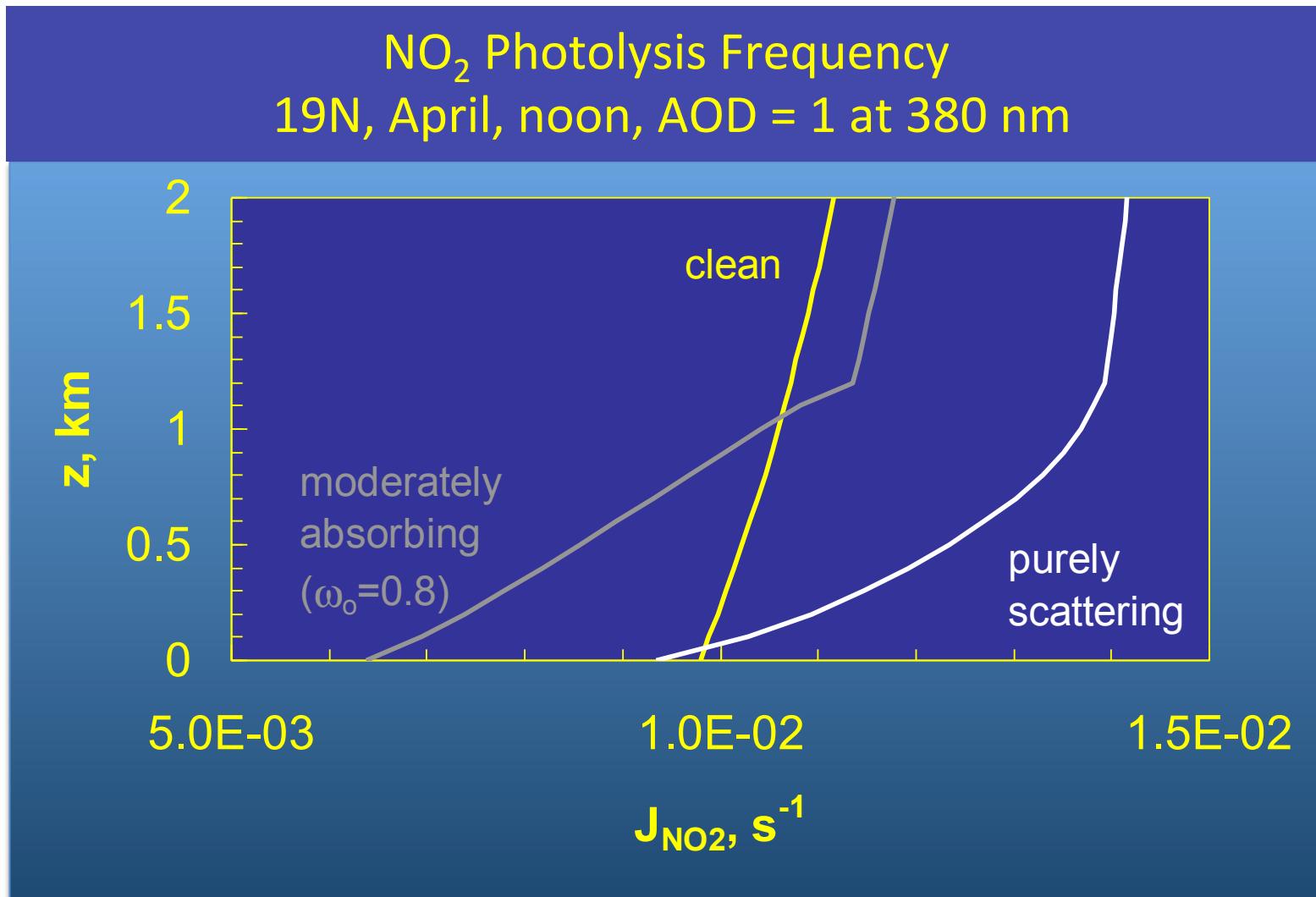
for aerosols:, typically 0.5-0.7

for clouds, typically 0.7-0.9

Mie theory for spherical particles: can compute $\Delta\tau$, w_o , g
from knowledge of λ , particle radius and complex index of refraction

