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Patryk's Notes

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## Chapter 1

# $G_0W_0$ Derivations: 11/9

## 1.1 Deriving Hedin's equations

### 1.1.1 Time-Domain Definition of the Green's Function

Start by considering the equation of motion for the field operators

$$i\frac{\partial}{\partial t}\hat{\psi}(\mathbf{x},t) = [\hat{\psi}(\mathbf{x},t), \hat{H}_{elec}] = [\hat{\psi}(\mathbf{x},t), \hat{H}_0 + \hat{H}_{int}] = [\hat{\psi}(\mathbf{x},t), \hat{H}_0] + [\hat{\psi}(\mathbf{x},t), \hat{H}_{int}]$$
(1.1)

For the non-interacting part, we have

$$[\hat{\psi}(\mathbf{x},t),\hat{H}_0] = [\hat{\psi}(\mathbf{x},t), \int d\mathbf{x}' \hat{\psi}^{\dagger}(\mathbf{x}',t) \hat{h}^0(\mathbf{x}') \hat{\psi}(\mathbf{x}',t)]$$
(1.2)

$$= \int d\mathbf{x}' [\hat{\psi}(\mathbf{x}, t), \hat{\psi}^{\dagger}(\mathbf{x}', t) \hat{h}^{0}(\mathbf{x}') \hat{\psi}(\mathbf{x}', t)]$$
(1.3)

$$= \int d\mathbf{x}' \left( \underbrace{\left[\hat{\psi}(\mathbf{x},t), \hat{\psi}^{\dagger}(\mathbf{x}',t)\right]}_{\delta(\mathbf{x}-\mathbf{x}')} \hat{h}^{0}(\mathbf{x}') \hat{\psi}(\mathbf{x}',t) + \hat{\psi}^{\dagger}(\mathbf{x}',t) \hat{h}^{0}(\mathbf{x}') \underbrace{\left[\hat{\psi}(\mathbf{x},t), \hat{\psi}(\mathbf{x}',t)\right]}_{0} \right)$$
(1.4)

$$= h^0(\mathbf{x})\hat{\psi}(\mathbf{x}, t). \tag{1.5}$$

For the interacting part, we have

$$[\hat{\psi}(\mathbf{x},t),\hat{H}_{int}] = \frac{1}{2} \int d\mathbf{x}' d\mathbf{x}'' v(\mathbf{x},\mathbf{x}') [\hat{\psi}(\mathbf{x},t),\hat{\psi}^{\dagger}(\mathbf{x}',t)\hat{\psi}^{\dagger}(\mathbf{x}'',t)\hat{\psi}(\mathbf{x}'',t)\hat{\psi}(\mathbf{x}'',t)]$$
(1.6)

$$= \frac{1}{2} \int d\mathbf{x}' d\mathbf{x}'' v(\mathbf{x}, \mathbf{x}') \left( \delta(\mathbf{x} - \mathbf{x}') \hat{\psi}^{\dagger}(\mathbf{x}'', t) \hat{\psi}(\mathbf{x}'', t) \hat{\psi}(\mathbf{x}', t) + \hat{\psi}^{\dagger}(\mathbf{x}', t) \delta(\mathbf{x} - \mathbf{x}'') \hat{\psi}(\mathbf{x}'', t) \hat{\psi}(\mathbf{x}', t) \right)$$

$$(1.7)$$

 $= \int d\mathbf{x}' \hat{\psi}^{\dagger}(\mathbf{x}', t) v(\mathbf{x}, \mathbf{x}') \hat{\psi}(\mathbf{x}', t) \hat{\psi}(\mathbf{x}, t)$ (1.8)

so overall we have

$$i\frac{\partial}{\partial t}\hat{\psi}(\mathbf{x},t) = \left(\hat{h}^0(\mathbf{x}) + \int d\mathbf{x}' v(\mathbf{x}, \mathbf{x}') \hat{\psi}^{\dagger}(\mathbf{x}', t) \hat{\psi}(\mathbf{x}', t)\right) \hat{\psi}(\mathbf{x}, t)$$
(1.9)

