# Autoregressive Spectral Estimation\*

# Emanuel Parzen

#### Introduction

The problem of spectral analysis of time series is clearly of great interest to the many applied scientists who use spectral analysis in their scientific research. It should be of great interest to statisticians because it embodies prototypes of two of the great problems of modern statistics: functional inference and modeling. A problem of statistical inference usually assumes three ingredients: a sample of observations, a parameter which indexes the family of possible joint probability densities of the sample, and a formula for the probability density of the sample

f(sample | parameter).

Classical statistical inference assumes the parameter  $\theta = (\theta_1, \dots, \theta_k)$ . Functional inference assumes the parameter is a function, such as  $f(\omega)$ ,  $0 \le \omega \le 1$ .

The parameter estimation problem seeks to form optimal estimators (denoted  $\hat{\theta}$ ) of the parameter. A typical model identification problem seeks to find the smallest number of significantly nonzero components  $\theta_i$  of the parameter  $\theta$ .

Estimation of a function often has similar features to model identification, since a function can be parametrized exactly by a countable infinity of parameters. However, in practice, one can only efficiently estimate a finite number of parameters. Therefore, to estimate a function, one must use the smallest finite number of parameters which provide an adequate approximation of the function.

The goals of functional inference and model identification are in my view best pursued simultaneously. One seeks methods of statistical inference which are finite-parametric and non-(or infinite-) parametric. One achieves this goal by using finite parameter models (which theory indicates might be exact

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methods) in ways that enable them to be interpreted also as approximating models.

Autoregressive spectral estimation is one of the new techniques for spectral analysis developed in the last two decades. Its theory and applications are extremely extensive. This article aims to provide an overview (rather than a detailed account) of the main ideas. A comprehensive bibliography guides the reader to articles in which case studies and comparisons of autoregressive spectral estimators are described.

The spectral density is defined in Section 1. Infinite-order AR and MA representations of a stationary time series are introduced in Section 2. Entropy as a motivation for autoregressive schemes is discussed in Section 3. Alternate parametrizations of an autoregressive scheme are outlined in Section 4. Algorithms for computing the coefficients of autoregressive spectral densities are stated in Section 5. Criteria for determining the orders of autoregressive schemes are mentioned in Section 6. Suggestions for empirical spectral analysis are outlined in Section 7. The final section provides a guide to the literature of autoregressive spectral estimation by listing the articles which correspond to some important developments.

# 1. Correlations and spectral functions of a stationary time series

The theory of time series discusses separately discrete-parameter time series  $\{Y(t), t=0, \pm 1, \pm 2, \ldots\}$  and continuous-parameter time series  $\{Y(t), -\infty < t < \infty\}$ . Only the former case is discussed in this article. Discrete-parameter series usually arise by observing a continuous-parameter time series at equispaced times. The frequency variable  $\omega$  of a pure harmonic

$$Y(t) = A\cos 2\pi\omega t + B\sin 2\pi\omega t \tag{1.1}$$

observed at  $t = 0, \pm 1, \ldots$  can be assumed to vary in either  $-0.5 \le \omega \le 0.5$  or  $0 \le \omega \le 1$ . The first interval is usually adopted, but the second interval will be adopted in this article because it is more convenient for developing isomorphisms between spectral analysis and nonparametric data modeling using quantile and density-quantile functions (see Parzen, 1979).

The definitions and notation we adopt for the functions used to describe a zero-mean *stationary* Gaussian discrete-parameter time series Y(t), t = 0,  $\pm 1$ ,... are as follows.

A 'time-domain' specification of the probability law of  $Y(\cdot)$  is provided by the covariance function

$$R(v) = E[Y(t)Y(t+v)], \quad v = 0, \pm 1, \pm 2, \dots;$$
 (1.2)

or by the variance R(0) and the correlation function

$$\rho(v) = \frac{R(v)}{R(0)} = \text{corr}[Y(t), Y(t+v)]. \tag{1.3}$$

To define spectral (frequency) domain specification of the probability law of  $Y(\cdot)$ , we first assume summability of  $R(\cdot)$  and  $\rho(\cdot)$ . The Fourier transforms of R(v) and  $\rho(v)$  are called the *power spectrum*  $S(\omega)$  and *spectral density function*  $f(\omega)$  respectively, and are defined by

$$S(\omega) = \sum_{v=-\infty}^{\infty} e^{-2\pi i v \omega} R(v), \quad 0 \le \omega \le 1; \tag{1.4}$$

$$f(\omega) = \sum_{v=-\infty}^{\infty} e^{-2\pi i v \omega} \rho(v), \quad 0 \le \omega \le 1.$$
 (1.5)

The spectral distribution function is defined by

$$F(\omega) = \int_0^{\omega} f(\omega') d\omega', \quad 0 \le \omega \le 1.$$
 (1.6)

For data analysis, one actually computes a modified spectral distribution function denoted  $F_+(w)$  and defined for  $0 \le \omega \le 0.5$ :

$$F_{+}(\omega) = 2F(\omega), \quad 0 \le \omega \le 0.5$$
 (1.7)

A correlation function  $\rho(v)$  has a mathematical property of being a positivedefinite function which, without assuming summability, guarantees the existence of: (1) a spectral distribution function  $F(\omega)$ , and (2) the spectral representation of the correlation function  $\rho(v)$  given by

$$\rho(v) = \int_0^1 e^{2\pi i v \omega} dF(\omega)$$

$$= \int_0^{0.5} \cos 2\pi v \omega dF_+(\omega). \qquad (1.8)$$

The notion of an *ergodic* time series is not given a rigorous definition in this article but its intuitive meaning is important for us. We call a time series ergodic when the parameters of its probability law possess consistent estimators (and thus can be determined with probability one, given a sample of infinite length). An example of a nonergodic stationary Gaussian zero-mean time series is

$$Y(t) = A \cos 2\pi\omega t + B \sin 2\pi\omega t$$

where A and B are independent  $N(0, \sigma^2)$  random variables. One can infer the values of A and B in the sample observed, but one cannot infer the value of

 $\sigma^2$ . This time series has correlation function

$$\rho(v) = \cos 2\pi \omega v, \quad v = 0, \pm 1, \pm 2, \dots, \tag{1.9}$$

and does not possess a spectral density.

Parzen (1982) proposes that it is useful in practice to distinguish qualitatively between three types of time series, which we call

no memory: white noise,

short memory: stationary and ergodic,

long memory: nonstationary or nonergodic.

A no-memory or white noise time series is a stationary Gaussian time series satisfying either of the equivalent conditions:

$$\rho(v) = 0 \text{ for } v > 0; \quad f(\omega) = 1, \ 0 \le \omega \le 1.$$
 (1.10)

A short-memory time series is a stationary time series possessing a summable correlation function  $\rho(v)$  and a spectral density  $f(\omega)$  which is bounded above and below in the sense that the dynamic range of  $f(\omega)$ 

$$DR(f) = \max_{0 \le \omega \le 1} f(\omega) \div \min_{0 \le \omega \le 1} f(\omega)$$
(1.11)

satisfies  $1 < DR(f) < \infty$ .