Aerosol Effects on the Radiation

- Aerosols either scatter or absorb radiation

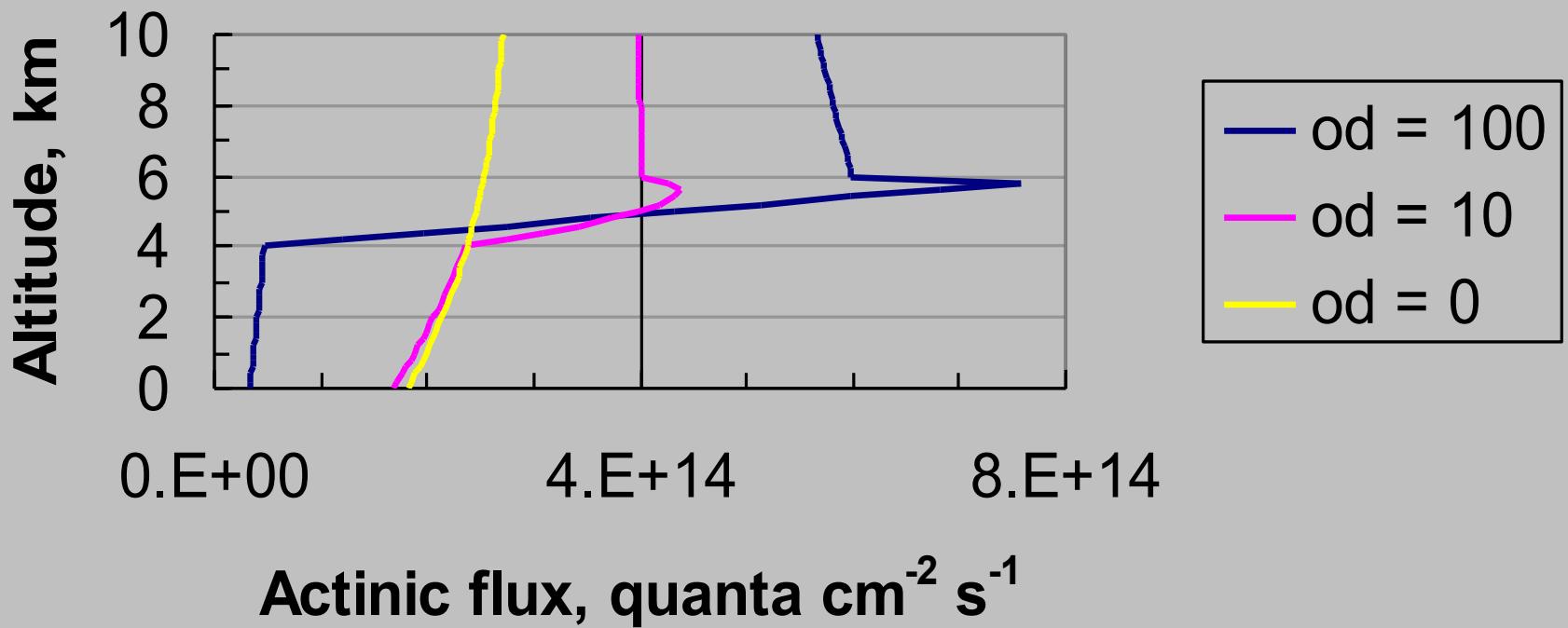


Uniform Cloud Layer

- **Above cloud:** - high radiation because of reflection
- **Below cloud:** - lower radiation because of attenuation by cloud
- **Inside cloud:** - complicated behavior
 - Top half: very high values (for high sun)
 - Bottom half: lower values

Effect Of Uniform Clouds On Actinic Flux

340 nm, sza = 0 deg.,
cloud between 4 and 6 km



Numerical Solutions To Radiative Transfer Equation

- Discrete ordinates
 - n-streams ($n = \text{even}$), angular distribution exact as n integrals but speed $\approx 1/n^2$
- Two-stream family
 - delta-Eddington, many others
 - very fast but not exact
- Monte Carlo
 - slow, but ideal for 3D problems
- Others
 - matrix operator, Feautrier, adding-doubling, successive orders, etc.

<http://cprm.acd.ucar.edu/Models/TUV/> Photolysis Rates

The screenshot shows a web browser window with the URL "cprm.acd.ucar.edu/Models/TUV/" in the address bar. The page title is "Chemical Processing and Regional Modeling" with the subtitle "a research group in Atmospheric Chemistry (ACD)". Below the title are links for UCAR, NCAR, ACD, CPRM Home, Models, and About Us. The Models menu is currently selected. At the bottom of the page, there are links for Models Home, MasterMech, WRF-Chem, and TUV.

Tropospheric Ultraviolet and Visible (TUV) Radiation Model

Tropospheric ultraviolet (UV) radiation is the driving force for all tropospheric photochemical processes. Photons in the UV wavelength have the potential to break usually fairly stable molecules into very reactive fragments (photolysis) and thus initiate reaction chains otherwise unlikely or even impossible. UV radiation is also harmful to living organisms and detrimental to human health. High doses of UV radiation are considered the major contributing factor for the development of skin cancer or cataracts. UV radiation can weaken the human immune system and can affect crop yields and phytoplankton activity (to only name a few effects).

Some questions of interest might be: What factors influence the amount of UV radiation available? What is the vertical structure of the radiative field? What sort of feedbacks (e.g., increased/decreased photolysis rates) can be expected from perturbations that - directly or indirectly - affect UV radiation? What are some of the health-related effects that can be expected from changes in atmospheric composition?

Climatology of Erythemal Ultraviolet Radiation, 1979-2000

The [monthly climatological distribution](#) for the period 1979-2000 of daily total erythemal (skin-reddening) ultraviolet radiation at Earth's surface, calculated with the TUV model using satellite-based (Nimbus-7, Meteor-3 and Earth Probe) TOMS (Total Ozone Mapping Spectrometer) observations of atmospheric ozone. The effects of clouds and scattering aerosols are accounted for using TOMS reflectivity at 380 nm. [Download erythemal data](#).

Downloads and Tools

- [TUV source code](#)
- [UV Climatologies](#) (NCAR Technical Note, figures, data)
- [Quick TUV calculator](#)

<http://cprm.acd.ucar.edu/Models/TUV/> Photolysis Rates

QUICK TUV CALCULATOR

[ACD](#) > [Models](#) > [TUV](#) > [Interactive TUV](#)

This web page runs the 4.1 version of the TUV model. You can run the model for a specified latitude, longitude and time (input option 1), or for a given solar zenith angle (input option 2). In either case, you must also specify the additional parameters in the second column. Also, you may select to print out the photolysis rates and/or the solar actinic flux spectrum at a given altitude above the surface (output option 1), or the erythemal UV and/or solar irradiance at that altitude (output option 2). For any problem, or to send comments, email [TUV administrators](#).