Now we can consider the equation of motion for the Green's function, defined as  $G(\mathbf{x}t, \mathbf{x}'t') = -i \langle N | \mathcal{T} \left( \hat{\psi}(\mathbf{x}, t) \hat{\psi}^{\dagger}(\mathbf{x}', t') \right) | N \rangle$ , where  $\mathcal{T}$  is the time-ordering operator.

$$\frac{\partial}{\partial t}G(\mathbf{x}t, \mathbf{x}'t') = -i \langle N | \frac{\partial}{\partial t} \mathcal{T} \left( \hat{\psi}(\mathbf{x}, t) \hat{\psi}^{\dagger}(\mathbf{x}', t') \right) | N \rangle$$
(1.10)

Now.

$$\frac{\partial}{\partial t} \mathcal{T} \left( \hat{\psi}(\mathbf{x}, t) \hat{\psi}^{\dagger}(\mathbf{x}', t') \right) = \frac{\partial}{\partial t} \left( \theta(t - t') \hat{\psi}(\mathbf{x}, t) \hat{\psi}^{\dagger}(\mathbf{x}', t') - \theta(t' - t) \hat{\psi}^{\dagger}(\mathbf{x}', t') \hat{\psi}(\mathbf{x}, t) \right) \tag{1.12}$$

$$= \underbrace{\frac{\partial \theta(t - t')}{\partial t}}_{\delta(t - t')} \hat{\psi}(\mathbf{x}, t) \hat{\psi}^{\dagger}(\mathbf{x}', t') + \theta(t - t') \frac{\partial}{\partial t} \hat{\psi}(\mathbf{x}, t) \hat{\psi}^{\dagger}(\mathbf{x}', t') - \underbrace{\frac{\partial \theta(t' - t)}{\partial t}}_{-\delta(t' - t)} \hat{\psi}^{\dagger}(\mathbf{x}', t') \hat{\psi}(\mathbf{x}, t) - \theta(t' - t) \frac{\partial}{\partial t} \hat{\psi}^{\dagger}(\mathbf{x}', t') + \underbrace{\frac{\partial}{\partial t} \hat{\psi}^{\dagger}(\mathbf{x}', t') \hat{\psi}(\mathbf{x}, t') \hat{\psi}(\mathbf{x}, t') \hat{\psi}(\mathbf{x}, t') \hat{\psi}(\mathbf{x}, t') \hat{\psi}(\mathbf{x}, t') \hat{\psi}^{\dagger}(\mathbf{x}', t')}_{(1.13)}$$

$$= \underbrace{\delta(t - t') \Big\{ \hat{\psi}(\mathbf{x}, t), \hat{\psi}^{\dagger}(\mathbf{x}', t') \Big\}}_{\delta(\mathbf{x} - \mathbf{x}') \delta(t - t')} + \mathcal{T} \left( \frac{\partial}{\partial t} \hat{\psi}(\mathbf{x}, t) \hat{\psi}^{\dagger}(\mathbf{x}', t') \right)$$

$$(1.14)$$

So now consider plugging in the equation of motion for  $\hat{\psi}(\mathbf{x},t)$  into the above expression

$$\mathcal{T}\left(\frac{\partial}{\partial t}\hat{\psi}(\mathbf{x},t)\hat{\psi}^{\dagger}(\mathbf{x}',t')\right) = \mathcal{T}\left(-i\left(\hat{h}^{0}(\mathbf{x}) + \int d\mathbf{x}''v(\mathbf{x},\mathbf{x}'')\hat{\psi}^{\dagger}(\mathbf{x}'',t)\hat{\psi}(\mathbf{x}'',t)\right)\hat{\psi}(\mathbf{x},t)\hat{\psi}^{\dagger}(\mathbf{x}',t')\right)$$

$$= -i\hat{h}^{0}(\mathbf{x})\mathcal{T}\left(\hat{\psi}(\mathbf{x},t)\hat{\psi}^{\dagger}(\mathbf{x}',t')\right) - i\int d\mathbf{x}''v(\mathbf{x},\mathbf{x}'')\mathcal{T}\left(\hat{\psi}(\mathbf{x},t)\hat{\psi}^{\dagger}(\mathbf{x}'',t)\hat{\psi}(\mathbf{x}'',t)\hat{\psi}^{\dagger}(\mathbf{x}',t')\right)$$

