

Dimensional analysis in relativity and in differential geometry

P.G.L. Porta Mana 

Kavli Institute, Trondheim, Norway <pgl@portamana.org>

13 December 2019; updated 11 November 2020

This note provides a short guide to dimensional analysis in Lorentzian and general relativity and in differential geometry. It tries to revive Dorgelo and Schouten's notion of 'intrinsic' or 'absolute' dimension of a tensorial quantity. The intrinsic dimension is independent of the dimensions of the coordinates and expresses the physical and operational meaning of a tensor. The dimensional analysis of several important tensors and tensor operations is summarized. In particular it is shown that the components of a tensor need not have all the same dimension, and that the Riemann (once contravariant and thrice covariant), Ricci (twice covariant), and Einstein (twice covariant) curvature tensors are dimensionless. The relation between dimension and operational meaning for the metric and stress-energy-momentum tensors, and the dimension of Einstein's constant are also discussed.

per la piccola Emma

1 Introduction

From the point of view of dimensional analysis, do all components of a tensor need to have the same dimension? What happens to these components if we choose coordinates that don't all have dimensions of length or time? And if the components of a tensor have different dimensions, then does it make sense to speak of "the dimension of the tensor"? What are the dimensions of the metric and of the curvature tensors? What is the dimension of the constant in the Einstein equations?

A sense of insecurity gets hold of many students (and possibly of some researchers) in relativity, when they have to discuss and answer this kind of questions. This is evident in many question & answer websites and wiki pages, where several incorrect or unfounded statements about dimensional analysis in relativity are in circulation. (For the record, the answers to the questions above are: No – See eq. (12) below – Yes – They are dimensionless – $M^{-1}L^{-1}T^2$, or $M^{-1}L$, or $M^{-1}L^3T^{-2}$, depending on

the dimensions you assign to the metric and stress-energy-momentum tensors, see §§ 9–11.)

Several factors contribute to these misconceptions and insecurity. Modern texts in Lorentzian and general relativity commonly use geometrized units. They say that, for finding the dimension of some constant in a tensorial equation, it's sufficient to compare the dimensions of the terms in the equation. But the application of this procedure is sometimes not so immediate, because some tensors don't have universally agreed dimensions – prime example the metric tensor. Many texts use four coordinates with dimension of length, and base their dimensional analyses on that specific choice¹, multiplying timelike tensorial components by appropriate powers of c . Such common practices can therefore give students the impression that coordinates ought to always be lengths, and that all components of a tensor ought to have the same dimension. Yet, students cannot find such rules explicitly stated anywhere. We'll see shortly that no such rules in fact exist, nor are they necessary.

Dimensional analysis is thus not very self-evident in relativity and in differential geometry. The present note wants to provide a short but exhaustive guide to it. Some important dimensional-analysis questions in general relativity are also consistently settled here; for example the dimension of the Riemann curvature tensor, or the effect of the covariant or Lie derivatives on dimensions.

The application of dimensional analysis in relativity is most straightforward and self-evident if we rely on the coordinate-free or intrinsic approach to differential geometry, briefly recalled below, and if we adopt the perhaps overlooked notion of *intrinsic dimension* of a tensor. The intrinsic dimension of a tensor was introduced under the name 'absolute dimension' by Schouten and Dorgelo² and used in Truesdell & Toupin³. As its name implies, this dimension is independent of the choice and dimensions of coordinate functions. It is distinct from the dimensions of the tensor's *components*, which instead depend on the dimensions of the coordinates. The intrinsic dimension of a tensor is determined by

¹ e.g. Tolman 1949 p. 71 eq. (37.1); Landau & Lifshitz 1996 p. 80 eq. (32.15); for exceptions with coordinates with mixed dimensions see e.g. Fock 1964 § V.55; McVittie 1965 § 4.1; Fokker 1965 § VII.1; Truesdell & Toupin 1960 § F.III.280; Kitano 2013 § X.

² Dorgelo & Schouten 1946; Schouten 1989 ch. VI. ³ Truesdell & Toupin 1960 Appendix II.

the latter's physical and operational⁴ meaning. It is therefore a natural notion for dimensional analysis in relativity.

Here is a synopsis of the rest of this note. A brief reminder of the intrinsic approach to differential geometry, with references, is given in the next section, together with some special notation necessary to our discussion. Section 3 gives a simple example of dimensional analysis in a two-dimensional spacetime. This example might be enough for most readers to get the basic way of reasoning; such readers can work out the rest for themselves whenever they need and don't need to read the rest of this note. Sections 4–7 offer a more systematic discussion and a synopsis of dimensional analysis for the main tensorial operations. The notion of intrinsic dimension is explained in § 5. The intrinsic dimensions of various curvature tensors, of the metric tensor, and of the stress-energy-momentum tensor are discussed in §§ 8–10. The contravariant and thrice covariant Riemann, fully covariant Ricci, and fully covariant Einstein curvature tensors, in particular, are found to have intrinsic dimension 1, that is, to be dimensionless. The operational motivation of several standard choices for the dimensions of the metric and the stress-energy-momentum tensors are also discussed. The dimension of the constant in the Einstein equations is finally derived in § 11.

This note obviously assumes familiarity with basic tensor calculus and related notions, for example of co- and contra-variance, tensor product, contraction. Some passages assume familiarity with the exterior calculus of differential forms. The general ideas, however, should be understandable even without such familiarity.

Finally, quoting Truesdell & Toupin⁵, “dimensional analysis remains a controversial and somewhat obscure subject. We do not attempt a complete presentation here”. References about recent developments in this subject are given in the summary of § 12.

2 Intrinsic view of differential-geometric objects: brief reminder and notation

From the intrinsic point of view, a tensor is defined by its geometric properties. For example, a vector field v is an object that operates on functions defined on the (spacetime) manifold, yielding new functions,

⁴ Bridgman 1958.

⁵ Truesdell & Toupin 1960 Appendix § 7 footnote 4.

with the properties $v(af + bg) = av(f) + bv(g)$ and $v(fg) = v(f)g + f v(g)$ for all functions f, g and reals a, b . A covector field (also called 1-form) ω is an object that operates on vector fields, yielding functions, with the property $\omega(fu + gv) = f\omega(u) + g\omega(v)$ for all vector fields u, v and functions f, g . The sum of vector or covector fields and their products by functions are defined in an obvious way. Tensors are constructed from these objects; see also the end of this section for a slightly different point of view.

A system of coordinates (x^i) is just a set of linearly independent functions. This set gives rise to a set of vector fields $(\frac{\partial}{\partial x^i})$ and to a set of covector fields (dx^i) by the obvious requirements that $\frac{\partial}{\partial x^i}(x^j) = \delta_i^j$ and $dx^i(\frac{\partial}{\partial x^j}) = \delta_j^i$. These two sets can be used as bases to express all other vectors and covectors as linear combinations. A vector field v can thus be written as

$$v \equiv \sum_i v^i \frac{\partial}{\partial x^i} \equiv v^i \frac{\partial}{\partial x^i}, \quad (1)$$

where the functions $v^i := dx^i(v)$ are its components with respect to the basis $(\frac{\partial}{\partial x^i})$. Analogously for a covector field.

For the presentation of the intrinsic view I recommend the excellent texts by Choquet-Bruhat et al. (1996), Boothby (2003), Abraham et al. (1988), Burke (1987), Bossavit (1991), and more on the general-relativity side Misner et al. (1973 ch. 9),ourgoulhon (2012 ch. 2), Penrose & Rindler (2003 ch. 4).

For the notation in dimensional analysis I use ISO conventions:⁶ $\dim(\mathbf{A})$ is the dimension of the quantity \mathbf{A} , and among the base quantities are mass M , length L , time T , temperature Θ , electric current I , and the dimensionless 1. Note that I don't discuss units – it doesn't matter here whether the unit for length is the metre or the foot, for example.

Throughout this note c denotes the speed of light, with $\dim(c) = LT^{-1}$. Its numerical quantity value $\{c\}$ depends on the chosen units of length and time.

The number, ordering, and symmetries of a tensor's covariant and contravariant “slots”⁷ will be important in our discussion. The traditional coordinate-free notation ‘ \mathbf{A} ’ omits this information. We thus need a coordinate-free notation that makes it explicit when needed. Penrose &

⁶ ISO 2009 § 5.

⁷ Misner et al. 1973 § 3.2.

Rindler⁸ propose an abstract-index notation where ‘ A_i^{jk} ’, for example, denotes a tensor covariant in its first slot and contravariant in its second and third slots. Every index in this notation is “a *label* whose sole purpose is to keep track of the type of tensor under discussion”⁹. So this notation doesn’t stand for the set of *components* of the tensor. For the latter set, Penrose & Rindler use **bold** indices instead: ‘ A_i^{jk} ’. But in our discussion the difference between a tensor and its set of components is crucial, and Penrose & Rindler’s abstract-index notation unfortunately lends itself to conceptual and typographic misunderstanding.

I shall therefore use a notation such as $\mathbf{A}_{\bullet}^{\bullet}$ to indicate that \mathbf{A} is covariant in its first slot and contravariant in its second and third slots. Its components would thus be (A_i^{jk}) . For brevity I’ll call this a ‘co-contravariant’ tensor, with an obvious naming generalization for other tensor types. A set of completely antisymmetric slots will be put within bars: thus the notation $\mathbf{A}_{\bullet\ldots\bullet}^{\bullet}$ means that \mathbf{A} is completely antisymmetric in its last two covariant slots. Finally, in accord with convenient modern terminology, a completely antisymmetric contravariant tensor of rank k will be called ‘ k -vector’; and a completely antisymmetric covariant tensor of rank k , a ‘ k -covector’. The terms ‘multi-vector’ and ‘multi-covector’ are used when k isn’t specified.

The only weak points of the notation just explained are the operations of transposition and contraction, which literal indices depict so well instead. Considering that transposition is a generalization of matrix transposition, and contraction a generalization of trace, I’ll use the following notation:

- $\mathbf{A}^{\top\alpha\beta}$ is the transposition (swapping) of the α th and β th slots. Its coordinate-free definition is

$$(\mathbf{A}^{\top\alpha\beta})(\dots, \underset{\beta \text{th slot}}{\zeta}, \dots, \underset{\alpha \text{th slot}}{\eta}, \dots) := \mathbf{A}(\dots, \underset{\beta \text{th slot}}{\eta}, \dots, \underset{\alpha \text{th slot}}{\zeta}, \dots) \quad (2)$$

for all ζ, η of appropriate variance type.

- $\text{tr}_{\alpha\beta} \mathbf{A}$ is the contraction of the α th and β th slots, which must have opposite variant types; note that we may have $\beta < \alpha$. Its coordinate-free

⁸ Penrose & Rindler 2003 § 2.2.

⁹ Penrose & Rindler 2003 p. 75.

definition is

$$(\text{tr}_{\alpha\beta} \mathbf{A})(\dots, \dots, \dots) := \sum_i \mathbf{A}(\dots, \overset{\alpha \text{th slot}}{u_i}, \dots, \underset{\beta \text{th slot}}{\omega^i}, \dots) \quad (3)$$

for any arbitrary complete and linearly independent sets $\{u_i\}, \{\omega^j\}$ such that $\omega^j(u_i) = \delta^j_i$.

