

# 2D gravity inversion with isostatic constraint applied to passive rifted margins

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**GEO-XXXX**

Running head: **2D gravity inversion for passive rifted margins**

## ABSTRACT

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## INTRODUCTION

Several methods have been proposed for using gravity and/or magnetic data to estimate the relief of basement under sedimentary basins and/or the Moho.

Here, we consider methods that represent the basement and/or Moho by interfaces separating layers with constant physical properties (density and/or magnetization).

In the case of basement, the interface separates a sedimentary package from the upper crust. In the case of Moho, these two layers are the lower crust and upper mantle.

These methods usually presume that an interface representing the basement or Moho oscillates around a known reference depth.

These methods can be divided into two groups.

The first one consists of methods estimating the geometry of a single interface separating two layers.

(Bott, 1960; Tanner, 1967; Cordell and Henderson, 1968; Dyrelius and Vogel, 1972; Pedersen, 1977; Richardson and MacInnes, 1989; Barbosa et al., 1997, 1999b,a; Silva et al., 2006; Chakravarthi and Sundararajan, 2007; Martins et al., 2010; Silva et al., 2010; Camacho et al., 2011; Lima et al., 2011; Martins et al., 2011; Barnes and Barraud, 2012; Silva et al., 2014; Santos et al., 2015; Silva and Santos, 2017)

(Oldenburg, 1974; Granser, 1987; Reamer and Ferguson, 1989; Guspí, 1993; Braitenberg et al., 1997; Braitenberg and Zadro, 1999).

Uses tesserooids (Uieda and Barbosa, 2017).

Uses the Vening Meinesz-Moritz approach (Bagherbandi and Eshagh, 2012)

(Sjöberg, 2009)

It is well known that, by using different combinations of reference depths and physical property values, it is possible to find different interfaces producing the same potential-field data.

To deal with this inherent ambiguity (Roy, 1962; Skeels, 1947) and obtain meaningful solutions, the interpreter must commonly use a priori information obtained from seismic data and/or boreholes in order to constrain the range of possible models.

The methods can be divided, according to their scale, into three groups.

Group 1: local scale, basement or Moho?, grav, mag or joint? space or Fourier domain?

Methods in Fourier domain use Parker's formula (Parker, 1973).

Satellite data (Reguzzoni et al., 2013).

(Sampietro, 2015) - pegar as refs citadas no trabalho "GOCE data to infer the Moho depth or the density contrast between crust and upper mantle at a regional/continental scale with extremely high accuracy (Braitenberg et al. 2010, 2011; Mariani et al. 2013; Van der Meijde et al. 2013; Sampietro et al. 2014; Bouman et al. 2015; Van der Meijde et al. 2015)"

(Salem et al., 2014)

Magnetic data (Pilkington and Crossley, 1986a,b)

Joint inversion of gravity and magnetic data (Gallardo-Delgado et al., 2003; Gallardo et al., 2005; Pilkington, 2006)

Two interfaces separating three layers with constant physical property (density and/or

magnetization).

## METHODOLOGY

### Forward problem

Let  $\mathbf{d}^o$  be the observed data vector, whose  $i$ -th element  $d_i^o$ ,  $i = 1, \dots, N$ , represent the observed gravity disturbance at the point  $(x_i, y_i, z_i)$ , on a profile located over a rifted passive margin. The coordinates are referred to a topocentric Cartesian system, with  $z$  axis pointing down,  $y$ -axis along the profile and  $x$ -axis perpendicular to the profile. We assume that the observed gravity disturbance is produced by an anomalous mass distribution defined as the difference between the actual mass distribution in the subsurface, which is schematically represented in Figure ??, and a reference mass distribution (Figure ??). In doing it, we implicitly assume that Figure ?? represents the outer layers of a global mass distribution producing the normal gravity field.

