Dimensional analysis in general relativity and differential geometry

P.G.L. Porta Mana <pgl@portamana.org>

13 December 2019; updated 11 January 2020

Some notes on dimensional analysis on differential manifolds, with an eye on general relativity and the Einstein equation.

Note: Dear Reader & Peer, this manuscript is being peer-reviewed by you. Thank you.

Dedicated to Emma

1 Introduction

From the point of view of dimensional analysis, do all components of a tensor need to have the same dimension? What is the dimension of the metric and the curvature? And what is the dimension of the constant in the Einstein equations?

There seem to be insecurity and incorrect notions, among some students and even some researchers in relativity, regarding the dimensions of tensors and of tensor components, the effect of tensor operators on dimensions, and the dimension of constants in field and differential-geometric equations. I've met, for example, with the statement that the components of a tensor should all have the same dimension; and with calculations of the dimensions of curvature tensors assuming that coordinates have dimensions of length. That statement is wrong; and that assumption is unnecessary.

Several factors probably cause or contribute to such difficulties. Modern texts in Lorentzian and general relativity commonly use geometrized units. They say that, for finding the dimension of some constant in a tensorial equation, it's sufficient to compare the dimensions of the tensors in the equation. But this is not so immediate, because some tensors don't have universally agreed dimensions – prime example the metric tensor. Older texts often use coordinates with dimension of length, and base their dimensional analysis on that specific choice. They even multiply coordinates with dimension of time, as well as some tensorial components, by powers of c; thus giving the impression that coordinates

ought to always be lengths, and that all components of a tensor ought to have the same dimension¹.

The main purpose of this Note is to illustrate an extremely simple and intuitive way of reasoning with which one can quickly and consistently settle any dimensional-analysis questions in general relativity and differential geometry; for example the dimension of the Riemann tensor, or the dimensional result of contraction, covariant derivative, raising an index. This way of reasoning is presented and illustrated in §§ 4–5. It's just an application of the usual rules of dimensional analysis: if two quantities are summed, then they must have the same dimension; the dimension of a product is the product of the dimensions; and so on. This way of reasoning relies on the coordinate-free, *intrinsic* view of tensors and other differential-geometrical objects, a brief reminder of which is given in § 2, with references.

Another important purpose of this Note is to revive the forgotten notion of *absolute dimension* of a tensor, and the distinction between it and the dimensions of that tensor's *components* in some coordinate system. This notion is explained in § 4. The absolute dimension, introduced by Schouten and Dorgelo² and used in Truesdell et al.³, is invariant with respect to the dimensions of the coordinates; it's an intrinsic property of a tensor. The dimensions of the components depend instead on the dimensions of the coordinate functions – which can be completely arbitrary, as discussed in § 3. The absolute dimension should thus be the primary focus in dimensional analysis in general relativity and in differential geometry.

The rest of this Note is mainly an application of the intuitive way of reasoning mentioned above and of the notion of absolute dimension. We'll reveal and rectify a couple of misconceptions; for example, the components of a tensor need not have the same dimensions. The results for the main tensor operations and operators are summarized in § 5. The results for the curvature tensors and other objects depending on a connection are presented in § 7. It is shown, in particular, that the absolute dimension of the (contravariant and thrice covariant) Riemann and (fully covariant) Ricci tensors is 1, that is, they are dimensionless. Some geometric

 $^{^1}$ e.g. Tolman 1949 § 37 eq. (37.8); Landau et al. 1996 § 32 eq. (32.15); Adler et al. 1975 § 10.1 eq. (10.15). 2 Dorgelo et al. 1946; Schouten 1989 ch. VI. 3 Truesdell et al. 1960 Appendix II.

objects or operators may be unfamiliar to researchers working in general relativity only; their discussion may simply be skipped by these readers.

The absolute dimension of the metric tensor is discussed in \S 8. The literature on general relativity presents two standard choices for it. The absolute dimension of the stress-energy-momentum tensor is discussed in \S 9. The dimension of the constant in the Einstein equation is derived in \S 10, and the two standard results from the literature are recovered.

The final § 11 gives a summary of the main results, with some additional comments.

Since the type of a tensor – that is, the number and ordering of its covariant and contravariant 'slots'⁴ – is important in dimensional analysis, I often denote it explicitly with a notation like A." to indicate that A is covariant in its first slot and contravariant in its second and third slots. Its components would thus be $(A_i^{\ jk})$. For brevity I'll call this a 'co-contra-contravariant' tensor, with an obvious naming generalization for other types.

I denote the contraction of the α th and β th slots (which must have opposite variant types) of a tensor \boldsymbol{A} by $\operatorname{tr}_{\alpha\beta}\boldsymbol{A}$, and the transposition (swapping) of the α th and β th slots with $\boldsymbol{A}^{\mathsf{T}_{\alpha\beta}}$. In index notation these operations are

(There doesn't seem to be an invariant notation for these two operations, so I propose this notation here.)

For the notation in dimensional analysis I use iso conventions: $\dim(A)$ is the dimension of the quantity A, and among the base quantities are mass M, length L, time T, temperature Θ , electric current I. Note that I don't discuss units – it doesn't matter here whether the unit for length is the metre or the centimetre, for example.

Finally, quoting Truesdell and Toupin⁵, 'dimensional analysis remains a controversial and somewhat obscure subject. We do not attempt a complete presentation here.'

⁴ Misner et al. 1973 § 3.2. ⁵ Truesdell et al. 1960 Appendix § 7 footnote 4.

