Midterm 1

Paul Fischer Department of Physics California State University Long Beach

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Abstract

The Runge-Kutta method RK4 is used to study a double pendulum (Problem 1) and a vibrating neutron star (Problem 2). We analyze the energies of the double pendulum for four different sets of initial conditions and observe their corresponding Poincare sections. We observe that a vibrating neutron star will have a nearly constant z-component while also nutating about the z-axis.

1 Introduction

Ever since Newton's 2^{nd} Law was published, Physics has been formulated in the language of differential equations. Many nonlinear ordinary differential equations that arise in physics have no analytical solutions in terms of elementary functions. Physicists often use numerical approximation methods to simulate such systems. Around the year 1900, Mathematicians Carl Runge and Wilhelm Kutta developed a series of implicit and explicit iterative methods for approximating solutions to ordinary differential equations [1]. At each iteration, the method uses a trial step in the middle of a determined interval to cancel lower-order terms. The code we will use takes advantage of the RK4 method, which is a fourth-order formula which is shown below [2]:

$$k_1 = hf(x_n, y_n)$$

 $k_2 = hf(x_n + \frac{1}{2}h, y_n + \frac{1}{2}k_1)$

$$k_3 = hf(x_n + \frac{1}{2}h, y_n + \frac{1}{2}k_2)$$

$$k_4 = hf(x_n + h, y_n + k_3)$$

$$y_{n+1} = y_n + \frac{1}{6}k_1 + \frac{1}{3}k_2 + \frac{1}{3}k_3 + \frac{1}{6}k_4 + O(h^5)$$

This midterm will be organized as follows. In section Problem 1 (Analysis), the answers to Problem 1 will be discussed qualitatively, followed by the code in section Problem 1 (Code) and the figures in the section Problem 1 (Figures). The length of the code and number of figures required for Problem 1 have warranted them getting their own sections. Section Problem 2 contains an analysis of the vibrating neutron star simulation. This will be followed by the Summary and Conclusions section.

2 Problem 1 (Analysis)

For the first problem we are asked to solve the problem of the double pendulum using the rk4 method. This is done using the subroutine rk4step. We create an array $y = (\theta_1, \omega_1, \theta_2, \omega_2)$ and insert parameters g = 1, $l_1 = 2$, $m_1 = 3$, $l_2 = 1$, and $m_2 = 1$. We must first determine the energies for different initial conditions, namely $E_1(0,0,0,0)$, $E_2(\pi,0,0,0)$, $E_3(0,0,\pi,0)$, and $E_4(\pi,0,\pi,0)$. To determine the energy of the system given the initial conditions, we must derive its kinetic and potential energies in plane-polar coordinates. We start with the parametrization of (x_1, y_1, x_2, y_2) in terms of θ_1 and θ_2 , where we set the resting position of m_2 as $(x_2, y_2) = (0, 0)$ (this is a personal choice, not a necessary one):

$$x_1 = l_1 sin\theta_1$$

$$y_1 = l_2 + l_1 (1 - cos\theta_1)$$

$$x_2 = l_1 sin\theta_1 + l_2 sin\theta_2$$

$$y_2 = l_1 (1 - cos\theta_1) + l_2 (1 - cos\theta_2)$$

We plug this parametrization into the kinetic and potential energies for the system:

$$T = \frac{1}{2}m_1(\dot{x}_1^2 + \dot{y}_1^2) + \frac{1}{2}m_2(\dot{x}_2^2 + \dot{y}_2^2)$$

= $\frac{1}{2}(m_1 + m_2)l_1^2\omega_1^2 + \frac{1}{2}m_2l_2^2\omega^2 + m_2l_1l_2\omega_1\omega_2\cos(\theta_1 - \theta_2)$

$$V = m_1 g y_1 + m_2 g y_2$$

= $m_1 g l_2 + (m_1 + m_2) g l_1 (1 - \cos \theta_1) + m_2 g l_2 (1 - \cos \theta_2)$

This reveals a Lagrangian:

$$\mathcal{L} = T - V$$

$$= \frac{1}{2}(m_1 + m_2)l_1^2\omega_1^2 + \frac{1}{2}m_2l_2^2\omega^2 + m_2l_1l_2\omega_1\omega_2\cos(\theta_1 - \theta_2)$$

$$- m_1gl_2 - (m_1 + m_2)gl_1(1 - \cos\theta_1) - m_2gl_2(1 - \cos\theta_2).$$

We can solve for the equations of motion using the Euler-Lagrange equations:

$$\frac{d}{dt}\frac{\partial \mathcal{L}}{\partial \dot{q}} - \frac{\partial \mathcal{L}}{\partial q} = 0$$