In index notation the two operations above are the familiar

$$\overset{\alpha \text{th slot}}{A}_{\dots}^i \dots \underset{\beta \text{th slot}}{j} \dots \mapsto \underset{\alpha \text{th slot}}{A}_{\dots}^i \dots \overset{\beta \text{th slot}}{j} \dots \quad \text{and} \quad \overset{\alpha \text{th slot}}{A}_{\dots}^i \dots \underset{\beta \text{th slot}}{j} \dots$$

For the sake of notation economy I'll denote the contraction of the closest slots of two tensor-multiplied tensors by simple juxtaposition. For example, if $\mathbf{A} \equiv \mathbf{A}_{\bullet}^{\bullet}$, $\mathbf{B} \equiv \mathbf{B}_{\bullet\bullet}$, and v is a vector, then

$$\mathbf{AB} := \text{tr}_{23}(\mathbf{A}_{\bullet}^{\bullet} \otimes \mathbf{B}_{\bullet\bullet}), \quad \mathbf{B}v := \text{tr}_{23}(\mathbf{B}_{\bullet\bullet} \otimes v^{\bullet}). \quad (4)$$

This notation makes sense considering tensors as linear operators.

Contraction and transposition will be discussed only sparsely, so I hope you won't find the notation above too uncomfortable.

It is possible to build the tensor-product architecture not on vectors and covectors, but on multi-vectors and multi-covectors, with their straight and twisted (also called 'even' and 'odd', or 'polar' and 'axial') orientations. This elegant and powerful geometric point of view leads to deeper physical insight and is gaining popularity in the literature. For its presentation I recommend the texts of Bossavit (1991 especially ch. 3), Burke (1983; 1987; 1980; 1995), de Rham (1984 ch. 2), Schouten (1954), Deschamps (1970; 1981) and Lindell (2004).

In the notation above, the bars identify k -vectors and k -covectors for $k > 1$. Thus $\mathbf{A}_{\bullet\bullet}^{\bullet}$ indicates that \mathbf{A} belongs to the tensor product of 1-vectors and 2-covectors; it's also called a 1-vector-valued 2-covector. To avoid burdening the notation I won't add symbols denoting straight or twisted orientation, but I'll explicitly say in the text when any object has a twisted orientation.

3 An introductory two-dimensional example

Let me first present a simple example of dimensional analysis in a two-dimensional spacetime. I provide very little explanation, letting the

analysis speak for itself. The next sections will give a longer discussion of the general point of view, of the assumptions, and of cases with more elaborate geometric objects.

In a region of a two-dimensional spacetime we use coordinates (x, y) . These coordinates allow us to uniquely label every event in the region (otherwise they wouldn't be coordinates). Let us say that coordinate x has dimension of temperature, and y of specific entropy:

$$\dim(x) = \Theta, \quad \dim(y) = s := L^2 T^{-2} \Theta^{-1}. \quad (5)$$

This choice may be possible for several reasons. For example, the region could be occupied by a heat-conducting material; in a specific spacetime foliation, its temperature increases along each 1-dimensional spacelike slice, and its entropy density is uniform on each slice but increases from slice to slice.¹⁰ Owing to this kind of monotonic behaviour for these quantities, if we are given a pair of temperature & specific-entropy values we can identify a unique event associated to them in this spacetime region. They can thus be used as a coordinate system. The point here is that coordinates can have any dimensions owing to physical reasons. In atmospheric and ocean dynamics, for example, pressure or mass density are sometimes used as coordinates for depth¹¹.

From these coordinates we construct two covector fields (dx, dy) , and two vector fields $(\frac{\partial}{\partial x}, \frac{\partial}{\partial y})$ that serve as bases for the spaces of tangent covectors, vectors, and tensors. Their dimensions are

$$\begin{aligned} \dim(dx) &= \Theta & \dim(dy) &= s, \\ \dim\left(\frac{\partial}{\partial x}\right) &= \Theta^{-1} & \dim\left(\frac{\partial}{\partial y}\right) &= s^{-1}. \end{aligned} \quad (6)$$

Consider a contra-co-variant tensor field $\mathbf{A} \equiv \mathbf{A}^\bullet$, in this region. Using the basis fields above it can be written as

$$\mathbf{A} = A^x_x \frac{\partial}{\partial x} \otimes dx + A^x_y \frac{\partial}{\partial x} \otimes dy + A^y_x \frac{\partial}{\partial y} \otimes dx + A^y_y \frac{\partial}{\partial y} \otimes dy, \quad (7)$$

where $A^x_x := \mathbf{A}(dx, \frac{\partial}{\partial x})$, $A^x_y := \mathbf{A}(dx, \frac{\partial}{\partial y})$, and so on, are the components of the tensor in the coordinate system (x, y) .

¹⁰ For general-relativistic thermomechanics see e.g. Eckart 1940; Maugin 1974; 1978a,b,c,d; Muschik & von Borzeszkowski 2014. ¹¹ Griffies 2004 ch. 6; Vallis 2006 § 2.6.2.

By the rules of dimensional analysis, the two sides of the expansion above must have the same dimension. The same holds for the four summands on the right side. Denoting $A := \dim(\mathbf{A})$, we thus find the four equations

$$\begin{aligned} A &= \dim(A^x_x) & A &= \dim(A^x_y) \Theta^{-1} s \\ A &= \dim(A^y_x) \Theta s^{-1} & A &= \dim(A^y_y) , \end{aligned}$$

or

$$\begin{aligned} \dim(A^x_x) &= A & \dim(A^x_y) &= A \Theta s^{-1} \equiv A L^{-2} T^2 \Theta^2 \\ \dim(A^y_x) &= A \Theta^{-1} s & \dim(A^y_y) &= A . \end{aligned} \tag{8}$$

The intrinsic dimension of the tensor \mathbf{A} is A . The expansion (7) shows that this dimension is independent of the coordinate system, by construction – such expansion could be done in any other coordinate system, and the left side would be the same. The effect of coordinate transformations is examined more in detail in § 5. The intrinsic dimension A is determined by the physical and operational meaning of the tensor; see §§ 9, 10 for concrete examples. Together with the dimensions of the coordinates it determines the dimensions of the components, eqs (8), which need not be all equal.

This simple example should have disclosed the main points of dimensional analysis on manifolds, which will now be discussed in more generality. In the derivation above we silently adopted a couple of natural conventions: for example, that the tensor product behaves similarly to multiplication with regard to dimensions. Such conventions are briefly discussed in § 12.

4 Coordinates

From a physical point of view a coordinate is just a function that associates values of some physical quantity with the events in a region (the domain of the coordinate chart) of spacetime. Together with the other coordinates such function allows us to uniquely identify every event within that region. Any physical quantity will do: the distance from something, the time elapsed since something, an angle, an energy density, the strength of a magnetic flux, a temperature, and so on. A coordinate can thus have

any dimension: length L , time T , angle 1 , temperature Θ , magnetic flux $\Phi := ML^2T^{-2}I^{-1}$, and so on.

The functional relation between two sets of coordinates must of course be dimensionally consistent. For example, if $\dim(x^0) = T$ and $\dim(x^1) = L$, and we introduce a new coordinate $\xi(x^0, x^1)$ with dimension Φ , additive in the previous two, then we must have $\xi = ax^0 + bx^1$ with $\dim(a) = \Phi T^{-1}$ and $\dim(b) = \Phi L^{-1}$.

5 Tensors: intrinsic dimension and components' dimensions

Consider a system of coordinates (x^i) with dimensions (X_i) , and the ensuing sets of covector fields (1-forms) dx^i and of vector fields $(\frac{\partial}{\partial x^i})$, bases for the cotangent and tangent spaces. Their tensor products are bases for the tangent spaces of higher tensor types.

The differential dx^i traditionally has the same dimension as x^i : $\dim(dx^i) = X_i$, and the vector $\frac{\partial}{\partial x^i}$ traditionally has the inverse dimension: $\dim(\frac{\partial}{\partial x^i}) = X_i^{-1}$.

For our discussion let's take a concrete example: a contra-co-variant tensor field $\mathbf{A} \equiv \mathbf{A}^\bullet$. The discussion generalizes to tensors of other types in an obvious way.

The tensor \mathbf{A} can be expanded in terms of the basis vectors and covectors, as in § 2 and in the example of § 3:

$$\mathbf{A} = A^i_j \frac{\partial}{\partial x^i} \otimes dx^j \equiv A^0_0 \frac{\partial}{\partial x^0} \otimes dx^0 + A^0_1 \frac{\partial}{\partial x^0} \otimes dx^1 + \dots \quad (9)$$

Each function

$$A^i_j := \mathbf{A}\left(dx^i, \frac{\partial}{\partial x^j}\right) \quad (10)$$

is a component of the tensor in this coordinate system.

To make dimensional sense, all terms in the sum (9) must have the same dimension. This is possible only if the generic component A^i_j has dimension

$$\dim(A^i_j) = A X_i X_j^{-1}, \quad (11)$$

where A is common to all components. In fact, the $X_i X_j^{-1}$ term cancels the $X_i^{-1} X_j$ term coming from $\frac{\partial}{\partial x^i} \otimes dx^j$ in the sum (9), and each summand therefore has dimension A .

The generalization of the formula above to tensors of other types is obvious:

$$\dim(A^{ij\dots}_{kl\dots}) = A X_i X_j \cdots X_k^{-1} X_l^{-1} \cdots \quad (12)$$

where the ordering of the indices doesn't matter.

Clearly the components can have different dimensions¹². But this isn't an issue. What matters is that the sum (9) be dimensionally consistent.

The dimension A , which is also the dimension of the sum (9), I'll call the *intrinsic dimension* of the tensor \mathbf{A} , and we write

$$\dim(\mathbf{A}) = A. \quad (13)$$

This dimension is independent of any coordinate system. It reflects the physical or operational¹³ meaning of the tensor. We shall see an example of such an operational analysis in §§ 9 and 10 for the metric and stress-energy-momentum tensors.

The notion of intrinsic dimension was introduced by Dorgelo and Schouten¹⁴ under the name 'absolute dimension'. I find the adjective 'intrinsic' more congruous to modern terminology and less prone to suggest spurious connections with absolute values.

In the following I'll drop the adjective 'intrinsic' when it is clear from the context.

Different coordinate systems lead to different dimensions of the *components* of a tensor \mathbf{A} , but the intrinsic dimension of the tensor remains the same. Formula (12) for the dimensions of the components is consistent under changes of coordinates. For example, in new coordinates (\bar{x}^k) with dimensions (\bar{X}_k) , the new components of \mathbf{A} are

$$\bar{A}^k_l = A^i_j \frac{\partial \bar{x}^k}{\partial x^i} \frac{\partial x^j}{\partial \bar{x}^l} \quad (14)$$

and a quick check shows that $\dim(\bar{A}^k_l) = A \bar{X}_k \bar{X}_l^{-1}$, consistently with the general formula (12).

If in eq. (14), relating intrinsic and component dimensions, we have $X_1 = X_2 = \dots = X$, then all components also have equal dimensions. So if

¹² cf. the discussion in Synge 1960b § IV.5 p. 179. ¹³ Bridgman 1958; see also Synge 1960a § A.2; Truesdell & Toupin 1960 §§ A.3–4. ¹⁴ Dorgelo & Schouten 1946; Schouten 1989 ch. VI.

we use a system of coordinates having equal dimensions, the components of any tensor must also have equal dimensions. This justifies common practice in the literature.