A long-memory time series is one which is neither no memory nor short memory; alternatively, a long-memory time series is one which is nonstationary or nonergodic. It usually has components representing cycles or trends. An example of a long-memory time series is (1.1) where A and B are independent  $N(0, \sigma^2)$  random variables.

For a short-memory time series, one can define the inverse-correlation function

$$\rho i(v) = \int_0^1 e^{2\pi i v \omega} f^{-1}(\omega) d\omega \div \int_0^1 f^{-1}(\omega) d\omega$$
 (1.12)

and the cepstral-correlation function

$$\gamma(v) = \int_0^1 e^{2\pi i v \omega} \log f(\omega) d\omega.$$
 (1.13)

It should be noted that the inverse-correlation function is positive definite. However, the cepstral-correlation function is not. These new types of correlation functions are introduced because they may provide more parsimonious parametrizations in the sense that they decay to 0 faster than does the correlation function. Statistical inference from a sample of the probability law of a time series often achieves greatest statistical efficiency by using the most

parsimonious parametrizations. Thus to form estimators  $\hat{f}(\omega)$  on the spectral density  $f(\omega)$  from a raw estimator  $\tilde{f}(\omega)$ , greater precision may be attained by first forming estimators  $\{f^{-1}(\omega)\}$  and  $\{\log f(\omega)\}$  of the inverse or logarithm of the spectral density. Autoregressive spectral estimation may be regarded as an approach to estimating  $f(\omega)$  by first estimating  $f^{-1}(\omega)$ .

# 2. Filter models of a short-memory stationary time series

A short-memory zero-mean Gaussian stationary time series Y(t) has a property of fundamental importance: it can be linearly transformed to a white noise time series, denoting  $Y^{\nu}(t)$  or  $\varepsilon(t)$ , by a filter that is invertible. The definition of  $Y^{\nu}(t)$  is provided by the theory of minimum mean-square error prediction.

The definitions and notation of prediction theory that we adopt are as follows:

$$Y^{\mu,m}(t) = E[Y(t) | Y(t-1), \dots, Y(t-m)]$$
 (2.1)

denotes the memory m one-step ahead predictor, while

$$Y^{\nu,m}(t) = Y(t) - Y^{\mu,m}(t) \tag{2.2}$$

is the prediction error, and

$$\sigma_m^2 = E[|Y^{\nu,m}(t)|^2] \div E[|Y(t)|^2] \tag{2.3}$$

is the mean-square prediction error measured in units of the variance R(0) of Y(t). Corresponding infinite-memory notation is

$$Y^{\mu}(t) = E[Y(t)|Y(t-1), Y(t-2), \ldots], \qquad (2.4)$$

$$Y^{\nu}(t) = Y(t) - Y^{\mu}(t), \qquad (2.5)$$

$$\sigma_{\infty}^{2} = E[|Y^{\nu}(t)|^{2}] \div E[|Y(t)|^{2}]. \tag{2.6}$$

Explicit formulas for the foregoing predictors, prediction errors, and meansquare prediction errors can be obtained from the correlation function  $\rho(v)$ . The autoregressive coefficients  $\alpha_m(1), \ldots, \alpha_m(m)$  of order m are defined by

$$-Y^{\mu,m}(t) = \alpha_m(1)Y(t-1) + \cdots + \alpha_m(m)Y(t-m), \qquad (2.7)$$

$$Y^{\nu,m}(t) = Y(t) + \alpha_m(1)Y(t-1) + \cdots + \alpha_m(m)Y(t-m).$$
 (2.8)

A predictor is characterized by the condition that the prediction error is orthogonal (normal) to the predictor variables:

$$E[Y^{\nu,m}(t)Y(t-k)] = 0, \quad k = 1, \dots, m.$$
 (2.9)

By substituting (2.8) into (2.9), one obtains the famous Yule-Walker equations, defining  $\alpha_m(0) = 1$ .

$$\sum_{j=0}^{m} \alpha_m(j) \rho(j-k) = 0, \quad k = 1, \dots, m.$$
 (2.10)

One obtains  $\sigma_m^2$  by

$$\sigma_m^2 = E[Y^{\nu,m}(t)Y(t)] \div E[|Y(t)|^2] = \sum_{j=0}^m \alpha_m(j)\rho(j).$$
 (2.11)

For a short-memory time series, these equations also hold with  $m = \infty$ .

The time series of infinite-memory prediction errors  $Y^{\nu}(t)$  is always a white noise series called the *innovations*. It provides a transformation of the time series Y(t) to a white noise time series  $Y^{\nu}(t)$ , which we write

$$Y(t) + \alpha_{\infty}(1)Y(t-1) + \cdots = Y^{\nu}(t),$$
 (2.12)

and call an AR( $\infty$ ) or infinite-order autoregressive scheme for Y(t). An MA( $\infty$ ) or infinite-order moving average scheme representation is

$$Y(t) = Y^{\nu}(t) + \beta_{\infty}(1)Y^{\nu}(t-1) + \cdots, \qquad (2.13)$$

whose coefficients  $\beta_{\infty}(k)$  can be determined recursively from  $\alpha_{\infty}(j)$  by  $\beta_{\infty}(0) = 1$ , and for k > 0

$$\beta_{\infty}(k) + \sum_{j=1}^{k} \alpha_{\infty}(j)\beta_{\infty}(k-j) = 0.$$
 (2.14)

The AR( $\infty$ ) and MA( $\infty$ ) representations have important implications for spectral analysis since they provide formulas for the spectral density function  $f(\omega)$  alternative to the formula that  $f(\omega)$  is the Fourier transform of  $\rho(v)$ . One can show that

$$f(\omega) = \sigma_{\infty}^2 |h_{\infty}(e^{2\pi i \omega})|^2, \qquad (2.15)$$

$$f^{-1}(\omega) = \sigma_{\infty}^{-2} |g_{\infty}(e^{2\pi i \omega})|^2,$$
 (2.16)

defining

$$h_{\infty}(z) = \sum_{j=0}^{\infty} \beta_{\infty}(j)z^{j}, \qquad g_{\infty}(z) = \sum_{k=0}^{\infty} \alpha_{\infty}(k)z^{k}.$$
 (2.17)

