## Bates College

### SENIOR THESIS

# Physics Simulations Using Javascript

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Presented to The Department of Physics, Bates College In Partial Fulfillment of the Requirements for the Degree of Bachelor of Science

> Lewiston, Maine December 1st, 2014



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## Abstract

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#### Physics Simulations Using Javascript

by Peter Krieg

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# Abbreviations

**HTML** HyperText Markup Language

JS JavaScript

**API** Application Programming Interface

Dedicated to my parents

## Introduction

### 0.1 What is a Physics Simulation?

The purpose of this thesis is to present a series of physics simulations, each modeling a specific problem of physics as realistically as possible. These simulations differ from animations, which can be seen as predictable representations that always display the same visual. Animations are analogous to a movie script: no matter how many times you watch the movie, it will always end in the same way. Simulations, on the other hand, need to adapt to variable conditions, and be based partly on random processes. This brings up the topic of dynamic vs. static animation. Most of the physics simulations in this thesis will be dyamic because they present a unique viewing each time they are run, and can also involve user feedback which influences the outcome of the simulation.

Any simulation requires creating the illusion of motion. Almost every form of projected motion media uses frames to accomplish this. Researchers have shown that to make the simulation look realistic, it must be presented at a rate of around 60-100 frames per second. Anything far slower than this, and the human eye will detect the "choppiness" of the simulation. People can't detect anything much faster than 100 frames per second, so there is no need to project media faster than that, with the exception of slow-motion videography.

#### 0.2 Methods of Producing a Simulation

The physics simulations in this thesis differ greatly from common animations. Movies and cartoons, for example, operate by displaying a series of images similar to one another, and displaying them as many frames per second to create the illusion of motion. My simulations, on the other hand, function by providing the *information* for each frame, and then providing the data for *how* the animation can be created. These instructions are passed onto the HTML5 canvas API, which creates the visual which can be seen in the web browser. Because physics simulations contain instructions instead of a series of images, the files of code take up far less space than a movie file would, for example. This is one primary advantage of coded simulations. Every simulation follows a similar set of steps, which can be simplified below:



Figure 1: The frames of a general simulation

The canvas API gets the initial state of the simulation, which could the position of a ball, for example. Then, the frame is rendered by applying rules to the canvas element, and changing the initial state of the simulation. Once the rules have been applied, and all conditions are satisfied, the frame is rendered, and then displayed on the canvas element, to be seen in the web browser. The canvas is embedded into a web page with the ¡canvas¡ tag, like any other HTML tag. The positioning of objects in the canvas element is specified with a coordinate system, which uses pixels as its unit. Figure 2 shows the orientation of the canvas, which differs from the traditional cartesian coordinate system.



FIGURE 2: The canvas coordinate system. The canvas is displayed on the screen as an invisible white rectangle by default. A sample point of (20,30) is shown for clarification.

To produce any realistic simulation, the steps in figure 1 must be repeated multiple times per second. In fact, these steps must be repeated 60 times per second to achieve the desired 60 frames per second outlined in the previous section. Luckily, the canvas API is capable of running the instructions very quickly to make this simulation possible.

#### 0.2.1 The Code

To program the simulations in this thesis, I chose to write the code in javascript. This scripting language is easy to view in any modern browser: therefore, all the simulations of this thesis can be viewed online. Javascript combines seamlessly with HTML5, which is why I mostly decided to use it for this thesis. The evolution of HTML (HyperText Markup Language) has progressed from simple web documents to complex web applications. For this thesis, every simulation utilizes the HTML5 <canvas>element, which has been used since around 2011. The HTML5 canvas API allows programmers to write javascript code that accesses the element and runs visual displays through a web browser The HTML needed to include a canvas can seen below:

10 | </html>

Listing 1: The bare bones code necessary for an HTML document to include the canvas element

The above code displays the most basic HTML combined with javascript necessary to begin any simulation. Lines 7-8 are the only ones that actually contain javascript: this is the simple step necessary for the canvas API to recognize the HTML document. These two steps are necessary for any physics simulation. The step on line 7 initializes a JS variable and sets it equal to the canvas element on the web document object. The second step

All web browsers include some form of javascript interpreter: whenever the browser encounters a <script >element, it "passes" the code onto the JS interpreter. In listing 1, the HTML and JS code are written in the same document for clarity. While this is an acceptable practice, all future simulations will involve the HTML referencing to external JS documents to keep the contents separate. The appendices in this thesis have full code files from all the simulations.