The anomalous mass distribution producing the observed data is approximated by an interpretation model (Figure ??) formed by  $N$  adjacent columns. For convenience, we presume that the observed data are regularly spaced, so that there is one observation at the centre of the top of each column forming the interpretation model. The  $i$ -th column is formed by four vertically adjacent layers, , which in turn are composed of vertically adjacent prisms having infinite length along the  $x$ -axis. The first and shallowest layer represents the water layer, is formed by a single prism, has thickness  $t_i^w$  and a constant density contrast  $\Delta\rho^w = \rho^w - \rho^r$ , where  $\rho^w$  and  $\rho^r$  represents, respectively, the densities of water and the reference mass distribution (Figure ??) at the same point. The third layer represents the crust, it is also formed by a single prism, has thickness  $t_i^c$  and density contrast  $\Delta\rho_i^c = \rho^c - \rho^r$ ,

with  $\rho^c$  being the crust density. For simplicity, we presume that the crust density  $\rho_i^c$  may be equal to  $\rho^{cc}$ , for  $y_i \leq y_{COT}$ , which represents continental crust, or equal to  $\rho^{oc}$ , for  $y_i > y_{COT}$ , which represents oceanic crust. The crust density depends on the position of the  $i$ -th column with respect to  $y_{COT}$ , which defines an abrupt Crust-Ocean Transition (COT). Consequently, the crust may have two possible density contrasts:  $\Delta\rho_i^c = \rho^{cc} - \rho^r$  or  $\Delta\rho_i^c = \rho^{oc} - \rho^r$ . The top of this layer defines the basement relief and its bottom the relief of the Moho. The fourth layer represents the mantle, it is divided into two parts, each one formed by a single prism having the same density  $\rho^m$  and, consequently, the same density contrast  $\Delta\rho^m = \rho^m - \rho^r$ . The shallowest portion of this layer has thickness  $t_i^m$ . Its top and bottom define, respectively, the depths of Moho and the planar isostatic compensation layer  $S_0$ . The deepest portion of the fourth layer has thickness  $\Delta S_0$ , top at the surface  $S_0$  and bottom at the planar surface  $S_0 + \Delta S_0$ , which defines the Moho in the reference mass distribution model (Figure ??). Finally, the second layer forming the  $t$ -th column of the interpretation model is defined by the interpreter, according to the geological environment to be studied and the a priori information availability. As a general rule, this layer can be defined by a set of  $Q$  vertically adjacent prisms, each one with thickness  $t_i^q$ , density  $\rho^q$  and density contrast  $\Delta\rho^q = \rho^q - \rho^r$ ,  $q = 1, \dots, Q$ .

Given the density contrasts, the COT position  $y_{COT}$ , the isostatic compensation surface  $S_0$ , the thickness of the water layer and of the  $Q - 1$  prisms forming the shallowest portion of the second layer, it is possible to describe the interpretation model in terms of an  $M \times 1$  parameter vector  $\mathbf{p}$ ,  $M = 2N + 1$ , defined as follows:

$$\mathbf{p} = \begin{bmatrix} \mathbf{t}^Q \\ \mathbf{t}^m \\ \Delta S_0 \end{bmatrix}, \quad (1)$$

where  $\mathbf{t}^Q$  and  $\mathbf{t}^m$  are  $N \times 1$  vectors whose  $i$ -th elements  $t_i^Q$  and  $t_i^m$  represent, respectively, the thickness of the prism forming the deepest portion of the second layer and the thickness of the prism forming the shallowest portion of the fourth layer of the interpretation model. In this case, the gravity disturbance produced by the interpretation model (the predicted gravity disturbance) at the position  $(x_i, y_i, z_i)$  can be written as the sum of the vertical component of the gravitational attraction exerted by the  $L$  prisms forming the interpretation model as follows:

$$d_i(\mathbf{p}) = k_g G \sum_{j=1}^L f_{ij}(\mathbf{p}) , \quad (2)$$

where  $f_{ij}(\mathbf{p})$  represents an integral over the volume of the  $j$ -th prism. Here, these volume integrals are computed with the expressions proposed by Nagy et al. (2000), by using the open-source Python package *Fatiando a Terra* (Uieda et al., 2013).

## Inverse problem

Let  $\mathbf{d}(\mathbf{p})$  be the predicted data vector, whose  $i$ -th element  $d_i(\mathbf{p})$  is defined by Equation 2. Estimating the particular parameter vector  $\mathbf{p} = \hat{\mathbf{p}}$  producing a predicted data vector  $\mathbf{d}(\mathbf{p})$  as close as possible to the observed data vector  $\mathbf{d}^o$  can be formulated as the problem of minimizing the goal function

$$\Gamma(\mathbf{p}) = \Phi(\mathbf{p}) + \mu \sum_{k=0}^3 \alpha_k \Psi_k(\mathbf{p}) , \quad (3)$$

subject to all elements of  $\hat{\mathbf{p}}$  be positive. In Equation 3,  $\mu$  represents the regularizing parameter,  $\Phi(\mathbf{p})$  represents the misfit function given by

$$\Phi(\mathbf{p}) = \frac{1}{N} \|\mathbf{d}^o - \mathbf{d}(\mathbf{p})\|_2^2 , \quad (4)$$

where  $\|\cdot\|_2^2$  represents the squared Euclidean norm,  $\alpha_k$  represent the weights assigned to the regularizing functions  $\Psi_k(\mathbf{p})$ , with define the constraints on the parameters to be estimated,

$k = 0, 1, 2, 3$ .