Choosing coordinates of different dimensions, however, has several advantages. First, it allows us to use dimensional analysis as a heuristic tool to determine the variance type of a tensor; we'll see an example in § 10. Second, it can lead to components with familiar dimensions. For example, if we use a timelike coordinate of dimension T and spacelike coordinates of dimension L , then the components of the (co-contravariant) stress-energy-momentum tensor have the familiar dimensions of energy density, surface energy-flux density, momentum density, and pressure, with no c factors involved; see again § 10.

6 Tensor operations

By the reasoning of the previous section, which simply applies standard dimensional considerations to the basis expansion (9), it's easy to find the resulting intrinsic dimension of various operations and operators on tensors and tensor fields.

Here is a summary of the dimensional rules for the main differential-geometric operations and operators, except for the covariant derivative, the metric, and related tensors, discussed more in depth in §§ 8–9 below. Some of these rules are actually definition or conventions, as briefly discussed in their description. The others can be proved; I only give proof for some of them, leaving the other proofs as an exercise. For reference, in brackets I give the section of Choquet-Bruhat et al. (1996) where these operations are defined.

- The *tensor product* [III.B.5] multiplies dimensions:

$$\dim(\mathbf{A} \otimes \mathbf{B}) = \dim(\mathbf{A}) \dim(\mathbf{B}). \quad (15)$$

This is actually a definition or convention. We tacitly used this rule already in the example of § 3 and in § 5 for the coordinate expansion (9). It is a natural definition, because for tensors of order 0 (functions) the tensor product is just the ordinary product, and the dimension of a product is the product of the dimensions. This definition doesn't lead to inconsistencies.

- The *contraction* [III.B.5] or trace of the α th and β th slots of a tensor has the same dimension as the tensor:

$$\dim(\text{tr}_{\alpha\beta} \mathbf{A}) = \dim(\mathbf{A}). \quad (16)$$

Note that the formula above only holds *without raising or lowering indices*; see § 9 for those operations.

This operation can be traced back to the duality of vectors and covectors mentioned in § 2: a covector field ω operates linearly on a vector field v to yield a function $f = \omega(v)$. Also in this case we have that $\dim(f) = \dim(\omega) \dim(v)$ by definition or convention, and the rule (16) follows from this convention. Also in this case this convention seems very natural, owing to the linearity properties of the trace, and doesn't lead to inconsistencies.

- The *transposition*¹⁵ of the α th and β th slots of a tensor has the same dimension as the tensor:

$$\dim(\mathbf{A}^{\top\alpha\beta}) = \dim(\mathbf{A}). \quad (17)$$

- The *Lie bracket* [III.B.3] of two vectors has the product of their dimensions:

$$\dim([u, v]) = \dim(u) \dim(v). \quad (18)$$

In fact, in coordinates (x^i) the bracket can be expressed as

$$[u, v] = \left(u^j \frac{\partial v^i}{\partial x^j} - v^j \frac{\partial u^i}{\partial x^j} \right) \frac{\partial}{\partial x^i}, \quad (19)$$

and equating the dimensions of the left and right sides, considering that

$$\dim(u^i) = \dim(u) X_i, \quad \dim(v^i) = \dim(v) X_i, \quad (20)$$

we find again that all X terms cancel out, leaving the result (18).

- The *pull-back* [III.A.2], *tangent map* [III.B.1] and *push-forward* of a map F between manifolds don't change the dimensions of the tensors they map. The reason, evident from their definitions, is that they all rest on the pull-back of functions: $F^*(f) := f \circ F$. Being a composition, the pull-back of a function has the same dimension of the function.

¹⁵ called "building an isomer" by Schouten 1954 § I.3 p. 13; 1989 § II.4 p. 20.

- The *Lie derivative* [III.C.2] of a tensor with respect to a vector field has the product of the dimensions of the tensor and of the vector:

$$\dim(\mathcal{L}_v \mathbf{A}) = \dim(v) \dim(\mathbf{A}). \quad (21)$$

Regarding operations and operators on differential forms:

- The *exterior product* [IV.A.1] of two differential forms multiplies their dimensions:

$$\dim(\omega \wedge \tau) = \dim(\omega) \dim(\tau). \quad (22)$$

- The *interior product* [IV.A.4] (also called ‘dual’ or ‘dot’ product) of a vector and a form multiplies their dimensions:

$$\dim(v \lrcorner \omega) = \dim(v) \dim(\omega). \quad (23)$$

This equation also holds for the generalized inner product¹⁶ of a multi-vector v and a multi-covector ω .

- The *exterior derivative* [IV.A.2] of a form has the same dimension of the form:

$$\dim(d\omega) = \dim(\omega). \quad (24)$$

This can be proven using the identity $d \circ (v \lrcorner) + (v \lrcorner) \circ d = \mathcal{L}_v$ or similar identities¹⁷ together with eqs (21) and (23).

- The *integral* [IV.B.1] of a form over a submanifold (or more generally a chain) M has the same dimension as the form:

$$\dim\left(\int_M \omega\right) = \dim(\omega). \quad (25)$$

The reason is that the integral of a form over a submanifold or chain ultimately rests on the standard definition of integration on the real line¹⁸, which satisfies the dimensional rule above. In fact, the integral is invariant with respect to reparameterizations of the chain; it depends only on its image (some texts¹⁹ even define chains as equivalence classes determined by their image).

All rules above extend in obvious ways to tensor densities, and apply regardless whether the objects have straight or twisted orientation.

¹⁶ Deschamps 1970; 1981 Appendices; Lindell 2004; Truesdell & Toupin 1960 § F.I.267; Misner et al. 1973 Box 4.1, p. 92, item 4; see also Porta Mana 2019a. ¹⁷ Curtis & Miller 1985 ch. 9 p. 180 Theorem 9.78; Abraham et al. 1988 § 6.4 Theorem 6.4.8. ¹⁸ e.g. Choquet-Bruhat et al. 1996 §§ IV.B.1–2; de Rham 1984 § 5 p. 21, § 6 p. 24; Abraham et al. 1988 § 7.1; Boothby 2003 § VI.2. ¹⁹ e.g. Martin 2004 § 10.4 p. 297; Fecko 2006 § 7.3.

7 Curves and integral curves, 4-velocity

Consider a curve into spacetime, $C: s \mapsto P(s)$, with the parameter s having dimension $\dim(s) = S$.

If we consider the events of the spacetime manifold as dimensionless quantities, then the dimension of the tangent or velocity vector \dot{C} to the curve is

$$\dim(\dot{C}) = S^{-1}, \quad (26)$$

owing to the definition²⁰

$$\dot{C} := \frac{\partial(x^i \circ C)}{\partial s} \frac{\partial}{\partial x^i}. \quad (27)$$

This has a quirky interesting consequence. Given a vector field v we say that C is an integral curve for it if

$$v = \dot{C} \quad (28)$$

at all events $C(s)$ in the image of the curve (or more precisely $v_{C(s)} = \dot{C}_{C(s)}$ in usual differential-geometric notation²¹). From the point of view of dimensional analysis this definition can only be valid if v has dimension S^{-1} . If v and s^{-1} have different dimensions – a case which could happen for physical reasons – the condition (27) must be modified into $v = k\dot{C}$, where k is a possibly dimensional constant. This is equivalent to considering an affine and dimensional reparameterization of C .

The worldline of a small body is a curve into spacetime parameterized by some time t (usually but not necessarily the proper time; see § 9), with $\dim(t) = T$. The 4-velocity of the body at a given spacetime event is the tangent vector to the worldline thus parameterized. From eq. (26) it follows that *the 4-velocity has intrinsic dimension T^{-1}* .

You might find this result counter-intuitive. Probably the reason is that you're thinking in terms of components. There is no contradiction: if we use a system of coordinates with dimensions L and T , then the components of the 4-velocity have dimensions L/T and 1 . But the fact that the intrinsic dimension is T^{-1} , with no lengths involved, is quite sensible. We have equipped the body with some kind of clock, so we can say how much time has passed between two spacetime events intersected by the

²⁰ Choquet-Bruhat et al. 1996 § III.B.1; Boothby 2003 § IV.(1.9).
et al. 1996 § III.B.1.

²¹ Choquet-Bruhat

body's worldline. But we cannot say what their distance is, because no metric has been introduced yet.

If we introduce a metric tensor, The *norm* of the 4-velocity, calculated using the metric tensor, has dimension LT^{-1} or 1, depending on the dimension chosen for the metric tensor; see § 9.

8 Connection, covariant derivative, curvature tensors

Consider an arbitrary connection²² with covariant derivative ∇ . For the moment we don't assume the presence of any metric structure.

The covariant derivative of the product fv of a function and a vector satisfies²³

$$\nabla(fv) = df \otimes v + f\nabla v. \quad (29)$$

The first summand, from formulae (24) and (15), has dimension $\dim(f) \dim(v)$; for dimensional consistency this must also be the dimension of the second summand. Thus

$$\dim(\nabla v) = \dim(v). \quad (30)$$

It follows that the *directional* covariant derivative ∇_u has dimension

$$\dim(\nabla_u v) = \dim(u) \dim(v), \quad (31)$$

and by its derivation properties²⁴ we see that formula (30) extends from vectors to tensors of arbitrary type.

In the coordinate system (x^i) , the action of the covariant derivative is carried by the *connection coefficients* or Christoffel symbols (Γ^i_{jk}) defined by

$$\nabla \frac{\partial}{\partial x^k} = \Gamma^i_{jk} dx^j \otimes \frac{\partial}{\partial x^i}. \quad (32)$$

From this equation and eqs (15), (30) it follows that these coefficients have dimensions

$$\dim(\Gamma^i_{jk}) = X_i X_j^{-1} X_k^{-1}. \quad (33)$$

²² Choquet-Bruhat et al. 1996 § V.B.

²³ Choquet-Bruhat et al. 1996 § V.B.1.

²⁴ Choquet-Bruhat et al. 1996 § V.B.1 p. 303.

The *torsion* $\tau^\bullet_{|\bullet|}$, *Riemann curvature* $\mathbf{Rie}^\bullet_{|\bullet|}$, and *Ricci curvature* $\mathbf{Ric}_{..}$ tensors are defined by²⁵

$$\tau(u, v) := \nabla_u v - \nabla_v u - [u, v] , \quad (34)$$

$$\mathbf{Rie}(w; u, v) := \nabla_u \nabla_v w - \nabla_v \nabla_u w - \nabla_{[u, v]} w , \quad (35)$$

$$\mathbf{Ric}_{..} := \text{tr}_{13} \mathbf{Rie}^\bullet_{|\bullet|} . \quad (36)$$

From these definitions and the results of § 6 we find the dimensional requirements

$$\dim(\tau^\bullet_{|\bullet|}) \dim(u) \dim(v) = \dim(u) \dim(v) , \quad (37)$$

$$\dim(\mathbf{Rie}^\bullet_{|\bullet|}) \dim(w) \dim(u) \dim(v) = \dim(w) \dim(u) \dim(v) , \quad (38)$$

$$\dim(\mathbf{Ric}_{..}) = \dim(\mathbf{Rie}^\bullet_{|\bullet|}) , \quad (39)$$

which imply that *torsion*, *Riemann curvature*, and *Ricci curvature* tensors are dimensionless:

$$\dim(\tau^\bullet_{|\bullet|}) = \dim(\mathbf{Rie}^\bullet_{|\bullet|}) = \dim(\mathbf{Ric}_{..}) = 1 . \quad (40)$$

This result is sensible, because the notion of local parallelism, which these tensors express, doesn't involve any notion of distance or angle²⁶. The exact contra- and co-variant primitive type of these tensors is very important in the equations above. If a metric tensor is also introduced and used to raise or lower any indices of these tensors, the resulting tensors will have different dimensions; see next section, especially eq. (48).