2 Intrinsic view of differential-geometric objects: brief reminder

From the intrinsic point of view, a tensor is defined by its geometric properties. For example, a vector field $v \equiv v(\cdot)$ is an object that operates on functions defined on the (spacetime) manifold, yielding new functions, with the properties v(af+bg) = av(f)+bv(g) and v(fg) = v(f)g+fv(g) for all functions f, g and reals a, b. A covector field (1-form) \boldsymbol{w} is an object that operates on vector fields, yielding functions ('duality'), with the property $\boldsymbol{w}(fu+gv) = f\boldsymbol{w}(u)+g\boldsymbol{w}(v)$ for all vector fields u, v and functions f, g. The sum of vector or covector fields, and their products by functions – let's call this 'linearity' – are defined in an obvious way. Tensors are constructed from these objects.

A system of coordinates (x^i) is just a set of linearly independent functions. This set gives rise to a set of vectors fields $\left(\frac{\partial}{\partial x^i}\right)$ and to a set of covector fields (dx^i) by the obvious requirements that $\frac{\partial}{\partial x^i}(x^j) = \delta_i^j$ and $dx^i\left(\frac{\partial}{\partial x^j}\right) = \delta_j^i$. These two sets can be used as bases to express all other vectors and covectors as linear combinations. A vector field v can thus be written as

$$v \equiv \sum_{i} v^{i} \frac{\partial}{\partial x^{i}} \equiv v^{i} \frac{\partial}{\partial x^{i}}, \tag{1}$$

where the *functions* $v^i := v(x^i)$ are its components with respect to the basis $\left(\frac{\partial}{\partial x^i}\right)$. Analogously for a covector field.

For the presentation of the intrinsic view I recommend the excellent texts by Choquet-Bruhat et al. (1996), Boothby (2003), Abraham et al. (1988), Bossavit (1991), Burke (1987; 1980 ch. 2), and more on the general-relativity side Misner et al. (1973 ch. 9), Gourgoulhon (2012 ch. 2), Penrose et al. (2003 ch. 4).

3 Coordinates

From a physical point of view, a coordinate is just a function that associates a value of a physical quantity with every event in a region (the domain of the coordinate chart) of spacetime. Together with the other coordinates, such function allows us to uniquely identify every event within that region. Any physical quantity will do: the distance from something, the time elapsed since something, an angle, an energy density, the strength of a magnetic flux, a temperature, and so on. A coordinate

can thus have any dimensions: length L, time T, angle 1, energy density $E := ML^{-1}T^{-2}$, magnetic flux $\Phi := ML^{2}T^{-2}I^{-1}$, temperature Θ , and so on.

The functional relation between two sets of coordinates must of course be dimensionally consistent. For example, if $\dim(x^0) = T$ and $\dim(x^1) = L$, and we introduce a coordinate $\xi(x^0, x^1)$ with dimension 1, additive in the previous two, then we must have $\xi = ax^0 + bx^1$ with $\dim(a) = T^{-1}$ and $\dim(b) = L^{-1}$.

4 Tensors: absolute dimension and components' dimensions

Consider a system of coordinates (x^i) with dimensions (X_i) , and the ensuing sets of covector fields (1-form) dx^i and of vector fields $\left(\frac{\partial}{\partial x^i}\right)$, bases for the cotangent and tangent spaces. Their tensor products are bases for the tangent spaces of higher tensor types.

The differential $\mathrm{d}x^i$ traditionally has the same dimension as x^i : $\mathrm{dim}(\mathrm{d}x^i) = \mathsf{X}_i$, and the operator (a vector) $\frac{\partial}{\partial x^i}$ traditionally has the inverse dimension: $\mathrm{dim}\,\frac{\partial}{\partial x^i} = \mathsf{X}_i^{-1}$. We'll see later that these conventions are self-consistent.

For our discussion let's take a concrete example: a contra-covariant tensor field $A \equiv A^*$. The discussion generalizes to tensors of other types in an obvious way.

The tensor \boldsymbol{A} can be expanded in terms of the basis vectors and covectors, as mentioned in § 2:

$$\mathbf{A} = A^{i}_{j} \frac{\partial}{\partial x^{i}} \otimes dx^{j} \equiv A^{0}_{0} \frac{\partial}{\partial x^{0}} \otimes dx^{0} + A^{0}_{1} \frac{\partial}{\partial x^{0}} \otimes dx^{1} + \cdots$$
 (2)

Each function

$$A^{i}_{j} := \mathbf{A} \left(\mathrm{d} x^{i}, \frac{\partial}{\partial x^{j}} \right) \tag{3}$$

is a *component* of the tensor in this coordinate system.

To make dimensional sense, all terms in the sum (2) must have the same dimension. This is possible only if the generic component $A^i_{\ j}$ has dimension

$$\dim(A_{j}^{i}) = A X_{i} X_{j}^{-1}, \tag{4}$$

where A is *common* to all components. In fact, the $X_i X_j^{-1}$ term cancels the $X_i^{-1} X_j$ term coming from $\frac{\partial}{\partial x^i} \otimes dx^j$ in the sum (2), and each summand therefore has dimension A.

For example, if we're using coordinates with dimensions

$$\dim(x^0) = \Theta$$
, $\dim(x^1) = L$, $\dim(x^2) = L$, $\dim(x^3) = ML^{-1}T^{-2}$, (5)

then the components of \boldsymbol{A} have dimensions

$$\left(\dim(A_{j}^{i})\right) = A \times \begin{pmatrix} 1 & L^{-1}\Theta & L^{-1}\Theta & M^{-1}LT^{2}\Theta \\ L\Theta^{-1} & 1 & 1 & M^{-1}L^{2}T^{2} \\ L\Theta^{-1} & 1 & 1 & M^{-1}L^{2}T^{2} \\ ML^{-1}T^{-2}\Theta^{-1} & ML^{-2}T^{-2} & ML^{-2}T^{-2} & 1 \end{pmatrix}. (6)$$

The dimension A, which is also the final dimension of the sum (2), is called the *absolute dimension*⁶ of the tensor A, and we write

$$\dim(\mathbf{A}) = A. \tag{7}$$

This is the intrinsic dimension of the tensor, independent of any coordinate system. It reflects the physical or operational 7 meaning of the tensor. We'll see an example of what this mean in § 8 with the metric tensor.