These infinite series converge in general in mean square on the unit interval. In order to guarantee pointwise convergence at each  $\omega$  in  $0 \le \omega \le 1$ , one must

make an additional smoothness condition such as a Lipshitz condition on  $f(\omega)$ , which is implied in turn by the condition

$$\sum_{v=-\infty}^{\infty} |v| |\rho(v)| < \infty. \tag{2.18}$$

In spectral analysis we usually assume at least the existence of a continuous second derivative, which is implied by the condition

$$\sum_{v=-\infty}^{\infty} |v|^2 |\rho(v)| < \infty. \tag{2.19}$$

The notion of a time series  $Y(\cdot)$  being an autoregressive scheme of order p, denoted AR(p), can be defined in terms of prediction theory as follows:  $Y(\cdot)$  is an AR(p) if and only if the memory p prediction errors  $Y^{\nu,p}(\cdot)$  is a white noise time series and  $\alpha_p(p) \neq 0$ . The spectral density of  $Y^{\nu,p}(\cdot)$  can be expressed in terms of the autoregressive transfer function of order p

$$g_p(z) = \alpha_p(0) + \alpha_p(1)z + \dots + \alpha_p(p)z^p$$
 (2.20)

by

$$f_{Y^{\nu,p}}(\omega) = \frac{1}{\sigma_p^2} |g_p(e^{2\pi i\omega})|^2 f(\omega).$$
 (2.21)

If the time series  $Y(\cdot)$  is in the fact AR(p), then its spectral density equals the function

$$f_p(\omega) = \sigma_p^2 |g_p(e^{2\pi i \omega})|^{-2}, \qquad (2.22)$$

which we call, in general, the approximating autoregressive spectral density of order p. A time series  $Y(\cdot)$  can be regarded as approximated by an AR(p) if

$$\bar{f}_p(\omega) = \frac{f(\omega)}{f_p(\omega)} \tag{2.23}$$

can be regarded as not 'significantly' different from the constant 1. In this way a test of the hypothesis that a time series  $Y(\cdot)$  is AR(p) can be converted to a test of the hypothesis that the prediction error time series is white noise.

The sequence of approximating autoregressive spectral densities  $f_m(\omega)$ ,  $m = 1, 2, \ldots$  may be shown to converge as m tends to  $\infty$  to  $f(\omega)$  at each  $\omega$  in  $0 \le \omega \le 1$  under suitable conditions (see especially Nevai, 1979). Sufficient conditions are that  $f(\omega)$  has finite dynamic range (and therefore is bounded above and below) and has a continuous derivative. When an estimator, denoted  $\hat{f}_m(\omega)$ , of  $f_m(\omega)$  is used as an estimator of  $f(\omega)$ , one has to take into account two kinds of errors, called respectively bias and variance. Bias is a measure of the deterministic difference between  $f_m(\omega)$  and  $f(\omega)$ , while variance is a measure of

the stochastic distance between  $\hat{f}_m(\omega)$  and  $f_m(\omega)$ . As m increases, bias decreases, while variance increases. This is an example of the fundamental problem of empirical spectral analysis which is how to achieve an optimal balance between bias and variance. When one uses autoregressive spectral estimation, this problem reduces to a question of determining the order m of the approximating autoregressive scheme, which is discussed in Section 6.

It should be noted that basic references for the mathematical properties (especially convergence) of autoregressive transfer functions  $g_m(z)$  are Geronimus (1960) and Grenander and Szegö (1958).

# 3. Entropy interpretation of autoregressive spectral densities

The use of autoregressive spectral densities as exact models, and as approximating models, for true spectral densities is often questioned by sceptical statisticians on the ground that their use in general is ad hoc and without theoretical justification. A possible answer to this criticism is provided by entropy concepts. This section reviews these concepts in order to state their application to spectral estimation.

The notion of entropy in statistics is usually first defined for a discrete distribution with probability mass function p(x). The entropy of this distribution, denoted H(p), is defined by

$$H(p) = -\sum_{x} p(x) \log p(x)$$
. (3.1)

For the distribution of a continuous real-valued random variable X, with probability density function f(x), entropy is defined (analogously or formally) by

$$H(f) = -\int_{-\infty}^{\infty} f(x) \log f(x) dx.$$
 (3.2)

A concept closely related to entropy is information divergence I(f; g) between two probability densities f(x) and g(x), defined by

$$I(f;g) = \int_{-\infty}^{\infty} \left\{ -\log \frac{g(x)}{f(x)} \right\} f(x) \, \mathrm{d}x.$$
 (3.3)

One should note that I(f; g) equals minus the generalized entropy H(f|g) defined by

$$H(f|g) = \int_{-\infty}^{\infty} \left\{ -\frac{f(x)}{g(x)} \log \frac{f(x)}{g(x)} \right\} g(x) dx.$$
 (3.4)

Another fundamental concept is cross-entropy defined by

$$H(f;g) = \int_{-\infty}^{\infty} \{-\log g(x)\} f(x) \, \mathrm{d}x.$$
 (3.5)

Note that H(f) = H(f; f).

Information divergence is expressed in terms of cross-entropy and entropy by

$$I(f;g) = H(f;g) - H(f)$$
. (3.6)

IMPORTANT INFORMATION INEQUALITY:

$$I(f;g) \ge 0 \tag{3.7}$$

with equality if and only if f = g; consequently

$$H(f) \le H(f;g). \tag{3.8}$$

Proof. Apply Jensen's inequality which states for an arbitrary function h(x)

$$\int_{-\infty}^{\infty} \{\log h(x)\} f(x) \, \mathrm{d}x \le \log \int_{-\infty}^{\infty} h(x) f(x) \, \mathrm{d}x \tag{3.9}$$

with equality if and only if h(x) = 1 for almost all x with respect to the probability density f(x).

Some applications of entropy in probability and statistical modeling are as follows.

The method of maximum likelihood parameter estimation can be described abstractly as follows. One introduces a parametric family of probability densities  $f_{\theta}(x)$ , indexed by a vector parameter  $\theta = (\theta_1, \dots, \theta_k)$ . Suppose there is a true parameter value  $\overline{\theta}$  in the sense that the true probability density  $f(x) = f_{\overline{\theta}}(x)$ . Then  $\overline{\theta}$  satisfies

$$H(f) = H(f; f_{\theta}) = \min_{\theta} H(f; f_{\theta}). \tag{3.10}$$

To estimate  $\bar{\theta}$  from data, one forms an estimator  $\tilde{H}(f; f_{\theta})$  of  $H(f; f_{\theta})$  and defines an estimator  $\hat{\theta}$  of  $\bar{\theta}$  by

$$\tilde{H}(f; f_{\hat{\theta}}) = \min_{\theta} \tilde{H}(f; f_{\theta}). \tag{3.11}$$

The estimator  $\tilde{H}(f; f_{\theta})$  could be of the form

$$\tilde{H}(f; f_{\theta}) = H(\tilde{f}; f_{\theta}) \tag{3.12}$$

for a suitable raw estimator  $\tilde{f}(x)$  of f(x).

The parametric families of probability densities  $f_{\theta}(x)$  are often derived axiomatically using a maximum entropy principle.

