

# PROBLEMS:

Physics 125 b

Midterm (Problem set number 5)

Due midnight Wednesday, February 7, 2024

READING: Section 18.5 in Shankar on the interaction of atoms with electromagnetic radiation.

This is the Ph 125a "midterm", aka problem set 5 .

Collaboration prohibited: Please do this problem set without consulting with or working with other students. This is the only restriction compared with our usual problem sets. You may use Piazza and attend office hours.

## 1

17. See if you can generalize the result for the first Born approximation:

$$\frac{d\sigma}{d\Omega'} = \frac{m^2}{(2\pi)^2} \left| \hat{V}(\mathbf{p}' - \mathbf{p}) \right|^2$$

to the case where the scattered particle (mass  $m_f$  ) may have a different mass than the incident particle (mass  $m_i$  ).

### 1.1

Looking over the note in 7.3 on approximate methods, I found the derivation for the Born approximation. Basically, this scattering rate is given by an expression of the form:

$$d\sigma = \frac{1}{\text{flux}} d\Gamma \tag{1}$$

and the idea is that the masses that go into this flux and rate are different. For the flux, we care about the mass of the incident particle  $m_i$ , and then for the differential rate, there is an expression of the form:

$$d\Gamma = \int_{\mathbf{p}' \in d\Omega'} \frac{L^3}{(2\pi)^3} d^3\mathbf{p}' 2\pi \frac{|\hat{V}(\mathbf{p}' - \mathbf{p})|^2}{L^6} \delta(E_f - E_i) \quad (2)$$

$$= \frac{d\Omega'}{L^3} \frac{m_f p_f}{(2\pi)^2} |\hat{V}(\mathbf{p}' - \mathbf{p})|^2 \quad (3)$$

The flux is given by:

$$\text{flux} = \frac{p_i}{m_i L^3} \quad (4)$$

And this term divided by the flux gives us the differential cross section in the first Born approximation, which for equal masses (and momenta) is expressed as:

$$\frac{d\sigma}{d\Omega'} = \frac{m^2}{(2\pi)^2} |\hat{V}(\mathbf{p}' - \mathbf{p})|^2 \quad (5)$$

This time the two masses will be different from each other, so we can write the formula as:

$$\frac{d\sigma}{d\Omega'_{\text{diffms}}} = \frac{m_f p_f m_i}{p_i (2\pi)^2} |\hat{V}(\mathbf{p}' - \mathbf{p})|^2 \quad (6)$$

## 2

18. We consider the potential (called the "Yukawa potential"):

$$V(\mathbf{x}) = \frac{K e^{-\mu r}}{r}, \quad r = |\mathbf{x}|$$

with real parameters  $K$  and  $\mu > 0$ . The parameter  $K$  can be regarded as the "strength" of the potential ("interaction"), and  $\frac{1}{\mu}$  is effectively the "range" of distance over which the potential is important.  $\mu$  itself has units of mass - note that as  $\mu \rightarrow 0$  we obtain the Coulomb potential:  $\mu$  can be thought of as the mass of an "exchanged particle" which mediates the force. In electromagnetism, this is the photon, hence  $\mu \rightarrow m_\gamma = 0$

(a) Find a condition on  $K$  and  $\mu$  which guarantees that there are at least  $n$  bound states in this potential. You will likely fashion and use some kind of "comparison" theorem in arriving at your result. You should give at least a "heuristically convincing" argument, if you don't actually prove it.

## 2.1

We know that for any bound states at all, we need  $K < 0$  or else we would just be talking about the free particle. Let us consider the Yukawa potential as a perturbation for our original Coulomb potential where we have  $H_0 = \frac{p^2}{2m} + \frac{K}{r}$  and the perturbation  $V = K \left( \frac{e^{-\mu r}}{r} - \frac{1}{r} \right)$ :

$$H = H_0 + V \quad (7)$$

It is always the case that  $K < 0$ , and the quantity inside of the parentheses will always be negative, so let's maximize the inside of the parentheses. We can expand it in a Taylor series:

$$e^{-\mu r} = 1 - \mu r + \frac{(\mu r)^2}{2} + \dots \quad (8)$$

So we can write the inside of the parentheses as:

$$\frac{e^{-\mu r}}{r} - \frac{1}{r} = -\mu + \frac{\mu^2 r}{2} + \dots \quad (9)$$