For $q \in [\theta_1, \theta_2]$, $\dot{q} \in [\omega_1, \omega_2]$ we obtain the equations of motion:

$$\dot{\omega}_{1} = \frac{-m_{2}l_{2}[\dot{\omega}_{2}cos(\theta_{1} - \theta_{2}) + w_{2}^{2}sin(\theta_{1} - \theta_{2})] - (m_{1} + m_{2})gsin\theta_{1}}{(m_{1} + m_{2})l_{1}}$$

$$\dot{\omega}_{2} = -\frac{1}{l_{2}}(l_{1}[\dot{\omega}_{1}cos(\theta_{1} - \theta_{2}) - \omega_{1}^{2}sin(\theta_{1} - \theta_{2})] + gsin\theta_{1})$$

For rk4, the differential equations must be uncoupled so they can be rearranged through the a substitution that will produce the following results [3]. We use the following variables to represent the method rk4 in our code:

$$y(1) = \theta_1$$

$$f(1) = y(2) = \dot{\theta}_1 = \omega_1$$

$$f(2) = \dot{\omega}_1$$

$$= \frac{-g(2m_1 + m_2)sin\theta_1 - m_2gsin(\theta_1 - 2\theta_2)}{l_1(2m_1 + m_2 - m_2cos(2\theta_1 - 2\theta_2))}$$

$$\frac{-2sin(\theta_1 - \theta_2)m_2(\omega_2^2l_2 + \omega_1^2l_1cos(\theta_1 - \theta_2))}{l_1(2m_1 + m_2 - m_2cos(2\theta_1 - 2\theta_2))}$$

$$y(3) = \theta_2$$

$$f(3) = y(4) = \dot{\theta}_2 = \omega_2$$

$$f(4) = \dot{\omega_2}$$

$$= \frac{2sin(\theta_1 - \theta_2)(\omega_1^2 l_1(m_1 + m_2) + g(m_1 + m_2)cos\theta_1 + \omega_2^2 l_2 m_2 cos(\theta_1 - \theta_2))}{l_2(2m_1 + m_2 - m_2 cos(2\theta_1 - 2\theta_2))}$$

The total energy, E, is the sum of the kinetic energy, T, and potential energy, V, which were derived above. When using this model for the double pendulum, we assume the system is closed and that therefore the energy will be conserved. This means we only need to calculate the energy at the initial conditions to have the value for the energy of the system at all times. We can notice that for all four given initial conditions, $\omega_1 = 0$. This means that before calculating the energy directly in our Fortran90 code we can simplify the calculation. Plugging in this condition reveals the following expression for the energy:

$$E_i = \frac{1}{2} m_2 l_2^2 \omega_2^2 + g[m_1(l_2 + l_1[1 - \cos\theta_1]) + m_2 g(l_1[1 - \cos\theta_1] + l_2[1 - \cos\theta_2])]$$

Plugging the initial conditions into this expression for calculating and using it to calculate the energies in the Fortran90 code reveals the values:

$$E_1(0,0,0,0) = 3$$

 $E_2(0,0,\pi,0) = 5$
 $E_3(\pi,0,0,0) = 19$

$$E_4(\pi, 0, \pi, 0) = 21$$

These values can be shifted based on where the position y = 0 is set for the potential energy, these values come from setting the resting position of m_2 at (x,y) = (0,0), setting the pivot point for the first pendulum as the y = 0 would result in energies $E_i \in [-9, -7, 7, 9], i \in [1, 2, 3, 4]$.

By noticing that $\omega_2 = 0$ in all of these initial conditions as well, the E_i term will cancel out all of the terms to the right of the ω_2 term in the E_i equation given above. Therefore, we can calculate the value of ω_2 in the code with the following expression:

$$\omega_2 = \frac{2\Delta E}{m_2 l_1^2}$$

We can notice that this equation is independent of the energies, so for all given initial conditions the ω_2 such that $E = E_i + \Delta E$, $i \in [1, 2, 3, 4]$ will reveal the following value calculated by the Fortran90 code:

$$\omega_2 = 0.14142135465680267$$

The code is shown in the following section. The Makefile is shown in Listing 1 and it contains flags to optimize the calculation and shows that we can run the code by typing pen into the terminal after it has been compiled by typing make. Listing 2 contains the numtype file, which defines the precision of the values we will use as well as our value of pi. The file dubpen.f90 is found in Listing 3 and begins with the module setup which defines the parameters and the length of the arrays for y and f required by rk4. The program begins with defining the initial conditions and time parameters. The code here has the initial conditions for energy E_4 and shows that we are simulating for a time of 600 s with a step dt = 0.1. Then we calculate and print out the initial energy and ω_2 value required by the problem. The do-loop first plots the angles θ_1 and θ_2 as functions of time as the files fort.1 and fort.2. A section to plot energy as a function of time was placed in the do loop to see if rk4 would conserve total energy, which should happen in this model with this closed system. Then the Poincare section is plotted, which is a graph of the values of ω_2 vs. θ_2 at each time where $\theta_1 = 0$ and $\omega_1 > 0$. Since this code is a numerical approximation, the Poincare section of the code bounds the angles θ_1 and θ_2 to be in the range $[-\pi, \pi)$ and looks for when $\theta_1 \in (-0.1, 0.1)$. To calculate the values of θ_1 and θ_2 at each iteration

of the do-loop, the function rk4step is called on. The y and f arrays are defined as above. Fig. 2, 2, 3, and 4 shown in section Problem 1 (Figures) represent the results of the code for the initial conditions set for E_1 . All of the graphs are made with giving the initial energy that small boost in ω_2 calculated earlier. Since the system starts nearly at rest in E_1 , the behavior is similar to a regular pendulum, and the graphs appear to be periodic and the Poincare section is fairly symmetric. The energy as a function of time changes by approximately 1\%, so it is not conserved with rk4 but stays close enough to being conserved to represent a fairly accurate simulation. Fig. 5, 6 and 10 are for the initial conditions given for E_2 . This still involves no initial angle given for θ_1 , so the motion is still similar to a single pendulum and not chaotic, giving a periodic graph for θ_1 and another well-behaved Poincare section. However, this extra energy allows us to see some behavior that creates some chaos as we see θ_2 start to cycle around different values. Fig. 8, 9, 10 are for E_3 and Fig. 12, 13 and 14 are for E_4 . The graphs for these two initial conditions start with a significant θ_1 value so the behavior is not periodic and chaotic. The spatial graph for E_3 shown in Fig. 11 was made to visually represent the spatial path of the masses in the double pendulum and it acts as a check that the code is working properly.

3 Problem 1 (Code)

Listing 1: Makefile

```
objs1 = numtype.o dubpen.o

prog1 = pen

f90 = gfortran

f90flags = -03

libs = -framework Accelerate

ldflags = $(libs)

all: $(prog1)
```

```
$\(\prog1\): $\(\p
```

Listing 2: Module numtype

```
module numtype

module numtype

save

integer, parameter :: &

dp = selected_real_kind(15,307)

real(dp), parameter :: pi = 4*atan(1._dp)

end module numtype
```