THEOREM 1. Fix k functions  $T_j(x)$ , j = 1, 2, ..., k, and k real numbers  $\tau_1$ ,  $\tau_2, ..., \tau_k$  such that there exists probability densities f(x) satisfying

$$\int_{-\infty}^{\infty} T_j(x)f(x) dx = \tau_j, \quad j = 1, \dots, k.$$
(3.13)

The density with maximum entropy H(f) among these densities is of the form

$$\log f_{\theta}(x) = \sum_{j=1}^{k} \theta_j T_j(x) - \Psi(\theta_1, \dots, \theta_k), \qquad (3.14)$$

where

$$\Psi(\theta_1,\ldots,\theta_k) = \log \int_{-\infty}^{\infty} \mathrm{d}x \, \exp \sum_{k=1}^k \theta_j T_j(x) \,, \tag{3.15}$$

and  $\theta_1, \ldots, \theta_k$  are chosen to satisfy

$$\int_{-\infty}^{\infty} T_j(x) f_{\theta}(x) \, \mathrm{d}x = \tau_j, \quad j = 1, \dots, k.$$
 (3.16)

PROOF. The theorem may be proved by calculus of variations arguments. A quick proof is to verify that for any f(x) satisfying the moment constraints (3.13)

$$H(f; f_{\theta}) = \Psi(\theta_1, \dots, \theta_k) - \sum_{j=1}^k \theta_j \tau_j = H(f_{\theta}),$$
 (3.17)

and therefore

$$H(f) \leq H(f; f_{\theta}) = H(f_{\theta}). \tag{3.18}$$

Thus the maximum entropy is achieved by  $f_{\theta}(x)$ .

Natural exponential models. A parametric family of probability densities  $f_{\theta}(x)$  is said to obey a natural exponential model when it is of the form (3.14). Thus natural exponential models are maximum entropy probability densities.

To extend entropy concepts to short-memory stationary zero-mean Gaussian time series, define the information divergence for a sample  $\{Y(t), t=1, 2, \ldots, T\}$  as a function of the true probability density f of the sample, and a model g for f. We define

$$I(f;g) = \lim_{T \to \infty} I_T(f;g), \qquad (3.19)$$

$$I_T(f;g) = \frac{-1}{T} E_f \left[ \log \frac{g(Y(1), \dots, Y(T))}{f(Y(1), \dots, Y(T))} \right].$$
 (3.20)

It should be noted that we are using the notation f and g with a variety of meanings. For a Gaussian zero-mean stationary time series, the probability density of the sample is specified by the spectral densities  $f(\omega)$  of the true distribution and  $g(\omega)$  of the model. The arguments of the information divergence I(f;g) indicate spectral densities in the following discussion. Pinsker (1963) derives the following very important formula:

$$I(f;g) = \frac{1}{2} \int_0^1 \left\{ \frac{f(\omega)}{g(\omega)} - \log \frac{f(\omega)}{g(\omega)} - 1 \right\} d\omega.$$
 (3.21)

Since  $u - \log u - 1 \ge 0$  for all u, I has two of the properties of a distance:  $I(f;g) \ge 0$ , I(f;f) = 0. However, I does not satisfy the triangle-inequality. We define the cross-entropy of spectral density functions  $f(\omega)$  and  $g(\omega)$  by

$$H(f;g) = \frac{1}{2} \int_0^1 \left\{ \log g(\omega) + \frac{f(\omega)}{g(\omega)} \right\} d\omega.$$
 (3.22)

The entropy of f is

$$H(f) = H(f; f) = \frac{1}{2} \int_0^1 \{\log f(\omega) + 1\} d\omega$$
 (3.23)

Information divergence can be expressed

$$I(f;g) = H(f;g) - H(f)$$
. (3.24)

Hence

$$H(f) \le H(f;g). \tag{3.25}$$

An approximating autoregressive spectral density of order m, denoted  $\bar{f}_m(\omega)$ , to a spectral density  $f(\omega)$  is defined by

$$H(f; \bar{f}_m) = \min_{f_m} H(f; f_m),$$
 (3.26)

where the minimization is over all  $f_m$  of the form

$$f_m(\omega) = \sigma_m^2 |g_m(e^{2\pi i\omega})|^{-2},$$
 (3.27)

$$g_m(z) = 1 + \alpha_m(1)z + \dots + \alpha_m(m)z^m$$
. (3.28)

One may verify that

$$H(f; f_m) = \frac{1}{2} \left\{ \log \sigma_m^2 + \frac{1}{\sigma_m^2} \int_0^1 g_m |(e^{2\pi i \omega})|^2 f(\omega) d\omega \right\}.$$
 (3.29)

The coefficients  $\hat{\sigma}_m^2$ ,  $\bar{\alpha}_m(1)$ , ...,  $\bar{\alpha}_m(m)$  of the minimum cross-entropy approximating autoregressive spectral density satisfy

$$\bar{\sigma}_{m}^{2} = \int_{0}^{1} |\bar{g}_{m}(e^{2\pi i\omega})|^{2} f(\omega) d\omega$$

$$= \int_{0}^{1} \bar{g}_{m}(e^{2\pi i\omega}) f(\omega) d\omega = \sum_{j=0}^{m} \hat{\alpha}_{m}(j) \rho(j)$$

$$= \min_{g_{m}} \int_{0}^{1} |g_{m}(e^{2\pi i\omega})|^{2} f(\omega) d\omega, \qquad (3.30)$$

$$\int_{0}^{1} \bar{g}_{m}(e^{2\pi i\omega}) e^{-2\pi ik\omega} f(\omega) d\omega = \sum_{j=0}^{m} \bar{\alpha}_{m}(j) \rho(j-k) = 0, \quad k = 1, 2, \dots, m.$$

$$(3.31)$$

Further,

$$H(f; \bar{f}_m) = \frac{1}{2} \{ \log \bar{\sigma}_m^2 + 1 \} = H(\bar{f}_m).$$
 (3.32)

The autoregressive spectral density  $\bar{f}_m(\omega)$  can be derived axiomatically using a maximum entropy principle.

THEOREM 2. The spectral density with maximum entropy among all spectral densities  $f(\omega)$  satisfying the constraints

$$\int_0^1 e^{2\pi i \omega j} f(\omega) d\omega = \rho(j), \quad j = 1, 2, \dots, m,$$
(3.33)

for m specified correlation coefficients  $\rho(1), \ldots, \rho(m)$  is  $\bar{f}_m(\omega)$ .

PROOF. It may be verified that  $\bar{f}_m(\omega)$  satisfies the constraints (3.33) and (3.32) holds for any  $f(\omega)$  satisfying (3.33). Since

$$H(f) \leq H(f; \bar{f}_m) = H(\tilde{f}_m), \tag{3.34}$$

it follows that  $\bar{f}_m$  has maximum entropy among all spectral densities satisfying the constraints (3.33).