Different coordinate systems lead to different dimensions of the *components* of a tensor A, but the absolute dimension of the tensor remains the same. Formula (4) for the dimensions of the components is consistent under changes of coordinates. For example, in new coordinates (x'^k) with dimensions (X'_k) , the new components of A are

$$A^{\prime k}_{l} = A^{i}_{j} \frac{\partial x^{\prime k}}{\partial x^{i}} \frac{\partial x^{j}}{\partial x^{\prime l}}$$
 (8)

and a quick check shows that $\dim(A'_l^k) = AX'_kX'_l^{-1}$, consistent with the general formula (4).

In the following I'll drop the adjective 'absolute' when it's clear from the context.

5 Tensor operations

By the reasoning of the previous section, which simply applies standard dimensional considerations to the basis expansion (2), it's easy to find

Dorgelo et al. 1946; Schouten 1989 ch. VI.
 Bridgman 1958; see also Synge 1960a
 A.2; Truesdell et al. 1960 §§ A.3–4.

out the resultant absolute dimension of various operations and operators on tensors and tensor fields.

The tensor product of A^{\bullet} and $B_{\bullet\bullet}$, for example, can be written as the sum

$$\mathbf{A} \otimes \mathbf{B} = A^{i}_{j} B_{kl}^{m} \frac{\partial}{\partial x^{i}} \otimes \mathrm{d}x^{j} \otimes \mathrm{d}x^{k} \otimes \mathrm{d}x^{l} \otimes \frac{\partial}{\partial x^{m}}$$
(9)

from which it follows that

$$\dim(A_{i}^{i} B_{kl}^{m}) = A B X_{i} X_{j}^{-1} X_{k}^{-1} X_{l}^{-1} X_{m}$$
(10)

with $A = \dim(\mathbf{A})$ and $B = \dim(\mathbf{B})$. The absolute dimension of $\mathbf{A} \otimes \mathbf{B}$ is therefore $AB \equiv \dim(\mathbf{A}) \dim(\mathbf{B})$.

Here is a summary of the dimensional results for the main differential-geometric operations and operators, except for the covariant derivative, the metric, and related tensors, discussed more in depth in §§ 7–8 below. I invite you to prove them for yourself. For reference, in brackets I give the section of Choquet-Bruhat et al. (1996) where the operations are defined.

• *Tensor multiplication* [III.B.5] multiplies dimensions:

$$\dim(\mathbf{A} \otimes \mathbf{B}) = \dim(\mathbf{A})\dim(\mathbf{B}). \tag{11}$$

• The *contraction* [III.B.5] of the α th and β th slots of a tensor has the same dimension as the tensor:

$$\dim(\operatorname{tr}_{\alpha\beta}\mathbf{A}) = \dim(\mathbf{A}). \tag{12}$$

Note that the formula above only holds *without* raising or lowering indices; see § 8 for those operations.

• The *transposition* of the α th and β th slots of a tensor has the same dimension as the tensor:

$$\dim(\mathbf{A}^{\mathsf{T}\alpha\beta}) = \dim(\mathbf{A}). \tag{13}$$

• The *Lie bracket* [III.B.3] of two vectors has the product of their dimensions:

$$\dim([u,v]) = \dim(u)\dim(v). \tag{14}$$

• The *Lie derivative* [III.C.2] of a tensor with respect to a vector field has the product of the dimensions of the tensor and of the vector:

$$\dim(\mathbf{L}_{v}\mathbf{A}) = \dim(v)\dim(\mathbf{A}). \tag{15}$$

Regarding operations and operators on differential forms:

• The *exterior product* [IV.A.1] of two differential forms multiplies their dimensions:

$$\dim(\boldsymbol{\omega} \wedge \boldsymbol{\tau}) = \dim(\boldsymbol{\omega}) \dim(\boldsymbol{\tau}). \tag{16}$$

• The *interior product* [IV.A.4] of a vector and a form multiplies their dimensions:

$$\dim(i_v \omega) = \dim(v) \dim(\omega). \tag{17}$$

• The *exterior derivative* [IV.A.2] of a form has the same dimension of the form:

$$\dim(\mathrm{d}\boldsymbol{\omega}) = \dim(\boldsymbol{\omega}). \tag{18}$$

This can be proven using the identity $d i_v + i_v d = L_v$ or similar identities⁸ together with eqs (15) and (17).

• The *integral* [IV.B.1] of a form over a submanifold has the same dimension as the form:

$$\dim(\int_{C} \boldsymbol{\omega}) = \dim(\boldsymbol{\omega}). \tag{19}$$

6 Curves and integral curves

Consider a curve into spacetime, $c: s \mapsto P(s)$, with the parameter s having dimension dim(s) = S.

If we consider the events of the spacetime manifold as dimensionless quantities, then the dimension of the tangent or velocity vector \dot{C} to the curve is

$$\dim(\dot{C}) = S^{-1},\tag{20}$$

owing to the definition9

$$\dot{C} \coloneqq \frac{\partial x^i [C(s)]}{\partial s} \frac{\partial}{\partial x^i}.$$
 (21)

This has a quirky but interesting consequence. Given a vector field \boldsymbol{v} we say that \boldsymbol{C} is an integral curve for it if

$$v = \dot{C} \tag{22}$$

(or more precisely $v_{C(s)} = \dot{C}_{C(s)}$ in usual differential-geometric notation¹⁰) at all events C(s) in the image of the curve. From the point of view of

 $^{^{\}mathbf{8}}$ Curtis et al. 1985 ch. 9 p. 180 Theorem 9.78; Abraham et al. 1988 § 6.4 Theorem 6.4.8.