INPUT OPTION 1

LATITUDE (deg):

LONGITUDE (deg):

TIME (hh:mm:ss, GMT):

INPUT OPTION 2

SOLAR ZENITH ANGLE
(deg):

OTHER INPUT PARAMETERS

DATE (YYMMDD):

OVERHEAD OZONE COLUMN
(du):

SURFACE ALBEDO (0-1):

GROUND ELEVATION (km asl):

MEASUREM. ALTITUDE (km
asl):

OUTPUT OPTION 1

(for Atmospheric Science)

MOLECULAR PHOTOLYSIS

FREQUENCIES (s-1)

ACTINIC FLUX, SPECTRAL (quanta s-1
cm-2 nm-1)

OUTPUT OPTION 2

(for Biology)

IRRADIANCE, WEIGHTED (W m-2)

IRRADIANCE, SPECTRAL (W m-2 nm-1)

RADIATION TRANSFER MODEL

Pseudo-spherical 2 streams (faster, less accurate)

Pseudo-spherical discrete ordinate 4 streams (slower, more accurate)

[GO!](#)

Photochemistry Summary

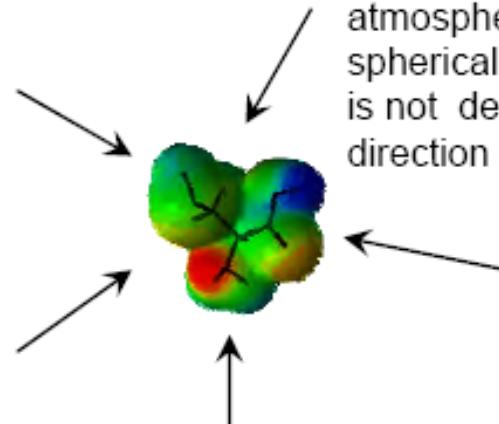
$$J \text{ (s}^{-1}\text{)} = \int \sigma(\lambda) \phi(\lambda) F(\lambda) d\lambda$$

$\sigma(\lambda)$ =absorption cross section, $\text{cm}^2 \text{ molec}^{-1}$

$\phi(\lambda)$ =photodissociation quantum yield, molec quanta^{-1}

$F(\lambda)$ = spectral actinic flux, $\text{quanta cm}^{-2} \text{ s}^{-1} \text{ nm}^{-1}$

- Clouds affect radiation – scattering
- Aerosols affect radiation – either scatter or absorb



Actinic flux: The photochemically active radiation flux in the earth's atmosphere. This flux is spherically integrated and is not dependent the direction of the radiation.

During Break

Go to the following web site

<http://cprm.acd.ucar.edu/Models/TUV/>

- 1) If you are able to compile and run fortran code, download the TUV code
- 2) Go to the TUV Quick Calculator web page

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1) Will need to create username and password; should get a zip file.

2)

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Some questions of interest might be: What factors influence the amount of UV radiation available? What is the vertical structure of the radiative field? What sort of feedbacks (e.g., increased/decreased photolysis rates) can be expected from perturbations that - directly or indirectly - affect UV radiation? What are some of the health-related effects that can be expected from changes in atmospheric composition?

Climatology of Erythemal Ultraviolet Radiation, 1979-2000

The [monthly climatological distribution](#) for the period 1979-2000 of daily total erythemal (skin-reddening) ultraviolet radiation at Earth's surface, calculated with the TUV model using satellite-based (Nimbus-7, Meteor-3 and Earth Probe) TOMS (Total Ozone Mapping Spectrometer) observations of atmospheric ozone. The effects of clouds and scattering aerosols are accounted for using TOMS reflectivity at 380 nm. [Download erythemal data](#).

Downloads and Tools

- [TUV source code](#)
- [UV Climatologies](#) (NCAR Technical Note, figures, data)
- [Quick TUV calculator](#)

Quick TUV Calculator to get Photolysis Rates

QUICK TUV CALCULATOR

[ACD](#) > [Models](#) > [TUV](#) > [Interactive TUV](#)

This web page runs the 4.1 version of the TUV model. You can run the model for a specified latitude, longitude and time (input option 1), or for a given solar zenith angle (input option 2). In either case, you must also specify the additional parameters in the second column. Also, you may select to print out the photolysis rates and/or the solar actinic flux spectrum at a given altitude above the surface (output option 1), or the erythemal UV and/or solar irradiance at that altitude (output option 2). For any problem, or to send comments, email [TUV administrators](#).

<p><input checked="" type="radio"/> INPUT OPTION 1</p> <p>LATITUDE (deg): <input type="text" value="0"/></p> <p>LONGITUDE (deg): <input type="text" value="0"/></p> <p>TIME (hh:mm:ss, GMT): <input type="text" value="12:00:00"/></p> <p><input type="radio"/> INPUT OPTION 2</p> <p>SOLAR ZENITH ANGLE <input type="text" value="0"/> (deg):</p>	<p>OTHER INPUT PARAMETERS</p> <p>DATE (YYMMDD): <input type="text" value="000630"/></p> <p>OVERHEAD OZONE COLUMN <input type="text" value="300"/> (du):</p> <p>SURFACE ALBEDO (0-1): <input type="text" value="0.1"/></p> <p>GROUND ELEVATION (km asl): <input type="text" value="0"/></p> <p>MEASUREM. ALTITUDE (km <input type="text" value="0"/> asl):</p>	<p><input checked="" type="radio"/> OUTPUT OPTION 1 (for Atmospheric Science)</p> <p><input checked="" type="checkbox"/> MOLECULAR PHOTOLYSIS FREQUENCIES (s-1)</p> <p><input type="checkbox"/> ACTINIC FLUX, SPECTRAL (quanta s-1 cm-2 nm-1)</p> <p><input type="radio"/> OUTPUT OPTION 2 (for Biology)</p> <p><input checked="" type="checkbox"/> IRRADIANCE, WEIGHTED (W m-2)</p> <p><input type="checkbox"/> IRRADIANCE, SPECTRAL (W m-2 nm-1)</p>
--	---	--

RADIATION TRANSFER MODEL

- Pseudo-spherical 2 streams (faster, less accurate)
 Pseudo-spherical discrete ordinate 4 streams (slower, more accurate)

[GO!](#)

TUV Fortran Program

-- code is compiled and ready to run --

```
bodhi:~/Documents/tuv/v5.0>tuv
```

TROPOSPHERIC ULTRAVIOLET VISIBLE (TUV) MODEL
(version 5.0)

S. Madronich et al., Atmospheric Chemistry Division
National Center for Atmospheric Research
P. O. Box 3000, Boulder, Colorado
tuv@acd.ucar.edu

Copyright (C) 1994-2010
University Corporation for Atmospheric Research

```
Type ?? for general information  
or <enter> to continue
```

??