The maximum entropy principle provides a motivation or justification for the use of autoregressive spectral estimators. However, the maximum entropy principle provides no insight into how to identify an optimal order m, or even what are the effects of different methods of estimating the parameters  $\sigma_m^2$ ,  $\alpha_m(1), \ldots, \alpha_m(m)$ . It provides no guidance for how to learn from the data whether the time series is nonstationary (long memory) or stationary (short memory), or whether the best time series model is AR, MA, or ARMA. It is a principle for deriving probability models, rather than statistically fitting models to data. Further, the maximum entropy principle justifies autoregressive esti-

mators only for short-memory time series. Autoregressive estimators are justified for long-memory time series by the fact that a pure harmonic  $Y(t) = A \cos 2\pi\omega t + B \sin 2\pi\psi t$  satisfies  $Y(t) - \phi Y(t-1) + Y(t-2) = 0$ , where  $\phi = 2\cos 2\pi\omega$ .

#### 4. Parametrizations of autoregressive spectral estimators

There are many ways for forming autoregressive spectral estimators, because there are four equivalent ways of parametrizing them.

# A. Autoregressive coefficients

Consider coefficients  $\alpha_m(1), \ldots, \alpha_m(m)$  such that  $g(z) = 1 + \alpha_m(1)z + \cdots + \alpha_m(m)z^m$  satisfies  $g(z) \neq 0$  for complex z such that  $|z| \leq 1$ . We call g(z) a minimum phase filter transfer function. The autoregressive coefficients determine  $\sigma_m^2$  by

$$1 = \sigma_m^2 \int_0^1 |g_m(e^{2\pi i\omega})|^{-2} d\omega.$$

One computes the correlation coefficients  $\rho(1), \ldots, \rho(m)$  by

$$\rho(j) = \int_0^1 \exp(2\pi i \omega j) \sigma_m^2 |g_m(e^{2\pi i \omega})|^{-2} d\omega.$$

#### B. Correlations

A set of m coefficients  $\rho(1), \ldots, \rho(m)$ , such that the matrix [with  $\rho(0) = 1$ ,  $\rho(-v) = \rho(v)$ ]

$$\begin{bmatrix} \rho(0) & \rho(-1) & \dots & \rho(1-m) \\ \rho(1) & \rho(0) & \dots & \rho(2-m) \\ \rho(m-1) & \rho(m-2) & \dots & \rho(0) \end{bmatrix}$$

is positive definite, is correlation coefficients of a time series. They determine autoregressive coefficients by solving the Yule-Walker equations [with  $\alpha_m(0) = 1$ ]

$$\sum_{j=0}^m \alpha_m(j)\rho(j-k)=0, \quad k=1,\ldots,m.$$

One computes  $\sigma_m^2$  by  $\sigma_m^2 = \sum_{j=0}^m \alpha_m(j)\rho(j)$ .

# C. Partial correlations

Consider coefficients  $\pi(1), \ldots, \pi(m)$  satisfying  $|\pi(j)| < 1, j = 1, 2, \ldots, m$ . They represent partial correlation coefficients defined theoretically by

$$\pi(j) = \operatorname{corr}[Y(t), Y(t-j)|Y(t-1), \dots, Y(t-j+1)].$$

Partial correlation coefficients determine autoregressive coefficients and residual variances by a recursive algorithm called the Levinson-Durbin recursion (see Levinson, 1947; Durbin, 1960): for k = 1

$$\alpha_1(1) = -\pi(1), \qquad \sigma_1^2 = 1 - \pi^2(1),$$

while for  $k = 2, 3, \ldots, m$ 

$$\alpha_{k}(k) = -\pi(k) = \frac{-1}{\sigma_{k-1}^{2}} \sum_{j=0}^{k-1} \alpha_{k-1}(j) \rho(k-j),$$

$$\sigma_{k}^{2} = \sigma_{k-1}^{2} \{1 - \pi^{2}(k)\},$$

$$\alpha_{k}(j) = \alpha_{k-1}(j) - \pi(k) \alpha_{k-1}(k-j).$$

Autoregressive coefficients determine partial correlation coefficients by the recursion (see Barndorff-Nielssen and Schon, 1973)

$$\alpha_{k-1}(j) = \{1 - \pi^2(k)\}^{-1}\{\alpha_k(j) + \pi(k)\alpha_k(k-j)\}.$$

#### D. Residual variances

Consider coefficients  $\sigma_1^2, \sigma_2^2, \ldots, \sigma_m^2$  satisfying

$$1 > \sigma_1^2 > \sigma_2^2 > \cdots > \sigma_m^2 > 0$$
,

and m coefficients representing sign  $\pi(1), \ldots$ , sign  $\pi(m)$ . The  $\sigma$ 's represent residual variances; they determine partial correlation coefficients by a formula noted by Dickenson (1978)

$$\pi(k) = \operatorname{sign} \pi(k) \left[ 1 - \frac{\sigma_k^2}{\sigma_{k-1}^2} \right]^{1/2}.$$

#### 5. Empirical autoregressive spectral estimation

Given a sample  $\{Y(t), t = 1, 2, ..., T\}$  of a zero-mean Gaussian stationary time series whose spectral density  $f(\omega)$  is to be estimated by an autoregressive spectral density estimator

$$\hat{f}_m(\omega) = \hat{\sigma}_m^2 |\hat{g}_m(e^{2\pi i \omega})|^{-2},$$

$$\hat{g}_m(z) = 1 + \hat{\alpha}_m(1)z + \dots + \alpha_m(m)z^m,$$

we define the order identification problem to be the choice of order m, and the parameter estimation problem to be the choice of algorithm for computing the coefficients  $\hat{\alpha}_m(1), \ldots, \hat{\alpha}_m(m)$  and the residual variance  $\hat{\alpha}_m^2$ .

For a sample  $Y(1), \ldots, Y(t)$  of a zero-mean Gaussian stationary time series, an approximation for the joint probability density function  $f_{\theta}(Y(1), \ldots, Y(T))$  indexed by a parameter  $\theta$  is obtained as follows. We assume that the time series Y(t) has been divided by  $\{R(0)\}^{1/2}$  so that its covariance function equals its correlation function. Then

$$-2\log f_{\theta}(Y(1),\ldots,Y(T)) = \log(2\pi)^{T} \det K_{\theta} + Y_{T}^{*}K^{-1}Y_{T},$$

where \* denotes complex conjugate transpose,  $Y_T^* = (Y(1), \ldots, Y(T))$ ,  $K_{\theta} = EY_TY_T^*$  is a covariance matrix with (s, t)-element equal to  $\rho_{\theta}(s - t)$ . The subscript  $\theta$  on  $\rho_{\theta}(v)$  and  $f_{\theta}(\omega)$  indicates that they are functions of the parameters  $\theta$ , which are to be estimated.