 $^{^9}$ Choquet-Bruhat et al. 1996 § III.B.1; Boothby 2003 § IV.(1.9). 10 Choquet-Bruhat et al. 1996 § III.B.1.

dimensional analysis this definition can only be valid if v has dimension S^{-1} . If v and s^{-1} have different dimensions – a case which can happen for physical reasons – the condition (21) must be modified into $v = k\dot{C}$, where k is a possibly dimensionful constant. This is equivalent to considering an affine and dimensional reparameterization of C.

7 Connection, covariant derivative, curvature tensors

Consider an arbitrary connection¹¹ with covariant derivative ∇ . For the moment we don't assume the presence of any metric structure.

The covariant derivative of the product fv of a function and a vector satisfies 12

$$\nabla(fv) = \mathrm{d}f \otimes v + f \nabla v. \tag{23}$$

The first summand, from formulae (18) and (11), has dimension $\dim(f)\dim(v)$; for dimensional consistency this must also be the dimension of the second summand. Thus

$$\dim(\nabla v) = \dim(v). \tag{24}$$

It follows that the *directional* covariant derivative ∇_u has dimension

$$\dim(\nabla_u v) = \dim(u)\dim(v), \tag{25}$$

and by its derivation properties¹³ we see that formula (24) extends from vectors to tensors of arbitrary type.

In the coordinate system (x^i) , the action of the covariant derivative is carried by the *connection coefficients* or Christoffel symbols (Γ^i_{jk}) defined by

$$\nabla \frac{\partial}{\partial x^k} = \Gamma^i_{jk} \, \mathrm{d} x^j \otimes \frac{\partial}{\partial x^i}. \tag{26}$$

From this equation and the previous ones it follows that these coefficients have dimensions

$$\dim(\Gamma_{jk}^{i}) = X_{i} X_{j}^{-1} X_{k}^{-1}.$$
(27)

¹¹ Choquet-Bruhat et al. 1996 § V.B. 12 Choquet-Bruhat et al. 1996 § V.B.1.

¹³ Choquet-Bruhat et al. 1996 § V.B.1 p. 303.

The torsion τ ..., Riemann curvature R..., and Ricci curvature Ric.. tensors are defined by 14

$$\tau(u,v) \coloneqq \nabla_u v - \nabla_v u - [u,v], \tag{28}$$

$$\mathbf{R}(u, v; w) := \nabla_u \nabla_v w - \nabla_v \nabla_u w - \nabla_{[u,v]} w, \tag{29}$$

$$Ric... := tr_{13} R^{\bullet}....$$
 (30)

From these definitions and the results of $\S\,5$ we find the dimensional requirements

$$\dim(\tau^*..)\dim(u)\dim(v) = \dim(u)\dim(v), \tag{31}$$

$$\dim(\mathbf{R}^{\bullet}_{...})\dim(\mathbf{u})\dim(\mathbf{v})\dim(\mathbf{w}) = \dim(\mathbf{u})\dim(\mathbf{v})\dim(\mathbf{w}), \quad (32)$$

$$\dim(\mathbf{Ric..}) = \dim(\mathbf{R}^{\bullet}...), \tag{33}$$

which imply that torsion, Riemann curvature, and Ricci curvature are dimensionless:

$$\dim(\boldsymbol{\tau}^{\bullet}...) = \dim(\boldsymbol{R}^{\bullet}...) = \dim(\boldsymbol{R}^{i}\boldsymbol{c}...) = 1. \tag{34}$$

The exact contra- and covariant type used above for these tensors is very important in these equations. If we raise any of their indices using a metric, their dimensions will generally change.

Misner et al. (1973 pp. 35, 407) say that 'curvature', the Riemann or Einstein tensors (for which see § 8 below), has dimension L^{-2} , a statement seemingly at variance with the dimensionless results (34). But I believe that they refer to the *components* of those tensors, in specific coordinates of dimension L, and using geometrized units. In such specific coordinates every *component* R^i_{jkl} does indeed have dimension L^{-2} , according to the general formula (4), if and only if the *absolute* dimension of R is unity, $\dim(R) = 1$. So I believe that there's no real contradiction with Mister et al.'s statement and the results (34). This possible misunderstanding shows that it's important to distinguish between absolute dimensions, which don't depend on any specific coordinate choice, and component dimensions, which do.

The formulae above are also valid if a metric is defined and the connection is compatible with it. The connection coefficients in this case are defined in terms of the metric tensor, but using the results of \S 8 it's easy to see that eqs (24), (25), (27), (34) still hold.

 $^{^{\}mathbf{14}}$ Choquet-Bruhat et al. 1996 § V.B.1.

8 Metric and related tensors and operations

Let's now consider a metric tensor $g_{...}$. What is its absolute dimension $\dim(g)$? There seem to be two choices in the literature; both can be derived from the operational meaning of the metric.