So, we can say the perturbation approximately goes like:

$$V \approx -\mu K \quad (10)$$

We know that the unperturbed energies in the Coulomb potential go like:

$$E_n = \frac{a}{n^2} \quad (11)$$

(b) Using the Born approximation for the differential cross section that we developed in our discussion of time-dependent perturbation theory, calculate the differential cross section,  $\frac{d\sigma}{d\Omega}$ , for scattering on this potential. Consider the limit  $\mu \rightarrow 0$  and compare with the classical Rutherford cross section (Is this calculation valid?).

## 2.2

We follow from 19.3.8 in Shankar. When the Yukawa potential is given by:

$$V(r) = \frac{K e^{-\mu r}}{r} \quad (12)$$

The scattering amplitude is given by:

$$f(\theta) = -\frac{2\mu_{reduced}K}{\hbar^2q} \int_0^\infty \frac{e^{iqr'} - e^{-iqr'}}{2i} e^{-\mu r'} dr' \quad (13)$$

where  $\mu_{reduced}$  is the reduced mass of the system. This simplifies to:

$$f(\theta) = -\frac{2\mu_{reduced}K}{\hbar(\mu^2 + q^2)} \quad (14)$$

The differential cross section in the Born approximation for the Yukawa potential is given by:

$$\frac{d\sigma}{d\Omega} = |f(\theta)|^2 \quad (15)$$

Now, we know that the square of the momentum  $q$  transferred to the particle is given by:

$$q^2 = 4k^2 \sin^2\left(\frac{\theta}{2}\right) \quad (16)$$

So we can write the differential cross section as:

$$\frac{d\sigma}{d\Omega} = \frac{4\mu_{reduced}^2 K^2}{\hbar^4(\mu^2 + q^2)^2} = \frac{4\mu_{reduced}^2 K^2}{\hbar^4(\mu^2 + 4k^2 \sin^2(\frac{\theta}{2}))^2} \quad (17)$$

To compare with the classical Rutherford cross section, we need to consider the limit  $\mu \rightarrow 0$  and let us set  $K = Ze^2$ . We can write the differential cross section as:

$$\frac{d\sigma}{d\Omega} = \frac{4\mu_{reduced}^2 (Ze^2)^2}{\hbar^4(4k^2 \sin^2(\frac{\theta}{2}))^2} = \frac{(Ze^2)^2}{16E^2 \sin^4(\frac{\theta}{2})} \quad (18)$$

where we have used the fact that the momentum  $p$  is given by:

$$p = \frac{\hbar k}{2} \quad (19)$$

and the energy  $E$  is given by:

$$E = \frac{p^2}{2\mu_{reduced}} \quad (20)$$

Although the result is the same as the Rutherford cross section, the calculation is not completely valid because the formulation developed is for a potential that vanishes faster than the inverse of the radius, which the

Coulomb does not. For a comprehensive discussion, see Shankar pg. 532.

(c) Integrate your differential cross section over all solid angles to obtain the "total cross section". Again, consider the limit  $\mu \rightarrow 0$ . Assuming the Rutherford cross section holds, what is the total cross section for scattering on a Coulomb potential?

## 2.3

Here we are interested in the total cross section, which is given by:

$$\sigma = \int \frac{d\sigma}{d\Omega} d\Omega \quad (21)$$

Plugging in our expression for the differential cross section after we took the limit  $\mu \rightarrow 0$ , we have:

$$\sigma = \int \frac{(Ze^2)^2}{16E^2 \sin^4\left(\frac{\theta}{2}\right)} d\Omega \quad (22)$$

We can write the differential solid angle as:

$$d\Omega = 2\pi \sin \theta d\theta \quad (23)$$

Pulling out all the terms that do not depend on  $\theta$  and inserting this identity, we have:

$$\sigma = \frac{\pi(Ze^2)^2}{8E^2} \int \frac{\sin \theta}{\sin^4\left(\frac{\theta}{2}\right)} d\theta \quad (24)$$

This integral diverges at  $\theta = 0, \pi$ . So, we are not able to calculate the total cross section for the Coulomb potential.