The covariance matrix  $K_{\theta}$  is a Toeplitz matrix. Asymptotically, as T tends to  $\infty$ , all T by T Toeplitz matrices have the same eigenvectors  $\exp(-2\pi itj/T)$ ,  $j=0,1,\ldots,T-1$ . The corresponding eigenvalues are  $f_{\theta}(j/T)$ . An approximation for likelihood function frequently adopted is therefore

$$-\frac{1}{T}\log f_{\theta}(Y(1), \ldots, Y(T)) = \frac{1}{2}\log 2\pi + \frac{1}{2}\int_{0}^{1} \left\{\log f_{\theta}(\omega) + \frac{\tilde{f}(\omega)}{f_{\theta}(\omega)}\right\} d\omega$$
$$= \frac{1}{2}\log 2\pi + H(\tilde{f}; f_{0}),$$

where  $\tilde{f}(\omega)$  is the sample spectral density defined by

$$\tilde{f}(\omega) = \sum_{t=1}^{T} |Y(t) \exp(-2\pi i t\omega)|^2 \div \sum_{t=1}^{T} Y^2(t).$$

Maximum likelihood estimators  $\hat{\theta}$  are asymptotically equivalent to the estimators  $\hat{\theta}$  minimizing the sample cross-entropy  $H(\tilde{f}; f_{\theta})$ .

If the parametric model  $f_{\theta}(\omega)$  is assumed to be the spectral density of an AR(p), then estimators  $\hat{\sigma}_{p}^{2}$ ,  $\hat{\alpha}_{p}(1)$ , ...,  $\hat{\alpha}_{p}(p)$  of the coefficients satisfy Yule–Walker equations corresponding to the sample correlation function

$$\hat{\rho}(v) = \int_0^1 e^{2\pi i\omega v} \tilde{f}(\omega) d\omega$$

$$= \sum_{t=1}^{T-v} Y(t)Y(t+v) \div \sum_{t=1}^T Y^2(t).$$

The sample correlation function  $\hat{\rho}(v)$  can be computed, using the fast Fourier transform, by

$$\hat{\rho}(v) = \frac{1}{Q} \sum_{k=0}^{Q-1} \exp\left(2\pi i \frac{k}{Q} v\right) \tilde{f}\left(\frac{k}{Q}\right),$$

which holds for  $0 \le v < Q - T$ .

It should be noted that we are assuming the time series Y(t) to be zero mean, or more generally to have been detrended by subtraction of  $\hat{\mu}(t)$ , an estimator of the mean-value function  $\mu(t) = E[Y(t)]$ . When  $\mu(t) = \mu$ , a constant, we take  $\hat{\mu} = \overline{Y}$ . When  $\mu(t)$  is a function with period d (as might be the case with d = 12 for monthly time series), one might take for  $\hat{\mu}(t)$  the mean of Y(s) for all s = t modulo d.

By recursively solving the Yule-Walker equations, one can determine sequences of

(1) estimated residual variances

$$1 > \hat{\sigma}_1^2 > \hat{\sigma}_2^2 > \cdots > \hat{\sigma}_m^2 \ldots$$

(2) estimated partial correlations

$$\hat{\pi}(1), \hat{\pi}(2), \ldots, \hat{\pi}(m), \ldots,$$

(3) estimated autoregressive coefficients

$$\hat{\alpha}_m(0) = 1, \ \hat{\alpha}_m(1), \ldots, \hat{\alpha}_m(m),$$

(4) autoregressive spectral density estimators

$$\hat{f}_m(\omega) = \hat{\sigma}_m^2 \left| \sum_{i=0}^m \hat{\alpha}_m(j) \exp 2\pi i j \omega \right|^{-2},$$

(5) residual spectral densities

$$\tilde{f}_m(\omega) = \frac{\tilde{f}(\omega)}{\hat{f}_m(\omega)}.$$

By forming a smoothed version  $\hat{f}_m(\omega)$  of  $\tilde{f}_m(\omega)$ , one can obtain a final estimator  $\hat{f}(\omega)$  of the unknown spectral density:

$$\hat{f}(\omega) = \hat{f}_m(\omega)\hat{f}_m(\omega).$$

When  $\tilde{f}(\omega)$  is tested for white noise, and found not to be significantly different from white noise, then

$$\hat{f}(\omega) = \hat{f}_m(\omega) \,,$$

and the autoregressive spectral density estimator is the final estimator.

The important question of criteria for choosing the orders of approximating spectral densities is discussed in the next section.

Computing estimators of autoregressive coefficients by solving Yule-Walker equations is called *stationary autoregression* because the autoregressive coefficients obtained are guaranteed-to correspond to a stationary time series. When  $\hat{\sigma}_m^2$  in the foregoing analysis is tending to approximate 0, we consider the time series to be long memory; experimental evidence indicates that more reliable estimators of the spectral density, and also of the autoregressive coefficients, are provided by *least-squares autoregression*, which solves the normal equations

$$\begin{bmatrix} \hat{K}(0,0) & \dots & \hat{K}(0,m) \\ \hat{K}(1,0) & \dots & \hat{K}(1,m) \\ \dots & \dots & \dots \\ \hat{K}(m,0) & \dots & \hat{K}(m,m) \end{bmatrix} \begin{bmatrix} 1 \\ \hat{\alpha}_m(1) \\ \dots \\ \hat{\alpha}_m(m) \end{bmatrix} = \begin{bmatrix} \hat{\sigma}_m^2 \\ 0 \\ \dots \\ 0 \end{bmatrix}$$

for a suitable estimator  $\hat{K}(i, j)$  of

$$K(i, j) = E[Y(t-i)Y(t-j)].$$

Possible estimators (for i, j = 0, 1, ..., m) are least-squares forward algorithm

$$\hat{K}(i,j) = \frac{1}{T-M} \sum_{t=0}^{T-m-1} Y(t+m-i)Y(t+m-j),$$

or least-squares forward and backward algorithm

$$\hat{K}(i,j) = \frac{1}{2(T-M)} \sum_{t=0}^{T-m-1} \{Y(t+m-i)Y(t+m-j) + Y(t+i)Y(t+j)\}.$$

When several harmonics are present in the data, whose frequencies are close together, least-squares autoregressive coefficient estimators are more effective than Yule-Walker autoregressive coefficient estimators in providing autoregressive spectral estimators which exhibit the split peaks one would like to see in the estimated spectral density.

An important and popular algorithm for estimation of AR coefficients was introduced by Burg in 1967 (see Burg, 1967, 1968). For references to descriptions of Burg's algorithm, see Kay and Marple (1981).