$$(1.17)$$

So we have

$$\left[i\frac{\partial}{\partial t} - \hat{h}^{0}(\mathbf{x})\right] G(\mathbf{x}t, \mathbf{x}'t') = \delta(\mathbf{x} - \mathbf{x}')\delta(t - t') - i \int d\mathbf{x}'' v(\mathbf{x}, \mathbf{x}'') \underbrace{G_{2}(\mathbf{x}t, \mathbf{x}''t, \mathbf{x}'t', \mathbf{x}'t')}_{\langle N|\mathcal{T}\left(\hat{\psi}(\mathbf{x},t)\hat{\psi}(\mathbf{x}'',t)\hat{\psi}^{\dagger}(\mathbf{x}'',t')\hat{\psi}^{\dagger}(\mathbf{x}'',t')\right)|N\rangle} (1.18)$$

and we notice that in order to compute the single-particle Green's function, we need to know the two-particle Green's function, which needs the three-particle Green's function, and so on. So to simplify we introduce a nonlocal, time-dependent self-energy  $\Sigma(\mathbf{x}t, \mathbf{x}'t')$  that satisfies

$$-i \int d\mathbf{x}'' v(\mathbf{x}, \mathbf{x}'') G_2(\mathbf{x}t, \mathbf{x}''t, \mathbf{x}'t^+, \mathbf{x}'t') \equiv \int dt'' \int d\mathbf{x}'' \overline{\Sigma}(\mathbf{x}t, \mathbf{x}''t'') G(\mathbf{x}''t'', \mathbf{x}'t')$$
(1.19)

and further define  $\Sigma = \overline{\Sigma} - v_H$  with

$$v_H(\mathbf{x},t) = \int d\mathbf{x}' v(\mathbf{x}, \mathbf{x}') \underbrace{\langle N | \hat{\psi}^{\dagger}(\mathbf{x}') \hat{\psi}(\mathbf{x}') | N \rangle}_{-\frac{1}{i}G(\mathbf{x}'t, \mathbf{x}'t)} = i \int d\mathbf{x}' v(\mathbf{x}, \mathbf{x}') G(\mathbf{x}'t, \mathbf{x}'t)$$
(1.20)

and we can rewrite the equation of motion as

$$\left[i\frac{\partial}{\partial t} - \hat{h}^{0}(\mathbf{x}) - v_{H}(\mathbf{x}, t)\right] G(\mathbf{x}t, \mathbf{x}'t') = \delta(\mathbf{x} - \mathbf{x}')\delta(t - t') + \int dt'' \int d\mathbf{x}'' \Sigma(\mathbf{x}t, \mathbf{x}''t'') G(\mathbf{x}''t'', \mathbf{x}'t')$$
(1.21)

Now consider defining the  $G_0$  of the non-interacting system

$$\left[i\frac{\partial}{\partial t} - \hat{h}^{0}(\mathbf{x}) - v_{H}(\mathbf{x}, t)\right] G_{0}(\mathbf{x}t, \mathbf{x}'t') = \delta(\mathbf{x} - \mathbf{x}')\delta(t - t')$$
(1.22)

So we can write equations 1.21 and 1.22 symbolically as

$$\hat{O}G = \delta + \Sigma G \quad \text{and} \quad \hat{O}G_0 = \delta$$
 (1.23)

$$\implies G_0 = \hat{O}^{-1} \implies G = G_0 + G_0 \Sigma G \tag{1.24}$$

$$\implies G(1,2) = G_0(1,2) + \int d3d4G_0(1,3)\Sigma(3,4)G(4,2) \tag{1.25}$$

where we use the space-time notation  $1 = (\mathbf{x}_1, t_1)$  etc.