Listing 3: The Program dubpen.f90.

```
module setup
3
       use numtype
       implicit none
       integer, parameter :: n_eq = 4
6
       real(dp), parameter :: g = 1._dp, 1_1 = 2._dp, &
           m_1 = 3._dp, 1_2 = 1._dp, m_2 = 1._dp
10
  end module setup
11
12
  program dubpen
13
14
       use setup
15
```

```
implicit none
16
       real(dp), dimension(n_eq) :: y
17
       real(dp) :: t, dt, tmax, w2, y1, y1, yu, &
18
           x_1, x_2, y_1, y_2, E, Et, eps, y_3
19
       real(dp), dimension(3) :: Ei
       real(dp), dimension(4) :: Ef
       t = 0._dp
23
       dt = 0.1_dp
24
       tmax = 600._dp
25
26
       ! Enter initial conditions
27
       y(1) = pi
                                     ! theta_1
       y(2) = 0._dp
                                     ! omega_1
       y(3) = pi
                                     ! theta_2
30
       y(4) = 0.14142135465680267 ! omega_2
31
32
       ! Energy terms to be summed
33
       Ei(1) = 1._dp / 2 * m_2 * 1_2**2 * y(4)**2
34
       Ei(2) = m_1 * g * (l_2 + l_1 &
           & * ( 1 - cos( y(1) ) )
       Ei(3) = m_2 * g * (l_1 * (l_1 - cos(y(1)))) &
37
           & + 1_2 * (1 - cos(y(3)))
38
39
       ! Energy for intial conditions
40
       E = sum(Ei)
41
       ! w2 such that E = Ei + 0.01
       w2 = sqrt(2 * (0.01) / (m_2 * 1_2 * * 2))
45
       print *, "E_{\square} = ", E, "w2_{\square} = ", w2
46
47
       ! For plotting Poincare section
48
       eps = 0.2
49
       do while ( t < tmax )</pre>
53
           ! Plot angles vs. time
54
           write(1,*) t, y(1)
55
```

```
write(2,*) t, y(3)
56
57
           ! Plot positions y vs. x
           x_1 = 1_1 * sin(y(1))
           y_1 = 1_2 + 1_1 * (1 - cos(y(1)))
           x_2 = x_1 + 1_2 * sin(y(3))
61
           y_2 = y_1 - 1_2 * cos(y(3))
62
           write(3,*) x_1, y_1
63
           write(4,*) x_2, y_2
64
65
           ! Plot energy vs. time
66
           Ef(1) = 1._dp / 2 * (m_1 + m_2) &
                & * 1_1 **2 * y(2) **2
68
           Ef(2) = 1._dp / 2 * m_2 * 1_2**2 * y(4)**2
69
           Ef(3) = m_2 * l_1 * l_2 * y(2) &
70
                & * y(4) * cos(y(1) - y(3))
71
           Ef(4) = m_1 * g * (y_1 + y_2)
72
           Et = sum(Ef)
73
           write (7,*) t, Et
74
           ! Plot Poincare Section
76
           y1 = y(1)
77
           y3 = y(3)
78
           if ( y1 >= pi ) then
79
                do\ while ( y1 >= pi )
80
                    y1 = y1 - 2 * pi
                end do
           elseif ( y1 < - pi ) then
83
                do\ while ( y1 < - pi )
84
                    y1 = y1 + 2 * pi
85
                end do
86
           end if
87
           if ( y3 >= pi ) then
                do\ while ( y3 >= pi )
                    y3 = y3 - 2 * pi
                end do
           elseif ( y3 < - pi ) then
92
                do\ while ( y3 < - pi )
93
                    y3 = y3 + 2 * pi
94
                end do
95
```

```
end if
96
            yl = y1 - eps
97
            yu = y1 + eps
98
            if ( yl < 0 . and. yu > 0 &
                 & and y(2) > 0 ) then
100
                      write(8,*) y3, y(4)
101
                      write(9,*) y(3), y(4)
102
                      ! PS with unbounded theta_2
103
            end if
104
105
             call rk4step( t, dt, y )
106
107
        end do
108
109
   end program dubpen
110
111
   subroutine rk4step(x, h, y)
112
113
        use setup
114
        implicit none
115
        real(dp), intent(inout) :: x
        real(dp), intent(in) :: h
117
        real(dp), intent(inout), dimension(n_eq) :: y
118
        real(dp), dimension(n_eq) :: k1, k2, k3, k4, dy
119
120
        k1 = kv (x, h, y)
121
        k2 = kv (x+h/2, h, y+k1/2)
122
        k3 = kv (x+h/2, h, y+k2/2)
        k4 = kv (x+h, h, y+k3)
124
125
        dy = (k1 + 2*k2 + 2*k3 + k4) / 6
126
127
        x = x + h
128
        y = y + dy
129
        contains
131
132
            function kv (t, dt, y) result(k)
133
134
                 use setup
135
```

```
implicit none
136
                real(dp), intent(in) :: t
137
                real(dp), intent(in) :: dt
138
                real(dp), intent(in), dimension(n_eq) :: y
139
                real(dp), dimension(n_eq) :: f, k
                real(dp), dimension(4) :: n_1
141
                real(dp), dimension(3) :: n_2
142
                real(dp) :: coeff, d
143
144
                f(1) = y(2)
145
                f(3) = y(4)
146
                ! Numerator terms for f(2)
148
                n_1(1) = -g * (2 * m_1 + m_2) &
149
                    & * sin(y(1))
150
                n_1(2) = -m_2 * g * sin (y(1) - 2 * y(3))
151
                coeff = 2 * sin ( y(1) - y(3) )
152
                n_1(3) = - coeff * m_2 * y(4)**2 * 1_2
153
                n_1(4) = - coeff * m_2 * y(2)**2 &
154
                    & * 1_1 * cos (y(1) - y(3))
156
                ! Numerator terms for f(4)
157
                n_2(1) = y(2)**2 * 1_1 * (m_1 + m_2)
158
                n_2(2) = g * (m_1 + m_2) * cos (y(1))
159
                n_2(3) = y(4)**2 * 1_2 * m_2 &
160
                    & * cos(y(1) - y(3))
161
                ! Denominator term
163
                d = 2 * m_1 + m_2 - m_2 &
164
                    & * cos(2 * (y(1) - y(3)))
165
166
                f(2) = sum(n_1) / (l_1 * d)
167
                f(4) = 2 * sin(y(1) - y(3)) &
168
                    & * sum(n_2) / (1_2 * d)
169
                k(1:n_eq) = h*f(1:n_eq)
171
172
173
           end function kv
174
   end subroutine rk4step
```