TUV Fortran Program

-- code is compiled and ready to run --

??

The TUV model calculates solar short-wave radiation in the Earth's atmosphere. Available output includes:

- Spectral irradiance, W m⁻² nm⁻¹
- Spectral actinic flux, quanta cm⁻² s⁻¹ nm⁻¹
- Spectrally integrated irradiance, unweighted or weighted by biological action spectra, W m⁻²
- Photolysis rate coefficients, s⁻¹

Type ?xxx (where xxx = variable name) for help on inputs,
e.g. type ?tmzone to get help on entering time zone.

Inputs and outputs can be saved in different files.

A log file (tuvlog) is also created.

Only some simple changes are possible in the interactive version. Additional changes (e.g. shapes of vertical profiles of ozone, clouds, aerosols; wavelength dependent albedo, etc.) can be made by obtaining and editing the Fortran source code.

The full TUV model is available at:

<http://www.acd.ucar.edu/TUV>

PAUSE

To resume execution, type go. Other input will terminate the job.

TUV Fortran Program

-- code is compiled and ready to run --

```
go  
RESUMED
```

```
select input file  
<enter>: usrinp (if created before)  
1: defin1 (default No. 1, optimized for surface UV)  
2: defin2 (default No. 2, optimized for photochem )  
3: defin3 (default No. 3, optimized for master mech)  
4: defin4 (default No. 4, sample of all outputs)  
file-name for others
```

- defin1 → good for biology erythemal information
- defin2 → gives photolysis rates of a few reactions
- defin3 → gives all photolysis rates with outputs every 15 minutes for 24 hours
 - good for using as input into box model simulations

TUV Fortran Program

-- code is compiled and ready to run --

```
go  
RESUMED
```

```
select input file  
<enter>: usrinp (if created before)  
1: defin1 (default No. 1, optimized for surface UV)  
2: defin2 (default No. 2, optimized for photochem )  
3: defin3 (default No. 3, optimized for master mech)  
4: defin4 (default No. 4, sample of all outputs)  
file-name for others
```

defin1 → good for biology erythemal information

defin2 → gives photolysis rates of a few reactions

defin3 → gives all photolysis rates with outputs every 15 minutes for 24 hours
→ good for using as input into box model simulations

Type 2

TUV Fortran Program

Name of input file	TUV inputs:		Name of output file		Number of “streams” for radiation calc.	
	inpfil =	defin2	outfil =	usrout	nstr =	4
	lat =	0.000	lon =	0.000	tmzone =	0.0
	iyear =	2002	imonth =	3	iday =	21
	zstart =	0.000	zstop =	120.000	nz =	121
	wstart =	120.000	wstop =	735.000	nwint =	-156
	tstart =	12.000	tstop =	20.000	nt =	5
	lzenit =	F	alsurf =	0.100	psurf =	-999.0
	o3col =	300.000	so2col =	0.000	no2col =	0.000
	taucl =	0.000	zbase =	4.000	ztop =	5.000
	tauaur =	0.235	ssaaer =	0.990	alpha =	1.000
	dirsun =	1.000	difdn =	1.000	difup =	1.000
	zout =	0.500	zaird =	-9.990E+02	ztemp =	-999.000
	lirrad =	F	laflux =	T	lmmech =	F
	lrates =	F	isfix =	0	nms =	0
	ljvals =	T	ijfix =	0	nmj =	7
	iwfix =	0	itfix =	0	izfix =	0

Type ?variable for help on a variable, or
<enter> = keep these settings, or
Type variable name to change (lower case):

TUV Fortran Program

TUV inputs:

Latitude	Longitude	Time Zone
inpfil =	defin2	
lat =	0.000	tmzone = 0.0
iyear =	2002	iday = 21
zstart =	0.000	nz = 121
wstart =	120.000	nwint = -156
tstart =	12.000	nt = 5
lzenit =	F	psurf = -999.0
o3col =	300.000	no2col = 0.000
taucl =	0.000	ztop = 5.000
tauaur =	0.235	alpha = 1.000
dirsun =	1.000	difup = 1.000
zout =	0.500	ztemp = -999.000
lirrad =	F	lmmech = F
lrates =	F	nms = 0
ljvals =	T	nmj = 7
iwfix =	0	izfix = 0

Type ?variable for help on a variable, or

<enter> = keep these settings, or

Type variable name to change (lower case):

TUV Fortran Program

TUV inputs:

Year	inpfil =	defin2	Month	Day
	lat =	0.000		
	iyear =	2002	imonth =	3
	zstart =	0.000	zstop =	120.000
	wstart =	120.000	wstop =	735.000
	tstart =	12.000	tstop =	20.000
	lzenit =	F	alsurf =	0.100
	o3col =	300.000	so2col =	0.000
	taucl =	0.000	zbase =	4.000
	tauaur =	0.235	ssaaer =	0.990
	dirsun =	1.000	difdn =	1.000
	zout =	0.500	zaird =	-9.990E+02
	lirrad =	F	laflux =	T
	lrates =	F	isfix =	0
	ljvals =	T	ijfix =	0
	iwfix =	0	itfix =	0

Type ?variable for help on a variable, or
<enter> = keep these settings, or
Type variable name to change (lower case):

