Observations of Waves Generated by a 20keV Electron Beam in a Laboratory

Plasma

UCLA Plasma Sciences **Technology Institute**

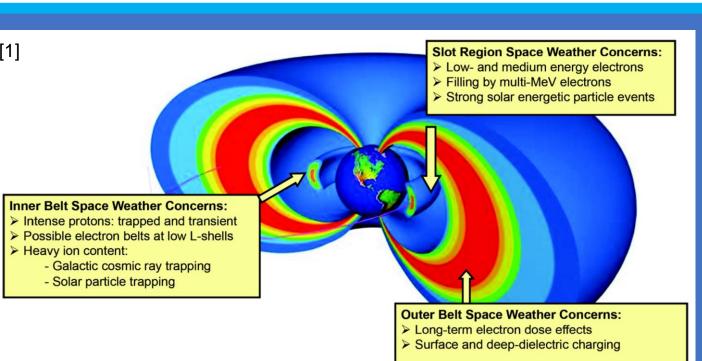
Jesus Perez ¹, Seth Dorfman ¹, Troy Carter ¹

¹University of California Los Angeles, CA

Key Points

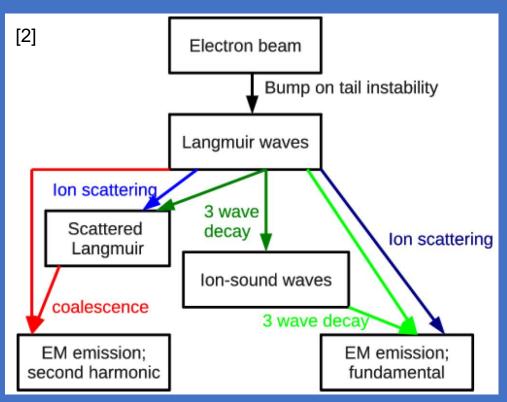
- Understanding fundamental plasma processes such as wave-particle interactions is of great importance to many of the subfields in plasma physics.
- Laboratory plasma experiments are a great way to gain insight on astrophysical phenomena.
- The proposed study aims to determine the efficacy of generating whistler waves in UCLA's Large Plasma Device(LAPD) in order to further our understanding of the wave-particle interactions present in our solar system.

Motivation



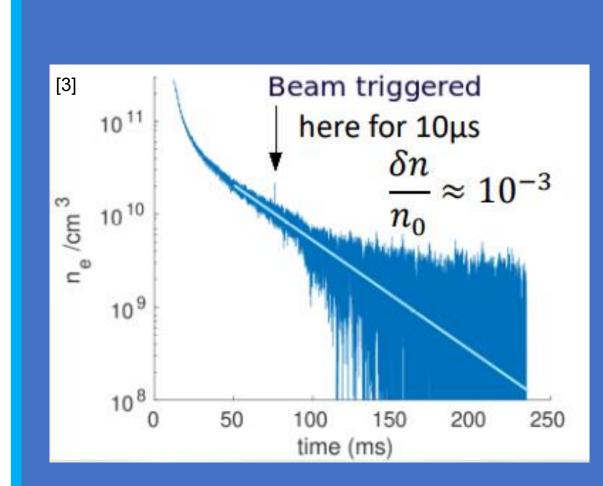
High energy electrons from either solar wind or from human caused high altitude nuclear explosions may become trapped inside the Van Allen radiation belts and persist there for long periods of time. Spacecrafts in the regions may be suspectable to damage from these trapped electrons. A proposed solution is using spacecraft to carry compact electron beams or antennas to remediate the trapped electrons.