### 1.1.2 Hedin's Equations

Schwinger chose to introduce a potential  $\varphi$  that we will later set to zero, in order to rewrite the two-particle Green's function as

$$G_2(1,3,2,3^+) = G(1,2) G(3,3^+) - \frac{\delta G(1,2)}{\delta \varphi(3)},$$
 (1.26)

So

$$\bar{\Sigma}(1,2) = -i \int d(3) \ v(1,3) G_2(1,3,2,3^+)$$
(1.27)

$$= -i \int d(3) v(1,3) \left[ G(1,2) G(3,3^{+}) - \frac{\delta G(1,2)}{\delta \varphi(3)} \right]$$
 (1.28)

$$= -i G(1,2) \underbrace{\int d(3) v(1,3) G(3,3^{+})}_{-iv_{H}(1)} + i \int d(3) v(1,3) \frac{\delta G(1,2)}{\delta \varphi(3)}. \tag{1.29}$$

Now because  $\delta G = -G(\delta G^{-1}) G$  we can write the identity

$$\frac{\delta G(1,2)}{\delta \varphi(3)} = -\int d(4)d(5) \ G(1,4) \frac{\delta G^{-1}(4,5)}{\delta \varphi(3)} G(5,2). \tag{1.30}$$

So the second term in Eq. (1.29) gives

$$i \int d(3) \ v(1,3) \frac{\delta G(1,2)}{\delta \varphi(3)} = -i \int d(3) \ v(1,3) \int d(4) d(5) \ G(1,4) \frac{\delta G^{-1}(4,5)}{\delta \varphi(3)} G(5,2)$$
$$= -i \int d(3,4,5) \ v(1,3) G(1,4) \frac{\delta G^{-1}(4,5)}{\delta \varphi(3)} G(5,2). \tag{1.31}$$

Now we can get rid of a G(1,2) dependence by multiplying with  $G^{-1}$ , yielding

$$\overline{\Sigma}(1,2) = -\delta(1,2) v_H(1) - i \int d(3,4) v(1,3) G(1,4) \frac{\delta G^{-1}(4,2)}{\delta \varphi(3)}.$$
 (1.32)

Introduce  $V(1) = \varphi(1) + v_H(1)$  as the total potential that electrons experience. Consider

$$\frac{\delta G^{-1}(1,2)}{\delta \varphi(3)} \equiv \underbrace{\frac{\delta G^{-1}(1,2)}{\delta V(5)}}_{-\Gamma(1,2,5)} \underbrace{\frac{\delta V(5)}{\delta \varphi(3)}}_{\varepsilon^{-1}(5,3)}.$$
(1.33)

So

$$\overline{\Sigma}(1,2) = -\delta(1,2) v_H(1) + i \int d(5) \underbrace{\int d(3)v(1,3) \varepsilon^{-1}(3,5)}_{W(1,5)} \int d(4)G(1,4) \Gamma(4,5,2). \quad (1.34)$$

and if we further make the GW approximation where  $\Gamma(4,5,2) \approx \delta(4,5)\delta(2,5)$  we get

$$\overline{\Sigma}(1,2) = -\delta(1,2) v_H(1) + iW(1,2)G(1,2)$$
(1.35)

and if we just care about the exchange-correlation part, we can define

$$\Sigma_{xc}(1,2) = \overline{\Sigma}(1,2) + \delta(1,2) v_H(1) = iW(1,2)G(1,2) \implies \Sigma_{xc}(\tau) = iW(\tau)G(\tau)$$
 (1.36)

where  $\tau = t_1 - t_2$ . Define  $G(\tau) = \int \frac{d\omega'}{2\pi} e^{-i\omega'\tau} G(\omega')$  and  $W(\tau) = \int \frac{d\omega''}{2\pi} e^{-i\omega''\tau} W(\omega'')$  to get

$$\Sigma_{xc}(\tau) = i \int \frac{d\omega'}{2\pi} \int \frac{d\omega''}{2\pi} e^{-i(\omega' + \omega'')\tau} G(\omega') W(\omega'')$$
(1.37)

Taking the inverse Fourier transform of  $\Sigma_{xc}(\tau)$  we get

$$\Sigma_{xc}(\omega) = \int \frac{d\tau}{2\pi} e^{i\omega\tau} \Sigma_{xc}(\tau) = i \int \frac{d\omega'}{2\pi} \int \frac{d\omega''}{2\pi} G(\omega') W(\omega'') \underbrace{\int d\tau e^{i(\omega - \omega' - \omega'')\tau}}_{2\pi\delta(\omega - \omega' - \omega'')} = i \int \frac{d\omega'}{2\pi} G(\omega') W(\omega - \omega')$$
(1.38)

Now in  $G_0W_0$  one applies the Cauchy residue theorem to solve this convolution integral, yielding the known form.