While this thesis can contain code excerpts, figures, and screen-shots of various simulations, it obviously can't contain the flow of images itself. Therefore, I have put the entire thesis and its simulations on my personal website, which can be found at: http://www.peterkrieg.com/thesis. You can navigate by each chapter and view the simulations outlined in thesis.

## Chapter 1

## Some Basic Simulations

While the introduction outlined the computer programming necessary to produce simulations in general, this chapter will start to deal with the physics necessary to make simulations seem realistic. In this chapter, I will outline some examples of simulations with balls bouncing, and discussing the basic mechanics involved through the code.

### 1.1 Basic Ball Bouncing

A ball bouncing will show the basic kinematic equations, and how they are used in the javascript code. The following example displays a ball being dropped with an initial  $v_x$ , and bouncing off the walls and floor of the canvas element. The full code is shown

below:

```
var canvas = document.getElementById('canvas');
   var context = canvas.getContext('2d');
    canvas.height = screen.height-200;
    canvas.width = screen.width -100;
   var radius = 20;
var color = ''red";
   var g = .1635; // acceleration due to gravity
10
    var x = 40; // initial horizontal position
   var y = 40; // initial vertical position
11
   var vx = parseFloat(prompt('what is the initial horizontal speed of ball you
        would like?(recommended values of 1-20')); // initial horizontal speed
   var vy = 0; // initial vertical speed
13
14
15
   window.onload = init;
```

```
17
    function init() {
18
     setInterval(onEachStep, 1000/60); // 60 fps
19
20
    function onEachStep() {
21
22
      vy += g; // gravity increases the vertical speed
23
      x += vx; // horizontal speed increases horizontal position
      y += vy; // vertical speed increases vertical position
24
25
      if (y > canvas.height - radius){ // if ball hits the ground
  y = canvas.height - radius; // reposition it at the ground
26
27
28
        vy *= -0.8; // then reverse and reduce its vertical speed
29
30
      if (x > canvas.width - radius){ // if ball hits right wall}
31
        x = canvas.width - radius; // reposition it right at wall
32
        vx *= -0.8; // then reduce and reverse horizontal speed
33
34
      if (x < radius) \{ // \text{ if ball hits left wall }
        x = radius; // reposition it right at wall
35
36
        vx *= -0.8 // then reverse and reduce horizontal speed
37
38
      drawBall(); // draw the ball
39
    };
40
41
    function drawBall() {
42
      with (context){
        clearRect(0, 0, canvas.width, canvas.height);
43
        fillStyle = color;
44
45
        beginPath();
        arc(x, y, radius, 0, 2*Math.PI, true);
46
47
        closePath();
48
        fill():
49
50
   };
```

Listing 1.1: A basic ball bouncing simulation

This code functions by first setting up the canvas to be an appropriate size, on lines 1-5. Then, the simple variables of radius, color, initial positions/velocities, and acceleration are initialized. As mentioned in the introduction, the canvas HTML element defines positions in terms of pixels, with the top left corner of the canvas being the origin. Therefore, the ball is initialized to appear at (40, 40) which is near the top left corner for any computer screen. The value of g, the gravitational constant, is set to .1635 to accurately represent its value near Earth's surface, of  $9.81 \frac{m}{s^2}$ . To understand why this value makes sense, it is necessary to understand the units of velocity on the HTML canvas. The position during the simulation is given in terms of pixels, which of course differs from the SI unit of meters. However, as long as g can be initialized to be  $9.81 \frac{px}{s^2}$ , the simulation will still look physically accurate. This can be explained with the equation below:

$$9.81 \frac{px}{s^2} = .1635 \frac{\frac{px}{s}}{frame} \times \frac{60 frame}{s}$$
 (1.1)

The value of g is calculated based on the fact that the simulation was run at 60 frames per second. Time is a central component of all physics, and for the simulations to behave realistically they must carefully take that into account.