Misner et al.²⁷ say that “curvature”, by which they seem to mean the Riemann tensor, has dimension L^{-2} . This statement is seemingly at variance with the dimensionless results (40). But I believe that Misner et al. refer to the *components* of the Riemann tensor in specific coordinates of dimension L and using geometrized units. In such specific coordinates every *component* R^i_{jkl} does indeed have dimension L^{-2} , according to the general formula (12), if and only if the intrinsic dimension of \mathbf{Rie} is unity, $\dim(\mathbf{Rie}) = 1$. So I believe that Misner et al.'s statement actually agrees with the results (40). This possible misunderstanding shows the importance of distinguishing between the intrinsic dimension, which doesn't depend on any specific coordinate choice, and component dimensions, which do.

²⁵ Choquet-Bruhat et al. 1996 § V.B.1. 1973 p. 35.

²⁶ cf. Porta Mana 2019b.

²⁷ Misner et al.

The formulae above are also valid if a metric is defined and the connection is compatible with it. The connection coefficients in this case are defined in terms of the metric tensor, but using the results of § 9 it's easy to see that eqs (30), (31), (33), (40) still hold.

9 Metric and related tensors and operations

Let us now consider a metric tensor $\mathbf{g}_{..}$. What is its intrinsic dimension $\dim(\mathbf{g})$? There seem to be two choices in the literature; both can be derived from the operational meaning of the metric.

Consider a (timelike) worldline $s \mapsto C(s)$, $s \in [a, b]$, between events $C(a)$ and $C(b)$. The metric tells us the *proper time* Δt elapsed for an observer having that worldline, according to the formula

$$\Delta t = \int_a^b \sqrt{|\mathbf{g}[\dot{C}(s), \dot{C}(s)]|} \, ds . \quad (41)$$

From the results of § 6 this formula implies that $T \equiv \dim(\Delta t) = \sqrt{\dim(\mathbf{g}_{..})}$, independently of the dimension of s , and therefore

$$\dim(\mathbf{g}_{..}) = T^2. \quad (42)$$

Most authors²⁸, however, prefer to include a dimensional factor $1/c$ in the definition (41):

$$\Delta t = \int_a^b \frac{1}{c} \sqrt{|\mathbf{g}[\dot{C}(s), \dot{C}(s)]|} \, ds , \quad (43)$$

thus obtaining

$$\dim(\mathbf{g}_{..}) = L^2. \quad (44)$$

The choice (44) seems also supported by the traditional expression for the “line element ds^2 ” as it appears in many works:

$$ds^2 = c^2 dt^2 - dx^2 - dy^2 - dz^2 , \quad (45)$$

possibly with opposite sign. If the coordinates (t, x, y, z) have the dimensions suggested by their symbols, this formula has dimension L^2 , so that if we interpret “ ds^2 ” as \mathbf{g} we find $\dim(\mathbf{g}) = L^2$. The line-element

²⁸ e.g. Fock 1964 § V.62 eq. (62.02); Curtis & Miller 1985 ch. 11 eq. (11.21); Rindler 1986 § 5.3 eq. (5.6); Hartle 2003 ch. 6 eq. (6.24).

expression above often has an ambiguous differential-geometric meaning, however, because it may also represent the metric applied to some *unspecified* vector, that is, $\mathbf{g}(\mathbf{v}, \mathbf{v})$, where \mathbf{v} is left unspecified²⁹. In this case we have

$$L^2 = \dim(\mathbf{g}) \dim(\mathbf{v})^2$$

and the dimension of \mathbf{g} is ambiguous or undefined, because the vector \mathbf{v} could have any dimension.

The standard choices for $\dim(\mathbf{g})$ are thus L^2 or T^2 ; the corresponding metric tensors differ by a factor c^2 .

The second choice has some merit for reasons discussed by Synge and Bressan³⁰. Synge gives a vivid summary:³¹

We are now launched on the task of giving physical meaning to the Riemannian geometry [...]. It is indeed a Riemannian *chronometry* rather than *geometry*, and the word *geometry*, with its dangerous suggestion that we should go about measuring *lengths* with *yardsticks*, might well be abandoned altogether in the present connection

In fact, to measure the proper time Δt defined above we only need to ensure that a clock has the worldline C , and then take the difference between the clock's final and initial times. On the other hand, consider the case when the curve C is *spacelike*. Its proper length is still defined by the integral (41) apart from a dimensional constant. Its measurement, however, is more involved than in the timelike case. It requires dividing the curve into very short pieces, and having specially-chosen observers (orthogonal to the pieces) measure each piece. But the measurement of each piece actually relies on the measurement of *proper time*³²: each observer uses radar distance, sending a lightlike signal which bounces back at the end of the piece, and measuring the time it takes to come back. Even if rigid rods are used, their calibration still relies on a measurement of time – this is also reflected in the current definition of the standard metre³³. McVittie, Synge, Kilmister³⁴ choose $\dim(\mathbf{g}) = T^2$.

The first choice, $\dim(\mathbf{g}) = L^2$, seems by far the most common, however. It has the merit that the projection of the metric onto a spacelike hypersurface also has dimension L^2 , which is sensible from a Newtonian point

²⁹ cf. Misner et al. 1973 Box 3.2 D p. 77. ³⁰ Synge 1960b §§ III.2–4; Bressan 1978 §§ 15, 18. ³¹ Synge 1960b § III.3 pp. 108–109. ³² Frankel 1979 ch. 2; Landau & Lifshitz 1996 § 84. ³³ BIPM 1983 p. 98; Giacomo 1984 p. 25. ³⁴ McVittie 1965 § 4.1; Synge 1960b § IV.5; Kilmister 1973 ch. II p. 25.

of view; and such projections are at the heart of the $3 + 1$ formulations of general relativity³⁵.  add refs to Wilson etc from commented text below

In the following we shall see how these choices affect some dimensional results.

The metric \mathbf{g} can be considered as an operator mapping vectors \mathbf{v} to covectors $\boldsymbol{\omega}$, which we can compactly write as $\boldsymbol{\omega} = \mathbf{g}\mathbf{v}$ instead of the cumbersome $\boldsymbol{\omega} = \text{tr}_{23}(\mathbf{g} \otimes \mathbf{v})$. The *inverse metric tensor* $\mathbf{g}^{-1\bullet\bullet}$ is then defined by the formula

$$\mathbf{g}\mathbf{g}^{-1} = \text{id}_{\bullet}^{\bullet}, \quad (46)$$

where $\text{id}_{\bullet}^{\bullet}: \boldsymbol{\omega} \mapsto \boldsymbol{\omega}$ is the dimensionless identity operator (also a tensor) on the cotangent space. Hence

$$\dim(\mathbf{g}^{-1}) = \dim(\mathbf{g})^{-1}. \quad (47)$$

The operation of *raising or lowering an index* of a tensor represents a contraction of the tensor product of that tensor with the metric or the metric inverse, for example $\underline{\mathbf{A}}_{\bullet\bullet} := \mathbf{g}\mathbf{A} \equiv \text{tr}_{23}(\mathbf{g}_{\bullet\bullet} \otimes \mathbf{A}^{\bullet\bullet})$ from the mixed tensor $\mathbf{A}^{\bullet\bullet}$, and similarly for tensors of other types. Therefore every lowering of a tensor's index multiplies its dimension by $\dim(\mathbf{g})$, and every rising divides it by $\dim(\mathbf{g})$:

$$\begin{aligned} \dim(\underline{\mathbf{A}}_{\bullet\bullet\bullet\bullet}) &= \dim(\mathbf{A}^{\bullet\bullet\bullet\bullet}) \dim(\mathbf{g}) \\ \dim(\bar{\mathbf{B}}^{\bullet\bullet\bullet\bullet}) &= \dim(\mathbf{B}_{\bullet\bullet\bullet\bullet}) \dim(\mathbf{g})^{-1}. \end{aligned} \quad (48)$$

A metric tensor on a four-dimensional spacetime determines a unique twisted 4-covector field or 4-form, called the volume element³⁶. It can be seen as a completely antisymmetric tensor of rank 4 (whose coordinate transformation includes the sign of the Jacobian), with only one non-zero component equal to the square root of the determinant of the components (g_{ij}) of the metric:

$$\sqrt{|\det(g_{ij})|} \, dx^0 \wedge dx^1 \wedge dx^2 \wedge dx^3,$$

where $dx^0 \wedge dx^1 \wedge dx^2 \wedge dx^3$ actually has a twisted orientation. Note that a non-vanishing volume element can be defined on orientable and non-orientable manifolds alike. From the results of § 6 it can be shown that

³⁵ Gourgoulhon 2012; Alcubierre 2008; Misner et al. 1973 ch. 21. I thank I. Bengtsson for this remark. ³⁶ De Rham 1984 § V.24; Choquet-Bruhat et al. 1996 § V.A.4; Abraham et al. 1988 § 6.2.

the 4-form above has intrinsic dimension $\dim(\mathbf{g})^2$ (in an n -dimensional spacetime it has dimension $\dim(\mathbf{g})^{n/2}$). It's convenient to multiply it by a power of c , and to define the *proper volume element* $\gamma_{|\dots|}$ as follows:

$$\gamma := \begin{cases} \frac{1}{c} \sqrt{|\det(g_{ij})|} dx^0 \wedge dx^1 \wedge dx^2 \wedge dx^3 \\ c^3 \sqrt{|\det(g_{ij})|} dx^0 \wedge dx^1 \wedge dx^2 \wedge dx^3 \end{cases} \quad \text{if } \dim(\mathbf{g}) := \begin{cases} L^2 \\ T^2 \end{cases} . \quad (49)$$

As a consequence we have

$$\dim(\gamma_{|\dots|}) = L^3 T \quad (50)$$

independently of whether $\dim(\mathbf{g})$ equals L^2 or T^2 . This convention has several advantages, and implies that the hypervolume of a four-dimensional region, given by the integral of γ (see § 6), also has dimension $L^3 T$ – which is a reasonable result for a *spacetime* region.

In general the metric \mathbf{g} induces volume, area, and line elements on three-, two-, and one-dimensional regions. It is convenient to multiply these elements by appropriate powers of c so that the region's volume has intuitive dimensions; such as L^3 for a spacelike three-dimensional region and $L^2 T$ for a timelike one. Indeed the definition of proper time (43) is doing exactly this, including a factor $1/c$ in the induced line element on a timelike curve.

The proper volume element appears in the definition of the *star operator* on covectors and forms. This operator acts by first rising all indices of a covector and then taking the generalized inner product (see § 6) with the proper volume element. For example, for a 2-covector $\omega \equiv \omega_{\bullet\bullet}$,

$$* \omega := (\mathbf{g}^{-1} \omega \mathbf{g}^{-1}) \rfloor \gamma . \quad (51)$$

From this definition it is clear that the star operator's effect on the dimension depends on the degree of the form it operates on³⁷.

The *inverse proper volume element* is the 4-vector field γ^{-1} , with twisted orientation, having unit generalized inner product with the proper volume element: $\gamma^{-1} \rfloor \gamma = 1$. Its intrinsic dimension is therefore

$$\dim(\gamma^{-1|\dots|}) = \dim(\gamma)^{-1} \equiv L^{-3} T^{-1} \quad (52)$$

again independently of whether $\dim(\mathbf{g})$ equals L^2 or T^2 .