Consider a (timelike) worldline $s \mapsto C(s)$, $s \in [a, b]$, between events C(a) and C(b). The metric tells us the *proper time* Δt elapsed for an observer having that worldline, according to the formula

$$\Delta t = \int_{a}^{b} \mathrm{d}s \, \sqrt{|\mathbf{g}[\dot{C}(s), \dot{C}(s)]|}. \tag{35}$$

From the results of § 5 this formula implies that $T \equiv \dim(\Delta t) = \sqrt{\dim(\mathbf{g_{..}})}$, (independently of the dimension of s) and therefore

$$\dim(\mathbf{g}_{\cdot\cdot\cdot}) = \mathsf{T}^2. \tag{36}$$

Many authors¹⁵, however, prefer to include a dimensional factor 1/c in front of the integral (35):

$$\Delta t = \frac{1}{c} \int_{a}^{b} \mathrm{d}s \, \sqrt{\left| \mathbf{g}[\dot{C}(s), \dot{C}(s)] \right|},\tag{37}$$

thus obtaining

$$\dim(\mathbf{g}_{\cdot \cdot}) = L^2. \tag{38}$$

The choice (38) seems also supported by the traditional expression for the 'line element ds^2 ' as it appears in many works (for an exception, with dimension T^2 , see Kilmister¹⁶),

$$ds^{2} = c^{2}dt^{2} - dx^{2} - dy^{2} - dz^{2},$$
(39)

possibly with opposite signature. If the coordinates (t, x, y, z) have the dimensions suggested by their symbols, this formula has dimension L^2 , so that if we interpret ' ds^2 ' as \boldsymbol{g} we find $\dim(\boldsymbol{g}) = L^2$. The line-element expression above often has an ambiguous differential-geometric meaning, however, because it may represent the metric applied to some unspecified vector, that is, $\boldsymbol{g}(v,v)$, where v is left unspecified v. In this case we have

$$L^2 = \dim(\mathbf{g})\dim(\mathbf{v})^2 \tag{40}$$

and the dimension of \boldsymbol{g} is ambiguous.

 $^{^{15}}$ e.g. Fock 1964 § V.62, eq. (62.02); Curtis et al. 1985 ch. 11 eq. (11.21); Rindler 1986 § 5.3 eq. (5.6); Hartle 2003 ch. 6 eq. (6.24). 16 Kilmister 1973 ch. II p. 25. 17 cf. Misner et al. 1973 Box 3.2 D, p. 77.

The standard choices for dim \mathbf{g} are thus T^2 or L^2 . My favourite choice is the first, (36), for reasons discussed by Synge and Bressan¹⁸. Synge gives a vivid summary:¹⁹

We are now launched on the task of giving physical meaning to the Riemannian geometry [...]. It is indeed a Riemannian *chronometry* rather than *geometry*, and the word *geometry*, with its dangerous suggestion that we should go about measuring *lengths* with *yardsticks*, might well be abandoned altogether in the present connection

In fact, to measure the proper time Δt defined above we only need to ensure that a clock has the worldline C, and then take the difference between the clock's final and initial times. On the other hand, consider the case when the curve C is *spacelike*. Its proper length is still defined by the integral (35) (apart from a dimensional constant). Its measurement, however, is more involved than the timelike case. It requires dividing the curve into very short pieces, and having specially-chosen observers (orthogonal to the pieces) measure each piece. But the measurement of each piece actually relies on the measurement of *proper time*: each observer uses 'radar distance' sending a lightlike signal which bounces back at the end of the piece and measuring the time it takes to come back. Even if rigid rods are used, their calibration still relies on a measurement of time – this is also reflected in the current definition of the standard metre²¹.

The metric \mathbf{g} can be considered as an operator mapping vectors to covectors, which we can compactly write as $\mathbf{\omega} = \mathbf{g} \mathbf{v}$ (instead of the cumbersome $\mathbf{\omega} = \operatorname{tr}_{23}(\mathbf{g} \otimes \mathbf{v})$). The *inverse metric tensor* \mathbf{g}^{-1} is then defined by the formula

$$\mathbf{g}^{-1}\mathbf{g} = \mathrm{id}^{\bullet}$$
, $\mathbf{g}\mathbf{g}^{-1} = \mathrm{id}^{\bullet}$, (41)

so that, obviously,

$$\dim(\mathbf{g}^{-1}) = \dim(\mathbf{g})^{-1}. \tag{42}$$

The *metric volume element*²² in spacetime is a 4-form γ , equivalent to a completely antisymmetric tensor γ, such that $\gamma(e_0, e_1, e_2, e_3) = 1$ for every set of positively-oriented orthonormal vector fields (e_k) , that is, such that $\mathbf{g}(e_k, e_l) = \pm \delta_{kl}$ (remember that the orientation is not

¹⁸ Synge 1960b §§ III.2–4; Bressan 1978 §§ 15, 18. ¹⁹ Synge 1960b § III.3 pp. 108–109.

²⁰ Landau et al. 1996 § 84. ²¹ вірм 1983 р. 98; Giacomo 1984 р. 25. ²² Abraham et al. 1988 § 6.2.

determined by the metric). It has only one non-zero component, given by the square root of the determinant of the (positively ordered) components (g_{ij}) of the metric:

$$\gamma = \sqrt{|\det(g_{ij})|} \, \mathrm{d}x^0 \wedge \mathrm{d}x^1 \wedge \mathrm{d}x^2 \wedge \mathrm{d}x^3. \tag{43}$$

From this expression and the results of § 5 it can be shown that, in spacetime,

$$\dim(\boldsymbol{\gamma}) = \dim(\boldsymbol{g})^2 \equiv \begin{cases} \mathsf{T}^4 & \text{if } \dim(\boldsymbol{g}) \coloneqq \begin{cases} \mathsf{T}^2 \\ \mathsf{L}^2 \end{cases} . \tag{44}$$

(This is also the dimension of the *density* $|\gamma|$, which, as opposed to the volume element, has the property that $|\gamma|(e_0, e_1, e_2, e_3) = 1$ for all sets of orthonormal vector fields, not only positively-oriented ones.)