TUV Fortran Program

TUV inputs:

	inpfil =	defin2	outfil =	usrout	Number of vertical levels
Surface elevation above sea level	lat =	0.000			
	iyear =	2002		Top of atmosphere	
	zstart =	0.000	zstop =	120.000	nz = 121
	wstart =	120.000	wstop =	735.000	nwint = -156
	tstart =	12.000	tstop =	20.000	nt = 5
	lzenit =	F	alsurf =	0.100	psurf = -999.0
	o3col =	300.000	so2col =	0.000	no2col = 0.000
	taucl =	0.000	zbase =	4.000	ztop = 5.000
	tauaur =	0.235	ssaaer =	0.990	alpha = 1.000
	dirsun =	1.000	difdn =	1.000	difup = 1.000
	zout =	0.500	zaird =	-9.990E+02	ztemp = -999.000
	lirrad =	F	laflux =	T	lmmech = F
	lrates =	F	isfix =	0	nms = 0
	ljvals =	T	ijfix =	0	nmj = 7
	iwfix =	0	itfix =	0	izfix = 0

Type ?variable for help on a variable, or

<enter> = keep these settings, or

Type variable name to change (lower case):

TUV Fortran Program

TUV inputs:

	inpfil =	defin2	outfil =	usrout	nstr =	4
	lat =	0.000	lon =	0.000	Number of wavelength intervals	
	iyear =	2002	Last wavelength			
	zstart =	0.000	2000.000			
Starting wavelength	wstart =	120.000	wstop =	735.000	nwint =	-156
	tstart =	12.000	tstop =	20.000	If nwint < 0, the standard atmosphere wavelength grid is used	
	lzenit =	F	alsurf =	0.100		
	o3col =	300.000	so2col =	0.000		
	taucl =	0.000	zbase =	4.000		
	tauaur =	0.235	ssaaer =	0.990	alpha =	1.000
	dirlsun =	1.000	difdn =	1.000	difup =	1.000
	zout =	0.500	zaird =	-9.990E+02	ztemp =	-999.000
	lirrad =	F	laflux =	T	lmmech =	F
	lrates =	F	isfix =	0	nms =	0
	ljvals =	T	ijfix =	0	nmj =	7
	iwfix =	0	itfix =	0	izfix =	0

Type ?variable for help on a variable, or

<enter> = keep these settings, or

Type variable name to change (lower case):

TUV Fortran Program

TUV inputs:

Starting time, local hours	inpfil =	defin2	outfil =	usrout	nstr =	4
	lat =	0.000	lon =	0.000	tmzone =	0.0
	iyear =	2002	imonth =	3	iday =	21
	zstart =	0.000				
	wstart =	120.000				
	tstart =	12.000	tstop =	20.000	Stopping time	Number of time steps
	lzenit =	F	alsurf =	0.100	nt =	5
	o3col =	300.000	so2col =	0.000	psurf =	-999.0
	taucl =	0.000	zbase =	4.000	no2col =	0.000
	tauaur =	0.235	ssaaer =	0.990	ztop =	5.000
	dirlsun =	1.000	difdn =	1.000	alpha =	1.000
	zout =	0.500	zaird =	-9.990E+02	difup =	1.000
	lirrad =	F	laflux =	T	ztemp =	-999.000
	lrates =	F	isfix =	0	lmmech =	F
	ljvals =	T	ijfix =	0	nms =	0
	iwfix =	0	itfix =	0	nmj =	7
					izfix =	0

Type ?variable for help on a variable, or

<enter> = keep these settings, or

Type variable name to change (lower case):

TUV Fortran Program

TUV inputs:

```
=====
inpfil =      defin2    outfil =      usrout   nstr =        4
lat =          0.000     lon =          0.000    tmzone =      0.0
iyear =         2002     imonth =       3        iday =        21
zstart =        0.000     zstop =       120.000   nz =        121
wstart =       120.000
tstart =       12.000
                           Surface albedo           Surface pressure
lzenit =        F
alsurf =        0.100
psurf =        -999.0
o3col =       300.000
so2col =        0.000
taucl =        0.000
zbase =         4.000
tauaur =       0.235
ssaur =         0.990
dirlsun =       1.000
difdn =         1.000
zout =         0.500
zaird =      -9.990E+02
ztemp =      -999.000
lirrad =        F
laflux =         T
lmmech =        F
lrates =        F
isfix =          0
nms =            0
ljvals =        T
ijfix =          0
nmj =            7
iwfix =          0
itfix =          0
izfix =          0
=====
```

False = use time
True = use solar
zenith angle

Type ?variable for help on a variable, or
<enter> = keep these settings, or
Type variable name to change (lower case):

TUV Fortran Program

TUV inputs:

Ozone column (DU)	inpfil =	defin2	outfil =	usrout	nstr =	4
	lat =	0.000	lon =	0.000	tmzone =	0.0
	iyear =	2002	imonth =	3	iday =	21
	zstart =	0.000	zstop =	120.000	nz =	121
	wstart =	120.000	wstop =	735.000	nwint =	-156
	tstart =	12.000				
	lzenit =	F	SO ₂ column (DU)		NO ₂ column (DU)	
	o3col =	300.000	so2col =	0.000	no2col =	0.000
	taucl =	0.000	zbase =	4.000	ztop =	5.000
	tauaur =	0.235	ssaaer =	0.990	alpha =	1.000
	dirsun =	1.000	difdn =	1.000	difup =	1.000
	zout =	0.500	zaird =	-9.990E+02	ztemp =	-999.000
	lirrad =	F	laflux =	T	lmmech =	F
	lrates =	F	isfix =	0	nms =	0
	ljvals =	T	ijfix =	0	nmj =	7
	iwfix =	0	itfix =	0	izfix =	0

Type ?variable for help on a variable, or

<enter> = keep these settings, or

Type variable name to change (lower case):

