

# Diffusion Maps

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## **Abstract**

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# Introduction

The contentual layout follows that of Coifman and Lafon (2006) and Belkin (2003). Some definitions, theorems and examples will be taken from these works as well.

## 0.1 Dimensionality Reduction

# Chapter 1

## Diffusion Maps

### 1.1 Diffusion Kernels

#### 1.1.1 Motivation

In this section it will be elaborated on why global structures need not be preserved and how this leads to diffusion processes.

#### Preservation of Global Structures

Suppose data points  $x_i \in \mathbb{R}^k$  are generated by a low dimensional parameter  $\theta_i \in \mathbb{R}^{k'}$ ,  $k' \ll k$  given a map  $\Phi : \mathbb{R}^{k'} \rightarrow \mathbb{R}^k$ . One problem is the (numerical) smoothness of  $\Phi$  which is not necessarily given. Another problem is that for large  $k$  the euclidian distance is no longer a meaningful measure as the volume of the  $k$ -dimensional unit ball  $\frac{\pi^{\frac{k}{2}}}{\Gamma(\frac{k}{2}+1)}$  converges to 0 as  $\lim_{k \rightarrow \infty}$  which is also known as the *curse of dimensionality*. One may conclude that large distances in the ambient space need not necessarily be preserved as they do not hold much information except that  $x_i$  is not “very close” to  $x_j$ .

Analogous to Riemannian manifolds (where metric tensors, inducing an inner product on the tangent space and a metric via the exponential map, define the manifold’s geometry) we focus solely on local distances in order to recover intrinsic global structures.

## A Dual Approach

From inverse problems in spectral geometry (e.g. “Can One Hear the Shape of a Drum?”) it is known that much<sup>1</sup> of the geometry of a given set  $\Gamma$  can be derived from the analysis of functions defined on  $\Gamma$ .

In this work eigenvalues and eigenfunctions of averaging operators, i.e., operators whose kernel corresponds to transition probabilities of a Markov process, will be studied in order to define a diffusion map which embeds the data into a Euclidian space where the Euclidian distance is just the diffusion metric.

### 1.1.2 Construction of a Random Walk on the Data

#### Definitions

Let  $(\Gamma, \mathcal{A}, \mu)$  be a measure space. In practical applications  $\Gamma$  is the given data set consisting of finitely many data points and  $\mu$  is the counting measure to represent the distribution of the points in the data set. In addition, suppose we are given a symmetric kernel  $k : \Gamma \times \Gamma \rightarrow \mathbb{R}^+$  which defines the local geometry of  $\Gamma$ .

#### Examples

Usually  $\Gamma$  is either a subset of the Euclidian space or a weighted graph.

In the first case it seems natural to write  $k$  as a function of the Euclidian distance  $\nu(\|x - y\|)$ .

In the second case let  $b(x, y)$  be the associated adjacency matrix, that is,  $b(x, y) = 1$  if there is an edge going from  $x$  to  $y$ , and  $b(x, y) = 0$  otherwise. The kernel  $b$  defines a notion of neighborhood for each point, and also a non-symmetric distance given by  $1 - b(x, y)$ . Clearly  $b$  is not symmetric in general, but we can consider

$$k_1(x, y) := \int_{\Gamma} b(x, u)b(y, u)d\mu(u)$$

$$k_2(x, y) := \int_{\Gamma} b(u, x)b(u, y)d\mu(u)$$

where  $k_1(x, y)$  counts the number of common neighbors to  $x$  and  $y$ , while  $k_2(x, y)$  counts the number of nodes for which  $x$  and  $y$  are common neighbors.

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<sup>1</sup>the most famous example being Weyl’s proof of  $\#\{\lambda_k : \lambda_k < \lambda\} \approx \frac{\text{area}(\Gamma)}{2\pi}\lambda$  as  $\lambda \rightarrow \infty$ . (Canzani (2013))

## Normalized Graph Laplacian Construction

Generally, such a kernel represents some notion of affinity between points of  $\Gamma$  and thus one can think of the data points as being the nodes of a symmetric graph whose weight function is specified by  $k$ . From the graph defined by  $(\Gamma, k)$ , one can construct a reversible Markov chain on  $\Gamma$ .

To normalize the kernel, define

**Definition 1.1.**

$$v^2(x) = \int_{\Gamma} k(x, y) d\mu(y)$$

and

**Definition 1.2.**

$$p(x, y) = \frac{k(x, y)}{v^2(x)}.$$

$p(x, y)$  is no longer symmetric, but inherited the positivity and now satisfies a conservation property:

$$\int_{\Gamma} p(x, y) d\mu(y) = 1$$

As a result the matrix  $P := (p(i, j))_{i,j}$  is stochastic and can be interpreted as the transition matrix of a homogeneous Markov process on  $\Gamma$ . In spectral graph theory  $\mathbb{I} - P$  is commonly referred to as normalized, weighted graph Laplacian. This naming is justified in A.2.1<sup>2</sup>.

To investigate the spectral properties of the corresponding integral operator  $P$  defined by  $Pf(x) = \int_{\Gamma} p(x, y)f(y)d\mu(y)$  it is beneficial to examine the symmetric, conjugated Operator  $A$ .

**Definition 1.3.** *Let*

$$a(x, y) = \frac{k(x, y)}{v(x)v(y)} = v(x)p(x, y)\frac{1}{v(y)},$$

*then the corresponding diffusion operator  $A$  is defined as*

$$Af(x) = \int_{\Gamma} a(x, y)f(y)d\mu(y)$$

Notice that by definition of  $a$  one obtains a symmetric form and thus a symmetric operator  $A$ .

---

<sup>2</sup>Let w.l.o.g.  $m_0 = 1$  and  $m_2 = 2$ . Now set  $L_{\varepsilon} = \frac{\varepsilon I - G_{\varepsilon}}{\varepsilon}$ ; for uniformly sampled data the potential term  $\omega$  vanishes and  $\lim_{\varepsilon \rightarrow 0} L_{\varepsilon} = \Delta$

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### 1.1.3 Diffusion Kernels

**Theorem 1.1.1** (Spectral Properties of the Diffusion Operator). *The diffusion operator  $A$  with kernel  $a$  is bounded from  $L^2(\Gamma, d\mu)$  into itself, symmetric and positive semi-definite.*

Moreover, its norm is

$$\|A\| = 1$$

and is taken by the eigenfunction

$$Av = v.$$

*Proof.* Let  $f \in L^2(\Gamma, d\mu)$ . We have:

$$\langle Af, f \rangle = \int_{\Gamma^2} k(x, y) \frac{f(x)}{v(x)} \frac{f(y)}{v(y)} d\mu(x) d\mu(y). \quad (1.1)$$