The operation of *raising or lowering an index* of a tensor represents a contraction of the tensor product of that tensor with the metric or the metric inverse, for example $A_{\cdot \cdot} \equiv \operatorname{tr}_{13} A^{\cdot} \cdot \otimes g_{\cdot \cdot}$ and similarly for tensors of other types. Therefore

$$\dim(\mathbf{A}_{\dots \cdot \dots}) = \dim(\mathbf{A}_{\dots \cdot \dots}) \dim(\mathbf{g}), \quad \dim(\mathbf{A}_{\dots \cdot \dots}) = \dim(\mathbf{A}_{\dots \cdot \dots}) \dim(\mathbf{g})^{-1}.$$
(45)

The formulae for the covariant derivative (24), connection coefficients (27), and curvature tensors (34) remain valid for a connection compatible with the metric. In this case the connection coefficients can be obtained from the metric by the formulae²³

$$\Gamma^{i}_{jk} = \frac{1}{2} \left(\frac{\partial}{\partial x^{k}} g_{jl} + \frac{\partial}{\partial x^{j}} g_{kl} - \frac{\partial}{\partial x^{l}} g_{jk} \right) g^{li}, \tag{46}$$

and it's easily verified that the dimensions of these coefficients (27) still holds. Also the result for the curvature tensors (34) still holds, since their expressions in terms of the connection coefficients is the same with or without a metric.

The scalar curvature ρ and the Einstein tensor **G**:

$$\rho := \operatorname{tr} \operatorname{Ric} : \equiv \operatorname{tr}_{23}(\operatorname{Ric} \otimes \operatorname{g}^{-1}), \qquad \operatorname{G} : = \operatorname{Ric} : -\frac{1}{2}\rho \text{ id}. \tag{47}$$

have therefore dimension

$$\dim(\rho) = \dim(\mathbf{G}.^{\bullet}) = \dim(\mathbf{g})^{-1} \equiv \begin{cases} \mathsf{T}^{-2} & \text{if } \dim(\mathbf{g}) := \mathsf{T}^{2} \\ \mathsf{L}^{2} & \text{.} \end{cases}$$
 (48)

²³ Choquet-Bruhat et al. 1996 § V.B.2.

9 Stress-energy-momentum tensor

To find the dimension of the stress-energy-momentum T, or '4-stress' for short, let's start with the analysis of the (3-)stress σ in Newtonian mechanics. The stress σ is the projection of the 4-stress T onto a spacelike tangent plane with respect to some observer²⁴. If we assume that such spatial projection preserves the absolute dimension, then the 4-stress and the stress have the same absolute dimension.

In Newtonian mechanics the stress σ is an object that, integrated over the boundary of a body, gives the total surface force acting on the body (such integration requires a flat connection). This means that it must be represented by a 'force-valued' 2-form. Force, in turn, can be interpreted as an object that, integrated over a (spacelike) trajectory, gives an energy – the work done by the force along the trajectory. It's therefore a 1-form. Putting these two requirements together, the stress turns out to be a covector-valued 2-form, equivalent to a tensor $\sigma_{\cdot\cdot\cdot\cdot}$ antisymmetric in its last two indices. Integrated over a surface, and then over a trajectory, it yields an energy. From § 5, integration of a form does not change the dimension of the form. Therefore

$$\dim(\sigma_{\bullet \bullet \bullet}) = E \equiv ML^2T^{-2}. \tag{49}$$

But usually the stress is represented by a co-contravariant tensor σ . This is obtained by contracting the last two slots of σ ... with the inverse of the volume element of the 3-metric – this is the duality²⁵ between k-vectors and (n-k)-covectors induced by the metric (and an orientation choice), where n is the geometric dimension of the manifold. If we assume the Newtonian 3-metric to have dimension L^2 , it can be shown similarly to § 8 that its volume element has dimension L^3 , and the inverse volume element has dimension L^{-3} . Thus we obtain

$$\dim(\boldsymbol{\sigma}_{\cdot}) = EL^{-3} \equiv ML^{-1}T^{-2}, \tag{50}$$

an energy density (or 'volumic energy' according to ISO²⁶).

 $^{^{24}}$ Gourgoulhon 2012 \S 3.4.1; Smarr et al. 1978; York 1979; Smarr et al. 1980; Wilson et al. 2007 \S 1.3; the projection doesn't need to be orthogonal: Marsden et al. 1994 \S 2.4; Hehl et al. 2003 \S B.1.4. 25 Bossavit 1991 \S 4.1.2. 26 iso 2009 item A.6.2.

Since the stress σ : is the projection of T: and the projection preserves the absolute dimension, we finally find that T: also has the dimension of an energy density:

$$\dim(\mathbf{T}_{\bullet}) = EL^{-3} \equiv ML^{-1}T^{-2}.$$
 (51)

Note that other co- or contravariant versions of the 4-stress will have different absolute dimension, because they're obtained by lowering or raising indices. For example, $\dim(T_{\cdot\cdot}) = \dim(T_{\cdot\cdot}) \dim(g) = \mathrm{ML}^{-1}$ if $\dim(g) := \mathrm{T}^2$.

Let me add a passing remark. Even though in most texts the 4-stress is represented by a tensor of order 2, as above, its most fitting geometrical nature is still shrouded by mystery from a kinematic and dynamical point of view. There are indications that it could be more properly represented by a covector-valued 3-form (equivalent to a tensor $T_{\cdot \cdot \cdot \cdot \cdot}$ antisymmetric in the last three slots), or by a 3-vector-valued 3-form (equivalent to a tensor $T^{\cdot \cdot \cdot \cdot \cdot}$ antisymmetric in the first three and last three slots), for reasons connected with integration, similar to those mentioned above for the stress $\sigma_{\cdot \cdot \cdot \cdot}$ See for example the discussion about '* T' by Misner et al. (1973 ch. 15), the works by Segev (2002; 1986; 1999; 2000a,b), the discussion by Burke (1987 § 41).