## Airy constraint

Consider that the interpretation model is in isostatic equilibrium according to the Airy model (Turcotte and Schubert, 2002; Hofmann-Wellenhof and Moritz, 2005; Lowrie, 2007).

In this case, the pressure (or lithostatic stress) exerted by the model is constant on the isostatic compensation surface  $S_0$ . The pressure per unit area exerted by the  $i$ -th column of the model on  $S_0$ , divided by gravity, is given by:

$$t_i^w \rho^w + t_i^1 \rho_i^1 + \dots + t_i^Q \rho_i^Q + t_i^c \rho_i^c + t_i^m \rho^m = \sigma_0, \quad (5)$$

where  $\sigma_0$  is an arbitrary positive constant. Rearranging terms in Equation 5 and using the relation

$$S_0 = t_i^w + t_i^1 + \dots + t_i^Q + t_i^c + t_i^m, \quad (6)$$

it is possible to show that:

$$(\rho_i^Q - \rho_i^c) t_i^Q + (\rho^m - \rho_i^c) t_i^m + (\rho^w - \rho_i^c) t_i^w + (\rho_i^1 - \rho_i^c) t_i^1 + \dots + (\rho_i^{Q-1} - \rho_i^c) t_i^{Q-1} + \rho_i^c S_0 = \sigma_0. \quad (7)$$

In order to describe the pressure exerted by all columns forming the interpretation model on the surface  $S_0$ , Equation 7 can be written, in matrix notation, as follows:

$$\mathbf{M}^Q \mathbf{t}^Q + \mathbf{M}^m \mathbf{t}^m + \mathbf{M}^w \mathbf{t}^w + \mathbf{M}^1 \mathbf{t}^1 + \dots + \mathbf{M}^{Q-1} \mathbf{t}^{Q-1} + \boldsymbol{\rho}^c S_0 = \sigma_0 \mathbf{1}, \quad (8)$$

where  $\mathbf{1}$  is an  $N \times 1$  vector with all elements equal to one,  $\mathbf{t}^\alpha$  are  $N \times 1$  vectors with  $i$ -th element defined by the thickness  $t_i^\alpha$  of a prism forming the  $i$ -th column,  $\alpha = w, 1, \dots, Q-1, Q, m$ , and  $\mathbf{M}^Q, \mathbf{M}^m, \mathbf{M}^w, \mathbf{M}^1, \dots, \mathbf{M}^{Q-1}$  are  $N \times N$  diagonal matrices with elements  $ii$  of main diagonal are given by density contrasts  $(\rho_i^Q - \rho_i^c), (\rho^m - \rho_i^c), (\rho^w - \rho_i^c), (\rho_i^1 - \rho_i^c)$  and  $\dots, (\rho_i^{Q-1} - \rho_i^c)$ , respectively, and  $\boldsymbol{\rho}^c$  is an  $N \times 1$  vector containing the densities of the

prisms representing the crust. By applying the first-order Tikhonov regularization (Aster et al., 2005) to the constant vector  $\sigma_0 \mathbf{1}$ , we obtain the following expression:

$$\mathbf{R}(\mathbf{C}\mathbf{p} + \mathbf{D}\mathbf{t}) = \mathbf{0}, \quad (9)$$

where  $\mathbf{0}$  is a vector with null elements and the remaining terms are given by:

$$\mathbf{C} = \begin{bmatrix} \mathbf{M}^Q & \mathbf{M}^m & \mathbf{0} \end{bmatrix}_{N \times M}, \quad (10)$$

$$\mathbf{D} = \begin{bmatrix} \mathbf{M}^w & \mathbf{M}^1 & \dots & \mathbf{M}^{Q-1} & \boldsymbol{\rho}^c \end{bmatrix}_{N \times (QN+1)}, \quad (11)$$

$$\mathbf{t} = \begin{bmatrix} \mathbf{t}^w \\ \mathbf{t}^1 \\ \vdots \\ \mathbf{t}^{Q-1} \\ S_0 \end{bmatrix}_{(QN+1) \times 1}, \quad (12)$$