Applying the Cauchy-Schwartz inequality we get:

$$\begin{aligned} \left| \int_{\Gamma} k(x, y) \frac{f(y)}{v(y)} d\mu(y) \right| &= \left( \int_{\Gamma} k(x, y) d\mu(y) \right)^{\frac{1}{2}} \left( \int_{\Gamma} k(x, y) \frac{f(y)^2}{v(y)^2} d\mu(y) \right)^{\frac{1}{2}} \\ &= v(x) \left( \int_{\Gamma} k(x, y) \frac{f(y)^2}{v(y)^2} d\mu(y) \right)^{\frac{1}{2}} \end{aligned}$$

Hence:

$$\langle Af, f \rangle \leq \int_{\Gamma} |f(x)| \left( \int_{\Gamma} k(x, y) \frac{f(y)^2}{v(y)^2} d\mu(y) \right)^{\frac{1}{2}} d\mu(x)$$

and by using the Cauchy-Schwartz inequality once again:

$$\langle Af, f \rangle \leq \|f\| \left( \int_{\Gamma^2} \frac{k(x, y)}{v(y)^2} f(y)^2 d\mu(y) d\mu(x) \right)^{\frac{1}{2}} = \|f\|^2$$

by symmetry of the kernel which, in combination with (1.1), also implies the positivity of  $A$ .

Plugging in  $v$  for  $f$  it follows immediately that the eigenvalue 1 is actually obtained and  $v$  is an eigenfunction.  $\square$

**Theorem 1.1.2** (Spectral Decomposition of the Diffusion Kernel). *Assuming  $A$  is compact<sup>3</sup> and  $A\phi_l = \lambda_l \phi_l$  we may write the kernel as*

$$a(x, y) = \sum_{l \geq 0} \lambda_l \phi_l(x) \phi_l(y)$$

with  $\lambda_0 = 1$  and  $\lim_{l \rightarrow \infty} \lambda_l = 0$  monotonically.

<sup>3</sup>which is no constraint in practice since data is finite

try to generalize lemma 3.4 of Teschl (2014) (guess it won't work though)



*Proof.* First, note that  $A$  being compact implies that the spectrum is discrete and the sum thus is well defined. By  $A$  being symmetric and compact the spectral theorem applies and we get that there exists a sequence of real eigenvalues  $\lambda_l$  converging to 0. The corresponding normalized eigenvectors  $\phi_l$  form an orthonormal set and every  $f \in L^2(\Gamma, d\mu)$  can be written as

$$f = \sum_{l \geq 0} \langle \phi_l, f \rangle \phi_l + h$$

where  $h \in \text{Ker}(A)$ .

It follows that

$$Af(x) = \int_{\Gamma} a(x, y) f(y) d\mu(y) = \sum_{l \geq 0} \lambda_l \int_{\Gamma} \phi_l(y) f(y) d\mu(y) \phi_l(x)$$

which, by linearity of the integral and comparison of components, is just what we were looking for.  $\square$

“Komponentenvergleich  
auf Englisch  
finden

From definition of  $a(x, y)$  we see that

$$p(x, y) = \sum_{l \geq 0} \lambda_l \underbrace{\frac{\phi_l(x)}{v(x)}}_{=: \psi_l(x)} \phi_l(y) v(y) \quad (1.2)$$

which enables us to efficiently compute  $t$ th powers  $p_t$  of  $p$ .

There are two ways to interpret  $p_t$ :

1.  $p_t$  has a probabilistic interpretation as the probability for a Markov chain with transition matrix  $P$  to reach  $y$  from  $x$  in  $t$  steps.
2. the dual point of view is that of the functions defined on the data. The kernel  $p_t$  can be viewed as a bump or more precisely, if  $x \in \Gamma$  is fixed, then  $p_t(x, \cdot)$  is a bump function centered at  $x$  and of width increasing with  $t$  which intuitively captures the idea of diffusion.

#### 1.1.4 Embedding in the Euclidian Space

**Definition 1.4.** Let

$$D_t(x, y)^2 = \|p_t(x, \cdot) - p_t(y, \cdot)\|_{L^2(\Gamma, d\mu/v)}^2 = \int_{\Gamma} (p_t(x, u) - p_t(y, u))^2 \frac{d\mu(u)}{v(u)}$$

be the family of diffusion distances parameterized by  $t$ .

For a fixed value of  $t$   $D_t$  defines a distance on the set  $\Gamma$  which is small only if there is a large number of small paths connecting  $x$  and  $y$  (i.e. if there is a large probability of getting from  $x$  to  $y$  in  $t$  steps). It thus emphasizes the notion of a cluster.

Another property following from the summation over all possible paths is that this distance is very robust to noise perturbation (in contrast to the geodesic distance).

**Theorem 1.1.3** (A Numerically Feasible Representation).

$$D_t(x, y) = \left( \sum_{l \geq 0} (\lambda_l^t (\psi_l(x) - \psi_l(y)))^2 \right)^{\frac{1}{2}}$$

*Proof.*  $\{\phi_l\}_{l \geq 0}$  forming an orthonormal basis for  $L^2(\Gamma, d\mu)$  implies that  $\{\phi_l v\}_{l \geq 0}$  is an orthonormal basis for  $L^2(\Gamma, d\mu/v)$  and thus for  $x$  fixed (1.2) may be seen as orthogonal expansion of the function  $y \mapsto p_t(x, y)$  into the basis  $\{\phi_l v\}_{l \geq 0}$ . The coefficients are given by  $\{\lambda_l^t \psi_l(x)\}_{l \geq 0}$ . The statement follows directly using the Pythagorean theorem.  $\square$

An immediate consequence is that the diffusion distance is well approximable and that it converges towards a function of (numerical) rank 1 as  $t \rightarrow \infty$  because of the vanishing influence of all eigenvectors with eigenvalues  $< 1$ .

One possible interpretation is that  $D_t(x, y)$  measures the distance between bumps of “magnitude”  $t$  being centered around two points  $x$  and  $y$ . As  $t$  gets larger so does the size of the supports and the number of eigenfunctions needed to calculate  $D_t(x, y)$  decreases. This number is related to the minimum number of bumps necessary to cover the set  $X$  (like in Weyl’s asymptotic law for the decay of the spectrum).

