Atmospheric Correction Tool for WISE MAN (ACWISE)

User Manual (v0.1)

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Introduction

ACWISE is a software package which is developed for atmospheric correction of WISE image. For the current version (v0.1), it can only handle the Level1G images which are geo-referenced radiance. To run it successfully, some other WISE data are also needed, like the navigation log and geo reference LUTs, etc. More details are given in section 4.

This package was written in pure Python, it runs in environment of Python>=3.5, but it may occur some warnings in Python 3.7 and 3.8. To run it, a bunch of 3rd part of libraries are required. To make it easier for user who is not very familiar with Python, this package is installed in a singularity image and all of the dependencies have been installed.

Quick start

Following the steps below to run it:

1. Go into the singularity sif image. Note that users don't have the write permission, so it needs to mount the external directory that contains the WISE images to be processed.

singularity shell --bind /data:/mnt acwise.sif

2. Set the run environment

Singularity> source /opt/env.sh

3. Run acwise to get help info

4. To run acwise, a configure file is mandatory.

The configure file

To make the software package more extensible, a bunch of options can be set in the configure file. But for the current version which only implemented the basic function, most of options are fixed. Here is a template for the configure file, the options marked as yellow are fixed, it can not be modified by now. The details of the options are given as below:

```
1 [SENSOR]
2 NAME=WISE
4 [WISE]
5 IMAGE_NAME = 190820_MC-16A-WI-1x1x1_v01
6 L1G_DIR = /mnt/WISE_IMAGE/L1G/
7 L1A_GLU_DIR = /mnt/WISE_IMAGE/L1A-GLU/
8 L1A DIR = /mnt/WISE IMAGE/L1A-GLU/
9 NAV_LOG_DIR = /mnt/WISE_IMAGE/NAV-LOG_SUM/
10
11 [AC]
12 #tile size, unit=km
13 TILE SIZE = 4
14 PROCEDURE = STANDARD
15 #ANCILLARY SOURCE = aer5
16 ANCILLARY_SOURCE = None
17 OZONE = 0.3
18 WATER_VAPOR = 2.0
19 WIND SPEED = 6.0
21 [STANDARD]
22 GAS CALCULATOR = SixS
23 GLINT CALCULATOR = CoxMunk
24 RAY_CALCULATOR = SixS
25 ADJ ALGRITHM = None
26 AERO ALGRITHM = DSF
27 AERO CALCULATOR = SixS
28
29 [DSF]
30 FITTING = 0
31 DARK_SPECTRUM_NUM = 5
32 DARK_PIXEL_NUM = 20
34 [OUTPUT]
35 #The ourput dir
36 DIR = /mnt/WISE_IMAGE
37 NAME = $NAME$ L2
39 [DATA]
40 DATA DIR = /mnt/ac4icw data
```

[WISE]

This is set for the input of the level-1 WISE image. **IMAGE_NAME** refers to the name of the image to be processed, for the pattern of the image name, please refers to the WISE flight document. L1G_DIR is for the directory that contains the L1G images. L1A_GLU_DIR and NAV_LOG_DIR are for the directories that contain the georeference LUTs (\$IMAGE_NAME\$-L1A.glu.hdr) and the navigation log files (\$IMAGE_NAME\$-Navcor sum.log), respectively.

[AC]

This section configures some general information about atmospheric correction. Currently, the entire image will be split into several tiles (squares) for processing. The **TILE_SIZE** option is for the size of each tile, the unit of the size is *km*. **PROCEDURE= STANDADRD** means the atmospheric correction will be performed in a standard procedure, step by step. Except for STANDARD, the other kinds of procedure is not implemented yet. The **STANDARD** procedure is further configured in the **STANDARD** section. The ancillary data mainly includes the total concentration of **OZONE** (m.atm), **WATER VAPOR** (g cm²) and **WIND SPEED** (m s¹¹). These data can only be fixed manually, since the automatically way is not implemented yet.

[STANDARD]

This section is set for the gas absorption, Rayleigh, glint, adjacency effect and aerosol correction. Note that adjacency effect correction algorithm was not implemented yet, so the option of ADJ ALGORIHTM should be set None. The calculation of gas absorption, Rayleigh and aerosol are all based on 6SV, so they are set as SixS. The calculation of glint based the Cox&Munk model, GLINT CALCULATION=CoxMunk. For the aerosol retrieval algorithm, only the DSF (Quinten et al,2018) algorithm is implemented in current version, so **AERO ALGORIHTM=DSF** (the details of DSF is further configured in the DSF section). If one of the options are set **None**, the related correction will be skipped, for example, no aerosol correction will be performed when AERO ALGORIHTM =None, but the Rayleigh correction is mandatory, mean RAY CALCULATOR should not be set None.

[DSF]

The darkest pixel could be determined directly or by linear fitting, which is controlled by the **FITTING** option, the value of 1 and 0 is a switch of using linear fitting or not. **DARK_SPECTRUM_NUM** determines the number of bands of the darkest pixel. For the hyperspectral WISE image, the default value is set 5, which means the aerosol reflectance spectra will be determined by darkest pixels in 5 bands. **DARK_PIXEL_NUM** is the number (the default value is 20) of real darkest pixels which are used to determine the ideal darkest pixel by linear fitting, it works only when **FITTING=1**.

[OUTPUT]

The level-2 file is save as GEOTIFF. Currently, only the Rrs product is saved in the level-2 file. The DIR option specifies the directory to save the level-2 file, and NAME specifies the name of the level-2 file. **\$NAME\$** will be replaced by **IMAGE NAME**.