4 Problem 1 (Figures)

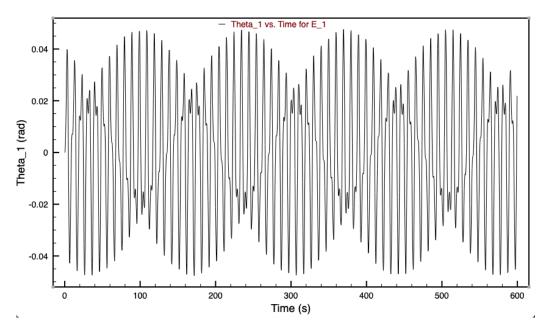


Figure 1: $\theta_1(t)$ vs. t for E_1 with \mathtt{dt} = 0.1 and \mathtt{tmax} = 600.

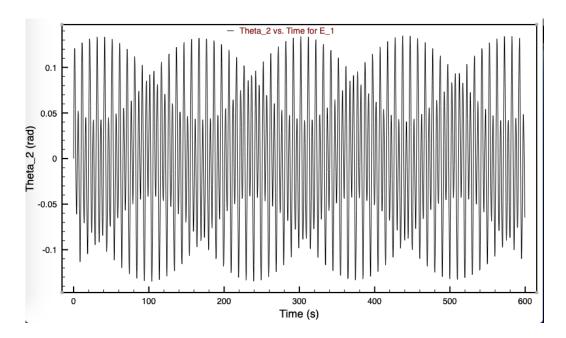


Figure 2: $\theta_2(t)$ vs. t for E_1 with \mathtt{dt} = 0.1 and \mathtt{tmax} = 600.

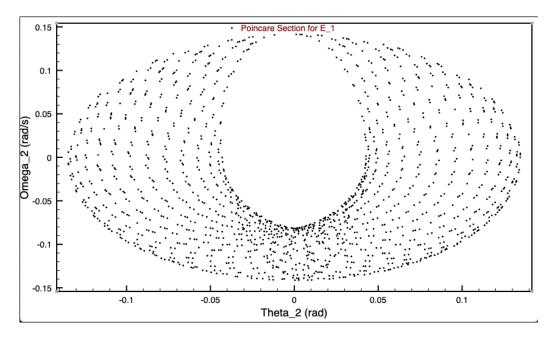


Figure 3: Poincare Section for E_1 with $\mathtt{dt} = 0.1$ and $\mathtt{tmax} = 600$.

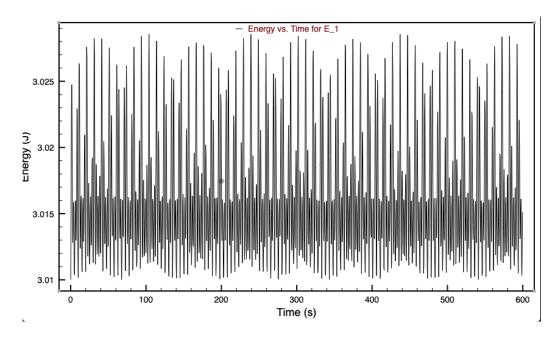


Figure 4: Energy vs. t for E_1 with $\mathtt{dt} = 0.1$ and $\mathtt{tmax} = 600$.

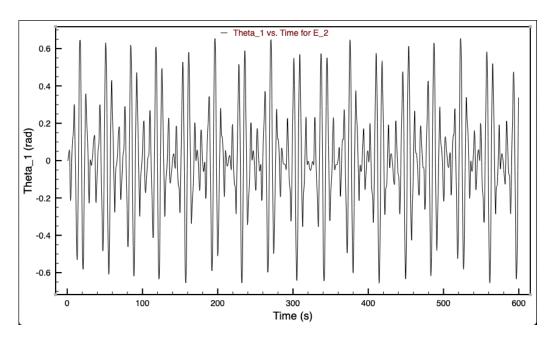


Figure 5: $\theta_1(t)$ vs. t for E_2 with dt = 0.1 and tmax = 600.

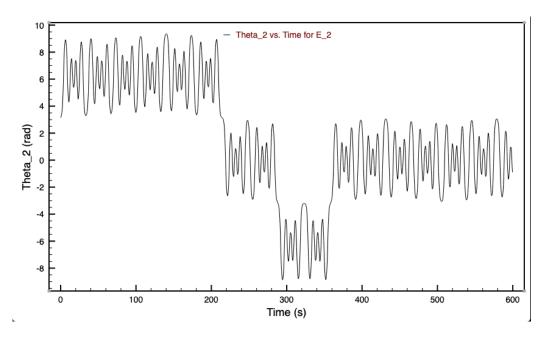


Figure 6: $\theta_2(t)$ vs. t for E_2 with \mathtt{dt} = 0.1 and \mathtt{tmax} = 600.