In order to calculate  $D_t(x, y)$  to a preset accuracy  $\delta > 0$  with a finite number of terms we set

$$s_t(\delta) = \max\{l \in \mathbf{N} : \lambda_l^t > \delta \lambda_1^t\}$$

so that, up to relative precision  $\delta$

$$D_t(x, y) = \left( \sum_{l=0}^{s_t(\delta)} (\lambda_l^t (\psi_l(x) - \psi_l(y)))^2 \right)^{\frac{1}{2}}. \quad (1.3)$$

**Definition 1.5.** Let  $\{\Psi_t\}_{t \in \mathbf{N}}$ ,

$$\Psi_t(x) = \begin{pmatrix} \lambda_1^t \psi_1(x) \\ \lambda_2^t \psi_2(x) \\ \vdots \\ \lambda_{s_t(\delta)}^t \psi_{s_t(\delta)}(x) \end{pmatrix}$$

be the family of diffusion maps. Each component of  $\Psi_t(x)$  is termed diffusion coordinate.

According to (1.3) diffusion maps embed data in a Euclidian space in such a way that the Euclidian distance equals the diffusion distance up to a relative error  $\delta$ .

## Chapter 2

# Connection with the Fokker-Planck Equation

### 2.1 A Family of Anisotropic Diffusion Maps

In this section we investigate the probability space  $(\Omega, \mathcal{A}, \mu)$  and therefore assume an infinite number of data points whose distribution is determined by the probability measure  $\mu(x) = e^{-U(x)}$ .  $U$  may be interpreted as a potential function. A motivation for the choice of  $\mu$  will be given with (2.8).

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SDEs, etc.

#### 2.1.1 Construction of the Family of Diffusions

Start with a Gaussian kernel  $k_\varepsilon(x, y) = e^{-\frac{\|x-y\|^2}{\varepsilon}}$  and let  $\alpha > 0$  being a parameter indexing this family of diffusions.

We can estimate the local density  $q_\varepsilon$  by

**Definition 2.1.**

$$q_\varepsilon(x) = \int_{\Omega} k_\varepsilon(x, y) d\mu(y)$$

Now consider the family of (generally) anisotropic kernels

**Definition 2.2.**

$$k_\varepsilon^{(\alpha)}(x, y) = \frac{k_\varepsilon(x, y)}{q_\varepsilon^\alpha(x) q_\varepsilon^\alpha(y)}$$

We compute the normalization factor<sup>1</sup>  $d_\varepsilon^{(\alpha)}$

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<sup>1</sup>Note that this factor is the continuous analogue to definition 1.1.

$$d_\varepsilon^{(\alpha)}(x) = \int_{\Omega} k_\varepsilon^{(\alpha)}(x, y) d\mu(y)$$

and define the forward and symmetric ( $p^{(\alpha)}$  and  $a^{(\alpha)}$  respectively) transition probability kernels:

**Definition 2.3.**

$$p_\varepsilon^{(\alpha)}(x|y) = \frac{k_\varepsilon^{(\alpha)}(x, y)}{d_\varepsilon^{(\alpha)}(y)} = Pr(x(t + \varepsilon) = x \mid x(t) = y)$$

$$a_\varepsilon^{(\alpha)}(x|y) = \frac{k_\varepsilon^{(\alpha)}(x, y)}{\sqrt{d_\varepsilon^{(\alpha)}(x)d_\varepsilon^{(\alpha)}(y)}}$$

Now we define the forward, backward and symmetric Chapman-Kolmogorov operators on functions defined on this probability space, as follows:

**Definition 2.4.**

$$T_{f,\varepsilon}^{(\alpha)}[\phi](x) = \int_{\Omega} p_\varepsilon^{(\alpha)}(x|y)\phi(y)d\mu(y)$$

$$T_{b,\varepsilon}^{(\alpha)}[\phi](x) = \int_{\Omega} p_\varepsilon^{(\alpha)}(y|x)\phi(y)d\mu(y)$$

$$T_{s,\varepsilon}^{(\alpha)}[\phi](x) = \int_{\Omega} a_\varepsilon^{(\alpha)}(x, y)\phi(y)d\mu(y)$$

If  $\phi(x)$  is the probability of finding the system at location  $x$  at time  $t = 0$ , then  $T_{f,\varepsilon}^{(\alpha)}[\phi]$  is the evolution of this probability to time  $t = \varepsilon$ . Similarly, if  $\psi(z)$  is some function on the space, then  $T_{b,\varepsilon}^{(\alpha)}[\psi](x)$  is the mean value of that function at time  $\varepsilon$  for a random walk that started at  $x$ , and so  $\left(T_{b,\varepsilon}^{(\alpha)}\right)^m[\psi](x)$  is the average value of the function at time  $t = m\varepsilon$ .

By definition, the operators  $T_{f,\varepsilon}^{(\alpha)}$  and  $T_{b,\varepsilon}^{(\alpha)}$  are adjoint under the inner product with weight  $\mu$ , while the operator  $T_{s,\varepsilon}^{(\alpha)}$  is self adjoint under this inner product.

Moreover, just like in the discrete case, since  $T_{s,\varepsilon}^{(\alpha)}$  is obtained via conjugation of the kernel  $p^{(\alpha)}$  all three operators share the same eigenvalues. The corresponding eigenfunctions can be found via conjugation by  $\sqrt{d_\varepsilon^{(\alpha)}}$ ; i.e.: If  $T_{s,\varepsilon}^{(\alpha)}\phi_s = \lambda\phi_s$ , then the corresponding eigenfunctions for  $T_{f,\varepsilon}^{(\alpha)}$  and  $T_{b,\varepsilon}^{(\alpha)}$  are  $\phi_f = \sqrt{q_\varepsilon}\phi_s$  and  $\phi_b = \frac{\phi_s}{\sqrt{q_\varepsilon}}$ , respectively.

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Since  $\sqrt{q_\varepsilon}$  is the first eigenfunction with  $\lambda_0 = 1$  of  $T_s^{(\alpha)}$ , the steady state of the forward operator is simply  $q_\varepsilon(x)$ , while for the backward operator it is the uniform density,  $\phi_b \equiv 1$ .

Furthermore, note that the eigenvalues and eigenvectors of the discrete Markov chain described in the previous section are discrete approximations to the eigenvalues and eigenfunctions of these continuous operators.