The remainder of the code involves 3 functions that call one another to create the flow of the simulation. The first function, init ("initialize") is called when the browser window is loaded(line 15). This function simply delays the next function, on Each Step, by 16.66 ms, meaning that the function essentially runs 60 times per second, producing the desired 60 frames per second. Line 18 accomplishes this in a crude method: simulations later will involve more sophisticated techniques. The on Each Step function contains the instructions for each frame of the simulation. It involves multiple conditional if-statement loops that create the illusion that the ball bounces off of the walls and floor.

All 3 conditional loops involve a coefficient of restitution, or  $C_r$ . This is a mechanical property, representing how "bouncy" the ball is, and measures the ratio of the kinetic energy after and before the impact. This is derived below:

$$C_r = \sqrt{\frac{KE_f}{KE_i}} = \sqrt{\frac{\frac{1}{2}mv_f^2}{\frac{1}{2}mv_i^2}} = \frac{v_f}{v_i}$$
 (1.2)

A  $C_r$  value of .8 was used for this simulation, which is comparable to that of a tennis ball<sup>[7]</sup>. The variable Cr represents this value in the code, and is simply multiplied by the velocity before impact, so the following equation results:

$$v_f = v_i * C_r \tag{1.3}$$

An essential part of any physics simulation involves *collision detection*. For the simple bouncing ball simulation, this is accomplished by conditional loops for if the ball's position exceeds the canvas constraints.

The last function of the program, drawBall, simply contains commands for the canvas API to draw. While these commands can be very complicated and intricate to create the exact visual aesthetic desired, the extent of these commands is not the purpose of this thesis. Basically, this function works by "erasing" the canvas of any previous graphics, and then creating a new visual with the arc() method.

The logic of the program can be summarized through the flow chart below:



FIGURE 1.1: The logic flow chart of the basic bouncing ball simulation

### 1.2 More Advanced Ball Bouncing

While the previous example realistically incorporated the basic kinematic equations into account, it still fails to recognize important fundamentals of physics. The simulation in this chapter will still be a simple ball bouncing, but will take into account air resistance.

#### 1.2.1 Background Physics

Drag is generally defined as the force on an object that resists its motion through a fluid. In the case of air resistance, the fluid is a gas, and therefore the process is called aerodynamic drag. Most of the drag force results as a response to the inertia of the fluid: the resistance it exerts to oppose being pushed aside. This can be expressed in the equation below:

$$f_{drag} = -\frac{1}{2}C_d\rho Av^2 \tag{1.4}$$

The equation involves a negative sign because the force of drag is always opposite the direction of motion.  $C_d$  is referred to as the drag coefficient, and is a dimensionless quantity that is used to model complex dependencies of shape, inclination, and flow conditions. While  $C_d$  is in general not an absolute constant for a given body shape, for the purpose of these simulations constant values were used. These values are typically determined experimentally: for example, the  $C_d$  of a sphere is approximately .47. In equation 1.4,  $\rho$  is the mass density of the fluid, in  $\frac{kg}{m^3}$ . Most of the simulations in this thesis occur in air, which has a density of 1.225  $\frac{kg}{m^3}$  (at sea level and 15 °C). Running the simulations in different fluids can be simulated by changing  $\rho$  to higher values (water, for example, would have  $\rho$  equal to  $1000 \frac{kg}{m^3}$ ). Lastly, A in equation 1.4 is the cross-sectional area of the object. A sphere, for example, would have a cross-sectional area of  $\pi r^2$ 

The basic kinematic equations can also be used to make the simulations more physically realistic.

$$d = v_i t + \frac{1}{2} a t^2 (1.5)$$

$$v_f = v_i + at \tag{1.6}$$

These equations are fundamental to any physical situation and can be used to make the ball bouncing example of the previous section more realistic