#### 6. Autoregressive order determination

The problem of determining the orders of approximating autoregressive schemes is an important example of the problem of estimating a function by using the smallest finite number of parameters which provide an adequate approximation of the function. The true spectral density is denoted  $f(\omega)$  or  $f_{\infty}(\omega)$ . An approximation  $\bar{f}(\omega)$  is defined by assuming a family of densities  $f_{\theta_1,\ldots,\theta_m}(\omega)$  which are functions of m scalar parameters  $\theta_1,\ldots,\theta_m$ . The parameter values  $\bar{\theta}_1,\ldots,\bar{\theta}_m$  which minimize the cross-entropy  $H(f;f_{\theta_1,\ldots,\theta_m})$  define a best approximating spectral density  $\bar{f}_m(\omega) = f_{\bar{\theta}_1,\ldots,\bar{\theta}_m}(\omega)$ . An estimator of  $f_m$  is  $\hat{f}_m(\omega) = f_{\bar{\theta}_1,\ldots,\bar{\theta}_m}(\omega)$ , where  $\hat{\theta}_1,\ldots,\hat{\theta}_m$  minimizes  $H(f;f_{\theta_1,\ldots,\theta_m})$ .

To evaluate the properties of  $\hat{f}_m(\omega)$  as an estimator of  $f_{\infty}(\omega)$ , one must distinguish two kinds of error. The model approximation or bias error is

$$B(m) = I(f_{\infty}; \overline{f}_m)$$
.

The parameter estimation error or variance is

$$V(m, T) = EI(\bar{f}_m; \hat{f}_m).$$

As m tends to  $\infty$ , B(m) tends to 0 and V(m, T) tends to  $\infty$ . The optimal value  $\hat{m}$  minimizes  $EI(f_{\infty}; \hat{f}_{m})$  as measured by

$$C(m) = B(m) + V(m, T).$$

In practice, one forms an estimator  $\hat{C}(m)$  of the terms in C(m) which depend on m.

One calls  $\hat{C}(m)$  a criterion function for order determination. It should be plotted, and interpreted as a function, not just examined for its minimum value. It is useful to define a best value of m (at which  $\hat{C}(m)$  is minimized) and a second best value of m (at which  $\hat{C}(m)$  achieves its lowest relative minimum).

One also has to define a value  $\hat{C}(0)$  of the criterion function at m = 0. If

$$\hat{C}(m) > \hat{C}(0)$$
, for  $m = 1, 2, ...,$ 

then the optimum order is 0, and the time series is considered to be not significantly different from white noise. Further research is required on the properties of order determining criteria as tests for white noise.

Tests for white noise provide an alternative approach to order determination since an autoregressive estimator  $\hat{f}_m(\omega)$  is regarded as an adequate fit (or smoother) if the residual spectral density  $\tilde{f}(\omega) \div \tilde{f}_m(\omega)$  is not significantly different from the sample spectral density of white noise.

A widely used order determining criterion is that introduced by Akaike (1974). It should be emphasized that Akaike's criterion had a different conceptual basis than the one outlined above; it seeks to determine the order of an

exact autoregressive scheme which the time series is assumed to obey. Then one can raise the objection against it that it does not consistently estimate the order, which is done by a criterion due to Hannan and Quinn (1979). Our point of view is that the approximating autoregressive scheme need only have the property that  $\tilde{f}(\omega) \div \hat{f}_m(\omega)$  is just barely not significantly different from the sample spectral density of white noise.

Akaike's order determining criterion AIC is defined by

$$AIC(m) = \log \hat{\sigma}_m^2 + \frac{2m}{T}, \quad m \ge 1.$$

Possible definitions for AIC(0) are 0 or -1/T. The Hannan and Quinn criterion is

AICHQ(m) = 
$$\log \hat{\sigma}_m^2 + \frac{m}{T} \log \log T$$
.

Parzen (1974, 1977) introduced an approximating autoregressive order criterion called CAT (criterion autoregressive transfer function), defined by

$$\begin{aligned} \mathrm{CAT}(m) &= \frac{1}{T} \sum_{j=1}^m \left( 1 - \frac{j}{T} \right) \hat{\sigma}_j^{-2} - \left( 1 - \frac{m}{T} \right) \hat{\sigma}_m^{-2}, \quad m \geq 1 \;, \\ \mathrm{CAT}(0) &= - \left( 1 + \frac{1}{T} \right). \end{aligned}$$

In practice, CAT and AIC lead in many examples to exactly the same orders. It appears reassuring that quite different conceptual foundations can lead to similar conclusions in practice.

#### 7. Suggestions for empirical spectral analysis

The basic aim of spectral analysis is to obtain an estimated spectral density which does not introduce spurious spectral peaks, and resolves close spectral peaks. To arrive at the final form of spectral estimator in an applied problem, autoregressive spectral estimators can be used to identify the memory type of a time series (long, short, or no memory) and the type of the whitening filter of a short-memory time series (AR, MA, or ARMA). An empirical time-series spectral analysis should involve the following stages.

# A. Preprocessing

To analyze a time-series sample Y(t), t = 1, ..., T, one will proceed in stages which often involve the subtraction of or elimination of strong effects in order to see more clearly weaker patterns in the time-series structure. The aim of

preprocessing is to transform  $Y(\cdot)$  to a new time series  $\bar{Y}(\cdot)$  which is short memory. Some basic preprocessing operations are memoryless transformation (such as square root and logarithm), detrending, 'high-pass' filtering, and differencing. One usually subtracts out the sample mean  $\bar{Y} = (1/T) \sum_{t=1}^{T} Y(t)$ ; then the time series actually processed is  $Y(t) - \bar{Y}$ .

# B. Sample Fourier transform by data windowing, extending with zeroes, and fast Fourier transform

Let Y(t) denote a preprocessed time series. The first step in the analysis could be to compute successive autoregressive schemes using operations only in the time domain. An alternative first step is the computation of the sample Fourier transform

$$\tilde{\psi}(\omega) = \sum_{t=1}^{T} Y(t) \exp(-2\pi i \omega t), \quad \omega = \frac{k}{Q}, \ k = 0, 1, ..., Q - 1,$$
 (7.1)

at an equispaced grid of frequencies in  $0 \le \omega \le 1$ . We call Q the spectral computation number. One should always choose  $Q \ge T$ , and we recommend  $Q \ge 2T$ . Prior to computing  $\tilde{\psi}(\omega)$ , one should extend the length of the time series by adding zeroes to it. Then  $\tilde{\psi}(\omega)$  can be computed using the fast Fourier transform.

If the time series may be long memory, one should compute in addition a sample 'tapered' or 'data-windowed' Fourier transform

$$\tilde{\psi}_{W}(\omega) = \sum_{t=1}^{T} Y(t)W\left(\frac{t}{T}\right) \exp(-2\pi i \omega t). \tag{7.2}$$

# C. Sample spectral density

The sample spectral density  $\tilde{f}(\omega)$  is obtained essentially by squaring and normalizing the sample Fourier transform:

$$\tilde{f}(\omega) = \frac{|\tilde{\psi}(\omega)|^2}{\frac{1}{Q} \sum_{k=0}^{Q-1} |\tilde{\psi}(\frac{k}{Q})|^2}, \quad \omega = \frac{k}{Q}, \quad k = 0, 1, \dots, Q-1.$$
 (7.3)

# D. Sample correlation function

The sample correlation function  $\hat{\rho}(v)$  is computed (using the fast Fourier transform).