³⁷ I personally prefer to avoid the star operator and explicitly use the inner product with the proper volume element; cf. Bossavit 1991 §§ 4.1–2.

Note that the inverse proper volume element is dimensionally and numerically different from the tensor $\bar{\gamma}$ obtained by raising all indices of γ . It can be proven, for example using an orthonormal basis, that the relation between the two is

$$\bar{\gamma} = -\frac{1}{c^2} \gamma^{-1}. \quad (53)$$

More generally the negative factor is replaced with $(-1)^s$, where s is the number of negative eigenvalues of the metric tensor.

For the case $\dim(\mathbf{g}) = L^2$, here is a summary of the components of the metric tensor \mathbf{g} and of the proper volume element γ , in a rectangular Cartesian coordinate system adapted to an inertial (freely falling) observer at a specific spacetime event. For coordinate with different dimensions:

$$\begin{aligned} \dim(t, x, y, z) &= (T, L, L, L), \\ (g_{ij}) &= \begin{pmatrix} -c^2 & 0 \text{ m s}^{-1} & 0 \text{ m s}^{-1} & 0 \text{ m s}^{-1} \\ 0 \text{ m s}^{-1} & 1 & 0 & 0 \\ 0 \text{ m s}^{-1} & 0 & 1 & 0 \\ 0 \text{ m s}^{-1} & 0 & 0 & 1 \end{pmatrix}, & \gamma_{txyz} &= 1, \end{aligned} \quad (54)$$

where the units of the zero timelike components above are arbitrary of course, as long as they have dimension LT^{-1} . For coordinates with equal dimensions:

$$\begin{aligned} \dim(x^0, x^1, x^2, x^3) &= (L, L, L, L), \\ (g_{ij}) &= \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}, & \gamma_{0123} &= \frac{1}{c}. \end{aligned} \quad (55)$$

Many authors use a metric tensor with opposite sign.

The formulae for the covariant derivative (30), connection coefficients (33), and curvature tensors (40) remain valid for a connection compatible with the metric. In this case the connection coefficients can be obtained from the metric by the formulae³⁸

$$\Gamma_{jk}^i = \frac{1}{2} \left(\frac{\partial}{\partial x^k} g_{jl} + \frac{\partial}{\partial x^j} g_{kl} - \frac{\partial}{\partial x^l} g_{jk} \right) g^{li}, \quad (56)$$

and it's easily verified that the dimensions of these coefficients given in eq. (33) still hold, as do the results for the curvature tensors (40).

³⁸ Choquet-Bruhat et al. 1996 § V.B.2.

The scalar curvature R and the co-co-variant Einstein tensor $\mathbf{G}_{..}$

$$R := \text{tr}(\mathbf{Ric} \, \mathbf{g}^{-1}) , \quad \mathbf{G}_{..} := \mathbf{Ric} - \frac{1}{2} R \, \mathbf{g} \quad (57)$$

have dimensions

$$\dim(R) = \dim(\mathbf{g})^{-1} \equiv \begin{cases} \text{T}^{-2} \\ \text{L}^{-2} \end{cases} \quad \text{if } \dim(\mathbf{g}) := \begin{cases} \text{T}^2 \\ \text{L}^2 \end{cases} , \quad (58)$$

$$\dim(\mathbf{G}_{..}) = 1 , \quad (59)$$

that is, the twice covariant Einstein tensor is dimensionless, independently of the dimension of the metric tensor.

10 Stress-energy-momentum tensor and four-momentum

Also in the case of the stress-energy-momentum tensor the literature offers two main choices of intrinsic dimension, independent of the choices for the metric tensor discussed in the previous section. Moreover, there seems to be no consensus yet on what the primitive variance type of the stress-energy-momentum tensor should be: its operational meaning is still surrounded by some mystery.

Let's try to find its dimension and variance type through a heuristic approach, which also shows the usefulness of intrinsic dimensional analysis on differential manifolds.

The stress-energy-momentum tensor for a material continuum at a spacetime event embodies the energy density (comprising rest energy, internal energy, kinetic energy) ϵ , surface energy-flux density q_r (comprising convected energy density, heating, and working), momentum density p_r , and stress (including convected momentum density) σ_{sr} of the material at that event. Here the indices $r, s \in \{x, y, z\}$, and the vertical positions of the indices do *not* denote any variance type. These quantities are measured by an inertial observer at that event using a system of one timelike and three spacelike coordinates (t, x, y, z) . If these coordinates have dimensions $(\text{T}, \text{L}, \text{L}, \text{L})$, then the dimensions of the quantities above are

$$\begin{aligned} \dim(\epsilon) &= \text{ML}^{-1}\text{T}^{-2} \equiv \text{EL}^{-3} , & \dim(q_r) &= \text{MT}^{-3} \equiv \text{EL}^{-2}\text{T}^{-1} , \\ \dim(p_r) &= \text{ML}^{-2}\text{T}^{-1} \equiv \text{EL}^{-4}\text{T} , & \dim(\sigma_{rs}) &= \text{ML}^{-1}\text{T}^{-2} \equiv \text{EL}^{-3} . \end{aligned} \quad (60)$$

Suppose we want to construct a spacetime tensor \mathcal{T} having these 16 independent quantities as components. What should the variance type and the intrinsic dimension of such a tensor be? I won't assume the symmetry of this tensor as an a-priori kinematic property, but leave it as a dynamical law enforced by the Einstein equations; it only needs to hold for the sum of the stress-energy-momentum tensors from all kinds of matter.

Since we have 16 components, this tensor should belong to the tensor product of two tangent spaces, each spanned by four basis elements. There are four such spaces: vectors, covectors, 3-vectors, and 3-covectors. Let's use shorthands such as $\partial_{tzy} := \partial_t \wedge \partial_z \wedge \partial_y$ and $dtzy := dt \wedge dz \wedge dy$. These four spaces then have the following coordinate-induced bases and corresponding dimensions:

$$(\partial_t, \partial_x, \partial_y, \partial_z) : \quad (T^{-1}, L^{-1}, L^{-1}, L^{-1}), \quad (61a)$$

$$(dt, dx, dy, dz) : \quad (T, L, L, L), \quad (61b)$$

$$(\partial_{xyz}, \partial_{tzy}, \partial_{txz}, \partial_{tyx}) : \quad (L^{-3}, L^{-2}T^{-1}, L^{-2}T^{-1}, L^{-2}T^{-1}), \quad (61c)$$

$$(dxyz, dtzy, dtxz, dtyx) : \quad (L^3, L^2T, L^2T, L^2T), \quad (61d)$$

the orderings chosen to minimize the minus signs appearing from inner products with a volume element. There are therefore 4×4 possible tensor-product spaces, each constructed by the product of two of the four spaces above; and thus 16 possible alternatives to represent our stress-energy-momentum tensor. Energy density is intuitively associated with the purely timelike component of this tensor, stress with the purely spacelike ones, and surface energy-flux density and momentum density with the mixed timelike-spacelike components.

Consider the following first alternative, obtained from the tensor product of the space (61a) with itself; omit y - and z -terms for brevity:

$$\mathcal{T} \stackrel{?}{=} \epsilon \partial_t \otimes \partial_t + q_x \partial_t \otimes \partial_x + p_x \partial_x \otimes \partial_t + \sigma_{xx} \partial_x \otimes \partial_x + \dots$$

The first and third summands of this expression have incompatible intrinsic dimensions $EL^{-3}T^{-2}$ and EL^{-5} . This alternative is therefore rejected because dimensionally inconsistent. Similar dimensional analyses on the remaining fifteen alternatives show that only four are dimensionally

consistent:

$$\mathcal{T}_{\bullet}^{\bullet} = -\epsilon \, dt \otimes \partial_t - q_x \, dt \otimes \partial_x + p_x \, dx \otimes \partial_t + \sigma_{xx} \, dx \otimes \partial_x + \cdots \quad (62a)$$

$$\begin{aligned} \mathcal{T}_{\bullet|\bullet\bullet\bullet}^{\bullet} = & -\epsilon \, dt \otimes dx yz - q_x \, dt \otimes dtzy + \\ & p_x \, dx \otimes dx yz + \sigma_{xx} \, dx \otimes dtzy + \cdots \end{aligned} \quad (62b)$$

$$\begin{aligned} \mathcal{T}^{\bullet|\bullet\bullet\bullet} = & -\epsilon \, \partial_{xyz} \otimes \partial_t - q_x \, \partial_{xyz} \otimes \partial_x + \\ & p_x \, \partial_{tzy} \otimes \partial_t + \sigma_{xx} \, \partial_{tzy} \otimes \partial_x + \cdots \end{aligned} \quad (62c)$$

$$\begin{aligned} \mathcal{T}^{\bullet|\bullet\bullet\bullet}_{|\bullet\bullet\bullet} = & -\epsilon \, \partial_{xyz} \otimes dx yz - q_x \, \partial_{xyz} \otimes dtzy + \\ & p_x \, \partial_{tzy} \otimes dyxz + \sigma_{xx} \, \partial_{tzy} \otimes dtzy + \cdots \end{aligned} \quad (62d)$$

where the particular signs of the components, which are not determined by dimensional analysis, will be explained later. Our analysis of the intrinsic dimensions therefore restricts the stress-energy-momentum tensor to be one of the four alternatives above. Note that their kind of orientation, straight or twisted, is still undetermined. To further restrict the possibilities let's consider a couple of additional, interrelated heuristic arguments.

First, the notions of energy and momentum density, surface energy-flux density, and stress imply some kind of integration over three-dimensional spacelike or timelike regions. Such integration needs a 3-form and thus excludes alternatives (62a) and (62c).

Second, the total energy measured by a specific observer within a *topologically* specified three-dimensional spatial region is considered to be independent of the metric extent of that region (just like charge, although the latter is also independent of the observer's motion). The energy *density* therefore does depend on the metric properties and must change accordingly. Similar arguments hold for the surface energy flux. Only the second tensor alternative (62b) above is consistent with these requirements. Two points are worth noting here: (i) This metric scaling argument doesn't quite work for the momentum-density part $p_x \, dx \otimes dx yz$, though, if it's interpreted as a mass flux $ML^{-2}T^{-1}$: the L^{-2} scaling in the component p_x and the L^4 scaling in the corresponding basis element $dx \otimes dx yz$ don't match. They do match, however, if we interpret the momentum-density part in terms of energy or action: $ML^{-2}T^{-1} = ETL^{-4}$. A similar situation occurs for the stress part. (ii) This

kind of scaling mismatch also occurs, even in worse forms, in all other 15 alternatives for the stress-energy-momentum tensor. Thus intrinsic dimensional analysis alone suggests connections between mass and energy, and seems to give the latter quantity a more fundamental status.

Finally, the value of the energy density should not change under a change in the orientation of the spacelike coordinates. This means that the 3-covector slot in alternative (62b) should have a twisted orientation.