#### 1.2.2 The Code

Using these basic mechanics equations, the previous ball bouncing example can be made more physically accurate. The code below shows a second simulation which incorporates

air resistance:

```
var x = 40:
   var y = 40;
   var vy = 0;
3
   var ay = 0;
4
    var m = 1;
   var r = 20;
   var rSI = r* 0.000230909; // radius in SI, converting px to m
7
    var C_r = .8; // Coefficient of restitution (tennis ball would be .8)
                     // density of air would be 1.2, water would be 1000
   var rho = 1.2:
9
10
   var dt = 60/1000; // Time Step
11
    var C_d = 0.47; //Coefficient of drag for sphere
12
   var A = Math.PI * rSI * rSI;
   var color = 'red';
13
14
15
   window.onload = init();
16
17
   function init(){
18
      setInterval(onEachStep, 1000/60);
19
20
21
   function onEachStep(){
22
     var fy = 0;
      fy += m * 9.81;
23
                         // weight force
24
      if (vy \ge 0) {
        fy -= 1* 0.5 *rho * C_d *A *vy *vy;
25
26
      }
27
      else {
        fy += 1*0.5 *rho *C_d *A *vy *vy;
28
     }
29
30
      ay = fy / m;
31
32
      vy += ay * dt;
33
     y += vy;
34
35
      // simple collision detection for floor only
36
      if (y + r > canvas.height){
37
        vy *= -C_r;
        y = canvas.height - r;
38
39
40
      drawBall();
   }
41
```

Listing 1.2: More advanced ball bouncing simulation

To eliminate redundancy, the code doesn't show previous functions used, such as drawBall(). The code also doesn't show the basic steps to initialize any simulation with the canvas and context commands. The code is very similar to the simulation in the

previous section, except it incorporates air resistance. Essentially, this simulation uses more kinematic equations, by calculating the net force, acceleration, and velocity for each frame. First, the net vertical force is caculated, by combining the force of gravity  $F_g = mg$  with the air drag from equation 1.4. This step involves a conditional loop for the cases of positive and negative velocity. Once the net force is calculated, the acceleration in the y-direction is found by using Newton's 2nd law of F = ma. From there, the velocity and vertical position of the ball are updated. Unlike the previous simulation, this example involves a variable dt, which is set to  $\sim 16$  ms for the same 60 frames per second.

This code involves interesting conversions between pixels and meters. Becuase the on screen simulation is presented eventually in terms of pixels, the physics equations must acknowledge this. The variable rSI on line 7 converts the radius of ball from pixels into meters. This is accomplished knowing the pixel density of the screen. This is commonly approximately 100 (dots per inch). The simulations were optimized for a macbook pro 15 inch model, which features 110 dpi. The calculation is shown below:

$$\frac{1m}{100cm} \; \frac{2.54cm}{1in} \; \frac{1in}{110px} \approx 0.00023091 \frac{m}{px} \tag{1.7} \label{eq:1.7}$$

Once the conversion is made, the physics equations use the radius of the ball in terms of meters instead of pixels, which would give erroneous answers.

### 1.3 Multiple Balls Bouncing

So far, this chapter has dealt with a single object in motion. However, physics rarely involves just one body in motion. To demonstrate how more than one object can be displayed simultaneously, this section will show the case of multiple bouncing balls.

There is no new physics introduced in this section, but the coding concepts will be used repeatedly in later chapters of this thesis.

#### 1.3.1 The Code

To generate more than one object, arrays can be used. The code below relies on arrays and object prototypes to create the effect:

```
var g = 0.1635;
    var balls;
    var numBalls = prompt('how many balls would you like to have bounce?');
3
    var C_d = .8;
5
    window.onload = init;
6
8
    function init() {
9
      balls = []; // creates empty array
10
      for (var i=0; i<numBalls; i++){</pre>
        radius = Math.random()*20+5;
11
12
        var ball = new Ball();
13
        ball.x = 50;
        ball.y = 75;
14
15
        ball.radius =
                        radius;
        ball.vx = Math.random()*15;
16
17
        ball.vy = (Math.random()-0.5)*10;
18
        ball.color = getRandomColor();
19
        ball.draw(context):
20
        balls.push(ball);
21
      setInterval(onEachStep, 1000/60); // 60 fps
22
23
   };
24
25
    function onEachStep() {
26
      context.clearRect(0, 0, canvas.width, canvas.height);
27
      for (var i=0; i<numBalls; i++){</pre>
28
        var ball = balls[i];
29
        ball.vy += g;
30
31
        if (ball.vx >0){ // while vx is positive, decrease to show friction/air drag
32
        ball.vx -= .001;
33
      } else{
34
        ball.vx === 0;// make sure ball stops moving appropriately
35
36
        ball.x += ball.vx;
37
        ball.y += ball.vy;
38
39
        if (ball.y > canvas.height - ball.radius){
40
          ball.y = canvas.height - ball.radius;
          ball.vy *= -C_d;
41
42
        if (ball.x + ball.radius > canvas.width){
  ball.x = canvas.width - ball.radius;
43
44
          ball.vx *= -C_d;
45
46
47
        if (ball.x < ball.radius){</pre>
          ball.x = ball.radius;
48
49
          ball.vx *= -C_d;
50
51
        ball.draw(context);
52
      }
53
   };
54
55
   function getRandomColor() {
        var letters = '0123456789ABCDEF'.split('');
```

```
57 | var color = '#';

58 | for (var i = 0; i < 6; i++) {

59 | color += letters[Math.floor(Math.random() * 16)];

60 | }

61 | return color;

62 |}
```