Type II and Type III radio burst are an important type of solar outputs as they are generated by beams of electrons and can be used to study electron acceleration, energy transport and the plasmas in which they travel through. However, the generation mechanism of these burst requires further understanding



Plasma Conditions

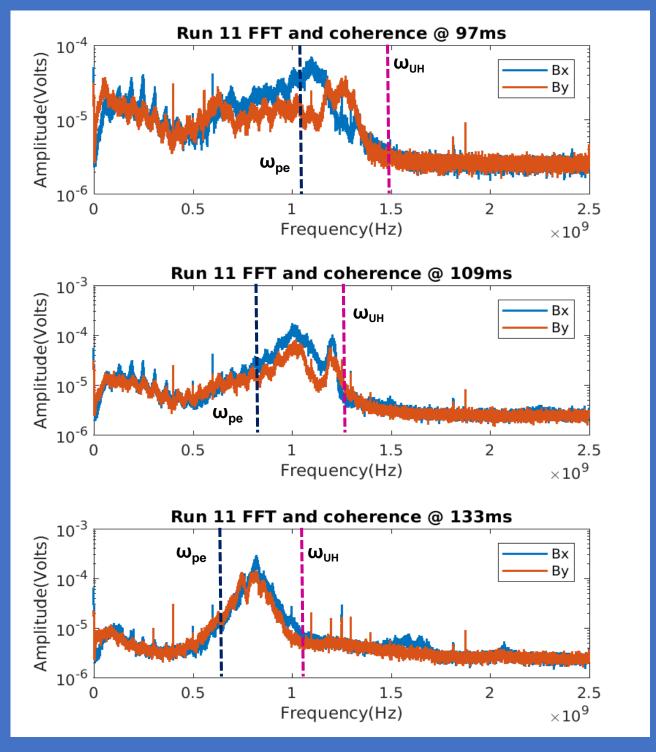




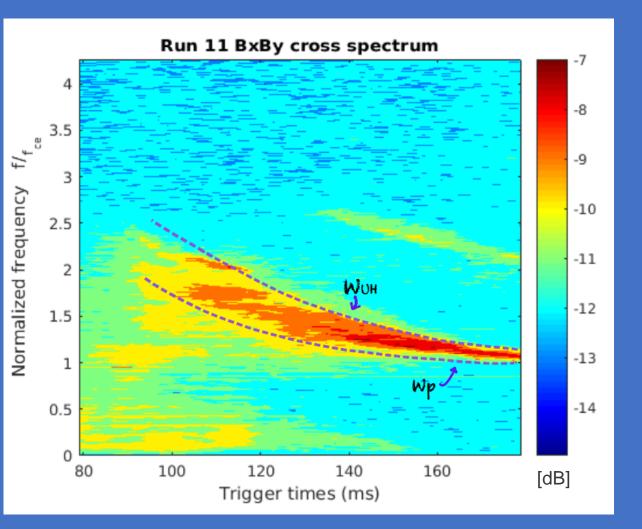
Electron beam is triggered at varying times after the plasma discharges as a way of varying the density of the plasma.

Experimental Setup LAPD ☐ 1 EB probe in, measuring Bx, By, Ex, and Ey LAPD cathode \square Probe is positioned at (2,7.75) \leftarrow (xy in cm) & and port $31 \rightarrow z = 7m$ **Electron** ☐ Electron gun triggering times range from 79gun 179ms EB probe @ port 31 (in 2ms increments)

Results & Discussion



Earlier gun trigger times show more activity between the whistler (low frequency) and X-mode region.



Looking at the cross spectrum of the Bx & By illuminates the activity between the whistler and X-mode frequency range. This activity suggest potential non-linear 3 wave coupling.

Bicoherence is used to measure the $\sum F_n(f_1)F_n(f_2)F_n^*(f_1+f_2)$ degree of three-wave coupling. $b(f_1,f_2) = rac{\prod_{n=1}^{n} |F_n(f_1)F_n(f_2)F_n^*(f_1+f_2)|}{\sum_{n=1}^{n} |F_n(f_1)F_n(f_2)F_n^*(f_1+f_2)|}$ (Take $F_n(f_i)$ to be the Fourier transform) Bicoherence for Bx @ 97ms Noise level ≈ 0.20 There appears to be some activity present in the low frequency region. Recall: $f = f_1 + f_2$ \leftarrow frequencies Normalized Frequency f/f Bicoherence for Bx @ 131ms Noise level ≈ 0.20

of the signal. One example from this grid patter is taking $f_1 \stackrel{\approx}{=} 1.2$, $f_2 \stackrel{\approx}{=} 0.2$, so f = 1.4. (These frequencies are plotted above)

the plasma frequency

frequencies mapping out to activity in the frequency space

Grid patterns pop up with some of the indicated

Normalized Frequency f/fce Signal in Frequency Domain @ 157ms Bicoherence for Bx @ 157ms Here in later trigger times there are indications of potential 3 wave coupling with the second harmonic of