#### E. Autoregressive analysis

The Yule-Walker equations are solved to estimate innovation variances  $\hat{\sigma}_m^2$ , to which are applied order determining criteria (AIC, CAT) to determine optimal orders  $\hat{m}$  and also to test for white noise. The value of  $\hat{\sigma}_m^2$  and the dynamic

range of the autoregressive spectral estimator  $\hat{f}_{\hat{m}}(\omega)$  are used to determine the memory type of the time series. Two orders (called the best  $\hat{m}$  and second best  $\hat{m}(2)$ ) are determined as candidates as optimal orders corresponding to the absolute minimum and lowest relative minimum of the criterion function.

#### F. ARMA analysis

When a time series is classified as short memory, an approximating AR scheme of order  $4\hat{m}$  is inverted to form  $MA(\infty)$  coefficients which are used to estimate covariance matrix of Y(t-j) and  $Y^{\nu}(t-k)$ . A subset regression procedure is then used to determine a 'best-fitting' ARMA scheme, and the corresponding ARMA spectral density estimator. One will be able to identify moving average schemes and ARMA schemes which are barely invertible, and require a long AR scheme for adequate approximation. The long AR spectral estimator introduces spurious spectral peaks when compared to the MA or ARMA estimator.

# G. Nonstationary autoregression

When a time series is classified as long memory, more accurate estimators of autoregressive coefficients are provided by minimizing a least-squares criterion or by Burg estimators. When several harmonics are present in the data, whose frequencies are close together, least-squares autoregressive coefficient estimators are more effective than Yule-Walker autoregressive coefficient estimators in providing autoregressive spectral estimators which exhibit the split peaks one would like to see in the estimated spectral density.

#### H. Long-memory analysis

In the long-memory case, one may want to represent Y(t) as S(t) + N(t), a long-memory signal plus short-memory noise. An approach to this problem may be provided by treating the sample spectral density values  $\tilde{f}(k/Q)$  as a data batch to be studied by nonparametric data modeling methods using quantile functions (see Parzen, 1979). The details of such methods are under development.

# I. Nonparametric kernel spectral density estimator

An estimator  $\hat{f}(\omega)$  of the spectral density is called *parametric* when it corresponds to a parametric model for the time series (such as an AR or ARMA model), *nonparametric* otherwise. A general form of the nonparametric estimator is the kernel estimator

$$\hat{f}(\omega) = \sum_{v=-\infty}^{\infty} k\left(\frac{v}{M}\right) \hat{\rho}(v) e^{-2\pi i \omega v}, \quad 0 \le \omega \le 1.$$

The problem of determining optimum truncations points M has no general

solution; one approach is to choose  $M = 4\hat{m}$  to obtain a preliminary smoothing of the sample spectral density.

# J. Inverse correlations and cepstral correlations

Estimators of  $\rho i(v)$  and  $\gamma(v)$  are computed and used to form nonparametric kernel estimators of  $f^{-1}(\omega)$  and  $\log f(\omega)$ , which may provide additional insights into the peaks and troughs to be given significance in the final estimator of the spectrum.

Extensive comparisons of different methods of spectral estimation are given in Pagano (1980), Beamish and Priestley (1981), and Kay and Marple (1981). It seems clear that autoregressive spectral estimators can give superior results when properly used. One should: determine two best orders; compute autoregressive coefficients by Yule-Walker equations and by least squares since when the time series is long-memory autoregressive spectral estimators are most accurate when based on least-squares estimators of autoregressive coefficients; use approximating autoregressive schemes to determine if an ARMA scheme fits better.

The end of the story of the search for the perfect spectral estimator seems attainable if one does not think of spectral estimation as a nonparametric procedure which can be conducted independently of model identification.

#### 8. A bibliography of autoregressive spectral estimation

The references aim to provide a comprehensive list of the publications in English which are directly concerned with developing the theory and methods of autoregressive spectral estimation.

This section lists some of the publications which contributed to the development of AR spectral estimation.

Yule (1972) introduces autoregressive schemes to model disturbed periodicities as an alternative to Schuster periodogram analysis and its spurious periodicities; Yule-Walker (1931) equations relate autoregressive coefficients and correlations of a stationary time series.

Wold (1938) introduces infinite-order autoregressive and moving average representations of a stationary time series; rigorous conditions are given by Akutowicz (1957).

Mann and Wald (1943) derive asymptotic distribution of estimators of autoregressive coefficients.

Levinson (1947) and Durbin (1960) derive recursive methods of solving Yule—Walker equations which subsequently lead to fast algorithms for calculation of high-order AR schemes.

Whittle (1954) seems to be the first to use autoregressive schemes to estimate a spectral density. He used a low-order model in a case where high-order models are indicated by order determining criterion (Akaike, 1974, p. 720).

- Grenander and Rosenblatt (1956) criticize attempts to apply low-order autoregressive schemes, and develop theory of nonparametric spectral density estimation, as do Bartlett, Parzen, and Tukey and Blackman.
- Parzen (1964), Schaerf (1964) and Parzen (1968) discuss autoregressive spectral estimation as a method for empirical time-series analysis; no theory is given.
- Burg (1967, 1968) publishes his pioneering work on MEM (maximum entropy method of spectral estimation) and his method of calculating their coefficients.
- Akaike (1969a,b, 1970) derives asymptotic variance formulas for autoregressive spectral estimators, and states FPE (final predictor error) criterion for order determination; precursor of FPE in Davisson (1965).
- Parzen (1969) derives heuristically a formula for the asymptotic variance of AR spectral estimators, confirmed by Kromer (1969) and Berk (1974); an order determining criterion is proposed.
- Kromer (1969) in an unpublished Ph.D. thesis presents first rigorous analysis of asymptotic distribution of autoregressive spectral estimators, especially their bias; consistency is proved only in an iterated limit mode of convergence.
- Berk (1974) provides first proof of consistency of autoregressive spectral estimators.
- Carmichael (1976) in an unpublished Ph.D. thesis provides alternative proof of consistency of autoregressive estimators, and extends technique to general problems of density estimation.
- Akaike (1973, 1974, 1977) introduces AIC for model order criterion and relates it to entropy maximization principles.
- Parzen (1974, 1977) introduces CAT for AR order determination based on concept of finite parameter AR schemes as approximations to infinite parameter AR schemes.
- Hannan and Quinn (1979) derive a modification of AIC which provides consistent estimators of the AR order, when exact model is assumed to be a finite order AR.
- Huzii (1977), Shibata (1977) and Bhansali (1980) discuss rigorously the convergence of AR spectral estimators and inverse correlations.
- Childers (1978) and Haykin (1979) contain very useful collections of papers.
- Pagano (1980), Beamish and Priestley (1981) and Kay and Marple (1981) provide illuminating reviews of AR spectral estimators and comparisons with alternative methods.

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