Listing 1.3: Multiple balls bouncing simulation

As with previous code listings, steps outlined in previous examples have been omitted to save space. This code differs mainly from previous examples because of its usage of prototypes, objects, and arrays. A separate javascript file, ball.js, contains the framework code for creating a ball. This will be used more in future chapters, so the code doesn't have to be repeated. This function is called a constructor function, becuase it allows other parts of code to reference the function when creating a new object. In the case of listing 1.3, an array holds an object for each different ball generated. The number of elements in the array is equal to the number of balls, which is selected by the user through the prompt() method on line 3. The object in each array element contains different properties for each ball: the radius, color, position, and velocities. For every frame of the simulation, a loop cycles through each element of the balls array, changing the properties of position and velocity, on lines 29 and 36-37. Exactly like in section 1, there is a conditional loop that controls the event of the ball colliding with a wall. The logic of this program can be visualized in the flow chart of figure 1.2.

Because of repeated for loops, this program involves a bit more complexity than the previous examples. However, the physics is very simple in this case. Future chapters will combine more complex physics with this complexity of coding to create more advanced simuations. These simulations can put a strain on a computer's performance: to simulate 20 balls bouncing, at 60 frames per second, 4,800 individual properties of objects need to be generated each second. Luckily, with the modern capabilities of computers, this isn't too difficult.



FIGURE 1.2: The logic flow chart of the balls bouncing simulation

This program involves each ball having a random color and radius, when the balls array is created in the init() function. The Math.rand() method is used for both of these, and is fundamental to the other simulations in this thesis. The random color generator function operates by creating a hex color by randomly assigning the 16 possible entries to each entry of the 6 character string. While completely unnecessary, this gives the program aesthetic appeal, and makes it easier to distinguish the balls.

## Chapter 2

# Simulating Orbits

In this chapter, a more advanced simulation of orbiting masses will be introduced. First, a simple orbit situation will be introduced, followed by more complex examples involving escape velocities. While the physics still isn't too advanced, the coding necessary is a bit more challenging.

#### 2.1 Basic Orbit Path

The first simulation will deal with an example of a planet orbiting another much more massive planet.

#### 2.1.1 Background Physics

This entire chapter is centered around Newton's law of universal gravitation:

$$F_g = G \frac{m_1 m_2}{r^2} (2.1)$$

Where  $F_g$  is the magnitude of the force acting on either mass, G is the gravitational constant (SI units of 6.67  $\frac{Nm^2}{kg^2}$ ),  $m_1$  is the mass of one object,  $m_2$  is the mass of the

other object, and r is the radius separating the two masses. By Newton's 3rd law, there is an equal and opposite force exerted on each mass.

This equation can be used to describe the orbiting paths of planets. For simple cases when one planet orbits another, the variables m and M can be used. For a first example, we will assume M >>m, that is, one planet has a much greater mass than the other. Therefore, while each planet exerts an equal force on the other, the acceleration on the massive planet will be negligible. So, the smaller planet will orbit around the stationary planet, without attracting the larger planet enough to move it.