### 2.1.2 Transition to Diffusion Processes

Interpreting  $\varepsilon$  as a timestep<sup>2</sup> and recalling the asymptotic expansions derived in the appendix it is instructive to look at the limit  $\varepsilon \rightarrow 0$ . In this case, the transition probability densities of the Markov chain (that is continuous in space, but discrete in time) converge to those of a diffusion process, whose time evolution is described by a differential equation

$$\frac{\partial \phi}{\partial t} = \mathcal{H}_f^{(\alpha)} \phi$$

where  $\mathcal{H}_f^{(\alpha)}$  is the infinitesimal generator or propagator of the forward operator, defined as

**Definition 2.5.**

$$\mathcal{H}_f^{(\alpha)} = \lim_{\varepsilon \rightarrow 0} \frac{T_{f,\varepsilon}^{(\alpha)} - I}{\varepsilon}$$

Similarly, the infinitesimal operator of the backward operator is given by

**Definition 2.6.**

$$\mathcal{H}_b^{(\alpha)} = \lim_{\varepsilon \rightarrow 0} \frac{T_{b,\varepsilon}^{(\alpha)} - I}{\varepsilon}$$

and as shown in the appendix (theorem A.2.2)

$$\mathcal{H}_b^{(\alpha)} \psi = \Delta \psi - 2(1 - \alpha) \nabla \psi \cdot \nabla U$$

as well as

$$\mathcal{H}_f^{(\alpha)} \psi = \Delta \psi - 2\alpha \nabla \psi \cdot \nabla U + (2\alpha - 1) \psi (\nabla U \cdot \nabla U - \Delta U)$$

Now we consider the following three choices of  $\alpha$ :

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<sup>2</sup>While in the case of a finite amount of data,  $\varepsilon$  must remain finite so as to have enough neighbors in a ball of radius  $\mathcal{O}(\varepsilon^{\frac{1}{2}})$  near each point  $x$ , as the number of samples goes to infinity, we can take smaller and smaller values of  $\varepsilon$ .

1.  $\alpha = 0$ :

$$\mathcal{H}_f^{(0)}\phi = \Delta\phi - (e^U \Delta e^{-U})\phi$$

Choosing  $\alpha = 0$ , “ignoring” the potential term (in the process of the normalization), we get the normalized graph Laplacian we already dealt with in the first chapter.

2.  $\alpha = \frac{1}{2}$ :

$$\mathcal{H}_f^{(\frac{1}{2})}\phi = \mathcal{H}_b^{(\frac{1}{2})}\phi = \Delta\phi - \nabla\phi \cdot \nabla U$$

First, note that the infinitesimal operator of the forward and backward operators coincide and reduce to the backward Fokker-Planck equation with potential  $U$ .

3.  $\alpha = 1$ :

$$\mathcal{H}_b^{(1)}\phi = \Delta\phi$$

The intention of choosing  $\alpha = 1$  is to “normalize the potential away” which is successful as the resulting infinitesimal generator is the Laplace-Beltrami operator and thus the corresponding diffusion only captures the geometry of the data, totally neglecting the underlying density  $e^{-U}$ . As such, it is optimal for tasks like manifold reconstruction.

### 2.1.3 Investigation of Stochastic Processes

First we start with the definition of a stochastic process:

**Definition 2.7.** *Let  $T$  be an ordered set. A stochastic process is a collection of random variables  $X = \{X_t, t \in T\}$  where, for each  $t \in T$ ,  $X_t : E \rightarrow \Omega$  is a random variable from a measurable space  $(E, \mathcal{F})$  to another one  $(\Omega, \mathcal{G})$ .*

From now on, let  $T = \mathbb{R}^+$  be the time and  $\Omega = \mathbb{R}^n$  the *state space*.<sup>3</sup>

Now we want to investigate the time evolution of a system that is described by its state  $x(t)$  at time  $t$ ,  $(x(t) \in \Omega)$  and is governed by the following stochastic differential equation

$$\dot{x} = -U(x) + \sqrt{2/\beta} \dot{w} \tag{2.1}$$

where  $U(x)$  is the potential energy of a state  $x$ ,  $\beta$  is some thermal factor which is inversely proportional to a given temperature and  $w(t)$  is the Brownian motion in  $n$  dimensions.

---

<sup>3</sup>It is important that  $x \in \Omega$  does *not* represent a particular particle, but a certain state of the entire system.

Integration of the SDE (2.1) produces random paths whose distribution defines a time dependent probability distributions on  $\Omega$ . To study the dynamics of the system, it is convenient to consider the time evolution of these probability distributions as it has been shown that the transition probability density  $p(x, t|x_0, 0)$  of finding the system at location  $x$  at time  $t$ , given an initial location  $x_0$  at time  $t = 0$  satisfies the forward Fokker-Planck equation

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$$\frac{\partial p}{\partial t} = \frac{1}{\beta} \Delta p + \nabla \cdot (p \nabla U) =: \mathcal{L}p \quad (2.2)$$

defined in  $(x, t) \in \Omega \times \mathbb{R}^+$  with reflecting boundary conditions on  $\partial\Omega$ , i.e.  $\frac{\partial p}{\partial n} = 0$  where  $\frac{\partial}{\partial n}$  is the derivative in the outward normal direction. Furthermore the initial condition is given by the Dirac-Delta function, i.e.  $\lim_{t \rightarrow 0} p(x, t|x_0, 0) = \delta(x - x_0)$ .

### Eigenvalues of the Fokker-Planck Operator

As seen in the first chapter one can compute the eigenvalues  $\lambda_{\varepsilon, i}^{(\alpha)}$  of  $T_{b, \varepsilon}^{(\alpha)}$  by computing those of  $T_{s, \varepsilon}^{(\alpha)}$ . From definition of  $\mathcal{H}_f^{(\alpha)}$  and  $\mathcal{H}_b^{(\alpha)}$  it follows that they share the eigenvalues  $\mu_i^{(\alpha)}$  which can be written as

$$\mu_i^{(\alpha)} = \lim_{\varepsilon \rightarrow 0} \frac{\lambda_{\varepsilon, i}^{(\alpha)} - 1}{\varepsilon}$$

and are discrete again. Remember that for  $\alpha = \frac{1}{2}$ ,  $\mathcal{H}_f^{(\alpha)}$  and  $\mathcal{H}_b^{(\alpha)}$  coincide and become the backward Fokker-Planck operator  $\mathcal{L}^*$

**Definition 2.8.**  $\mathcal{L}^*g = \frac{1}{\beta} \Delta g - \nabla g \cdot \nabla U$

**Lemma 2.1.1.**  $\mathcal{L}$  and  $\mathcal{L}^*$  are adjoint under the standard inner product.

*Proof.*

$$\begin{aligned} \langle \mathcal{L}^*f, g \rangle &= \left\langle \frac{1}{\beta} \Delta f - \nabla f \cdot \nabla U, g \right\rangle \\ &= \left\langle \frac{1}{\beta} \Delta f, g \right\rangle + \langle -\nabla f \cdot \nabla U, g \rangle \\ &= \left\langle f, \frac{1}{\beta} \Delta g \right\rangle + \langle -\nabla f, g \cdot \nabla U \rangle \\ &= \left\langle f, \frac{1}{\beta} \Delta g \right\rangle + \langle f, \operatorname{div}(g \cdot \nabla U) \rangle \\ &= \left\langle f, \frac{1}{\beta} \Delta g \right\rangle + \langle f, \nabla \cdot (g \cdot \nabla U) \rangle = \langle f, \mathcal{L}g \rangle \end{aligned}$$

□



$\mathcal{L}$  and  $\mathcal{L}^*$  being adjoint implies that  $\mu_i^{(\frac{1}{2})}$  are not only the eigenvalues of  $\mathcal{L}$ , but also of  $\mathcal{L}^*$ .