## 3

19. Suppose we have a system consisting of two spin- $\frac{1}{2}$ 's ( $\mathbf{S}_1$  and  $\mathbf{S}_2$ ), with an interaction Hamiltonian  $a(t)\mathbf{S}_1 \cdot \mathbf{S}_2$ . Assume that  $a(\pm\infty) = 0$  and that  $a(t)$  is significantly different from zero in an interval of order  $\tau$  in width about  $t = 0$ . (a) Suppose at very early times ( $t \rightarrow -\infty$ ), the system is in state

$$|\psi(-\infty)\rangle = |+-\rangle$$

where the state is labeled by the  $z$  components of  $\mathbf{S}_1$  and  $\mathbf{S}_2$ , respectively. Without making any approximations, what is the state of the system at  $t = \infty$ ? What is the probability,  $P(+ - \rightarrow - +)$ , that the state is observed to be  $| - + \rangle$  at  $t = \infty$ ? Show that this probability depends only on  $\int_{-\infty}^{\infty} a(t) dt$ .

### 3.1

We know that:

$$(\mathbf{S})^2 = (S_1 + S_2)^2 \rightarrow S^2 = S_1^2 + S_2^2 + 2\mathbf{S}_1 \cdot \mathbf{S}_2 \quad (25)$$

so we can write the Hamiltonian as:

$$H = a(t)\mathbf{S}_1 \cdot \mathbf{S}_2 = \frac{1}{2}a(t) (S^2 - S_1^2 - S_2^2) \quad (26)$$

We want to transform current basis into the basis of the total spin and its  $z$ -component. Our current basis is:

$$|+-\rangle = \left| S_1^z = +\frac{1}{2}, S_2^z = -\frac{1}{2} \right\rangle \quad (27)$$

and:

$$|-+\rangle = \left| S_1^z = -\frac{1}{2}, S_2^z = +\frac{1}{2} \right\rangle \quad (28)$$

The total  $S_z$  of 0 must be conserved, so we have the singlet:

$$|S^2 = 0, S^z = 0\rangle = \frac{1}{\sqrt{2}} (|+-\rangle - |-+\rangle) \quad (29)$$

and one of the triplets:

$$|S^2 = 1, S^z = 0\rangle = \frac{1}{\sqrt{2}} (|+-\rangle + |-+\rangle) \quad (30)$$

We can express our initial state in terms of these states:

$$|+-\rangle = \frac{1}{\sqrt{2}} (|S^2 = 1, S^z = 0\rangle + |S^2 = 0, S^z = 0\rangle) \quad (31)$$

and

$$|-+\rangle = \frac{1}{\sqrt{2}} (|S^2 = 1, S^z = 0\rangle - |S^2 = 0, S^z = 0\rangle) \quad (32)$$

We seek to figure out the eigenvalues of the Hamiltonian in the basis of the total spin and its  $z$ -component. We can write the Hamiltonian as:

$$H = a(t) \mathbf{S}_1 \cdot \mathbf{S}_2 = \frac{1}{2} a(t) (S^2 - S_1^2 - S_2^2) \quad (33)$$

So, it operates on the states as:

$$H |S^2 = 0, S^z = 0\rangle = \frac{1}{2} a(t) \left(0 - \frac{3}{4} - \frac{3}{4}\right) |S^2 = 0, S^z = 0\rangle = -\frac{3}{4} a(t) |S^2 = 0, S^z = 0\rangle \quad (34)$$

and:

$$H |S^2 = 1, S^z = 0\rangle = \frac{1}{2} a(t) \left(2 - \frac{3}{4} - \frac{3}{4}\right) |S^2 = 1, S^z = 0\rangle = \frac{1}{4} a(t) |S^2 = 1, S^z = 0\rangle \quad (35)$$