TUV Fortran Program

TUV inputs:

```
=====
inpfil =      defin2    outfil =      usrout   nstr =          4
lat =           0.000    lon =           0.000    tmzone =        0.0
iyear =         2002    imonth =       3        iday =         21
zstart =        0.000    zstop =        120.000   nz =            121
wstart =       120.000   wstop =       735.000   nwint =       -156
tstart =       12.000    tstop =        20.000   nt =            5
lzenit =          F
o3col =       300.000
                Cloud base height      Cloud top height (km)
Cloud Optical Depth
taucl =        0.000    zbase =        4.000    ztop =        5.000
tauaur =       0.235    ssaaur =      0.990    alpha =       1.000
dirsun =       1.000    difdn =        1.000    difup =       1.000
zout =         0.500    zaird =     -9.990E+02  ztemp =     -999.000
lirrad =          F    laflux =          T    lmmech =          F
lrates =          F    isfix =          0    nms =            0
ljvals =          T    ijfix =          0    nmj =            7
iwfix =          0    itfix =          0    izfix =          0
=====
```

Type ?variable for help on a variable, or
<enter> = keep these settings, or
Type variable name to change (lower case):

TUV Fortran Program

TUV inputs:

inpfil =	defin2	outfil =	usrout	nstr =	4
lat =	0.000	lon =	0.000	tmzone =	0.0
iyear =	2002	imonth =	3	iday =	21
zstart =	0.000	zstop =	120.000	nz =	121
wstart =	120.000	wstop =	735.000	nwint =	-156
tstart =	12.000	tstop =	20.000	nt =	5
lzenit =	F			psurf =	-999.0
o3col =	300.000				
taucld =	0.000				
Aerosol Optical Depth	tauauer = 0.235	ssaaer = 0.990	alpha = 1.000	Angstrom Coef.	
	dirsun = 1.000	difdn = 1.000	difup = 1.000		
	zout = 0.500	zaird = -9.990E+02	ztemp = -999.000		
	lirrad = F	laflux = T	lmmech = F		
	lrates = F	isfix = 0	nms = 0		
	ljvals = T	ijfix = 0	nmj = 7		
	iwfix = 0	itfix = 0	izfix = 0		

Type ?variable for help on a variable, or

<enter> = keep these settings, or

Type variable name to change (lower case):

TUV Fortran Program

TUV inputs:

inpfil =	defin2	outfil =	usrout	nstr =	4
lat =	0.000	lon =	0.000	tmzone =	0.0
iyear =	2002	imonth =	3	iday =	21
zstart =	0.000	zstop =	120.000	nz =	121
wstart =	120.000	wstop =	735.000	nwint =	-156
tstart =	12.000	tstop =	20.000	nt =	5
lzenit =	F	alsurf =	0.100	psurf =	-999.0
o3col =	300.000			no2col =	0.000
taucl =	0.000	Diffuse Down Radiation		Diffuse Up Radiation	
tauaer =	0.235				
Direct Sun Radiation	dirsun = 1.000	difdn = 1.000		difup = 1.000	
	zout = 0.500	zaird = -9.990E+02		ztemp = -999.000	

dirsun = difdn = 1.0, difup = 0 for total down-welling irradiance

dirsun = difdn = difup = 1.0 for actinic flux from all directions

dirsun = difdn = 1.0, difup = -1 for net irradiance

These numbers are different in each of the defin# files

<enter> = keep these settings, or

Type variable name to change (lower case):

TUV Fortran Program

TUV inputs:

```
=====
inpfil =      defin2    outfil =      usrout   nstr =      4
lat =          0.000     lon =          0.000    tmzone =    0.0
iyear =        2002      imonth =      3         iday =      21
zstart =       0.000     zstop =       120.000   nz =        121
wstart =       120.000   wstop =       735.000   nwint =    -156
tstart =       12.000    tstop =       20.000   nt =        5
lzenit =       F          alsurf =      0.100    psurf =    -999.0
o3col =        300.000
taucl =        0.000
tauaur =       0.235
dirsun =       1.000
zout =         0.500
zaird =        -9.990E+02
ztemp =        -999.000
lirrad =       F
lrates =       F
ljvals =       T
iwfix =        0
=====
Air density (molec
cm-3) of output
altitude
Temperature (K) of
output altitude
```

Altitude (km) for
desired output

If zaird or ztemp < 0, then US Standard
Atmosphere at zout is used

Type ?variable for help on a variable, or
<enter> = keep these settings, or
Type variable name to change (lower case):

TUV Fortran Program

TUV inputs:

inpfil =	defin2	outfil =	usrout	nstr =	4
lat =	0.000	lon =	0.000	tmzone =	0.0
iyear =	2002	imonth =	3	iday =	21
zstart =	0.000	zstop =	120.000	nz =	121
wstart =	120.000	wstop =	735.000	nwint =	-156
tstart =	12.000	tstop =	20.000	nt =	5
lzenit =	F	alsurf =	0.100	psurf =	-999.0
o3col =	300.000	so2col =	0.000	no2col =	0.000
taucl =	0.000	zbase =	4.000	ztop =	5.000
tauaur =	0.235				
dirsun =	1.000				
zout =	0.500				

Output includes
actinic flux

Output includes data
for box model

True or False for
whether spectral
irradiance is
included in output

lirrad =	F	laflux =	T	lmmech =	F
lrates =	F	isfix =	0	nms =	0
ljvals =	T	ijfix =	0	nmj =	7
iwfix =	0	itfix =	0	izfix =	0

Type ?variable for help on a variable, or
<enter> = keep these settings, or
Type variable name to change (lower case):