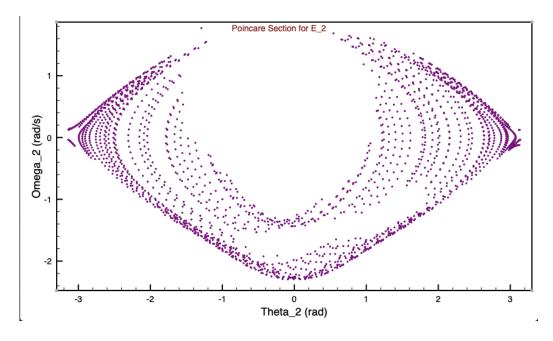


Figure 7: Poincare Section for E_2 with $\mathtt{dt} = 0.1$ and $\mathtt{tmax} = 600$.

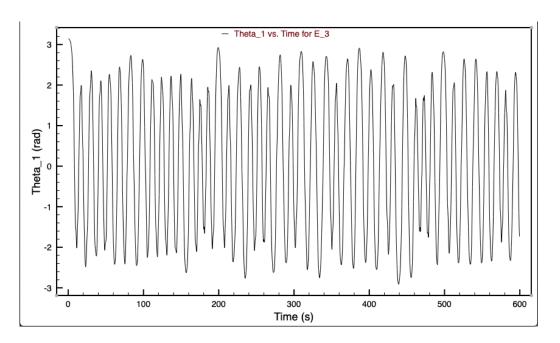


Figure 8: $\theta_1(t)$ vs. t for E_3 with dt = 0.1 and tmax = 600.

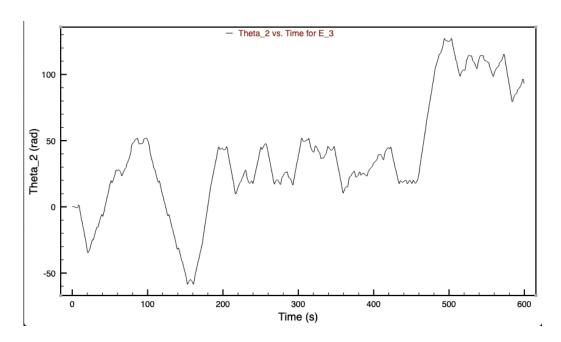


Figure 9: $\theta_2(t)$ vs. t for E_3 with dt = 0.1 and tmax = 600.

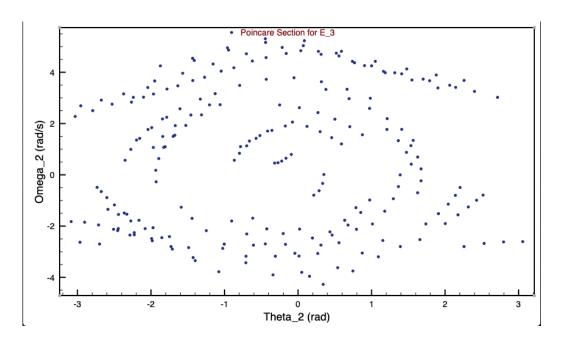


Figure 10: Poincare Section for E_3 with $\mathtt{dt} = \mathtt{0.1}$ and $\mathtt{tmax} = \mathtt{600}$.

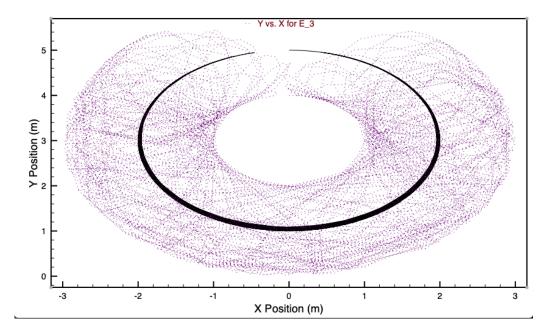


Figure 11: Spatial path for E_3 with $\mathtt{dt} = 0.1$ and $\mathtt{tmax} = 600$.

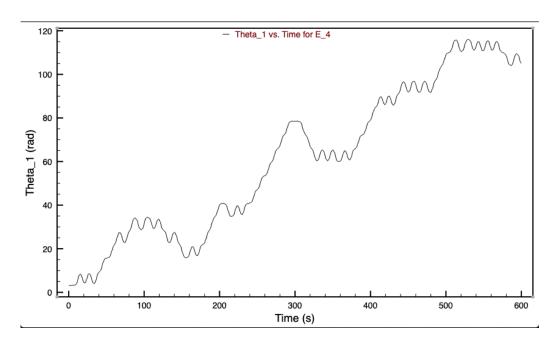


Figure 12: $\theta_1(t)$ vs. t for E_4 with dt = 0.1 and tmax = 600.