Noise level ≈ 0.20

Normalized Frequency f/

Conclusions

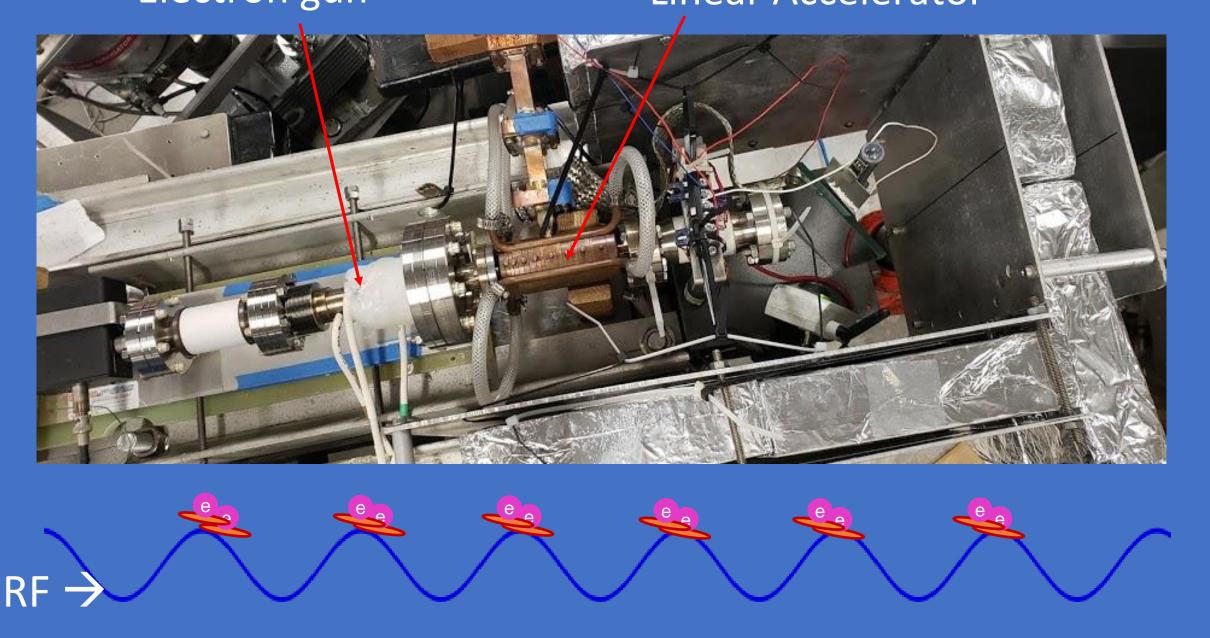
- ☐ Robust wave emissions in the X-band mode.
- ☐ Activity in the low frequency region suggestive of Whistler wave presence.
- ☐ Results indicate that 3 wave coupling between the whistler and X-mode region are possible.
- ☐ Bicoherence analysis does suggest some interactions between waves at the plasma frequency may lead to emissions at the second harmonic.

Future Work

- ☐ Further bicoherence analysis is needed. Observations of the By and the electric field underway.
- ☐ Further investigation into information about the wave vector at the second harmonic.
- ☐ For practical application in radiation belt remediation, a 1MeV electron beam is required. Thus early stages of this beam in LAPD is currently under way.

Electron gun

Linear Accelerator



Acknowledgments

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References

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- [2] Baker, D.N., Erickson, P.J., Fennell, J.F. et al. Space Weather Effects in the Earth's Radiation Belts. Space Sci Rev 214, 17 (2018). https://doi.org/10.1007/s11214-017-0452-7
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