### Computing the Transition Probability Density

For simplicity's sake the calculations will be done in one dimension for the forward Fokker-Planck operator.

First of all, notice that in one dimension (2.2) may be written as

$$\begin{aligned}\frac{\partial p}{\partial t} &= \frac{\partial^2 p}{\partial x^2} \frac{1}{\beta} + \frac{\partial}{\partial x} \left( p \frac{\partial U}{\partial x} \right) \\ &= \frac{\partial^2 p}{\partial x^2} \underbrace{\frac{1}{\beta}}_{=:f(x)} + \frac{\partial p}{\partial x} \underbrace{\frac{\partial U}{\partial x}}_{=:g(x)} + p \underbrace{\frac{\partial^2 U}{\partial x^2}}_{=:h(x)}\end{aligned}\tag{2.3}$$

and the reflecting boundary conditions simplify to  $p(a, t) = p(b, t) = 0$ .

Now apply separation of variables  $p(x, t) = X(x)T(t)$  and rewrite (2.3) as

$$\frac{\mathcal{L}X(x)}{X(x)} = \frac{\frac{d}{dt}T(t)}{T(t)}.\tag{2.4}$$

Since both sides of (2.4) only depend on one variable each, they must be equal to a constant, i.e.

$$\mathcal{L}X(x) = \lambda X(x)\tag{2.5}$$

and

$$\frac{d}{dt}T(t) = \lambda T(t).$$

(2.5) is *almost* a regular Sturm-Liouville problem. In order get it into the form

$$\frac{d}{dx} \left[ \tilde{f}(x) \frac{d\varphi_n}{dx} \right] + \tilde{h}(x)\varphi_n = \tilde{f}(x) \frac{d^2\varphi_n}{dx^2} + \frac{d\tilde{f}(x)}{dx} \frac{d\varphi_n}{dx} + \tilde{h}(x)\varphi_n = -\lambda_n w(x)\varphi_n$$

we consider the space  $L^2([a, b], w)$  with

$$w = \frac{e^{\int \frac{g}{f}}}{f} = \beta e^{\beta U}\tag{2.6}$$

and

$$\langle \phi, \psi \rangle_w = \int_a^b \phi(x) \psi(x) w(x) dx.$$

(2.6) is not only the eigenfunction to the first eigenvalue  $\lambda_0 = 0$  (plug it into (2.3)), but, if normalized to absolute value 1, is also known as the Boltzmann equilibrium distribution.

We also get  $X(a) = X(b) = 0$ . Let  $\varphi_n$  and  $\lambda_n$  be the eigenfunctions<sup>4</sup> and eigenvalues of  $\mathcal{L}$  and analogously  $\frac{d}{dt}T_n(t) = \lambda_n T(t)$ . Note that, as a Sturm-Liouville operator in  $L^2([a, b], w(x)dx)$ ,  $\mathcal{L}$  is self-adjoint and its eigenfunctions  $\varphi_n$  are orthogonal and maximal. Without loss of generality assume  $\varphi_n$  to be orthonormal.

Now we can write  $T_n(t)$  as

$$T_n(t) = a_n e^{-\lambda_n t} \quad (2.7)$$

for some constants  $a_n$  that can be written as

$$a_n = \langle \varphi_n, p(\cdot, 0) \rangle_w$$

Note that  $a_n$  is a generalized fourier series of the initial distribution  $p(\cdot, 0)$ , i.e.  $p(\cdot, 0) = \sum_k a_k \varphi_k$  on  $[a, b]$  because the set of orthonormal eigenfunctions forms a basis. It is chosen such that (2.7) holds true for  $t = 0$ . As a result

$$p(x, t) = \sum_{n \geq 0} a_n \varphi_n(x) e^{-\lambda_n t} \quad (2.8)$$

From that representation we get that, as  $t \rightarrow \infty$ , only the first term with  $\lambda_0 = 0$  survives which justifies the naming of (2.6) as Boltzmann *equilibrium* distribution.<sup>5</sup>

Analogously, for every smooth  $f \in L^2(\Omega)$ , one can consider  $g(x, t) = \mathbb{E}\{f(x(t)) | x(0) = x\}$  which satisfies the backward Fokker-Planck equation

$$\frac{\partial g}{\partial t} = \frac{1}{\beta} \Delta g - \nabla g \cdot \nabla U = \mathcal{L}^* g$$

in the domain  $(x, t) \in \Omega \times \mathbb{R}^+$ , with initial conditions

<sup>4</sup>There is no general analytic (exact) solution to Sturm-Liouville problems, but solving a Sturm-Liouville problem is a easier than solving the whole PDE.

<sup>5</sup>Under the assumption that the system is ergodic, the ergodic theorem tells us that the time average and the space average are the same almost everywhere; as such (and because  $\varphi_0$  is continuous), regardless of the initial configuration  $x_0 \in \Omega$ ,  $\lim_{t \rightarrow \infty} p(x, t | x_0, 0) = \varphi_0(x)$

find something to cite that this holds for SL problems with appropriate boundary conditions

not in this case(?) since  $p(x, 0) = \delta(x_0 - x)$  and the Dirac-Delta function is not a smooth function - any way to res-cue this ansatz?

$$g(x, 0) = f(x)$$

Remember that, as the adjoint<sup>6</sup> of  $\mathcal{L}$ ,  $\mathcal{L}^*$  shares the eigenvalues  $\lambda_n$ . It is easy to see that  $\psi_0(x) = 1$  is the eigenfunction corresponding to the eigenvalue  $\lambda_0 = 0$ .

