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Mechanics Simulations With JavaScript

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Overview - Why Did I Choose This Topic?

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- I hope to use programming as a lens to view physics
- Examine mechanics in more detail
- Solve physics problems through simulations
- JavaScript high level language viewable easily in web browser

What is a simulation?

Overview

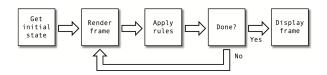
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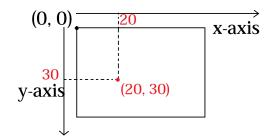
- Animation vs. Simulation
- Frames per second
- File size



Method of Basic Simulation

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- HTML5 canvas application programming interface (API)
- Timer for each frame



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Simulation #

Chapter 1: Basic kinematics and aerodynamic drag

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- Three simulations
- Simulation #1: Basic bouncing ball
- Simulation #2: Bouncing ball with aerodynamic drag
- Simulation #3: Multiple bouncing balls

Simulation #1: Basic Bouncing Ball

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Realistic g value

• 9.81
$$\frac{px}{s^2} = .1635 \frac{\frac{px}{s}}{frame} \times \frac{60 frame}{s}$$

• Coefficient of restitution (C_r)

$$\bullet \ \ \textit{C_r} = \sqrt{\frac{\textit{KE_f}}{\textit{KE_i}}} = \sqrt{\frac{\frac{1}{2}\textit{mv_f^2}}{\frac{1}{2}\textit{mv_i^2}}} = \frac{\textit{v_f}}{\textit{v_i}}$$

•
$$v_f = v_i * C_r$$

Simulation #2: Bouncing Ball With Aerodynamic Drag

Simulation #2

• $f_{drag} = -\frac{1}{2}C_d\rho Av^2$

- F_D = force of drag
- $\rho = \text{density of fluid}$
- v =speed of object relative to fluid
- C_d = drag coefficient (affected by texture, shape, viscosity, lift, etc)
- A = cross-sectional area of object

Simulation #3: Multiple Balls Bouncing

- Simulation #3

- Same physics as simulation #1
- Array of ball objects
- Each object has properties
- Each frame cycles through array, updating properties of each object

Chapter 2: Planetary Motion

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Simulation #4: Orbits

• Simulation #5: Escape velocity

• Simulation #6: Kepler's 2nd law

Simulation #4: Orbits

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Newton's Law of universal gravitation

$$\bullet \ F_g = G \frac{m_1 m_2}{r^2}$$

- Euler's Method to update velocity
- $x(t + dt) = x(t) + \frac{dx}{dt}(t) dt$

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 $\bullet \ K_i + U_{g_i} = K_f + U_{g_f}$

•
$$\frac{1}{2}mv_{esc}^2 - \frac{GMm}{r} = 0 + 0$$

•
$$v_{esc} = \sqrt{\frac{2GM}{r}}$$

•
$$v_{esc} = \sqrt{\frac{2*1\frac{px^3}{s^2}*1000000}{410px}} \approx 69.843\frac{px}{s}$$

 Used bigger canvas, and plotted velocities during planet's travel

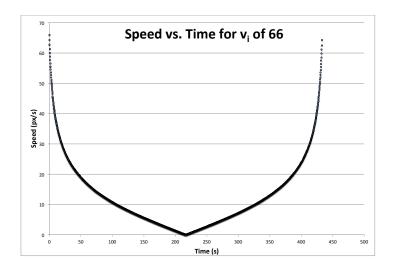
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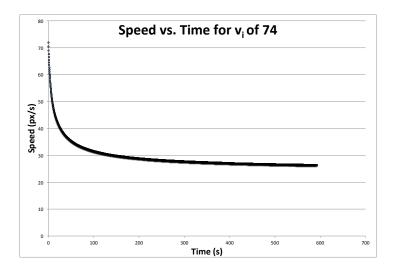
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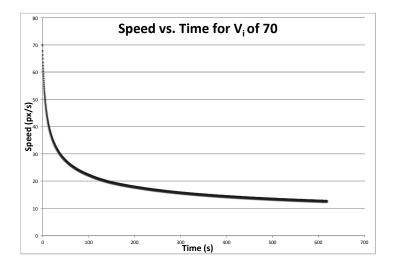
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Simulation #6: Kepler's 2nd law

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- Early 1600's Johannes Kepler proposed laws explaining how planets orbit the sun
- Law #2: "The radius vector drawn from the Sun to a planet sweeps out equal areas in equal time intervals"
- Simulation shows constant $\frac{dA}{dt}$

Derivation of Kepler's 2nd Law

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$$\bullet \ \vec{\tau} = \vec{r} \times \vec{F_g} = \frac{d\vec{L}}{dt}$$

$$\bullet \ \vec{L} = \vec{r} \times \vec{p} = M_p \vec{r} \times \vec{v}$$

•
$$L = M_p |\vec{r} \times \vec{v}|$$

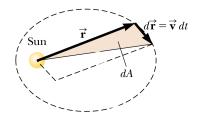


Figure : Relationship between \vec{r} and $d\vec{r}$

Derivation of Kepler's 2nd Law (Continued)

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•
$$|\vec{r} \times d\vec{r}|$$
 area of parallelogram

•
$$dA = \frac{1}{2}|\vec{r} \times d\vec{r}| = \frac{1}{2}|\vec{r} \times \vec{v}dt| = \frac{1}{2}|\vec{r} \times \vec{v}|dt$$

•
$$dA = \frac{1}{2} \left(\frac{L}{M_p} \right) dt$$

• From before,
$$|\vec{r} \times \vec{v}| = \frac{L}{M_P}$$

• L and M_p are constants

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Thank You