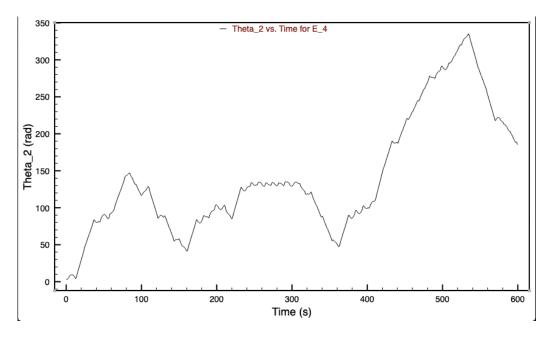


Figure 13: $\theta_2(t)$ vs. t for E_4 with dt = 0.1 and tmax = 600.

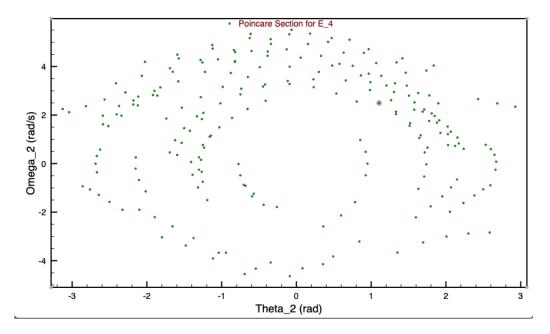


Figure 14: Poincare Section for E_4 with dt = 0.1 and tmax = 600.

5 Problem 2

For problem 2 we are given a neutron star that is vibrating and are tasked to show that the z-component of its spin, $\Omega(t)$, is nearly constant while the vector $\Omega(t)$ nutates about the z-axis. We are given the oscillating moments of inertia:

$$I_{zz} = I_0(1 + \epsilon cos(\omega t))$$

$$I_{xx} = I_{yy} = I_0(1 - \frac{\epsilon}{2}cos(\omega t)), \epsilon \ll 1, I_0 = \frac{2}{5}mr^2.$$

We can take advantage of the Euler equations for the body-fixed system [4]:

$$N = \dot{L} + \Omega \times L$$

N is equal to zero for this system since there are no external forces. This yields the equations of motion (for $g = \epsilon cos(\omega t)$):

$$\dot{L}_x = \frac{-3L_y L_z g}{2I_0(1+g)(1-\frac{g}{2})}$$

$$\dot{L}_y = \frac{3L_x L_z g}{2I_0(1+g)(1-\frac{g}{2})}$$

$$\dot{L}_z = 0$$

The code is found in Listing 5. The numtype module is identical to the one used before in Listing 2. The Makefile is very similar, with the only differences being contained in Listing 4. This gives instructions for the compiler when typing make into the terminal. The code can then be executed by typing star. The program starvibe begins with the module setup which inputs the parameters for the problem. For our calculation of I_0 , the value $\epsilon = 10^{-8}$ was chosen. The time parameters can be inputted, in this case the graphs are plotted for a time of 20 s with dt=0.01. The elements of the vector y(1:3) exist so that the length of the arrays y and f are equal. The equations of motion above represent a coupled first-order differential equation for the angular momentum L. y(4:6) are the x, y, and z-components respectively of the angular momentum vector L. The elements of the array f(1:3) = y(4:6) and its derivatives are f(4:6). rk4 was used instead of rkf45 because the adaptive-step size kept going to 0. Since we were given a magnitude for $\Omega(t)$, it was evenly split between its three components. Since the y array is in terms of the angular momentum, O is used to divide by the moments of inertia to get the components of the angular velocity, which are to be graphed. The files fort.1, fort.2 and fort.3 are the files written for the x, y and z-components of the angular momentum respectively vs. time. The three are plotted on the same graph in Fig. 15. This graph shows that the z-component of Ω remains nearly constant, and the sinusoidal nature of the plots for Ω_x and Ω_y show that $\Omega(t)$ nutates about the z-axis with nutation frequency approximately 0.3 Hz.