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<sup>6</sup>under the standard inner product

## Appendix A

# Asymptotics for Laplacian Operators

In the following we will deal with a compact manifold  $\mathcal{M}$  that is  $C^\infty$ . Let  $x$  be a fixed point, not on the boundary,  $T_x\mathcal{M}$  be the tangent space to  $\mathcal{M}$  at  $x$  and  $(e_1, \dots, e_d)$  be a fixed orthonormal basis of  $T_x\mathcal{M}$ . Furthermore two systems of local coordinates will be introduced:

1. (*Normal coordinates*) The exponential map  $\exp_x$  generates a set of orthogonal geodesics  $(\gamma_1, \dots, \gamma_d)$  intersecting at  $x$  with initial velocity  $(e_1, \dots, e_d)$ . Now every point  $y \in \mathcal{M}$  in a sufficiently small neighborhood of  $x$  has a set of *normal coordinates*  $(s_1, \dots, s_d)$  along these geodesics.
2. (*Tangent coordinates*) Considering the orthogonal projection  $u$  of  $y$  on  $T_x\mathcal{M}$ , where  $u_i = \langle y - x, e_i \rangle$  in  $(e_1, \dots, e_d)$ , we get a system of *tangent coordinates*. The submanifold is now locally parameterized as  $y = (u, g(u))$ , where  $g : \mathbb{R}^d \rightarrow \mathbb{R}^{n-d}$ . Since  $u = (u_1, \dots, u_d)$  are tangent coordinates, we must have that  $\partial g(0) = 0$ .

Notice that, locally, any function  $f$  on  $\mathcal{M}$  may be viewed as  $\tilde{f}$  of  $(s_1, \dots, s_d)$  and thus we may write  $\Delta f(x) = -\sum_{i=1}^d \frac{\partial^2 \tilde{f}}{\partial s_i^2}(0, \dots, 0)$ , where  $\Delta$  is the Laplace-Beltrami operator on  $\mathcal{M}$ .

### A.1 Comparison of the Geodesic and the Local Projection

In this section we will compute asymptotic expansions for the changes of

write introduction + motivation (later being the justification for naming it graph “Laplacian” earlier)

The following is basically a citation of Coifman and Lafon (2006).

variable  $u \mapsto (s_1, \dots, s_d)$  and  $u \mapsto y$ .

In the following,  $Q_{x,m}(u)$  denotes a generic homogeneous polynomial of degree  $m$  of the variable  $u = (u_1, \dots, u_d)$ , whose coefficient depends on  $x$ .

**Lemma A.1.1.** *If  $y \in \mathcal{M}$  is in a Euclidean ball of radius  $\varepsilon^{\frac{1}{2}}$  around  $x$ , then, for  $\varepsilon$  sufficiently small, there exists:*

$$s_i = u_i + Q_{x,3}(u) + \mathcal{O}(\varepsilon^2) \quad (\text{A.1})$$

*Proof.* Let  $\gamma$  be the geodesic connection  $x$  and  $y$  parameterized by arclength. We have  $\gamma(0) = x$  and let  $s$  be such that  $\gamma(s) = y$ . If  $y$  has normal coordinates  $(s_1, \dots, s_d)$ , then we have  $s\gamma'(0) = (s_1, \dots, s_d)$ . A Taylor expansion yields

$$\gamma(s) = \gamma(0) + s\gamma'(0) + \frac{s^2}{2}\gamma''(0) + \frac{s^3}{6}\gamma^{(3)}(0) + \mathcal{O}(\varepsilon^2).$$

By definition of a geodesic, the covariant derivative of the velocity is zero, which means that  $\gamma''(0)$  is orthogonal to the tangent plane at  $x$ . Now since the parameter  $u_i$  is defined by  $u_i = \langle \gamma(s) - \gamma(0), e_i \rangle$ , we obtain that  $u_i = s_i + \frac{s^3}{6}\langle \gamma^{(3)}(0), e_i \rangle + \mathcal{O}(\varepsilon^2)$ . Iterating this equation yields the result.  $\square$

**Lemma A.1.2.** *Again, let  $y \in \mathcal{M}$  be in a Euclidean ball of radius  $\varepsilon^{\frac{1}{2}}$  around  $x$ ; we have*

$$\|x - y\|^2 = \|u\|^2 + Q_{x,4}(u) + Q_{x,4}(u) + \mathcal{O}(\varepsilon^3) \quad (\text{A.2})$$

and

$$\det \left( \frac{dy}{du} \right) = 1 + Q_{x,2}(u) + Q_{x,3}(u) + \mathcal{O}(\varepsilon^2). \quad (\text{A.3})$$

*Proof.* The submanifold is locally parameterized as  $u \mapsto (u, g(u))$ , where  $g : \mathbb{R}^d \rightarrow \mathbb{R}^{n-d}$ . Writing  $g = (g_{i+1}, \dots, g_n)$  and applying Pythagore's theorem, we obtain

$$\|x - y\|^2 = \|u\|^2 + \sum_{i=d+1}^n g_i(u)^2.$$

Using that, by definition,  $g_i(0) = 0$  and, as noted before,  $\frac{\partial g}{\partial u_i}(0) = 0$ . As a consequence  $g_i(u) = b_{i,x}(u) + c_{i,x}(u) + \mathcal{O}(\varepsilon^2)$ , where  $b_{i,x}$  is the Hessian quadratic form of  $g_i$  at  $u = 0$  and  $c_{i,x}$  is the cubic term. This proves (A.2) with

$$Q_{x,4}(u) = \sum_{i=d+1}^n b_{i,x}^2(u) \text{ and } Q_{x,5}(u) = 2 \sum_{i=d+1}^n b_{i,x}(u)c_{i,x}(u).$$

To prove (A.3), observe that  $\frac{\partial g}{\partial u_i}(0) = 0$  implies that  $\frac{\partial g}{\partial u_i}(0) = \tilde{b}_{i,x}(u) + \tilde{c}_{i,x}(u) + \mathcal{O}(\varepsilon^{\frac{3}{2}})$ , where  $\tilde{b}_{i,x}(u)$  and  $\tilde{c}_{i,x}(u)$  are the linear and quadratic terms in the Taylor expansion of  $\frac{\partial g}{\partial u_i}(0)$  at 0. We thus have:

$$\begin{aligned} \frac{\partial y}{\partial u_i}(u) &= \left( v_i, \frac{\partial g}{\partial u_i}(u) \right), \text{ where } v_i = (0, \dots, 0, 1, 0, \dots, 0) \in \mathbb{R}^d \\ &= (v_i, \tilde{b}_{i,x}(u) + \tilde{c}_{i,x}(u) + \mathcal{O}(\varepsilon^{\frac{3}{2}})). \end{aligned}$$

The squared volume generated by these  $d$  vectors is the determinant of their Gram matrix, i.e.,

$$\left| \det \left( \frac{dy}{du} \right) \right|^2 = \sum_{i,j=1}^d E_{ij}(u) + \sum_{i,j=1}^d F_{ij}(u) + \mathcal{O}(\varepsilon^2),$$

where

$$E_{ij}(u) = \langle \tilde{b}_{i,x}(u), \tilde{b}_{j,x}(u) \rangle \text{ and } F_{ij}(u) = \langle \tilde{b}_{i,x}(u), \tilde{c}_{j,x}(u) \rangle + \langle \tilde{c}_{i,x}(u), \tilde{b}_{j,x}(u) \rangle.$$

Defining

$$Q_{x,2}(u) = \sum_{i,j=1}^d E_{ij}(u) \text{ and } Q_{x,3}(u) = \sum_{i,j=1}^d F_{ij}(u),$$

we obtain the last result.  $\square$

## A.2 Infinitesimal Operators for a Family of Graph Laplacians

In this section we present the calculation of the infinitesimal generators for the different diffusion maps characterized by a parameter  $\alpha$ .