TUV Fortran Program

TUV inputs:

inpfil =	defin2	outfil =	usrout	nstr =	4
lat =	0.000	lon =	0.000	tmzone =	0.0
iyear =	2002	imonth =	3	iday =	21
zstart =	0.000	zstop =	120.000	nz =	121
wstart =	120.000	wstop =	735.000	nwint =	-156
tstart =	12.000	tstop =	20.000	nt =	5
lzenit =	F	alsurf =	0.100	psurf =	-999.0
o3col =	300.000	so2col =	0.000	no2col =	0.000
taucl =	0.000	Output includes tabulated dose rates for different times and altitudes			ztop = 5.000
tauaur =	0.235				alpha = 1.000
dirsun =	1.000				
zout =	0.500				
lirrad =	F				
lrates =	F	isfix =	0	nms =	0
ljvals =	T	ijfix =	0	nmj =	7
iwfix =	0	itfix =	0	izfix =	0

True or False for
whether dose
rates are included
in output

Type ?variable for help on a variable, or
<enter> = keep these settings, or
Type variable name to change (lower case):

TUV Fortran Program

TUV inputs:

```
=====
inpfil =      defin2    outfil =      usrout   nstr =      4
lat =          0.000     lon =          0.000    tmzone =    0.0
iyear =        2002      imonth =      3         iday =      21
zstart =       0.000     zstop =       120.000   nz =        121
wstart =       120.000   wstop =       735.000   nwint =    -156
tstart =       12.000    tstop =       20.000   nt =        5
lzenit =       F         alsurf =      0.100    psurf =     -999.0
o3col =        300.000   so2col =      0.000    no2col =     0.000
taucl =        0.000     zbase =       4.000    ztop =       5.000
tauauer =      0.235
dirsun =        1.000
zout =         0.500
lirrad =       F
lrates =       F
ljvals =       T
iwfix =        0
ijfix =        0
itfix =        0
nmj =          7
izfix =        0
=====
```

True or False for
whether
photolysis rates
are included in
output

Output photolysis
rates for reaction
ijfix at different
times and altitudes

Number of photolysis
rates to be reported

Type ?variable for help on a variable, or
<enter> = keep these settings, or
Type variable name to change (lower case):

TUV Fortran Program

TUV inputs:

inpfil =	defin2	outfil =	usrout	nstr =	4
lat =	0.000	lon =	0.000	tmzone =	0.0
iyear =	2002	imonth =	3	iday =	21
zstart =	0.000	zstop =	120.000	nz =	121
wstart =	120.000	wstop =	735.000	nwint =	-156
tstart =	12.000	tstop =	20.000	nt =	5
lzenit =	F	alsurf =	0.100	psurf =	-999.0
o3col =	300.000	o32col =	0.000	no2col =	0.000
taucl =	0.000			ztop =	5.000
tauaur =	0.235				
dirsun =	1.000				
zout =	0.500				
lirrad =	F				
lrates =	F				
ljvals =	T				
iwfix =	0	itfix =	0	izfix =	0

Output spectral
irradiance or
spectral actinic
flux at wavelength
= iwfix for
different times
and altitudes

Output spectral
irradiance or
spectral actinic flux
at time = itfix for
different altitudes
and wavelengths

Output spectral
irradiance or spectral
actinic flux at altitude =
izfix for different times
and wavelengths

Type ?variable for help on a variable, or
<enter> = keep these settings, or
Type variable name to change (lower case):

TUV Fortran Program

Make one change to input:

Type: outfil

Type: screen

This will give output on screen instead of in an output file.

```
Type ?variable for help on a variable, or
<enter> = keep these settings, or
Type variable name to change (lower case):
outfil
  write new value for outfil
screen
```

Then push “return” (or “enter”)

And “enter” again (that is, do not save input file

The program calculates and prints output.

TUV Output

done: loading inputs
calculating....
equally spaced z-grid

Five time steps and solar zenith angles

step = 1 sza = 1.829 Earth-sun factor = 1.0076792
step = 2 sza = 28.200 Earth-sun factor = 1.0076326
step = 3 sza = 58.205 Earth-sun factor = 1.0075837
step = 4 sza = 88.211 Earth-sun factor = 1.0075371
step = 5 sza = 118.216 Earth-sun factor = 1.0074904

Spectral actinic flux, quanta cm⁻² s⁻¹ nm⁻¹

values at z = 0.500 km

Actinic Flux at z=0.5 km, different wavelengths and 5 times

Columns: wavelength (nm), times (hrs)

wc, nm	12.000	14.000	16.000	18.000	20.000
sza =	1.829	28.200	58.205	88.211	118.216
120.7000	0.000E+00	0.000E+00	0.000E+00	0.000E+00	0.000E+00
121.6500	0.000E+00	0.000E+00	0.000E+00	0.000E+00	0.000E+00
122.1000	0.000E+00	0.000E+00	0.000E+00	0.000E+00	0.000E+00
122.7000	0.000E+00	0.000E+00	0.000E+00	0.000E+00	0.000E+00
123.4500	0.000E+00	0.000E+00	0.000E+00	0.000E+00	0.000E+00
124.2000	0.000E+00	0.000E+00	0.000E+00	0.000E+00	0.000E+00
125.0000	0.000E+00	0.000E+00	0.000E+00	0.000E+00	0.000E+00

TUV Output

Photolysis rate coefficients, s⁻¹

- 1 = $\text{O}_3 \rightarrow \text{O}_2 + \text{O(1D)}$
 - 2 = $\text{H}_2\text{O}_2 \rightarrow 2 \text{ OH}$
 - 3 = $\text{NO}_2 \rightarrow \text{NO} + \text{O(3P)}$
 - 4 = $\text{NO}_3 \rightarrow \text{NO} + \text{O}_2$
 - 5 = $\text{NO}_3 \rightarrow \text{NO}_2 + \text{O(3P)}$
 - 6 = $\text{CH}_2\text{O} \rightarrow \text{H} + \text{HCO}$
 - 7 = $\text{CH}_2\text{O} \rightarrow \text{H}_2 + \text{CO}$

values at z = 0.500 km

Columns: time, sza, photo_reactions

Exercises

Modify input section:

- 1) Change latitude and longitude to location of the dust storm
Lat = 28N Lon = 73 E
- 2) Change date to time of dust storm
2010-04-21
- 3) Change start and stop time to go from morning to late afternoon
tstart=0, tstop=14
You could also change the number of times to print out
- 4) Change the aerosol optical depth (tauuaer) to the value for the dust storm
tauuaer = 1.5
- 5) Change the Angstrom component (alpha) to the value for the dust storm
alpha = 1.0
- 6) Change the altitude to compare results from near the surface to near the top of the troposphere
zout = 12 km

How much do the mid-day photolysis rates change from one step to the next?

