Name: SN:

Useful Equations:

Beer's law for absorption and scattering (no emission) given slant path $s = dz/\mu$, volume extinction coefficient b, gas density ρ_g (Stull chapter 2)

$$d\omega = \sin\theta d\theta d\phi = -d\mu d\phi \tag{1}$$

$$S = S_0 \cos(\theta) \tag{2}$$

$$S = S_0 \cos(v) \tag{2}$$

$$a_{\lambda} + r_{\lambda} + t_{\lambda} = 1 \tag{3}$$

$$L_{\lambda} = \frac{\Delta E}{\Delta \omega \Delta \lambda} \tag{4}$$

$$E_{\lambda BB} = \pi B_{\lambda} \tag{5}$$

where B_{λ} ($W m^{-2} sr^{-1} \mu m^{-1}$) is the monochromatic blackbody radiance from an infinite flat surface.

$$d\tau = n \, b \, dz \tag{12}$$

$$d\tau = \rho_a k \, dz \tag{13}$$

$$ds = dz/\mu \tag{14}$$

$$\tau = \int_{z_1}^{z_2} \rho_g k dz \tag{15}$$

$$\frac{dL_{\lambda}}{L_{\lambda}} = -\frac{d\tau}{\mu} \tag{16}$$

Hydrostatic equation:

$$dp = -\rho_d g dz \tag{17}$$

$$E_{BB} = \sigma T^4$$
 (6) Equation of state

$$\frac{dT}{dt} = \frac{-1}{\rho c_{nd}} \frac{dE_n}{dz} \tag{7}$$

$$p = R_d \rho_d T \tag{18}$$

Slant Transmission

$$E_{\lambda}(T) \approx E_{\lambda 0} + \left. \frac{dE_{\lambda}}{dT} \right|_{T_0, \lambda} (T - T_0) + \dots$$
 (8)

$$\exp(x) = 1 + x + \frac{x^2}{2} + \frac{3^2}{3!} + \dots$$
 (9)

$$\sin(\theta) = \theta - \frac{\theta^3}{3!} + \frac{\theta^5}{5!} + \dots \tag{10}$$

$$\cos(\theta) = 1 - \frac{\theta^2}{2!} + \frac{\theta^4}{4!} + \dots \tag{11}$$

$$\hat{t} = \exp(-\tau/\mu) \tag{19}$$

Schwarzchild's equation

$$dL = -L d\tau/\mu + B_{\lambda}(z)d\tau/\mu \qquad (20)$$

(11) Vertical radiance
$$L_{\lambda}$$
 with $\mu = 1$

$$L_{\lambda}(\tau) = B_{\lambda}(T_{skin})(\exp(-\tau) + \int_{0}^{\tau} \exp(-(\tau - \tau')) B_{\lambda}(T) d\tau'$$
 (21)

Radiance along a slant path:

$$L_{\lambda}(\tau/\mu) = B_{\lambda}(T_{skin})(\exp(-\tau/\mu) + \int_{0}^{\tau} \exp(-(\tau - \tau')/\mu) B_{\lambda}(T) d\tau'/\mu$$
(22)

Flux from radiance:

$$E_{\lambda} = \int_0^{2\pi} \int_0^1 \mu L_{\lambda} \, d\mu \, d\phi \tag{23}$$

Vertical radiance with change of variables to transmittance:

$$L_{\lambda}(\tau/\mu) = B_{\lambda}(T_{skin})\hat{t}_{tot} + \int_{0}^{\tau} B_{\lambda}(T)d\hat{t} \quad (24)$$

For constant temperature:

$$L_{\lambda}(\tau/\mu) = B_{\lambda}(T_{skin})\hat{t}_{tot} + (1 - \exp(-\tau/\mu))B_{\lambda}(T)$$
 (25)

Integrating over a hemisphere for flux, using the simple flux diffusivity approximation:

$$E_{\lambda} = \pi B_{\lambda}(T_{skin}) \exp(-1.66\tau_{tot}) + \pi B_{\lambda}(T)(1 - \exp(-1.66\tau))$$
 (26)

and integrated over all wavelengths:

$$E_{\uparrow} = \sigma T_{skin}^4 \exp(-1.66\overline{\tau}_{\lambda T}) + \sigma T_{layer}^4 (1 - \exp(-1.66\overline{\tau}_{\lambda T}))$$
 (27)

Radar

Rayleigh scattering

$$\frac{L}{L_0} \propto \frac{D^6}{\lambda^4} \tag{28}$$

Returned power
$$\propto |k^2| \frac{Z}{r^2}$$
 (29)

(23) where $|k^2| \approx 0.2$ for ice and 0.9 for liquid water.

$$Z = \int_0^\infty n(D)D^6 dD \tag{30}$$

Z-R relationship

$$Z = 300RR^{1.4} \tag{31}$$

(rain, RR in mm/hr, Z in mm^6/m^3)

$$MUR = \frac{c}{2 \cdot PRF} \tag{32}$$

Useful constants:

$$c_{pd} = 1004 \text{ J kg}^{-1} \text{ K}^{-1},$$

 $\sigma = 5.67 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}$
 $k_b = 1.381 \times 10^{-23} \text{ J K}^{-1}$
 $c = 3 \times 10^8 \text{ m s}^{-1}$
 $h = 6.626 \times 10^{-34} \text{ J s}$
 $\pi \approx 3.14159$
 $R_d = 287 \text{ J kg}^{-1} \text{ K}^{-1}$