To start with we first show an asymptotic expansion for diffusion operators  $G_\varepsilon$ .

Let  $k_\varepsilon(x, y)$  be an isotropic kernel, i.e.:

$$k_\varepsilon(x, y) = h \left( \frac{\|x - y\|^2}{\varepsilon} \right),$$

where  $h$  is assumed to have an exponential decay and let  $G_\varepsilon$  be the corresponding operator

$$G_\varepsilon f(x) = \frac{1}{\varepsilon^{\frac{d}{2}}} \int_{\mathcal{M}} k_\varepsilon(x, y) f(y) dy.$$

The idea is that, using the previous lemmata, for small  $\varepsilon$  integrating  $f$  against the kernel on the manifold is approximately like integrating on the tangent space.

**Theorem A.2.1.** *Let  $f \in C^3(\mathcal{M})$  and let  $0 < \gamma < 1/2$ . Then we have, uniformly for all  $x \in \mathcal{M}$  at distance larger than  $\varepsilon^\gamma$  from  $\partial\mathcal{M}$ ,*

$$G_\varepsilon f(x) = m_0 f(x) + \varepsilon \frac{m_2}{2} (\omega(x) f(x) - \Delta f(x)) + \mathcal{O}(\varepsilon^2),$$

where

$$m_0 = \int_{\mathbb{R}^d} h(|u|^2) du \text{ and } m_2 = \int_{\mathbb{R}^d} u_1^2 h(|u|^2) du$$

and  $\omega$  is a potential term depending on the embedding of  $\mathcal{M}$ .

*Proof.* Because of the exponential decay of  $h$ , the domain of integration can be restricted to the intersection of  $\mathcal{M}$  with the ball of radius  $\varepsilon^\gamma$  around  $x$ . In doing so we generate an error of order

$$\begin{aligned} \left| \frac{1}{\varepsilon^{\frac{d}{2}}} \int_{\substack{y \in \mathcal{M} \\ \|x-y\| > \varepsilon^\gamma}} h\left(\frac{\|x-y\|^2}{\varepsilon}\right) f(y) dy \right| &\leq \|f\|_\infty \frac{1}{\varepsilon^{\frac{d}{2}}} \int_{\substack{y \in \mathcal{M} \\ \|x-y\| > \varepsilon^\gamma}} \left| h\left(\frac{\|x-y\|^2}{\varepsilon}\right) \right| dy \\ &\leq \|f\|_\infty \int_{\substack{y \in \mathcal{M} \\ \|y\| > \varepsilon^{\gamma-1/2}}} |h(|y|^2)| dy \\ &\leq C \|f\|_\infty Q(\varepsilon^{1/2-\gamma}) e^{-\varepsilon^{\gamma-1/2}}, \end{aligned}$$

where we have used the exponential decay of the kernel and where  $Q$  is a polynomial. Since  $0 < \gamma < 1/2$ , this term is exponentially small and is bounded by  $\mathcal{O}(\varepsilon^{\frac{3}{2}})$ . Therefore,

$$G_\varepsilon f(x) = \frac{1}{\varepsilon^{\frac{d}{2}}} \int_{\substack{y \in \mathcal{M} \\ \|x-y\| < \varepsilon^\gamma}} h\left(\frac{\|x-y\|^2}{\varepsilon}\right) f(y) dy + \mathcal{O}(\varepsilon^{\frac{3}{2}}).$$

Now that things are localized around  $x$ , we can Taylor-expand the function  $(s_1, \dots, s_d) \mapsto f(y(s_1, \dots, s_d))$ :

$$f(y) = f(x) + \sum_{i=1}^d s_i \frac{\partial \tilde{f}}{\partial s_i}(0) + \frac{1}{2} \sum_{i,j=1}^d s_i s_j \frac{\partial^2 \tilde{f}}{\partial s_i \partial s_j}(0) + Q_{x,3}(s_1, \dots, s_d) + \mathcal{O}(\varepsilon^2),$$

where  $\tilde{f}(s_1, \dots, s_d) = f(y(s_1, \dots, s_d))$ . Invoking (A.1), we obtain

$$f(y) = \tilde{f}(0) + \sum_{i=1}^d u_i \frac{\partial \tilde{f}}{\partial s_i}(0) + \frac{1}{2} \sum_{i,j=1}^d u_i u_j \frac{\partial^2 \tilde{f}}{\partial s_i \partial s_j}(0) + Q_{x,3}(u) + \mathcal{O}(\varepsilon^2).$$

Likewise, because of (A.2), the Taylor expansion of the kernel is

$$h\left(\frac{\|x - y\|^2}{\varepsilon}\right) = h\left(\frac{\|u\|^2}{\varepsilon}\right) + \left(\frac{Q_{x,4}(u) + Q_{x,5}(u)}{\varepsilon}\right) h'\left(\frac{\|u\|^2}{\varepsilon}\right) + \mathcal{O}(\varepsilon^2).$$

Using (A.3) to change the variable  $s \mapsto u$  in the previous integral defining  $G_\varepsilon f(x)$  yields:

$$\begin{aligned} \varepsilon^{\frac{d}{2}} G_\varepsilon f(x) &= \int_{\|u\| < \varepsilon^\gamma} \left[ h\left(\frac{\|u\|^2}{\varepsilon}\right) + \left(\frac{Q_{x,4}(u) + Q_{x,5}(u)}{\varepsilon}\right) h'\left(\frac{\|u\|^2}{\varepsilon}\right) \right] \\ &\times \left[ \tilde{f}(0) + \sum_{i=1}^d u_i \frac{\partial \tilde{f}}{\partial s_i}(0) + \frac{1}{2} \sum_{i,j=1}^d u_i u_j \frac{\partial^2 \tilde{f}}{\partial s_i \partial s_j}(0) + Q_{x,3}(u) \right] \\ &\times (1 + Q_{x,2}(u) + Q_{x,3}(u)) du + \mathcal{O}\left(\varepsilon^{\frac{d}{2}+2}\right). \end{aligned}$$