WRF-Chem Sensitivity Simulations

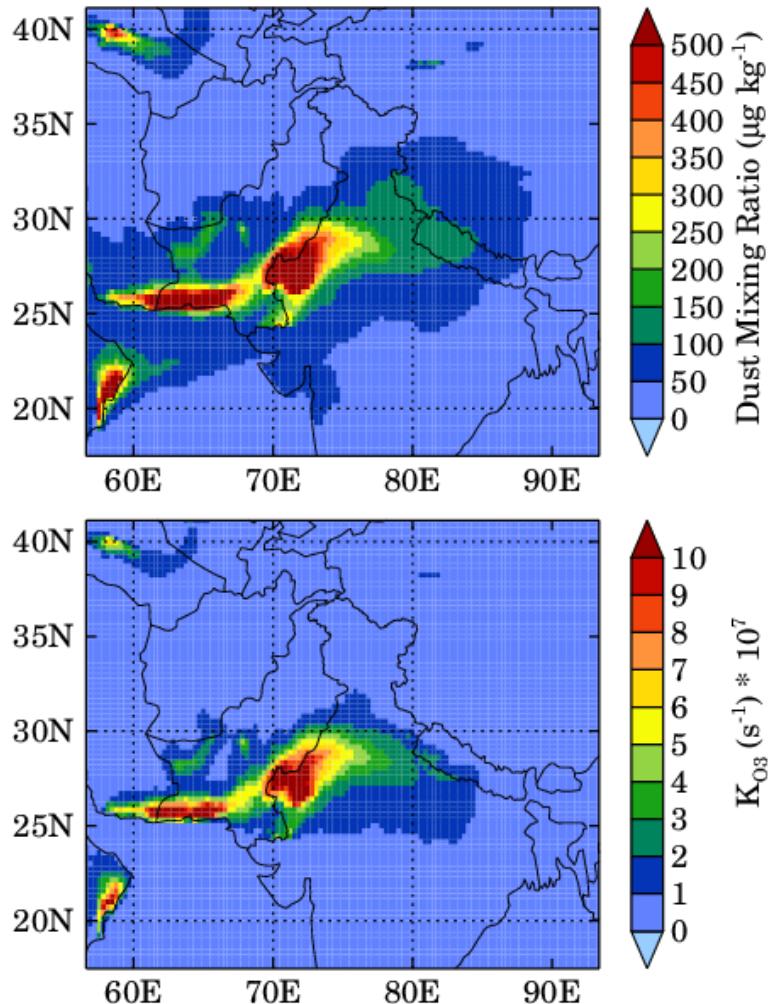
1. Base Case – with Dust emissions and j-values affected = Dust_J
2. No Dust emissions Case = No Dust
3. Dust emissions with J-values and Heterogeneous reactions (no relative humidity effect) = Dust_JH_NoRH
4. Dust emissions with J-values and Heterogeneous reactions (with relative humidity effect) = Dust_JH

Add 12 Heterogeneous Reactions in WRF-Chem

Reaction	γ_{dry}	RH dependence
$\text{O}_3 + \text{Dust} \rightarrow \text{P}$	2.7×10^{-5}	Cwiertny et al. (2008)
$\text{HNO}_3 + \text{Dust} \rightarrow 0.5 \text{ NO}_x + \text{P}$	2.0×10^{-3}	Liu et al. (2008)
$\text{NO}_2 + \text{Dust} \rightarrow \text{P}$	2.1×10^{-6}	-
$\text{NO}_3 + \text{Dust} \rightarrow \text{P}$	0.1	-
$\text{N}_2\text{O}_5 + \text{Dust} \rightarrow \text{P}$	0.03	-
$\text{OH} + \text{Dust} \rightarrow 0.05 \text{ H}_2\text{O}_2 + \text{P}$	0.18	Bedjanian et al. (2013a)
$\text{HO}_2 + \text{Dust} \rightarrow 0.1 \text{ H}_2\text{O}_2 + \text{P}$	6.42×10^{-2}	Bedjanian et al. (2013b)
$\text{H}_2\text{O}_2 + \text{Dust} \rightarrow \text{P}$	2×10^{-3}	Pradhan et al. (2010)
$\text{SO}_2 + \text{Dust} \rightarrow \text{P}$	3.0×10^{-5}	Preszler Prince et al. (2007)
$\text{CH}_3\text{COOH} + \text{Dust} \rightarrow \text{P}$	1×10^{-3}	-
$\text{CH}_3\text{OH} + \text{Dust} \rightarrow \text{P}$	1×10^{-5}	-
$\text{CH}_2\text{O} + \text{Dust} \rightarrow \text{P}$	1×10^{-5}	-

Heterogeneous Chemistry in WRF-Chem

example of O_3 + dust reaction rate constant



Dust mixing ratio
17-22 April 2010

Reaction rate constant for $O_3 + \text{dust}$

$$k_g = \sum_{i=1}^5 \frac{4\pi r_i D_g V N_i}{1 + K_n [\chi + \frac{4(1-\gamma)}{3\gamma}]}$$

[Heikes and Thompson, 1983]

WRF-Chem trace gases compared to observations at Nainital -- represents regional-scale concentrations



WRF-Chem reproduces observed variations in ozone and NOy at Nainital with effects of dust aerosols.

