

## Z Background Formulae

Paper - HIG-13-030, PASs: HIG-13-013, HIG-13-018,  
HIG-13-028

D. Colling, G. Davies P. Dunne, A.M. Magnan, A. Nikitenko  
and Joao Pela: Imperial with  
R. Aggleton, J. Brooke: Bristol  
C.Asawangtrakuldee, Q.Li: Peking  
P. Srimanobhas: Chulalongkorn  
S. Kumar, K. Mazumdar: Mumbai

## Definitions

$$N_S(Z \rightarrow \nu\nu) = (N_C^{Data} - N_C^{Bkg}) .$$

$$\frac{\sigma(Z \rightarrow \nu\nu, \text{EWK}) \frac{N_S^{MC}(\text{EWK})}{N_{\text{Gen}}(60 < m_Z^{\text{Gen}} < 120 \text{ GeV}, \text{EWK})} + \sigma(Z \rightarrow \nu\nu, \text{QCD}) \frac{N_S^{MC}(\text{QCD})}{N_{\text{Gen}}(60 < m_Z^{\text{Gen}} < 120 \text{ GeV}, \text{QCD})}}{\sigma(Z/\gamma^* \rightarrow \mu\mu, \text{EWK}) \frac{N_C^{MC}(\text{EWK})}{N_{\text{Gen}}(\text{EWK})} + \sigma(Z/\gamma^* \rightarrow \mu\mu, \text{QCD}) \frac{N_C^{MC}(\text{QCD})}{N_{\text{Gen}}(\text{QCD})}}$$

### Regions

- ▶ Control Region: Reco dimuon, with  $60 < m_{\mu\mu}^{\text{reco}} < 120 \text{ GeV}$  passing VBF selections
  - $N_C^{MC}$  is measured in  $Z \rightarrow \mu\mu + \text{Jets}$  MC with a generator level cut of  $m_Z^{\text{Gen}} > 50 \text{ GeV}$
- ▶ Signal Region: VBF selections and no veto leptons
  - We use the same  $Z \rightarrow \mu\mu$  sample as for  $N_C^{MC}$  and ignore the leptons to approximate a  $Z \rightarrow \nu\nu$  sample, this will be denoted  $N_S^{MC}$
  - For the efficiencies to be the same for the  $Z \rightarrow \mu\mu$  and  $Z \rightarrow \nu\nu$  samples a generator level mass window of  $60 < m_Z^{\text{Gen}} < 120$  must be applied, this will be denoted  $N_S^{MC}[60, 120]$ .

## Derivation of formula

- Basic formula for data driven estimate

$$N_S^{\nu\nu Data} = \frac{N_C^{Data} - N_C^{BKG}}{N_C^{MC}} \cdot N_S^{\nu\nu MC}$$

- To use  $Z \rightarrow \mu\mu$  MC we use the formula:

$$N_S^{\nu\nu MC} = N_S^{MC}[60, 120] \cdot \underbrace{\frac{\sigma(Z \rightarrow \nu\nu)}{\sigma(Z/\gamma^* \rightarrow \mu\mu, 60 < m_Z^{Gen} < 120\text{GeV})}}_{R[60,120]}$$

- The cross-section ratio that we have calculated is:

$$R[50, \infty] = \frac{\sigma(Z \rightarrow \nu\nu)}{\sigma(Z/\gamma^* \rightarrow \mu\mu, m_Z^{Gen} > 50)}$$

- We therefore use:

$$\begin{aligned} R[60, 120] &= \frac{\sigma(Z/\gamma^* \rightarrow \mu\mu, m_Z^{Gen} > 50\text{GeV})}{\sigma(Z/\gamma^* \rightarrow \mu\mu, 60 < m_Z^{Gen} < 120\text{GeV})} \cdot R[50, \infty] \\ &= \frac{N(Z/\gamma^* \rightarrow \mu\mu, m_Z^{Gen} > 50\text{GeV})}{N(Z/\gamma^* \rightarrow \mu\mu, 60 < m_Z^{Gen} < 120\text{GeV})} \cdot R[50, \infty] \end{aligned}$$

## Derivation of formula (2)

- Substituting our expression for  $N_S^{\nu\nu MC}$  into the original formula gives:

$$\begin{aligned}
 N_S^{\nu\nu Data} &= \frac{N_C^{Data} - N_C^{BKG}}{N_C^{MC}} \cdot N_S^{MC}[60, 120] \cdot R[60, 120] \\
 &= \frac{N_C^{Data} - N_C^{BKG}}{N_C^{MC}} \cdot N_S^{MC}[60, 120] \cdot \frac{N(Z/\gamma^* \rightarrow \mu\mu, m_Z^{Gen} > 50\text{GeV})}{N(Z/\gamma^* \rightarrow \mu\mu, 60 < m_Z^{Gen} < 120\text{GeV})} \cdot R[50, \infty]
 \end{aligned}$$