#### 2.1.2 The Code

The full code is shown in the listing below:

```
var canvas = document.getElementById('canvas');
   var context = canvas.getContext('2d');
3
   var canvas_bg = document.getElementById('canvas_bg');
   var context_bg = canvas_bg.getContext('2d');
5
6
   var planet;
   var sun;
   var m = 1; // planet's mass
8
9
    var M = 1000000; // heavy planet's mass
10
   var G = 1;
11
   var t0,dt;
12
13
   window.onload = init;
14
15
    function init() {
16
            // create a stationary large planet
17
            sun = new Ball(70, 'orange', M);
18
            sun.pos2D = new Vector2D(275,200);
19
            sun.draw(context_bg);
20
            // create a moving planet
21
            planet = new Ball(10,'blue',m);
            planet.pos2D = new Vector2D(200,50);
22
23
            planet.draw(context);
            // make the planet orbit the large planet
24
25
            t0 = new Date().getTime();
26
            animFrame();
27
   };
28
29
   function animFrame(){
30
            animId = requestAnimationFrame(animFrame, canvas);
31
            onTimer();
32
33
   function onTimer(){
34
            var t1 = new Date().getTime();
            dt = 0.001*(t1-t0);
35
36
            t0 = t1;
37
            if (dt>0.1) {dt=0;};
38
            move();
39
40
   function move(){
            moveObject(planet);
41
            calcForce();
```

```
43
            updateAccel();
44
            updateVelo(planet);
   }
45
46
47
   function moveObject(obj){
48
            obj.pos2D = obj.pos2D.addScaled(obj.velo2D,dt);
49
            context.clearRect(0, 0, canvas.width, canvas.height);
50
            obj.draw(context);
51
52
   function calcForce(){
            force = Forces.gravity(G,M,m,planet.pos2D.subtract(sun.pos2D));
53
54
   function updateAccel(){
55
56
            acc = force.multiply(1/m);
57
   function updateVelo(obj){
58
59
            obj.velo2D = obj.velo2D.addScaled(acc,dt);
60
```

Listing 2.1: Basic planet orbiting simulation

This program differs from previous ones used so far in that it uses two canvases instead of one. This is essential for having the large planet remain stationary and not being erased every frame. Instead, there can be a constant "background" canvas containing the stationary planet. The code begins by initializing the variables planet and sun, where sun simply refers to any large planet that has much more mass. The gravitational constant G is intialized as a formality just to a value of 1. G in this simulation isn't necessary, because the constant simply is used for unit conversion. This will become clear later in this section. When the web page is loaded, it calls the init function, just as in previous simulations. The init function creates the sun and planet as objects from the Ball constructor function, exactly as in chapter 1. Instead of having separate variables X and Y in the previous examples, the position information can be stored into a property of each object, which is created using a different constructor function Vector 2D.

The next function, animFrame, functions simply by initializing the javascript animation frame, and then calling the next function, on Timer. This next function creates a variable dt by converting the unit javascript operates in (ms) to SI units of s. It then passes the flow of the program onto the next function, move. This function involves

calling 4 functions, the first of which simply updates the position of the planet, erases the foreground canvas, and then draws the updated canvas. This step can be analyzed through a physics kinematics equation:

$$x(t+dt) = x(t) + v_x(t)dt (2.2)$$

This is essentially analogous to Euler's method, by understanding that  $v_x = \frac{dx}{dt}$ .

$$x(t+dt) = x(t) + \frac{dx}{dt}(0) dt$$
(2.3)

By using constructor functions, with 2 different properties for the x and y position, the planet's location can be updated without updating variables and taking up more space. The location has to be updated for every frame, but so does the force, acceleration, and velocity. These next 3 steps are the remaining functions of the move function. The calcForce function updates the force of gravity acting on the planet, using equation 2.1. This equation calculates r by calculating the displacement vector between the two planets, and finding the magnitude of that vector. The updateAccel function simply takes the updated force vector and scales it by a certain "k" value which is represented by dividing by the mass. This is the step that incorporates Newton's 2nd law of  $a = \frac{F}{m}$ . The last function updates the velocity of the planet, similar to how the position was updated. Using Euler's method like before, we come to the following equation

$$v(t+dt) = v(t) + a(t)dt (2.4)$$

Essentially, what makes this program more complicated is that it uses many other functions to accomplish the overall simulation. However, this method of programming makes future simulations easier—the same functions can be used, with changed variables.