## 1.2 Final expressions

## 1.2.1 Fully analytic

I follow the notation of Tianyu's paper throughout this section [?]. We want to solve for the self-energy whose form along the real axis is:

$$\Sigma(\mathbf{r}, \mathbf{r}', \omega) = \frac{i}{2\pi} \int_{-\infty}^{\infty} e^{i\omega'\eta} d\omega' G_0(\mathbf{r}, \mathbf{r}', \omega + \omega') W_0(\mathbf{r}, \mathbf{r}', \omega')$$
(1.39)

In the molecular brutal basis, the self energy is given as:

$$\Sigma_{nn'}(\mathbf{k},\omega) = \iint d\mathbf{r} d\mathbf{r}' \psi_{n\mathbf{k}}^*(\mathbf{r}) \Sigma(\mathbf{r}, \mathbf{r}', \omega) \psi_{n'\mathbf{k}}(\mathbf{r}')$$
(1.40)

Also, recall that the Lehmann representation of the noninteracting Green's function is:

$$G_0(\mathbf{r}, \mathbf{r}', \omega) = \sum_{o\mathbf{q}} \frac{\psi_{o\mathbf{q}}(\mathbf{r})\psi_{o\mathbf{q}}^*(\mathbf{r}')}{\omega - \epsilon_{o\mathbf{q}} + i\eta \operatorname{sgn}(\epsilon_{o\mathbf{q}} - \mu)}$$
(1.41)

Now plugging both of these back into the original expression, we find:

$$\Sigma_{nn'}(\mathbf{k},\omega) = \frac{i}{2\pi} \sum_{o\mathbf{q}} \int_{-\infty}^{\infty} d\omega' \frac{e^{i\omega'\eta}}{\omega + \omega' - \epsilon_{o\mathbf{q}} + i\eta \operatorname{sgn}(\epsilon_{o\mathbf{q}} - \mu)}$$

$$\times \iint d\mathbf{r} d\mathbf{r}' \psi_{n\mathbf{k}}^{*}(\mathbf{r}) \psi_{o\mathbf{q}}(\mathbf{r}) W_{0}(\mathbf{r}, \mathbf{r}', \omega') \psi_{o\mathbf{q}}^{*}(\mathbf{r}') \psi_{n'\mathbf{k}}(\mathbf{r}')$$

$$= \frac{i}{2\pi} \sum_{o\mathbf{q}} \int_{-\infty}^{\infty} d\omega' \frac{e^{i\omega'\eta}}{\omega + \omega' - \epsilon_{o\mathbf{k}-\mathbf{q}} + i\eta \operatorname{sgn}(\epsilon_{o\mathbf{k}-\mathbf{q}} - \mu)} (n_{\mathbf{k}} o_{\mathbf{k}-\mathbf{q}} | W_{0} | o_{\mathbf{k}-\mathbf{q}} n_{\mathbf{k}}')$$
(1.43)

Where we have used the fact that the momentum index  $\mathbf{q}$  is the same as  $\mathbf{k} - \mathbf{q}$ , given that we are looping over both  $\mathbf{k}$  and  $\mathbf{q}$  anyways.