## Formulae from paper

$$N_S^{Data} = (N_C^{Data} - N_C^{BKG}) \cdot R[50, \infty] \cdot \frac{\epsilon_S^{VBF}}{\epsilon_C^{VBF} \epsilon_{\mu\mu}}$$

- ▶  $\epsilon_{\mu\mu} = \frac{N(Z/\gamma^* \rightarrow \mu\mu, \text{reco dimuon}, 60 < m_{\mu\mu}^{\text{reco}} < 120 \text{ GeV})}{N(Z/\gamma^* \rightarrow \mu\mu, m_Z^{\text{Gen}} > 50 \text{ GeV})}$
- ▶  $\epsilon_C^{VBF} = \frac{N_C^{MC}}{N(Z/\gamma^* \rightarrow \mu\mu, \text{reco dimuon}, 60 < m_{\mu\mu}^{\text{reco}} < 120 \text{ GeV})}$
- ▶  $\epsilon_S^{VBF} = \frac{N_S^{MC} [60, 120]}{N(Z/\gamma^* \rightarrow \mu\mu, 60 < m_Z^{\text{Gen}} < 120 \text{ GeV})}$
- ▶ n.b. efficiencies are not defined in the paper, so the differences in the denominator between  $\epsilon_{\mu\mu}$  and  $\epsilon_S^{VBF}$  are not apparent

## Simplifications

- ▶ Numerator of  $\epsilon_{\mu\mu}$  and denominator of  $\epsilon_C^{VBF}$  cancel so they should not be included in the error calculation
  - Currently stat, lepton ID, JES, JER and UES uncertainties are considered on all terms
- ▶  $\epsilon_C^{VBF} \cdot \epsilon_{\mu\mu} = \frac{N_C^{MC}}{N(Z/\gamma^* \rightarrow \mu\mu, m_Z^{Gen} > 50 \text{ GeV})}$
- ▶  $\frac{\epsilon_S^{VBF}}{\epsilon_C^{VBF} \cdot \epsilon_{\mu\mu}} = \frac{N_S^{MC}[60,120]}{N_C^{MC}} \cdot \frac{N(Z/\gamma^* \rightarrow \mu\mu, m_Z^{Gen} > 50 \text{ GeV})}{N(Z/\gamma^* \rightarrow \mu\mu, 60 < m_Z^{Gen} < 120 \text{ GeV})}$

### Final formula

$$N_S^{\nu\nu \text{ Data}} = \frac{N_C^{Data} - N_C^{BKG}}{N_C^{MC}(Z^{Gen} \rightarrow \mu\mu)} \cdot N_S^{MC}[60, 120](Z^{Gen} \rightarrow \mu\mu) \\ \times \frac{N(Z/\gamma^* \rightarrow \mu\mu, m_Z^{Gen} > 50 \text{ GeV})}{N(Z/\gamma^* \rightarrow \mu\mu, 60 < m_Z^{Gen} < 120 \text{ GeV})} \cdot R[50, \infty]$$

- ▶ This is the same as the formula derived above

## Preliminary Results - No Systematics

### Components

- ▶  $N_C^{Data} : 12 \pm 3.4641(stat)$
- ▶  $N_C^{Bkg} : 0.225755 \pm 0.118559(stat)$
- ▶  $N_S^{MC}[60, 120] : 40.9211 \pm 1.86195(stat)$
- ▶  $N_C^{MC}(Z^{Gen} \rightarrow \mu\mu) : 26.4646 \pm 1.46436(stat)$
- ▶  $N(Z/\gamma^* \rightarrow \mu\mu, m_Z^{Gen} > 50 GeV) : 3.3216e + 07 \pm 5763.33(stat)$
- ▶  $N(Z/\gamma^* \rightarrow \mu\mu, 60 < m_Z^{Gen} < 120 GeV) : 3.1915e + 07 \pm 5649.33(stat)$

### Preliminary Result

- ▶ Analysis B -  $98.4151 \pm 28.41(stat.) \pm 13.6326(MCstat.)$
- ▶ Analysis A -  $103.452 \pm 30.4366(stat.) \pm 14.2746(MCstat.)$

## What about QCD/EWK cross-section ratio difference

- ▶ For QCD the ratio of  $\nu\nu$  and  $\mu\mu$  cross-sections is 5.6
- ▶ For EWK Sasha has calculated that it is 1.6
- ▶  $N_S^{MC}$  is made up of 23 QCD events and 18 EWK events
- ▶ To my mind this means that the background estimation decreases considerable as a result of the much lower ratio for 40% of the signal events



## Conclusions

- ▶ Method does seem consistent
- ▶ Is there a reason not to calculate the cross-section ratio with the mass window?
  - It would remove the need for the additional event ratio.
- ▶ Preliminary results from analysis B are compatible at at least the same level as the W estimates

## Backup