[DATA]

DATA_DIR specifies the directory that contains the internal data that will be used by the package. By default, the data will be located in /opt/acwise data.

Methodology

1. Radiative transfer and simulations

The basic theory atmospheric radiative transfer and the simulations of each term (except for the glint reflectance) are based 6SV. The glint reflectance was simulated using the Cox&Monk model.

The apparent reflectance ρ^* at the sensor level is expressed as (6SV):

2.4.2. Airborne sensor simulation

In the case of a sensor inside the atmosphere (aircraft simulation), Eq (5) is modified as follows:

$$\rho * (\theta_s, \theta_v, \phi_s - \phi_v, z) = T_g(\theta_s, \theta_v, z) \left[\rho_r(z) + \rho_a(z) + \frac{\rho_t}{1 - S(z)\rho_t} T(\theta_v, z) T(\theta_s, z) \right]$$
(19)

1.1 Relative Spectral Response (RSR) of WISE

The RSR of WISE is calculated based on the center wavelength, the Full Width at Half Maximum (FWHM) recorded in the header file of WISE and the assumption of Gaussian distribution. For each channel, *RSR* is calculated using Equation 1:

$$RSR = \exp\left[-\frac{(x-b)^2}{2(\frac{FWHM}{\sqrt{2*\log(2)}})^2}\right]$$
(1)

where, x refers to the full wavelength, for example, from 400 to 1200 nm with interval of 0.5 nm; b refers to the center wavelength of the channel.

1.2 LUT of transmittance due to gas absorption $T_q(\theta_s, \theta_v, z)$

The downwelling and upwelling of transmittance are expressed as $T_g(\theta_s)$ and $T_g(\theta_v)$. The values of θ_s and θ_v are set as:

$$\theta_s = 15, 25, 35, 45, 55, 65$$

 $\theta_v = 0, 5, 10, 15, 25, 35, 45$

The atmosphere profile was defined by water vapor (wv) and O₃, the values are set as

$$wv = 1, 2, 4, 6$$
 ()
o₃ = 0.2, 0.3, 0.4, 0.5 ()

1.3 LUT of the Rayleigh scattering reflectance ho_r ,diffuse transmittance T_r and spherical albedo S_r due to Rayleigh scattering

The values θ_s and θ_v refer to section 1.1. The values relative azimuth ϕ are set as:

$$\phi = 0.15,30,45,60,75,90,105,120,135,150,165,180$$

The atmosphere profile is set as "SubarcticSummer".

1.4 LUTs of the aerosol scattering reflectance ho_a ,diffuse transmittance T_a and spherical albedo S_a due to aerosol scattering

The settings of θ_s , θ_v and ϕ refer to section 1.2.

The aerosol type is defined by the ratio of the four compositions, soot, water, oceanic and dust. The ratio interval was set as 0.2, therefore, 36 aerosol types in total were generated.

The aerosol optical thickness in 550 nm $\tau_a(550)$ are set as: $\tau_a(550) = 0.005, 0.01, 0.02, 0.04, 0.08, 0.16, 0.32$

1.5 LUTs of the path reflectance ρ_P , the diffuse irradiance E_{diff} reflected by the water surface ρ_{sky_w} , the total diffuse transmittance due to Rayleigh and aerosol scattering T, the total spherical albedo S.

Note that ρ_{sky_w} is calculated based on the diffuse irradiance E_{diff} at the ground level using Eq. 1-2:

Type equation here.

$$\rho_{sky_w} = \frac{E_{diff}}{F_0 \cos(\theta_s)} r(1)$$

$$r = 0.0256 + 0.00039w + 0.000034w^2 (2)$$

1.6 LUT of glint reflectance ρ_G

Glint reflectance is simulated using Cox&Monk model, and only the wind speed but not the wind direction is taken into account.

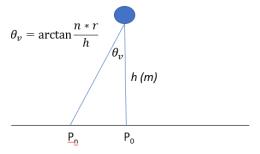
2. Image Processing

2.1 Pre-processing

2.1.1 Calculation of observing geometry

The solar zenith angle θ_s and azimuth ϕ_s is calculated based on the local timezone, local time, longitude and latitude.

The viewing zenith angle θ_v is calculated based on the height (h), spatial resolution (r) and the pixel offset from the center. Note that the Curvature of the earth is ignored since the swath WISE is less than 10 km. In addition, the *pitch* and *roll* of the plan is assumed to be 0 and without change during measuring. Figure 1. Shows the geometry.



 P_0 : the central pixel of the scan line P_0 : the n^{th} pixel offset P_0 r: spatial resolution of one pixel h: heigh of the plane

Figure 1. The basic geometry for θ_{v} calculation

The geometry for calculating the viewing azimuth angle ϕ_v is shown in Figure 2. Note that the pitch and roll of the plane are assumed to be 0, and there is no change during the measurement, and the heading is also assumed to be no change.

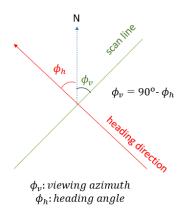


Figure 2. The basic geometry for ϕ_{v} calculation

2.1.2 Convert TOA radiance to reflectance

2.2 The "STANDDARD" atmospheric correction

2.3 The "PATHAC" atmospheric correction

This AC algorithm assumes the path reflectance are spatial homogenous, and the path reflectance and upward and downward transmittance due to scattering are determined based on the DSF algorithm proposed by Vanhellemont (2019). LUTs that support this AC method are generated using 6SV, details about the simulation are found in section 1.5.