So the Green's function will bring poles at  $\omega' = \epsilon_{0\mathbf{k}-\mathbf{q}} - \omega + i\eta \operatorname{sgn}(\mu - \epsilon_{0\mathbf{k}-\mathbf{q}})$ . Now, we know that the screened Coulomb interaction has the expansion in terms of the bare Coulomb potential v and the density response function  $\chi_0$  as  $W_0 = v + v\chi_0 v + v\chi_0 v\chi_0 v + \cdots = v\left(1 + \chi_0 v + \chi_0 v\chi_0 v + \cdots\right) = v\left(1 - \chi_0 v\right)^{-1}$ , where we recognize the dielectric function as  $\epsilon_0 = 1 - \chi_0 v$  so we can express the screened Coulomb interaction as

$$W_0(\mathbf{r}, \mathbf{r}', \omega) = \frac{v(\mathbf{r}, \mathbf{r}')}{1 - (\chi_0 v)(\mathbf{r}, \mathbf{r}', \omega)}$$
(1.44)

recalling that the bare Coulomb interaction should be independent of frequency. A discussion of how to compute the screened Coulomb interaction can be found in this old work [?]. To simplify notation let us define a polarizability  $\Pi(\mathbf{r}, \mathbf{r}', \omega) = (\chi_0 v)(\mathbf{r}, \mathbf{r}', \omega)$ , so that we can rewrite the screened Coulomb interaction as:

$$(n_{\mathbf{k}}o_{\mathbf{k}-\mathbf{q}}|W_0|o_{\mathbf{k}-\mathbf{q}}n_{\mathbf{k}}') = \iint d\mathbf{r}d\mathbf{r}'\psi_{n\mathbf{k}}^*(\mathbf{r})\psi_{o\mathbf{q}}(\mathbf{r})W_0(\omega)\psi_{o\mathbf{q}}^*(\mathbf{r}')\psi_{n'\mathbf{k}}(\mathbf{r}')$$
(1.45)

At this point, we recognize the decomposition of the ERIs with the Cholesky vectors as:

$$\left(p_{\mathbf{k}_{p}}q_{\mathbf{k}_{q}}\right|\frac{1}{|\mathbf{r}-\mathbf{r}'|}|r_{\mathbf{k}_{r}}s_{\mathbf{k}_{s}}) = \sum_{PQ} v_{P\mathbf{q}}^{p\mathbf{k}_{\mathbf{p}}q\mathbf{k}_{\mathbf{q}}} v_{Q(-\mathbf{q})}^{r\mathbf{k}_{\mathbf{r}}s\mathbf{k}_{\mathbf{s}}}$$
(1.46)

so each Cholesky brings a factor of  $\mathbf{J}^{\frac{1}{2}}$ . Each Cholesky is defined as:

$$v_{P\mathbf{q}}^{p\mathbf{k}_{\mathbf{p}}q\mathbf{k}_{\mathbf{q}}} = \sum_{R} \mathbf{J}_{RP}^{-\frac{1}{2}}(\mathbf{q}) \left( R\mathbf{q} \mid p\mathbf{k}_{p}q\mathbf{k}_{q} \right)$$
(1.47)

where

$$\mathbf{J}_{PQ}(\mathbf{k}) = \iint d\mathbf{r} d\mathbf{r}' \phi_{P(-\mathbf{k})}(\mathbf{r}) \frac{1}{|\mathbf{r} - \mathbf{r}'|} \phi_{Q\mathbf{k}}(\mathbf{r}')$$

$$(Q\mathbf{k}_{rs} \mid r\mathbf{k}_{r} s \mathbf{k}_{s}) = \iint d\mathbf{r} d\mathbf{r}' \phi_{Q\mathbf{k}_{rs}}(\mathbf{r}) \frac{1}{|\mathbf{r} - \mathbf{r}'|} \phi_{r\mathbf{k}_{r}}^{*}(\mathbf{r}') \phi_{s\mathbf{k}_{s}}(\mathbf{r}')$$
(1.48)

So the simplest thing now will be to derive an expression for the columb interaction in terms of an auxiliary basis:

$$W_{0,PQ}(\omega) = \left[ \mathbf{J} (\mathbf{I} - \mathbf{\Pi}(\mathbf{q}, \omega))^{-1} \right]_{PQ}$$
 (1.49)

and then we need to contract with the Choleskies to get the matrix element:

$$(n_{\mathbf{k}}o_{\mathbf{k}-\mathbf{q}}|W_0|o_{\mathbf{k}-\mathbf{q}}n_{\mathbf{k}}') = \sum_{PQ} v_P^{nm} \left[ \mathbf{I} - \mathbf{\Pi} \left( \mathbf{q}, \omega \right) \right]_{PQ}^{-1} v_Q^{mn'}$$
(1.50)