10 Einstein equation and Einstein's constant

We finally arrive at the Einstein equation,

$$G = \kappa T \tag{52}$$

where κ (sometimes seen with a minus²⁷ depending on the signature of the metric or on how the orientation of the stress is chosen) is Einstein's constant. For the dimension of κ we thus find

$$\dim(\kappa) = \dim(\mathbf{G}_{\cdot}) \dim(\mathbf{T}_{\cdot})^{-1} \equiv \begin{cases} M^{-1}L & \text{if } \dim(\mathbf{g}) := \begin{cases} T^2 \\ L^2 \end{cases}.$$
(53)

This constant can be obtained from the dimensions of Newton's gravitational constant $\dim(G) = M^{-1}L^3T^{-2}$ (this is not the Einstein tensor G!)

 $^{^{27}}$ e.g. Tolman 1949 § 78 eq. (78.3); Fock 1964 § 52 eq. (52.06); Rindler 2006 § 14.2 eq. (14.8).

and of the speed of light $dim(c) = LT^{-1}$ only in the following ways, with an 8π factor coming from the Newtonian limit:

$$\kappa = \begin{cases} 8\pi G/c^2 \\ 8\pi G/c^4 \end{cases} \text{ if } \dim(\mathbf{g}) \coloneqq \begin{cases} \mathsf{T}^2 \\ \mathsf{L}^2 \end{cases} . \tag{54}$$

The second choice is by far the most common, consistently with the most common choice of $dim(\mathbf{g}) = L^2$ discussed before. The first choice appears for example in Fock²⁸ and Adler et al.²⁹.

11 Summary and conclusions

We have seen that dimensional analysis in general relativity and differential geometry can be seamlessly done, following its usual standard rules, if we adopt a coordinate-free or coordinate-invariant view, typical of more modern texts. In this view, each tensor has an absolute dimension that doesn't depend on the dimensions of the coordinates. It would therefore more profitable to focus on the absolute dimension of a tensor rather than the dimensions of its components.

We found some important results for general relativity, in particular that the Riemann R^* ... and Ricci Ric.. tensors are dimensionless, and that the Einstein tensor G. has the inverse dimension of the metric.

Since the dimensions of the components are usually different from the absolute dimension and depend on the coordinates, I recommend to avoid statements such as 'the tensor $A_i^{\ jk}$ has dimension X', because it leaves unclear whether ' $A_i^{\ jk}$ ' is meant to represent the tensor in general, or to represent its set of components, or to represent just a specific component.

Finally, the above discussion specific to general relativity could of course be made along the same lines also for Newtonian mechanics.

Thanks

... to Mari, Miri, Emma for continuous encouragement and affection, and to Buster Keaton and Saitama for filling life with awe and inspiration. To the developers and maintainers of LaTeX, Emacs, AUCTeX, Open

²⁸ Fock 1964 § 55 eqs (55.15) and (52.06).
²⁹ Adler et al. 1975 § 10.5 eq. (10.98).

Science Framework, R, Python, Inkscape, Sci-Hub for making a free and impartial scientific exchange possible.

Bibliography

- ('de X' is listed under D, 'van X' under V, and so on, regardless of national conventions.)
- Abraham, R., Marsden, J. E., Ratiu, T. (1988): *Manifolds, Tensor Analysis, and Applications*, 2nd ed. (Springer, New York). First publ. 1983.
- Adler, R., Bazin, M., Schiffer, M. (1975): *Introduction to General Relativity*, 2nd ed. (McGraw-Hill, Tokyo). First publ. 1965.
- BIPM (Bureau international des poids et mesures), ed. (1983): 17^e conférence générale des poids et mesures (CGPM). (BIPM, Sèvres). https://www.bipm.org/en/worldwide-metrology/cgpm/resolutions.html#comptes-rendus.
- Boothby, W. M. (2003): An Introduction to Differentiable Manifolds and Riemannian Geometry, rev. 2nd ed. (Academic Press, Orlando, USA). First publ. 1975.
- Bossavit, A. (1991): Differential Geometry: for the student of numerical methods in Electromagnetism. https://www.researchgate.net/publication/200018385_Differential_Geometry for the student of numerical methods in Electromagnetism.
- Bressan, A. (1978): Relativistic Theories of Materials. (Springer, Berlin).
- Bridgman, P. W. (1958): *The Logic of Modern Physics*, eight printing. (Macmillan, New York). First publ. 1927.
- Burke, W. L. (1980): *Spacetime, Geometry, Cosmology*. (University Science Books, Mill Valley, USA).
- (1987): Applied Differential Geometry, reprint. (Cambridge University Press, Cambridge).
 First publ. 1985.
- Choquet-Bruhat, Y., DeWitt-Morette, C., Dillard-Bleick, M. (1996): *Analysis, Manifolds and Physics. Part I: Basics*, rev. ed. (Elsevier, Amsterdam). First publ. 1977.
- Curtis, W. D., Miller, F. R. (1985): *Differential Manifolds and Theoretical Physics*. (Academic Press, Orlando, USA).
- Dorgelo, H. B., Schouten, J. A. (1946): *On unities and dimensions. I.* Verh. Kon. Akad. Wetensch. Amsterdam **49**^{2, 3, 4}, 123–131, 282–291, 393–403.
- Flügge, S., ed. (1960): Handbuch der Physik: Band III/1: Prinzipien der klassischen Mechanik und Feldtheorie [Encyclopedia of Physics: Vol. III/1: Principles of Classical Mechanics and Field Theory]. (Springer, Berlin).
- Fock [Fok], V. A. (1964): *The Theory of Space, Time and Gravitation,* second revised ed. (Pergamon, Oxford). Transl. by N. Kemmer. First publ. in Russian 1955.
- Giacomo, P. (1984): News from the bipm. Metrologia **20**¹, 25–30.
- Gourgoulhon, É. (2012): 3+1 Formalism in General Relativity: Bases of Numerical Relativity. (Springer, Heidelberg). First publ. 2007 as arXiv:gr-qc/0703035.
- Hartle, J. B. (2003): Gravity: An Introduction to Einstein's General Relativity. (Addison-Wesley, San Francisco).
- Hehl, F. W., Obukhov, Y. N. (2003): Foundations of Classical Electrodynamics: Charge, Flux, and Metric. (Birkhäuser, Boston).
- iso (International Organization for Standardization) (2009): *ISO 80000-1:2009: Quantities and units 1: General.* International Organization for Standardization.
- Kilmister, C. W. (1973): General Theory of Relativity. (Pergamon, Oxford).