Listing 4: Module numtype

```
objs1 = numtype.o starvibe.o rkf45step.o

prog1 = star
```

Listing 5: Module numtype

```
module setup

module setup
```

```
use numtype
4
       implicit none
5
       integer, parameter :: n_eq = 6
6
       real(dp), parameter :: m = 40.e30_dp, &
           r = 10.e3_dp, nu = 40.e3_dp, eps = 10.e-8_dp
10
  end module setup
11
12
  program starvibe
13
14
       use setup
15
       implicit none
       real(dp), dimension(n_eq) :: y
17
       real(dp), dimension(3) :: 0
18
       real(dp) :: t, dt, tmax, omega, w, IO, &
19
           Ix, Iy, Iz
20
       integer :: i
21
22
       t = 0._dp
       dt = 0.01_dp
       tmax = 20._dp
25
26
       w = 2 * pi * nu
27
       omega = 100 * w
28
       I0 = 2._dp / 5 * m * r**2
29
       y(1) = 0._dp
       y(2) = 0._dp
32
       y(3) = 0._dp
33
       y(4) = I0 * omega / sqrt(3._dp) ! L_x
34
       y(5) = I0 * omega / sqrt(3._dp) ! L_y
35
       y(6) = I0 * omega / sqrt(3._dp) ! L_z
36
       do while ( t < tmax )</pre>
           Ix = I0 * (1 - eps / 2 * cos(w * t))
40
           Iy = Ix
41
           Iz = I0 * (1 + eps * cos(w * t))
42
43
```

```
0(1) = y(4) / Ix
44
           0(2) = y(5) / Iy
45
           0(3) = y(6) / Iz
46
47
           do i = 1 , 3
                write(i,*) t, O(i)
           end do
50
51
            call rk4step( t, dt, y )
52
53
       end do
54
   end program starvibe
57
   subroutine rk4step(x, h, y)
58
59
       use setup
60
       implicit none
61
       real(dp), intent(inout) :: x
62
       real(dp), intent(in) :: h
       real(dp), intent(inout), dimension(n_eq) :: y
       real(dp), dimension(n_eq) :: k1, k2, k3, k4, dy
66
       k1 = kv (x, h, y)
67
       k2 = kv (x+h/2, h, y+k1/2)
       k3 = kv (x+h/2, h, y+k2/2)
       k4 = kv (x+h, h, y+k3)
       dy = (k1 + 2*k2 + 2*k3 + k4) / 6
73
       x = x + h
74
       y = y + dy
75
76
       contains
77
           function kv (t, dt, y) result(k)
80
                use setup
81
                implicit none
82
                real(dp), intent(in) :: t
83
```

```
real(dp), intent(in) :: dt
84
                real(dp), intent(in), dimension(n_eq) :: y
85
                real(dp), dimension(n_eq) :: f, k
86
                real(dp) :: I0, w, g
87
                I0 = 2._dp / 5 * m * r**2
89
                w = 2 * pi * nu
90
                g = eps * cos (w * t)
91
92
                f(1:3) = y(4:6)
93
94
                f(4) = -3 / (2 * I0) * y(5) * y(6) &
                    & * g / ( ( 1 + g ) * ( 1 - g / 2 ) )
97
                f(5) = 3 / (2 * I0) * y(4) * y(6) &
98
                    & * g / ( ( 1 + g ) * ( 1 - g / 2 ) )
99
100
                f(6) = 0
101
102
                k(1:n_eq) = h*f(1:n_eq)
103
104
           end function kv
105
106
   end subroutine rk4step
107
```

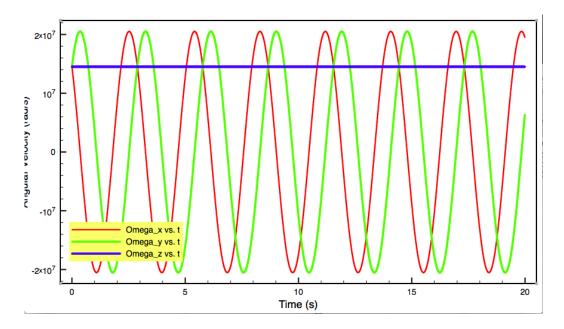


Figure 15: Components of angular velocity with respect to time for dt = 0.01 and tmax = 20.

6 Summary and conclusions

We presented our Fortran90 codes for the solutions of Problem 1 and 2 from Midterm 1. We can see that the motion for a double pendulum is approximately harmonic for $\theta_1 \approx 0$, but as θ_1 increases the behavior becomes more chaotic. We also observed that a neutron star with the given vibrating moments of inertia will maintain a nearly-constant Ω_z while nutating about the z-axis. Ultimately, we have been shown the incredible power of the Runge-Kutta method, specifically RK4, for allowing us to study interesting physics with simulations of nonlinear phenomena that could otherwise not be solved for analytically.

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