So in our quest to find poles of  $W_0$ , we are really just looking for poles of the  $\chi_0$ .  $\chi_0$  is given by:

$$\chi_0(\mathbf{r}, \mathbf{r}', \omega) = \sum_{r\mathbf{k}s\mathbf{k}'} (f_{r\mathbf{k}} - f_{s\mathbf{k}'}) \frac{\psi_{r\mathbf{k}}(\mathbf{r})\psi_{r\mathbf{k}}^*(\mathbf{r}')\psi_{s\mathbf{k}'}(\mathbf{r}')\psi_{s\mathbf{k}'}^*(\mathbf{r})}{\omega - (\epsilon_{r\mathbf{k}} - \epsilon_{s\mathbf{k}'}) + i\eta \operatorname{sgn}(\epsilon_{r\mathbf{k}} - \epsilon_{s\mathbf{k}'} - \mu)}$$
(1.51)

where the occupations of the KS states  $r\mathbf{k}(s\mathbf{k}')$  with energies  $\epsilon_{r\mathbf{k}}(\epsilon_{s\mathbf{k}'})$  are given by the Fermi-Dirac distribution  $f_{r\mathbf{k}}(f_{s\mathbf{k}'})$ , which is just a step function at zero temperature. Notice that the occupation factor will always be 0 unless rs form an occupied-virtual pair. So we can separate the density response into two terms, one where  $\delta_{ri}$  and  $\delta_{sa}$  and the other with  $\delta_{ra}$  and  $\delta_{si}$ , where i and a are occupied and virtual indices, respectively. This allows us to now combine with the bare Coulomb potential in order to form the polarizability  $\Pi \equiv \chi_0 v$  as:

$$\Pi\left(\mathbf{r}, \mathbf{r}', \omega\right) = \sum_{i\mathbf{k}a\mathbf{k}'} \frac{\psi_{i\mathbf{k}}(\mathbf{r})\psi_{i\mathbf{k}}^{*}\left(\mathbf{r}'\right)\frac{1}{|\mathbf{r}-\mathbf{r}'|}\psi_{a\mathbf{k}'}\left(\mathbf{r}'\right)\psi_{a\mathbf{k}'}^{*}(\mathbf{r})}{\omega + (\Omega_{i\mathbf{k}a\mathbf{k}'}) - i\eta} - \sum_{a\mathbf{k}i\mathbf{k}'} \frac{\psi_{a\mathbf{k}}(\mathbf{r})\psi_{a\mathbf{k}}^{*}\left(\mathbf{r}'\right)\frac{1}{|\mathbf{r}-\mathbf{r}'|}\psi_{i\mathbf{k}'}\left(\mathbf{r}'\right)\psi_{i\mathbf{k}'}^{*}(\mathbf{r})}{\omega - (\Omega_{i\mathbf{k}a\mathbf{k}'}) + i\eta},$$

$$(1.52)$$

where we define the KS eigenvalue differences as  $\Omega_{i\mathbf{k}a\mathbf{k}'} = \epsilon_{a\mathbf{k}} - \epsilon_{i\mathbf{k}'}$ , which will eventually become the excitation energies from RPA. So sandwiching this operator in between the molecular or brutal bases gives:

$$\langle n\mathbf{k}|\Pi(\omega)|n'\mathbf{k}\rangle = \sum_{iaib\mathbf{k}\mathbf{k}'} \frac{(ia\mid jb)}{(\omega + \Omega_{\mathbf{k}}^{\mu}) - i\eta} - \sum_{aibi\mathbf{k}\mathbf{k}'} \frac{(ai\mid bj)}{(\omega - \Omega_{\mathbf{k}}^{\mu}) + i\eta}$$
(1.53)