A heuristic application of intrinsic dimensional analysis, combined with a integration, rescaling, and reorientation arguments, thus tells us that the stress-energy-momentum tensor has variance type $\mathcal{T}_{\cdot|\dots|}$, or equivalently that it's a covector-valued 3-covector, or a four-times-covariant tensor completely antisymmetric in three slots. The 3-covector or antisymmetric slots have a twisted orientation. This tensor has the dimension of an *action*:

$$\dim(\mathcal{T}_{\cdot|\dots|}) = \text{ET} \equiv \text{ML}^2\text{T}^{-1} . \quad (63)$$

This form of the stress-energy-momentum tensor appears for example in Einstein's original work³⁹, Cartan's⁴⁰, Brillouin's⁴¹, Fokker's⁴², Truesdell & Toupin's⁴³, who try to find an expression universally valid in Newtonian, Lorentzian, and general-relativistic mechanics. It also appears in more recent works⁴⁴ which approach its definition from diverse points of view. This is also the form obtained with variational principles from an action Lagrangean⁴⁵, from which it is easily seen that this tensor has the intrinsic dimension of an action.

Some authors⁴⁶ think about this tensor in terms of mass rather than energy (Fock⁴⁷ calls it the "mass tensor"), and therefore assign to its covector-valued 3-covector form $\mathcal{T}_{\cdot|\dots|}$ the dimension of mass-time, that is, an action per squared velocity:

$$\dim(\mathcal{T}_{\cdot|\dots|}) = \text{MT} \equiv \text{EL}^{-2}\text{T}^3 . \quad (64)$$

³⁹ Einstein 1914 § C.9. ⁴⁰ Cartan 1923 § 13. ⁴¹ cf. Brillouin 1924 § 7. ⁴² Fokker 1965 § VII.2. ⁴³ Truesdell & Toupin 1960 § F.IV.288. ⁴⁴ Misner et al. 1973 ch. 14 Exercise 14.18, ch. 15; Hehl & McCrea 1986; Gronwald & Hehl 1997; cf. also Segev 1986; Segev & Rodnay 1999; Segev 2000a,b; 2002; Kanso et al. 2007. ⁴⁵ Misner et al. 1973 § 21.3; Hawking & Ellis 1994 § 3.3; Pauli 1958 § IV.55; see also Fokker 1965 § VII.8; Gotay & Marsden 1992. ⁴⁶ e.g. Fock 1964 § V.55; McVittie 1965 § 4.1; Adler et al. 1975 § 10.1.

⁴⁷ Fock 1964 § II.31.

With this intrinsic dimension, however, not all components of the stress-energy-momentum tensor have an intuitive meaning and dimensions, when coordinates with dimensions (T, L, L, L) are used.

The two choices (63), (64) differ by a factor c^2 .

The two dimensional choices for the metric, eqs (42) and (44), and for the stress-energy-momentum tensor, eqs (63) and (64), appear in all four combinations in the literature. For example, $\dim(\mathbf{g}) = T^2$ & $\dim(\mathcal{T}) = ET$ is used by Synge⁴⁸; $\dim(\mathbf{g}) = L^2$ & $\dim(\mathcal{T}) = MT$ is used by Fock⁴⁹ and Adler et al.⁵⁰; $\dim(\mathbf{g}) = T^2$ & $\dim(\mathcal{T}) = MT$ is used by McVittie⁵¹ and possibly Kilmister⁵². Most other works use $\dim(\mathbf{g}) = L^2$ & $\dim(\mathcal{T}) = ET$. These combinations lead to three possible values for Einstein's constant, as we'll see in the next section.

Note that some of the works cited speak of a tensor of variance type $\mathcal{T}_{\cdot|\dots|}$, others of a once covariant and once contravariant "V(olume)-tensor" or "tensor density", which has variance type $\mathcal{T}_{\cdot}^{\cdot|\dots|}$. These two objects are geometrically equivalent, in the sense that their independent components have the same transformation law under changes of coordinates⁵³ (this is why I chose a calligraphic letter to denote this tensor). They are related by an inner product with $\omega \otimes \omega^{-1}$, where ω is an arbitrary non-vanishing 4-form; this last object has only one independent component, which transforms as the scalar 1.

The signs of the components of \mathcal{T} depend on the signature of the metric \mathbf{g} . If the latter has signature $(-, +, +, +)$, then the energy components have negative sign, as in eq. (62b). If the metric has opposite sign, that is, signature $(+, -, -, -)$, then \mathcal{T} has opposite sign as well, and its momentum components are negative.

To obtain a twice covariant tensor to be used in the Einstein equations, we first take the inner product of the 3-covector, or antisymmetric, part by the inverse proper volume element, obtaining a co-contravariant tensor equivalent to (62a) above. In general coordinates this operation corresponds to dividing the components by $\frac{1}{c} \sqrt{|\det(g_{ij})|}$ if $\dim(\mathbf{g}) = L^2$, according to eq. (49), and so can be interpreted as passing from a tensor

⁴⁸ Synge 1960b §§ IV.4–5.

⁴⁹ Fock 1964 §§ V.54–55.

⁵⁰ Adler et al. 1975 § 10.1.

⁵¹ McVittie 1965 § 4.1.

⁵² Kilmister 1973 chs II–III, he seems to shift to natural units

at some point. ⁵³ Schouten 1989 § II.8 p. 30.

density to an ordinary tensor. Then we lower the new contravariant slot using the metric tensor:

$$\mathbf{T}_{..} := (\mathcal{T}_{\bullet, \dots, \bullet} \lrcorner \gamma^{-1}) \mathbf{g} . \quad (65)$$

According to this definition and eqs (64), (42), (44), (50), there are three possible dimensional choices for the co-co-variant tensor \mathbf{T} , dependent on the choices of dimensions of \mathbf{g} and \mathcal{T} :

$$\dim(\mathbf{T}_{..}) = \dim(\mathcal{T}_{\bullet, \dots, \bullet}) \dim(\mathbf{g}) \dim(\gamma)^{-1}$$

$$= \begin{cases} \text{MLT}^{-2} \equiv \text{EL}^{-1} \\ \text{ML}^{-1} \equiv \text{EL}^{-3}\text{T}^2 \\ \text{ML}^{-3}\text{T}^2 \equiv \text{EL}^{-5}\text{T}^4 \end{cases} \quad \text{if} \quad \begin{cases} \dim(\mathbf{g}) := \text{L}^2 \ \& \ \dim(\mathcal{T}) := \text{ET} \\ \dim(\mathbf{g}) := \text{T}^2 \ \& \ \dim(\mathcal{T}) := \text{ET} \ \text{or} \\ \dim(\mathbf{g}) := \text{L}^2 \ \& \ \dim(\mathcal{T}) := \text{MT} \\ \dim(\mathbf{g}) := \text{T}^2 \ \& \ \dim(\mathcal{T}) := \text{MT} \end{cases} \quad (66)$$

All three possibilities, which differ by factors c^2 , appear in the literature: see the examples cited above for the combinations of metric and stress-energy-momentum.

11 The constant in the Einstein equations

We finally arrive at the Einstein equations,

$$\mathbf{G} = \kappa \mathbf{T} \quad (67)$$

sometimes seen with a minus sign, depending on the signature of the metric and of alternative definitions of the curvature tensors⁵⁴. κ is Einstein's constant.

According to our definitions of $\mathbf{G}_{..}$, eq. (59), and $\mathbf{T}_{..}$, eq. (65), the equations above are considered in their co-co-variant form. This form is convenient because the left side is then dimensionless (its intrinsic dimension is 1), independently of the dimension of the metric tensor, as

⁵⁴ see the *Table of sign conventions* on the final pages of Misner et al. 1973.

explained at the end of § 9, eq. (59). We therefore find as a general rule that

$$\dim(\kappa) = \dim(\mathbf{T}_{..})^{-1} \quad (68)$$

always, for all choices of dimensions for the metric and stress-energy-momentum tensors. Combining this equation with the results for \mathbf{T} , eq. (66), we find three possible conventions:

$$\dim(\kappa) =$$

$$\left\{ \begin{array}{l} M^{-1}L^{-1}T^2 \equiv E^{-1}L \\ M^{-1}L \equiv E^{-1}L^3T^{-2} \\ M^{-1}L^3T^{-2} \equiv E^{-1}L^5T^{-4} \end{array} \right. \quad \text{if} \quad \left\{ \begin{array}{l} \dim(\mathbf{g}) := L^2 \ \& \ \dim(\mathcal{T}) := ET \\ \dim(\mathbf{g}) := T^2 \ \& \ \dim(\mathcal{T}) := ET \ \text{or} \\ \dim(\mathbf{g}) := L^2 \ \& \ \dim(\mathcal{T}) := MT \\ \dim(\mathbf{g}) := T^2 \ \& \ \dim(\mathcal{T}) := MT \end{array} \right. \quad (69)$$

Einstein's constant κ can therefore be obtained from Newton's gravitational constant $\dim(G) = M^{-1}L^3T^{-2}$ (this is not the Einstein tensor \mathbf{G} !) and the speed of light $\dim(c) = LT^{-1}$ only in the following three ways, with the 8π factor coming from the Newtonian limit:

$$\kappa = \left\{ \begin{array}{l} 8\pi G/c^4 \\ 8\pi G/c^2 \\ 8\pi G \end{array} \right. \quad \text{if} \quad \left\{ \begin{array}{l} \dim(\mathbf{g}) := L^2 \ \& \ \dim(\mathcal{T}) := ET \\ \dim(\mathbf{g}) := T^2 \ \& \ \dim(\mathcal{T}) := ET \ \text{or} \\ \dim(\mathbf{g}) := L^2 \ \& \ \dim(\mathcal{T}) := MT \\ \dim(\mathbf{g}) := T^2 \ \& \ \dim(\mathcal{T}) := MT \end{array} \right. . \quad (70)$$

From the discussion of the literature given in § 10, the first convention is the most common. The second convention appears for example in Fock⁵⁵ and Adler et al.⁵⁶. The third convention would appear in McVittie if he didn't cheat a factor $1/c^2$ into κ by writing the Einstein equations⁵⁷ as “ $\mathbf{G} = \kappa c^2 \mathbf{T}$ ”.

12 Summary and conclusions

We have seen that dimensional analysis, with its familiar rules, can be seamlessly performed in Lorentzian and general relativity and in

⁵⁵ Fock 1964 § 55 eqs (55.15) and (52.06).

⁵⁶ Adler et al. 1975 § 10.5 eq. (10.98).

⁵⁷ McVittie 1965 § 4.2 eq. (4.107).

differential geometry if we adopt the coordinate-free approach typical of modern texts. In this approach each tensor has an *intrinsic* dimension (a notion introduced by Schouten and Dorgelo). This dimension doesn't depend on the dimensions of the coordinates, and is determined by the physical and operational meaning of the tensor. It is therefore generally more profitable to focus on the intrinsic dimension of a tensor rather than on the dimensions of its components. The dimension of each specific component is easily found by formula (12): it's the product of the intrinsic dimension by the dimension of the i th coordinate function for each contravariant index i , by the inverse of the dimension of the j th coordinate function for each covariant index j . Intrinsic dimensional analysis seems to rest on two main conventions: the tensor product and the action of covectors on vectors behave analogously to usual multiplication for the purposes of dimensional analysis. Alternative, equivalent sets of conventions could perhaps be considered.

Intrinsic dimensional analysis can help us determine or at least constrain the variance type of candidate tensors, as we saw for the stress-energy-momentum tensor in § 10.

We found or re-derived some essential results for general relativity, in particular that the Riemann ***Rie***^{*}..., Ricci ***Ric***..., and twice-covariant Einstein ***G***... curvature tensors are dimensionless. Maybe these results can be of importance for some current research involving scales and conformal factors⁵⁸. We also discussed the operational reasons behind two common choices of dimensions for the metric and the stress-energy-momentum tensors.