This identity can be dramatically simplified by identifying odd functions and setting their integral to zero. One is left with

$$\begin{aligned} \varepsilon^{\frac{d}{2}} G_\varepsilon f(x) &= \tilde{f}(0) \int_{\mathbb{R}^d} h\left(\frac{\|u\|^2}{\varepsilon}\right) du + \frac{1}{2} \left( \sum_{i=1}^d \frac{\partial^2 \tilde{f}}{\partial s_i^2}(0) \right) \int_{\mathbb{R}^d} u_1^2 h\left(\frac{\|u\|^2}{\varepsilon}\right) du \\ &+ \tilde{f}(0) \int_{\mathbb{R}^d} \left[ \frac{Q_{x,4}(u)}{\varepsilon} h'\left(\frac{\|u\|^2}{\varepsilon}\right) + \tilde{Q}_{x,2}(u) h\left(\frac{\|u\|^2}{\varepsilon}\right) \right] du + \mathcal{O}\left(\varepsilon^{\frac{d}{2}+2}\right), \end{aligned}$$



where the domain of integration has been extended to  $\mathbb{R}^d$  (exponential decay of  $h$ ). Changing the variable according to  $u \mapsto \varepsilon^{\frac{1}{2}}u$ ,

$$\begin{aligned} G_\varepsilon f(x) &= \tilde{f}(0) \int_{\mathbb{R}^d} h(\|u\|^2) du + \frac{\varepsilon}{2} \left( \sum_{i=1}^d \frac{\partial^2 \tilde{f}}{\partial s_i^2}(0) \right) \int_{\mathbb{R}^d} u_1^2 h(\|u\|^2) du \\ &\quad + \varepsilon \tilde{f}(0) \int_{\mathbb{R}^d} (Q_{x,4}(u) h'(\|u\|^2) + Q_{x,2}(u) h(\|u\|^2)) du + \mathcal{O}(\varepsilon^2), \end{aligned}$$

where we have used the homogeneity of  $Q_{x,2}$  and  $Q_{x,4}$ . Finally, observing that

$$\tilde{f}(0) = f(x) \text{ and } \sum_{i=1}^d \frac{\partial^2 \tilde{f}}{\partial s_i^2}(0) = -\Delta f(x),$$

we end up with

$$G_\varepsilon f(x) = m_0 f(x) + \varepsilon \frac{m_2}{2} (\omega(x) f(x) - \Delta f(x)) + \mathcal{O}(\varepsilon^2),$$

where

$$\omega(x) = \frac{2}{m_2} \int_{\mathbb{R}^d} (Q_{x,4}(u) h'(\|u\|^2) + Q_{x,2}(u) h(\|u\|^2)) du.$$

Finally, the uniformity follows from the compactness and smoothness of  $\mathcal{M}$ .  $\square$

One can also show the same result holds for all  $x \in \mathcal{M}$  at distance smaller than  $\varepsilon^\gamma$  from  $\partial\mathcal{M}$ . For this proof we refer to Coifman and Lafon (2006) as this would go beyond the scope of the discussion.

Suppose that the data set  $X$  consists of a Riemannian manifold with a density  $p(x) = e^{-U(x)}$  and let  $k_\varepsilon(x, y)$  be a Gaussian kernel (which clearly satisfies the requirements for A.2.1).

**Theorem A.2.2.** *The infinitesimal generator  $\mathcal{H}_b \phi$  of the backward operator  $T_{b,\varepsilon}^{(\alpha)} \phi = \int_{\Gamma} \frac{k_\varepsilon^{(\alpha)}(x, y)}{d_\varepsilon^{(\alpha)}(x)} \phi(y) p(y) dy$  is  $\Delta \phi - 2(1 - \alpha) \nabla \phi \cdot \nabla U$ .*

*Proof.* From A.2.1 we see that<sup>1</sup>

$$p_\varepsilon(x) = p(x) + \varepsilon(\Delta p(x) + \omega(x)p(x)) + \mathcal{O}(\varepsilon^2)$$

---

<sup>1</sup>w.l.o.g. we assume  $m_0 = 1$  and  $m_2 = 2$

and consequently,

$$p_\varepsilon^{-\alpha} = p^{-\alpha} \left( 1 - \alpha\varepsilon \left( \frac{\Delta p}{p} + \omega \right) \right) (1 + \mathcal{O}(\varepsilon^2)).$$

Let

$$k_\varepsilon^{(\alpha)}(x, y) = \frac{k_\varepsilon(x, y)}{p_\varepsilon^\alpha(x) p_\varepsilon^\alpha(y)}$$

Then, the normalization factor  $d_\varepsilon^{(\alpha)}$  is given by

$$d_\varepsilon^{(\alpha)}(x) = \int_\Gamma k_\varepsilon^{(\alpha)}(x, y) p(y) dy = p_\varepsilon^{-\alpha}(x) p_\varepsilon^{1-\alpha}(x) \left[ 1 + \varepsilon \left( (1 - \alpha)\omega - \alpha \frac{\Delta p}{p} + \frac{\Delta p^{1-\alpha}}{p^{1-\alpha}(x)} \right) \right].$$

Therefore, the asymptotic expansion of the backward operator gives

$$T_b^{(\alpha)} \phi = \int_\Gamma \frac{k_\varepsilon^{(\alpha)}(x, y)}{d_\varepsilon^{(\alpha)}(x)} \phi(y) p(y) dy = \phi(x) + \varepsilon \left( \frac{\Delta(\phi p^{1-\alpha})}{p^{1-\alpha}} - \phi \frac{\Delta p^{1-\alpha}}{p^{1-\alpha}} \right)$$

and its infinitesimal generator is

$$\mathcal{H}_b \phi = \lim_{\varepsilon \rightarrow 0} \frac{T_b - I}{\varepsilon} \phi = \frac{\Delta(\phi p^{1-\alpha})}{p^{1-\alpha}} - \phi \frac{\Delta p^{1-\alpha}}{p^{1-\alpha}}.$$

Plugging in  $p = e^{-U}$  into the last equation gives the desired result.  $\square$

Similarly, one can show that the infinitesimal generator  $\mathcal{H}_f \psi^{(\alpha)}$  of the forward operator  $T_{f, \varepsilon}^{(\alpha)} \psi$  is  $\Delta \psi - 2\alpha \nabla \psi \cdot \nabla U + (2\alpha - 1) \psi (\nabla U \cdot \nabla U - \Delta U)$ .

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