The code listing below shows these "tool" functions that are used in future chapters of this thesis:

```
function Vector2D(x,y) {
2
            this.x = x;
3
            this.y = y;
4
5
   Vector2D.prototype = {
6
            lengthSquared: function(){
7
                     return this.x*this.x + this.y*this.y;
8
g
            length: function(){
10
                     return Math.sqrt(this.lengthSquared());
11
            }.
12
            add: function(vec) {
                     return new Vector2D(this.x + vec.x,this.y + vec.y);
13
            },
14
15
            subtract: function(vec) {
16
                     return new Vector2D(this.x - vec.x,this.y - vec.y);
17
18
            multiply: function(k) {
19
                     return new Vector2D(k*this.x,k*this.y);
20
21
            addScaled: function(vec,k) {
22
                     return new Vector2D(this.x + k*vec.x, this.y + k*vec.y);
23
24
25
            function Forces(){
26
   Forces.gravity = function(G,m1,m2,r){
27
            \texttt{return r.multiply(-G*m1*m2/(r.lengthSquared()*r.length()));}\\
28
29
```

Listing 2.2: Various tools functions used for orbit simulation

### 2.2 Escape Velocity

The previous section tested situations where the planet orbited the sun continuously. However, if the speed is great enough, the orbiting body is capable of "escaping" from the larger planet's influence. The minimum speed necessary for this is called the escape velocity.

This can be derived by understanding conversation of energy. When an object leaves the surface of a planet, it will have an initial kinetic energy, and potential gravitational energy. This will equal the final potential energy, defined as a condition when the final kinetic and gravitational potential energy is 0. This relationship is shown in the equation below:

$$K_i + U_{g_i} = K_f + U_{g_f}$$

Knowing that the final kinetic and gravitational energy is 0, this equation becomes:

$$\frac{1}{2}mv_{esc}^2 - \frac{GMm}{r} = 0 + 0$$

Solving for  $v_{esc}$  yields the following:

$$v_{esc} = \sqrt{\frac{2GM}{r}} \tag{2.5}$$

Where G is the gravitational constant, M is the mass of the planet the object is escaping from, and r is the starting distance from the center of mass of the planet.

To test the physics behind the escape velocity, a slightly different scenario can be created with a different program. This simulation will have the object begin right at the surface of the larger planet, to emulate the process of "escaping" from the planet's gravity influence. To visualize this, a much larger canvas will be used, and some code changes will be utilized, seen below:

```
1 sun = new Ball(400,'orange',M);
2 sun.pos2D = new Vector2D(
3
4 planet = new Ball(10,'blue',m);
5 planet.pos2D = new Vector2D(500,2490);
6
7 planet.velo2D = new Vector2D(0, -80);
```

Listing 2.3: New conditions for escape velocity simulation

These changes made the larger planet look visually bigger, to simulate the effect of a massive planet. It also positioned the object to begin right on the surface of the larger planet (in this case, 410 pixels above the center of mass of the larger planet). To calculate the escape velocity for the situation above, equation 2.5 can be used, but understanding some key factors:

- 1. The escape velocity calculated will be in  $\frac{px}{s}$  instead of SI unit  $\frac{m}{s}$
- 2. The masses of each planet don't need to contain units, it can more just represent a ratio between the large and small planet. Therefore, each mass will be a unitless quantity, just used as a test of the escape velocity equation.
- 3. Similarly to #2, the gravitational constant G doesn't have to include units, since this test is only in a more theoretical sense, and doesn't use actual units of mass. However, using dimensional analysis, for the equation below to make sense, G could be viewed as having units of  $\frac{px^3}{s^2}$ .

Proceeding with these conditions in mind, the escape velocity for the simulation of code listing ?? can be calculated as shown below:

$$v_{esc} = \sqrt{\frac{2 * 1 \frac{px^3}{s^2} * 1000000}{410px}} \approx 69.843 \frac{px}{s}$$
 (2.6)

Therefore, with the program simulation, any initial speed greater than this value will escape the gravitational influence of the larger planet. To test this, I used the following code to print out values of the velocity continuously:

```
var i =0;
ti++;
if(i%15 ===0){console.log(planet.velo2D.length());}
```

Listing 2.4: Code for printing out values of speed

This code operates by printing out the value of the speed of object 4 times per second, understanding that i is incremented by 1 for each frame, and there are the usual 60 frames per second. The data is outputted through the console.log() method, which prints it onto the web browser. This data was then plotted for different initial speeds, and the results are shown below:

### 2.3 Kepler's Laws

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