Since the dimensions of the components are usually different from the intrinsic dimension and depend on the coordinates, I recommend to avoid statements such as “the tensor A_i^{jk} has dimension X ”, which leave it unclear whether “ A_i^{jk} ” is meant to represent the tensor in general (as in Penrose & Rindler's notation), or to represent its set of components, or to represent just a specific component.

Dimensional analysis remains a controversial, obscure, but fascinating subject still today, 60 years from Truesdell & Toupin's remark quoted in the Introduction. For an overview of some recent and creative approaches to it, going beyond Bridgman's text⁵⁹ (whose point of view

⁵⁸ e.g. Röhr & Uggla 2005; Cadoni & Tuveri 2019.

⁵⁹ Bridgman 1963.

is in many respects at variance with modern developments: see the following references), I recommend for example the works by Mari et al.⁶⁰, Domotor and Batitsky⁶¹, Kitano⁶², the extensive analysis by Dybkaer⁶³, the historical review by de Boer⁶⁴, and references therein.

Thanks

... to (chronologically) Mariano Cadoni, Ingemar Bengtsson, Iván Davidovich, Claudia Battistin for valuable comments on previous drafts. To the staff of the NTNU library for their always prompt support. To Mari, Miri, Emma for continuous encouragement and affection, and to Buster Keaton and Saitama for filling life with awe and inspiration. To the developers and maintainers of L^AT_EX, Emacs, AUC_TE_X, Open Science Framework, R, Python, Inkscape, Sci-Hub for making a free and impartial scientific exchange possible. ☒

Note: this paper was rejected by arXiv without any kind of scientific motivation. It's sad and alarming that the possibility of pre-print feedback is curtailed like that.

This work is financially supported by the Kavli Foundation and the Centre of Excellence scheme of the Research Council of Norway (Roudi group).

Bibliography

("de X" is listed under D, "van X" under V, and so on, regardless of national conventions.)

Abraham, R., Marsden, J. E., Ratiu, T. (1988): *Manifolds, Tensor Analysis, and Applications*, 2nd ed. (Springer, New York). First publ. 1983. DOI:[10.1007/978-1-4612-1029-0](https://doi.org/10.1007/978-1-4612-1029-0).

Adler, R., Bazin, M., Schiffer, M. (1975): *Introduction to General Relativity*, 2nd ed. (McGraw-Hill, Tokyo). First publ. 1965.

Alcubierre, M. (2008): *Introduction to 3 + 1 Numerical Relativity*. (Oxford University Press, Oxford).

BIPM (Bureau international des poids et mesures) (1983): *17^e conférence générale des poids et mesures (CGPM)*. (BIPM, Sèvres). <https://www.bipm.org/en/worldwide-metrology/cgpm/resolutions.html#comptes-rendus>.

Bird, R. B., Stewart, W. E., Lightfoot, E. N. (2002): *Transport Phenomena*, 2nd ed. (Wiley, New York). First publ. 1960.

Boothby, W. M. (2003): *An Introduction to Differentiable Manifolds and Riemannian Geometry*, rev. 2nd ed. (Academic Press, Orlando, USA). First publ. 1975.

⁶⁰ Mari & Giordani 2012; Frigerio et al. 2010. ⁶¹ Domotor 2017; Domotor & Batitsky 2016; Domotor 2012. ⁶² Kitano 2013. ⁶³ Dybkaer 2010. ⁶⁴ De Boer 1995.

- Bossavit, A. (1991): *Differential Geometry: for the student of numerical methods in electromagnetism*. https://www.researchgate.net/publication/200018385_Differential_Geometry_for_the_student_of_numerical_methods_in_Electromagnetism.
- Bressan, A. (1978): *Relativistic Theories of Materials*. (Springer, Berlin). doi:10.1007/978-3-642-81120-3.
- Bridgman, P. W. (1958): *The Logic of Modern Physics*, 8th pr. (Macmillan, New York). First publ. 1927.
- (1963): *Dimensional Analysis*, rev. ed. (Yale University Press, New Haven). First publ. 1922. <http://archive.org/details/dimensionalanaly00bridrich>.
- Brillouin, L. (1924): *Les lois de l'élasticité en coordonnées quelconques*. Proc. Int. Congress Mathematicians **1924-II**, 73–97.
- Burke, W. L. (1980): *Spacetime, Geometry, Cosmology*. (University Science Books, Mill Valley, USA).
- (1983): *Manifestly parity invariant electromagnetic theory and twisted tensors*. J. Math. Phys. **24**¹, 65–69. doi:10.1063/1.525603.
- (1987): *Applied Differential Geometry*, repr. (Cambridge University Press, Cambridge). First publ. 1985. doi:10.1017/CB09781139171786.
- (1995): *Div, Grad, Curl Are Dead*. http://people.ucsc.edu/~rmont/papers/Burke_DivGradCurl.pdf; “preliminary draft II”. See also <http://www.ucolick.org/~burke/>.
- Cadoni, M., Tuveri, M. (2019): *Galactic dynamics and long-range quantum gravity*. Phys. Rev. D **100**², 024029. doi:10.1103/PhysRevD.100.024029.
- Cartan, É. (1923): *Sur les variétés à connexion affine et la théorie de la relativité généralisée (première partie)*. Annales scientifiques de l'École Normale Supérieure **40**, 325–412. doi:10.24033/asens.751. See also Cartan (1924; 1925); transl. in Cartan (1986).
- (1924): *Sur les variétés à connexion affine et la théorie de la relativité généralisée (première partie) (Suite)*. Annales scientifiques de l'École Normale Supérieure **41**, 1–25. doi:10.24033/asens.753. See also Cartan (1923; 1925); transl. in Cartan (1986).
- (1925): *Sur les variétés à connexion affine et la théorie de la relativité généralisée (deuxième partie)*. Annales scientifiques de l'École Normale Supérieure **42**, 17–88. doi:10.24033/asens.761. See also Cartan (1923; 1924); transl. in Cartan (1986).
- (1986): *On Manifolds with an Affine Connection and the Theory of General Relativity*. (Bibliopolis, Napoli). Transl. of Cartan (1923; 1924; 1925) by Anne Magnon and Abhay Ashtekar.
- Choquet-Bruhat, Y., DeWitt-Morette, C., Dillard-Bleick, M. (1996): *Analysis, Manifolds and Physics. Part I: Basics*, rev. ed. (Elsevier, Amsterdam). First publ. 1977.
- Curtis, W. D., Miller, F. R. (1985): *Differential Manifolds and Theoretical Physics*. (Academic Press, Orlando, USA).
- de Boer, J. (1995): *On the history of quantity calculus and the International System*. Metrologia **31**⁶, 405–429. doi:10.1088/0026-1394/31/6/001.
- de Rham, G. (1984): *Differentiable Manifolds: Forms, Currents, Harmonic Forms*. (Springer, Berlin). Transl. by F. R. Smith. First publ. in French 1955. doi:10.1007/978-3-642-61752-2.
- Deschamps, G. A. (1970): *Exterior differential forms*. In: Deschamps, de Jager, John, Lions, Moiseev, Sommer, Tihonov, Tikhomirov, et al. (1970): p. III:111–161.
- (1981): *Electromagnetics and differential forms*. Proc. IEEE **69**⁶, 676–696.
- Deschamps, G. A., de Jager, E. M., John, F., Lions, J. L., Moiseev, N., Sommer, F., Tihonov, A. N., Tikhomirov, V., et al. (1970): *Mathematics applied to physics*. (Springer, Berlin). Ed. by É. Roubine.

- Dixon, W. G. (1975): *On the uniqueness of the Newtonian theory as a geometric theory of gravitation*. Commun. Math. Phys. **45**², 167–182. doi:[10.1007/BF01629247](https://doi.org/10.1007/BF01629247).
- Domotor, Z. (2012): *Algebraic frameworks for measurement in the natural sciences*. Meas. Sci. Rev. **12**⁶, 213–233. doi:[10.2478/v10048-012-0032-7](https://doi.org/10.2478/v10048-012-0032-7).
- (2017): *Torsor theory of physical quantities and their measurement*. Meas. Sci. Rev. **17**⁴, 152–177. doi:[10.1515/msr-2017-0019](https://doi.org/10.1515/msr-2017-0019).
- Domotor, Z., Batitsky, V. (2016): *An algebraic approach to unital quantities and their measurement*. Meas. Sci. Rev. **16**³, 103–126. doi:[10.1515/msr-2016-0014](https://doi.org/10.1515/msr-2016-0014).
- Dorgelo, H. B., Schouten, J. A. (1946): *On unities and dimensions*. [I, II, III]. Verh. Kon. Akad. Wetensch. Amsterdam **49**^{2,3,4}, 123–131, 282–291, 393–403.
- Dybkaer, R. (2010): *An Ontology on Property: For Physical, Chemical and Biological Systems*, 2nd ed. (IUPAC, Research Triangle Park, USA). First publ. 2004. doi:[10.1351/978-87-990010-1-9](https://doi.org/10.1351/978-87-990010-1-9).
- Eckart, C. (1940): *The thermodynamics of irreversible processes*. III. Relativistic theory of the simple fluid. Phys. Rev. **58**¹⁰, 919–924. doi:[10.1103/PhysRev.58.919](https://doi.org/10.1103/PhysRev.58.919).
- Einstein, A. (1914): *Die formale Grundlage der allgemeinen Relativitätstheorie*. Sitzungsber. K. Preußische Akad. Wiss. **1914**^{XLI}, 1030–1085. Transl. with notes in Einstein (1997), Doc. 9 pp. 30–84.
- (1996): *The Collected Papers of Albert Einstein*. Vol. 6: *The Berlin Years: Writings, 1914–1917*. (Princeton University Press, Princeton). Ed. by A. J. Kox, Martin J. Klein, Robert Schulmann. Transl. in Einstein (1997).
- (1997): *The Collected Papers of Albert Einstein*. Vol. 6: *The Berlin Years: Writings, 1914–1917*. English translation of selected texts. (Princeton University Press, Princeton). Transl. of Einstein (1996) by Alfred Engel and Engelbert Schucking.
- Fecko, M. (2006): *Differential Geometry and Lie Groups for Physicists*. (Cambridge University Press, Cambridge). doi:[10.1017/CB09780511755590](https://doi.org/10.1017/CB09780511755590).
- Flügge, S., ed. (1960): *Handbuch der Physik: Band III/1: Prinzipien der klassischen Mechanik und Feldtheorie* [Encyclopedia of Physics: Vol. III/1: Principles of Classical Mechanics and Field Theory]. (Springer, Berlin). doi:[10.1007/978-3-642-45943-6](https://doi.org/10.1007/978-3-642-45943-6).
- Fock [Fok], V. A. (1964): *The Theory of Space, Time and Gravitation*, 2nd rev. ed. (Pergamon, Oxford). Transl. by N. Kemmer. First publ. in Russian 1955.
- Fokker, A. D. (1965): *Time and Space, Weight and Inertia: A chronogeometrical introduction to Einstein's theory*. (Pergamon, Oxford). Transl. by D. Bijl, translation ed. by D. Field. First publ. in Dutch 1960.
- Frankel, T. (1979): *Gravitational Curvature: An Introduction to Einstein's Theory*. (W. H. Freeman and Company, San Francisco).
- Frigerio, A., Giordani, A., Mari, L. (2010): *Outline of a general model of measurement*. Synthese **175**², 123–149. doi:[10.1007/s11229-009-9466-3](https://doi.org/10.1007/s11229-009-9466-3).
- Giacomo, P. (1984): *News from the BIPM*. Metrologia **20**¹, 25–30. doi:[10.1088/0026-1394/17/2/007](https://doi.org/10.1088/0026-1394/17/2/007).
- Gotay, M. J., Marsden, J. E. (1992): *Stress-energy-momentum tensors and the Belinfante-Rosenfeld formula*. Contemp. Math. **132**, 367–392. <http://www.cds.caltech.edu/~marsden/bib/1992.html>.
- Gourgoulhon, É. (2012): *3+1 Formalism in General Relativity: Bases of Numerical Relativity*. (Springer, Heidelberg). First publ. 2007 as arXiv:[gr-qc/0703035](https://arxiv.org/abs/gr-qc/0703035). doi:[0.1007/978-3-642-24525-1](https://doi.org/10.1007/978-3-642-24525-1).
- Griffies, S. M. (2004): *Fundamentals of Ocean Climate Models*. (Princeton University Press, Princeton).