So, the action on the initial state is:

$$H |+-\rangle = \frac{1}{\sqrt{2}} \left(-\frac{3}{4} a(t) + \frac{1}{4} a(t)\right) (|S^2 = 1, S^z = 0\rangle + |S^2 = 0, S^z = 0\rangle) = -\frac{1}{2} a(t) |+-\rangle \quad (36)$$

For the state  $|-+\rangle$ , we have:

$$H |-+\rangle = \frac{1}{\sqrt{2}} \left(-\frac{3}{4} a(t) - \frac{1}{4} a(t)\right) (|S^2 = 1, S^z = 0\rangle - |S^2 = 0, S^z = 0\rangle) = -a(t) |-+\rangle \quad (37)$$

With the time-dependent Schrödinger equation, we can write:

$$i \frac{\partial}{\partial t} |\psi(t)\rangle = H |\psi(t)\rangle = a(t) \mathbf{S}_1 \cdot \mathbf{S}_2 |\psi(t)\rangle \quad (38)$$

We can express  $|\psi(t)\rangle$  as a two-component vector in the basis of the total spin and its  $z$ -component:

$$|\psi(t)\rangle = \begin{pmatrix} c_1(t) \\ c_2(t) \end{pmatrix} \quad (39)$$

where  $c_1(t)$  and  $c_2(t)$  are the probability amplitudes for the states  $|S^2 = 0, S^z = 0\rangle$  and  $|S^2 = 1, S^z = 0\rangle$ , respectively. Then, we have the time-dependent Schrödinger equation with only one component:

$$i \partial_t c_1(t) = -\frac{3}{4} a(t) c_1(t) \quad (40)$$

Dividing through by  $c_1(t)$  and multiplying by  $-i$  gives us:

$$\frac{\partial_t c_1(t)}{c_1(t)} = +i\frac{3}{4}a(t) \quad (41)$$

We recognize the left-hand side as the derivative of the natural logarithm of  $c_1(t)$ , so we can write:

$$\ln(c_1(t)) = \frac{3i}{4} \int_{-\infty}^t a(t') dt' \quad (42)$$

We can exponentiate both sides to solve for  $c_1(t)$ :

$$c_1(t) = e^{\frac{3i}{4} \int_{-\infty}^t a(t') dt'} \quad (43)$$

We can do the same for  $c_2(t)$ :

$$c_2(t) = e^{-\frac{i}{4} \int_{-\infty}^t a(t') dt'} \quad (44)$$

Now, we that we have the propagators for the eigenstates of the Hamiltonian, we can write the state at  $t = \infty$  as:

$$|\psi(\infty)\rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} e^{\frac{3i}{4} \int_{-\infty}^{\infty} a(t') dt'} \\ e^{-\frac{i}{4} \int_{-\infty}^{\infty} a(t') dt'} \end{pmatrix} \quad (45)$$

We can write the probability of the state  $|+-\rangle$  transitioning to the state  $|-+\rangle$  as the projection of the state at  $t = \infty$  onto the state  $|-+\rangle$ :

$$P(+ - \rightarrow - +) = |\langle -+ | \psi(\infty) \rangle|^2 \quad (46)$$

In the basis of the total spin and its  $z$ -component, we can write the state  $|-+\rangle$  as:

$$|-+\rangle = \frac{1}{\sqrt{2}} \left( \begin{pmatrix} 1 \\ 0 \end{pmatrix} - \begin{pmatrix} 0 \\ 1 \end{pmatrix} \right) = \frac{1}{\sqrt{2}} \begin{pmatrix} -1 \\ 1 \end{pmatrix} \quad (47)$$

So we can write the probability as:

$$P(+ - \rightarrow - +) = \left| \frac{1}{\sqrt{2}} \begin{pmatrix} -1 & 1 \end{pmatrix} \cdot \frac{1}{\sqrt{2}} \begin{pmatrix} e^{\frac{3i}{4} \int_{-\infty}^{\infty} a(t') dt'} \\ e^{-\frac{i}{4} \int_{-\infty}^{\infty} a(t') dt'} \end{pmatrix} \right|^2 \quad (48)$$

which simplifies to:

$$P(+ - \rightarrow - +) = \left| -\frac{1}{2} e^{\frac{3i}{4} \int_{-\infty}^{\infty} a(t') dt'} + \frac{1}{2} e^{-\frac{i}{4} \int_{-\infty}^{\infty} a(t') dt'} \right|^2 \quad (49)$$

(b) Now calculate  $P(+ - \rightarrow - +)$  in first-order time-dependent perturbation theory. What is a condition for the validity of first-order time-dependent perturbation theory?