$\mathbf{p}$  is the parameter vector (Equation 1) and  $\mathbf{R}$  is an  $(N-1) \times N$  matrix, whose element  $ij$  is defined as follows:

$$[\mathbf{R}]_{ij} = \begin{cases} 1 & , \quad j = i \\ -1 & , \quad j = i + 1 \\ 0 & , \quad \text{otherwise} \end{cases}. \quad (13)$$

Finally, from Equation 9, it is possible to define the regularizing function  $\Psi_0(\mathbf{p})$  (Equation 3):

$$\Psi_0(\mathbf{p}) = \|\mathbf{R}(\mathbf{C}\mathbf{p} + \mathbf{D}\mathbf{t})\|_2^2. \quad (14)$$

We call this function as *Airy constraint*. Notice that minimizing this function imposes smoothness on the pressure exerted by the interpretation model on the isostatic compensation surface  $S_0$ .



## Smoothness constraint

This constraint imposes smoothness on the adjacent thickness of the prisms forming the deepest portion of the second layer and the shallowest part of the fourth layer of the interpretation model by applying the first-order Tikhonov regularization (Aster et al., 2005) to the vectors  $\mathbf{t}^Q$  and  $\mathbf{t}^m$  (Equation 1). Mathematically, this constraint is represented by the regularizing function  $\Psi_1(\mathbf{p})$  (Equation 3):

$$\Psi_1(\mathbf{p}) = \|\mathbf{S}\mathbf{p}\|_2^2, \quad (15)$$

where  $\mathbf{S}$  is an  $(N - 1) \times M$  matrix given by:

$$\mathbf{S} = \begin{bmatrix} \mathbf{R} & \mathbf{R} & \mathbf{0} \end{bmatrix}, \quad (16)$$

where  $\mathbf{R}$  is defined by Equation 13 and  $\mathbf{0}$  is a vector with all elements equal to zero.

## Equality constraint

### *Equality constraint on basement depths*

Let  $\mathbf{a}$  be a vector whose  $k$ -th element  $a_k$ ,  $k = 1, \dots, A$ , is the known basement depth at the horizontal coordinate  $y_k^A$  of the profile. These known basement depth values are used to define the regularizing function  $\Psi_2(\mathbf{p})$  (Equation 3):

$$\Psi_2(\mathbf{p}) = \|\mathbf{A}\mathbf{p} - \mathbf{a}\|_2^2, \quad (17)$$

where  $\mathbf{A}$  is an  $A \times M$  matrix whose  $k$ -th line has one element equal to one and all the remaining elements equal to zero. The location of the single non-null element in the  $k$ -th line of  $\mathbf{A}$  depends on the coordinate  $y_k^A$  of the known basement depth  $a_k$ . Let us consider, for example, an interpretation model formed by  $N = 10$  columns. Consider also that the

basement depth at the coordinates  $y_1^A = y_4$  and  $y_2^A = y_9$  of the profile are equal to 25 and 35.7 km, respectively. In this case,  $A = 2$ ,  $\mathbf{a}$  is a  $2 \times 1$  vector with elements  $a_1 = 25$  and  $a_2 = 35.7$  and  $\mathbf{A}$  is a  $2 \times M$  matrix ( $M = 2N + 1 = 21$ ). The element 4 of the first line and the element 9 of the second line of  $\mathbf{A}$  are equal to 1 and all its remaining elements are equal to zero.

#### *Equality constraint on Moho depths*

Let  $\mathbf{b}$  be a vector whose  $k$ -th element  $b_k$ ,  $k = 1, \dots, B$ , is the difference between the isostatic compensation depth  $S_0$  and the known Moho depth at the horizontal coordinate  $y_k^B$  of the profile. These differences, which must be positive, are used to define the regularizing function  $\Psi_3(\mathbf{p})$  (Equation 3):

$$\Psi_3(\mathbf{p}) = \|\mathbf{B}\mathbf{p} - \mathbf{b}\|_2^2, \quad (18)$$

where  $\mathbf{B}$  is a  $B \times M$  matrix whose  $k$ -th line has one element equal to one and all the remaining elements equal to zero. This matrix is defined in the same way as matrix  $\mathbf{A}$  (Equation 17).

## CONCLUSIONS

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## ACKNOWLEDGMENTS

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