- Landau, L. D., Lifshitz [Lifsic], E. M. (1996): The Classical Theory of Fields, corr. repr. of the 4th English ed. (Butterworth-Heinemann, Oxford). Transl. from the 1987 seventh Russian edition by Morton Hamermesh. First publ. 1939.
- Marsden, J. E., Hughes, T. J. R. (1994): *Mathematical Foundations of Elasticity*, unabridged corr. republ. (Dover, New York). http://resolver.caltech.edu/CaltechBOOK:1983.002. First publ. 1983.
- Misner, C. W., Thorne, K. S., Wheeler, J. A. (1973): *Gravitation*, repr. (W. H. Freeman and Company, New York). https://archive.org/details/Gravitation 201803.
- Penrose, R., Rindler, W. (2003): Spinors and Space-Time. Vol. 1: Two-spinor calculus and relativistic fields, corr. repr. (Cambridge University Press, Cambridge). First publ. 1984.
- Rindler, W. (1986): Essential Relativity: Special, General, and Cosmological, rev. second ed. (Springer, New York). First publ. 1969.
- (2006): Relativity: Special, General, and Cosmological, 2nd ed. (Oxford University Press, Oxford). First publ. 2001.
- Schouten, J. A. (1989): *Tensor Analysis for Physicists*, corr. second ed. (Dover, New York). First publ. 1951.
- Segev, R. (1986): Forces and the existence of stresses in invariant continuum mechanics. J. Math. Phys. 27¹, 163–170. http://www.bgu.ac.il/~rsegev/Papers/JMP86S.pdf.
- (2000a): The geometry of Cauchy's fluxes. Arch. Rational Mech. Anal. 154³, 183–198. htt p://www.bgu.ac.il/~rsegev/Papers/TheGeometryofCauchyFluxes.pdf.
- (2000b): Notes on stresses for manifolds. Rend. Sem. Mat. Univ. Pol. Torino 58², 199–206.
- (2001): A correction of an inconsistency in my paper "Cauchy's theorem on manifolds". J. Elasticity 63¹, 55–59. Ssee Segev, Rodnay (1999).
- (2002): Metric-independent analysis of the stress-energy tensor. J. Math. Phys. 43⁶, 3220–3231. http://www.bgu.ac.il/~rsegev/Papers/MetricIndependent.pdf.
- Segev, R., Rodnay, G. (1999): Cauchy's theorem on manifolds. J. Elasticity 56², 129–144. h ttp://www.bgu.ac.il/~rsegev/Papers/CauchyTheoremonManifolds.pdf. See also erratum Segev (2001).
- Smarr, L. L., ed. (1979): Sources of Gravitational Radiation. (Cambridge University Press, Cambridge).
- Smarr, L., Taubes, C., Wilson, J. R. (1980): General relativistic hydrodynamics: the comoving, Eulerian, and velocity potential formalisms. In: Tipler (1980), ch. 11, 157–183.
- Smarr, L., York Jr., J. W. (1978): *Kinematical conditions in the construction of spacetime*. Phys. Rev. D 17¹⁰, 2529–2551.
- Synge, J. L. (1960a): Classical Dynamics. In: Flügge (1960), I-VII, 1-225, 859-902.
- (1960b): Relativity: The General Theory. (North-Holland, Amsterdam).
- Tipler, F. J., ed. (1980): Essays in General Relativity: A Festschrift for Abraham Taub. (Academic Press, New York).
- Tolman, R. C. (1949): *Relativity, Thermodynamics and Cosmology*. (Oxford University Press). First publ. 1934.
- Truesdell III, C. A., Toupin, R. A. (1960): *The Classical Field Theories*. In: Flügge (1960), I–VII, 226–902. With an appendix on invariants by Jerald LaVerne Ericksen.
- Whitney, H. (1968a): The mathematics of physical quantities. Part I: Mathematical models for measurement. Am. Math. Monthly 75², 115–138. See also Whitney (1968b).
- (1968b): *The mathematics of physical quantities. Part II: Quantity structures and dimensional analysis.* Am. Math. Monthly 75³, 227–256. See also Whitney (1968a).
- Wilson, J. R., Mathews, G. J. (2007): *Relativistic Numerical Hydrodynamics*, reprint. (Cambridge University Press, Cambridge). First publ. 2003.

York Jr., J. W. (1979): Kinematics and dynamics of general relativity. In: Smarr (1979), 83–126.