### 3.2

For the total Hamiltonian, we can choose:

$$H = H_0 + V_t = 0 + a(t)\mathbf{S}_1 \cdot \mathbf{S}_2 \quad (50)$$

Ordinarily, in first-order time-dependent perturbation theory, we would write the transition probability as:

$$P(+ - \rightarrow - +) = \left| \frac{1}{i\hbar} \int_{-\infty}^{\infty} e^{i\omega_{+-}t} \langle f | V_t | i \rangle dt \right|^2 \quad (51)$$

But here the unperturbed Hamiltonian is 0, so the exponential factor is just 1, and the transition probability is just:

$$P(+ - \rightarrow - +) = \left| \frac{1}{i\hbar} \int_{-\infty}^{\infty} \langle f | V_t | i \rangle dt \right|^2 \quad (52)$$

The initial state is  $|i\rangle = |+-\rangle$  and the final state is  $|f\rangle = |-+\rangle$ . We can write the transition probability as:

$$P(+ - \rightarrow - +) = \left| \frac{1}{i\hbar} \int_{-\infty}^{\infty} \langle -+ | a(t)\mathbf{S}_1 \cdot \mathbf{S}_2 | +- \rangle dt \right|^2 \quad (53)$$

We found the action of the Hamiltonian on the state  $|+-\rangle$  in the previous part to be:

$$H |+-\rangle = -\frac{1}{2}a(t) |+-\rangle \quad (54)$$

So we can write the transition probability as:

$$P(+ - \rightarrow - +) = \left| \frac{1}{i\hbar} \int_{-\infty}^{\infty} -\frac{1}{2}a(t) \langle -+ | +- \rangle dt \right|^2 \quad (55)$$

Now, we want to compute the overlap of the states  $|-+\rangle$  and  $|+-\rangle$ :

$$\langle -+ | +- \rangle = \frac{1}{2} \begin{pmatrix} -1 & 1 \end{pmatrix} \cdot \begin{pmatrix} 1 \\ 1 \end{pmatrix} = 0 \quad (56)$$

## 4

20. With reference to the previous problem, consider the "Zeeman effect", with the addition of a uniform, static magnetic field of strength  $B_0$  in the direction of the  $z$ -axis. Thus, we have another term in the Hamiltonian:

$$H_0 = -B_0 (g_1 S_{1z} + g_2 S_{2z}),$$

where  $g_1$  and  $g_2$  are the gyromagnetic ratios of the two spins. We assume a gaussian form for  $a(t)$  :

$$a(t) = a(0)e^{-(t/\tau)^2}$$

Repeat your first order perturbation theory calculation of the previous problem in the presence of the magnetic field and with this explicit form for  $a(t)$ , and determine  $P(+ - \rightarrow - +)$ . Discuss the variation of this probability, for given  $a(0)$  and  $\tau$ , as a function of  $B_0$ .

### 4.1

Now, the unperturbed Hamiltonian gives non-zero energies for the states  $|+-\rangle$  and  $|-+\rangle$ :

$$H_0 |+-\rangle = -B_0 \left( g_1 \frac{1}{2} + g_2 \left( -\frac{1}{2} \right) \right) |+-\rangle = -\frac{1}{2} B_0 (g_1 - g_2) |+-\rangle \quad (57)$$

and:

$$H_0 |-+\rangle = -B_0 \left( g_1 \left( -\frac{1}{2} \right) + g_2 \frac{1}{2} \right) |-+\rangle = -\frac{1}{2} B_0 (g_2 - g_1) |-+\rangle \quad (58)$$