- Gronwald, F., Hehl, F. W. (1997): *Stress and hyperstress as fundamental concepts in continuum mechanics and in relativistic field theory*. arXiv:gr-qc/9701054.
- Hartle, J. B. (2003): *Gravity: An Introduction to Einstein's General Relativity*. (Addison-Wesley, San Francisco).
- Hawking, S. W., Ellis, G. F. R. (1994): *The Large Scale Structure of Space-Time*, repr. (Cambridge University Press, Cambridge).
- Hehl, F. W., McCrea, J. D. (1986): *Bianchi identities and the automatic conservation of energy-momentum and angular momentum in general-relativistic field theories*. *Found. Phys.* **16**³, 267–293. doi:10.1007/BF01889477.
- Hehl, F. W., Obukhov, Y. N. (2003): *Foundations of Classical Electrodynamics: Charge, Flux, and Metric*. (Birkhäuser, Boston). doi:10.1007/978-1-4612-0051-2.
- ISO (International Organization for Standardization) (2009): *ISO 80000-1:2009: Quantities and units 1: General*. International Organization for Standardization.
- Kanso, E., Arroyo, M., Tong, Y., Yavari, A., Marsden, J. G., Desbrun, M. (2007): *On the geometric character of stress in continuum mechanics*. *Z. Angew. Math. Phys.* **58**⁵, 843–856. <http://www.geometry.caltech.edu/pubs.html>, doi:10.1007/s00033-007-6141-8.
- Kilmister, C. W. (1973): *General Theory of Relativity*. (Pergamon, Oxford).
- Kitano, M. (2013): *Mathematical structure of unit systems*. *J. Math. Phys.* **54**⁵, 052901. doi:10.1063/1.4802876.
- Landau, L. D., Lifshitz [Lifšic], E. M. (1996): *The Classical Theory of Fields*, corr. repr. of the 4th English ed. (Butterworth-Heinemann, Oxford). Transl. from the 1987 seventh Russian edition by Morton Hamermesh. First publ. 1939.
- Lindell, I. V. (2004): *Differential Forms in Electromagnetics*. (IEEE Press and Wiley, Piscataway, USA). doi:10.1002/0471723096.
- Mari, L., Giordani, A. (2012): *Quantity and quantity value*. *Metrologia* **49**⁶, 756–764. doi:10.1088/0026-1394/49/6/756.
- Marsden, J. E., Hughes, T. J. R. (1994): *Mathematical Foundations of Elasticity*, unabridged corr. republ. (Dover, New York). <http://resolver.caltech.edu/CaltechB00K:1983.002>. First publ. 1983.
- Martin, D. (2004): *Manifold Theory: Introduction for Mathematical Physicists*, corr. repr. (Horwood, Chichester). First publ. 1991.
- Maugin, G. A. (1974): *Constitutive equations for heat conduction in general relativity*. *J. Phys. A* **7**⁴, 465–484. doi:10.1088/0305-4470/7/4/010.
- (1978a): *On the covariant equations of the relativistic electrodynamics of continua. I. General equations*. *J. Math. Phys.* **19**⁵, 1198–1205. doi:10.1063/1.523785.
- (1978b): *On the covariant equations of the relativistic electrodynamics of continua. II. Fluids*. *J. Math. Phys.* **19**⁵, 1206–1211. doi:10.1063/1.523786.
- (1978c): *On the covariant equations of the relativistic electrodynamics of continua. III. Elastic solids*. *J. Math. Phys.* **19**⁵, 1212–1219. doi:10.1063/1.523787.
- (1978d): *On the covariant equations of the relativistic electrodynamics of continua. IV. Media with spin*. *J. Math. Phys.* **19**⁵, 1220–1226. doi:10.1063/1.523788.
- McVittie, G. C. (1965): *General Relativity and Cosmology*, 2nd ed. (Chapman & Hall, London). First publ. 1956.
- Misner, C. W., Thorne, K. S., Wheeler, J. A. (1973): *Gravitation*, repr. (W. H. Freeman and Company, New York). First publ. 1970. https://archive.org/details/Gravitation_201803.

- Muschik, W., von Borzeszkowski, H.-H. (2014): *Exploitation of the dissipation inequality in general relativistic continuum thermodynamics*. Arch. Appl. Mech. **84**⁹, 1517–1531. doi:10.1007/s00419-014-0866-6.
- Pauli, W. (1958): *Theory of Relativity*. (Pergamon, London). Transl. by G. Field; with supplementary notes by W. Pauli. First publ. in German 1921.
- Penrose, R., Rindler, W. (2003): *Spinors and Space-Time. Vol. 1: Two-spinor calculus and relativistic fields*, corr. repr. (Cambridge University Press, Cambridge). First publ. 1984. doi:10.1017/CB09780511564048.
- Porta Mana, P. G. L. (2019a): *Answer to “MTW’s Gravitation. Factor of contraction of p-vector with p-form”*. <https://math.stackexchange.com/a/3208204/455507>.
- (2019b): *Affine and convex spaces: blending the analytic and geometric viewpoints*. Open Science Framework doi:10.31219/osf.io/jw9f6, arXiv:1104.0032. First publ. 2011.
- Rindler, W. (1986): *Essential Relativity: Special, General, and Cosmological*, rev. 2nd ed. (Springer, New York). First publ. 1969. doi:10.1007/978-1-4757-1135-6.
- (2006): *Relativity: Special, General, and Cosmological*, 2nd ed. (Oxford University Press, Oxford). First publ. 2001.
- Röhr, N., Uggla, C. (2005): *Conformal regularization of Einstein’s field equations*. Class. Quantum Grav. **22**¹⁷, 3775–3787. doi:10.1088/0264-9381/22/17/026.
- Samohýl, I. (1987): *Thermodynamics of Irreversible Processes in Fluid Mixtures (Approached by Rational Thermodynamics)*. (Teubner, Leipzig). Partially rewritten as Samohýl, Pekař (2014).
- Samohýl, I., Pekař, M. (2014): *The Thermodynamics of Linear Fluids and Fluid Mixtures*. (Springer, Cham). First published as Samohýl (1987).
- Schouten, J. A. (1954): *Ricci-Calculus: An Introduction to Tensor Analysis and Its Geometrical Applications*, 2nd ed. (Springer, Berlin). First publ. in German 1924. doi:10.1007/978-3-662-12927-2.
- (1989): *Tensor Analysis for Physicists*, corr. 2nd ed. (Dover, New York). First publ. 1951.
- Segev, R. (1986): *Forces and the existence of stresses in invariant continuum mechanics*. J. Math. Phys. **27**¹, 163–170. <http://www.bgu.ac.il/~rsegev/Papers/JMP86S.pdf>, doi:10.1063/1.527406.
- (2000a): *The geometry of Cauchy’s fluxes*. Arch. Rational Mech. Anal. **154**³, 183–198. <http://www.bgu.ac.il/~rsegev/Papers/TheGeometryofCauchyFluxes.pdf>, doi:10.1007/s002050000089.
- (2000b): *Notes on stresses for manifolds*. Rend. Sem. Mat. Univ. Pol. Torino **58**², 199–206.
- (2001): *A correction of an inconsistency in my paper “Cauchy’s theorem on manifolds”*. J. Elasticity **63**¹, 55–59. See Segev, Rodnay (1999). doi:10.1023/A:1013085312562.
- (2002): *Metric-independent analysis of the stress-energy tensor*. J. Math. Phys. **43**⁶, 3220–3231. <http://www.bgu.ac.il/~rsegev/Papers/MetricIndependent.pdf>, doi:10.1063/1.1475347.
- Segev, R., Rodnay, G. (1999): *Cauchy’s theorem on manifolds*. J. Elasticity **56**², 129–144. See also erratum Segev (2001). doi:10.1023/A:1007651917362.
- Smarr, L., Taubes, C., Wilson, J. R. (1980): *General relativistic hydrodynamics: the comoving, Eulerian, and velocity potential formalisms*. In: Tipler (1980): ch. 11:157–183.
- Smarr, L., York Jr., J. W. (1978): *Kinematical conditions in the construction of spacetime*. Phys. Rev. D **17**¹⁰, 2529–2551. doi:10.1103/PhysRevD.17.2529.
- Smarr, L. L., ed. (1979): *Sources of Gravitational Radiation*. (Cambridge University Press, Cambridge).
- Synge, J. L. (1960a): *Classical Dynamics*. In: Flügge (1960): I–VII, 1–225, 859–902.

- Synge, J. L. (1960b): *Relativity: The General Theory*. (North-Holland, Amsterdam).
- Tipler, F. J., ed. (1980): *Essays in General Relativity: A Festschrift for Abraham Taub*. (Academic Press, New York).
- Tolman, R. C. (1949): *Relativity, Thermodynamics and Cosmology*. (Oxford University Press). First publ. 1934.
- Trautman, A. (1964): *Foundations and current problems of general relativity*. In: Trautman, Pirani, Bondi (1964): 1–248. Notes by G. Dixon, P. Florides, G. Lemmer.
- Trautman, A., Pirani, F. A. E., Bondi, H. (1964): *Lectures on General Relativity*. (Prentice-Hall, Englewood Cliffs, USA). Notes by G. Dixon, P. Florides, G. Lemmer, J. J. Marek, J. Hyde; ed. by S. Deser and K. W. Ford.
- Truesdell III, C. A. (1991): *A First Course in Rational Continuum Mechanics. Vol. 1: General Concepts*, 2nd ed. (Academic Press, New York). First publ. 1977.
- Truesdell III, C. A., Toupin, R. A. (1960): *The Classical Field Theories*. In: Flügge (1960): I–VII, 226–902. With an appendix on invariants by Jerald LaVerne Ericksen. doi:[10.1007/978-3-642-45943-6_2](https://doi.org/10.1007/978-3-642-45943-6_2).
- Vallis, G. K. (2006): *Atmospheric and Oceanic Fluid Dynamics: Fundamentals and Large-scale Circulation*. (Cambridge University Press, Cambridge). doi:[10.1017/9781107588417](https://doi.org/10.1017/9781107588417).
- Wilson, J. R., Mathews, G. J. (2007): *Relativistic Numerical Hydrodynamics*, repr. (Cambridge University Press, Cambridge). First publ. 2003. doi:[10.1017/CB09780511615917](https://doi.org/10.1017/CB09780511615917).
- York Jr., J. W. (1979): *Kinematics and dynamics of general relativity*. In: Smarr (1979): 83–126.