

Low Reynolds number gravitational settling of a sphere through a fluid-fluid interface: Modelling using a boundary integral method

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Abstract

1 Introduction

2 Theoretical Development

The problem being modelled is the low Reynolds number gravitational settling of a sphere towards an initially horizontal interface separating two density stratified, immiscible, semi-infinite fluids (figure 1). The fluids are characterised by the velocity $\mathbf{u}_l(\mathbf{x})$ and pressure

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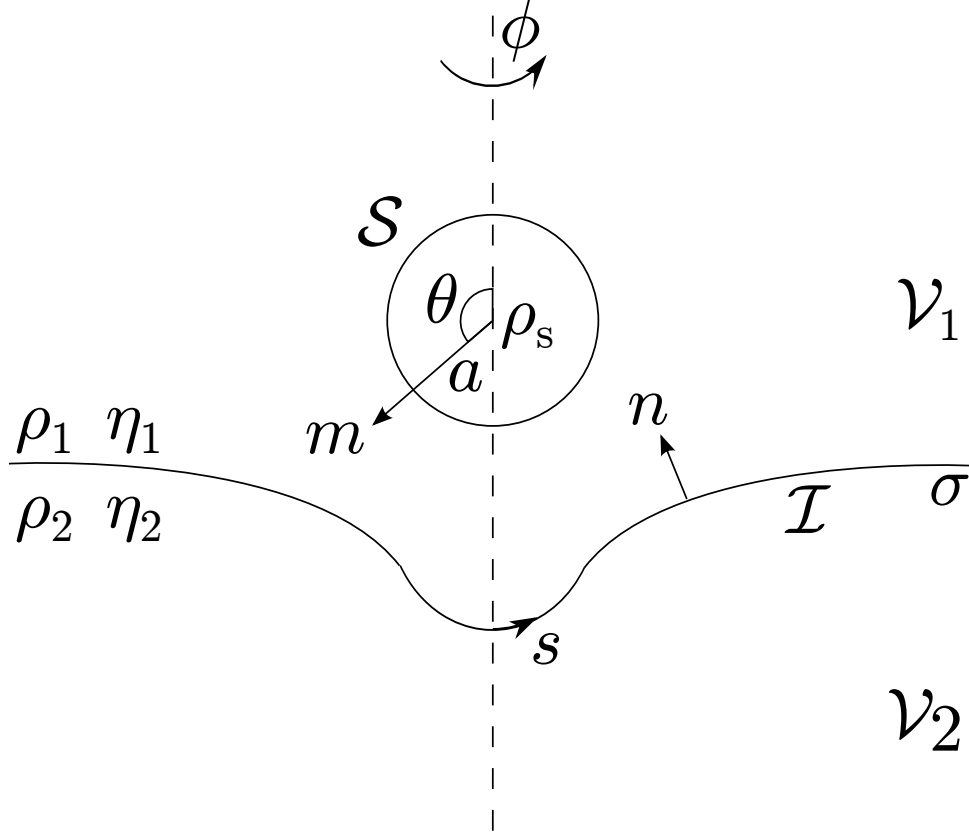


Figure 1: Diagrammatic representation of the system. A sphere falls under gravity, at low Reynolds number, towards an initially horizontal interface between two density stratified, immiscible semi-infinite fluids. See table 1 for definition of symbols.

$p_l(\mathbf{x})$ fields where $l = 1, 2$ denotes the fluid and \mathbf{x} is a position vector. The dynamic pressure is defined as

$$p_{d,l}(\mathbf{x}) = p_l(\mathbf{x}) - \rho_l \mathbf{g} \cdot \mathbf{x}, \quad (1)$$

where \mathbf{g} is acceleration due to gravity.

Table 1: Definition of symbols.

Symbol	Definition
a	Sphere radius
$\mathbf{g} = (-9.81 \text{ m s}^{-2}) \hat{\mathbf{z}}$	Acceleration due to gravity
\mathcal{I}	Surface of interface

$l = 1, 2$	Fluid label
m	Outward normal to sphere surface
n	Normal to interface (points into fluid 1)
$p_l(\mathbf{x})$	Pressure field of fluid l
$p_{d,l}(\mathbf{x})$	Dynamic pressure of fluid l
s	Arc length along interface measured from axis
\mathcal{S}	Surface of sphere
$\mathbf{u}_l(\mathbf{x})$	Velocity field of fluid l
\mathcal{V}_l	Volume of fluid l
\mathbf{x}	Position vector
$\hat{\mathbf{z}}$	Unit vector in the upward vertical direction
η_l	Viscosity of fluid l
θ	Polar angle with respect to sphere centre
ρ_l	Density of fluid l
ρ_s	Sphere density
σ	Interfacial Tension
ϕ	Azimuhtal angle with respect to axis of motion