So we see that we can get the poles of the screened Coulomb interaction by the poles of the polarizability, which are  $\omega = \Omega_{\mathbf{k}}^{\mu} - i\eta$  and  $\omega = \Omega_{\mathbf{k}}^{\mu} + i\eta$ , suggesting that they are in the upper complex plane for excitations and vice versa for deexcitations. See the figure 1.1 for a picture. For a more comprehensive picture, this should be juxtaposed with the figure from Tianyu's paper for CD. In the literature, they talk about approximating the dielectric function by a multiple one or a single pole approximation, so which one would I want to implement? This suggests that the notation in the  $G_0W_0$  literature is confusing because they always say that to solve for the  $\chi_0$  in the RPA, but if we are actually dealing with  $\chi_0$ , which is the Kohn-Sham density response function, then we don't use the RPA, where the density response function is solved for using a Dyson-like equation [?]:



Figure 1.1: Contour for the complex frequency integral. The poles are denoted by the various  $\omega$ . The Fermi energy is denoted by  $\varepsilon_F$ . The integration contour  $D_+$  is the semicircle in the upper complex plane, while  $D_-$  is the semicircle in the lower complex plane.

$$\chi^{\lambda}\left(\mathbf{r},\mathbf{r}',i\omega\right) = \chi^{0}\left(\mathbf{r},\mathbf{r}',i\omega\right) + \int d\mathbf{r}_{1}d\mathbf{r}_{2}\chi^{0}\left(\mathbf{r},\mathbf{r}_{1},i\omega\right) \left[\frac{\lambda}{|\mathbf{r}_{1}-\mathbf{r}_{2}|} + f_{xc}^{\lambda}\left(\mathbf{r}_{1},\mathbf{r}_{2},i\omega\right)\right]\chi^{\lambda}\left(\mathbf{r}_{2},\mathbf{r}',\omega\right)$$

$$(1.54)$$

where the parameter  $\lambda$  controls the amount of interaction in the system, ranging from  $\lambda=0$  for the KS reference system to  $\lambda=1$  for the fully interacting system. The  $f_{\rm xc}^{\lambda}$  is the exchange-correlation kernel, which is set to zero for the RPA. But we will proceed with an RPA calculation anyways in order to solve for the excitation energies and their corresponding eigenvectors. So it makes sense that the numerator of the expression for the screened Coulomb interaction should be given a construction of the ERIs with the excitation factors in a transition density defined as:

$$w_{pq}^{\mu} = \sum_{ia} (pq|ia) \left( X_{ia}^{\mu} + Y_{ai}^{\mu} \right) \tag{1.55}$$

where we have defined the excitation and de-excitation vectors at the excitation index  $\mu$  as  $X_{ia}^{\mu}$  and  $Y_{ai}^{\mu}$ , respectively. I am not sure how to connect this with the known expression  $v\epsilon^{-1}$ ; I see the similarities given that we are contracting an ERI with what we get from the RPA calculation that is connected to the polarizability, but can't connect exactly. We want to figure out how this matches with my previous  $O(N^6)$  expression, which was

$$\Sigma_{pp}^{\text{corr}}(\omega) = \sum_{\mu}^{\text{RPA}} \left( \sum_{i}^{\text{occupied}} \frac{w_{pi}^{\mu} w_{ip}^{\mu}}{\omega - (\epsilon_{i} - \Omega_{\mu})} + \sum_{a}^{\text{virtual}} \frac{w_{pa}^{\mu} w_{ap}^{\mu}}{\omega - (\epsilon_{a} + \Omega_{\mu})} \right)$$
(1.56)

for the molecular case. Today I want us to dissect how this equation came about, so that I can understand for my k-point version.

### 1.2.2 Analytic continuation

We start with the original form for the self-energy along the real axis:

$$\Sigma(\mathbf{r}, \mathbf{r}', \omega) = \frac{i}{2\pi} \int_{-\infty}^{\infty} d\omega' e^{i\omega'\eta} G_0(\mathbf{r}, \mathbf{r}', \omega + \omega') W_0(\mathbf{r}, \mathbf{r}', \omega')$$
(1.57)

But to avoid the poles, we need to evaluate along the imaginary axis, so the problem becomes:

$$\Sigma(\mathbf{r}, \mathbf{r}', i\omega) = -\frac{1}{2\pi} \int_{-\infty}^{\infty} d\omega' G_0(\mathbf{r}, \mathbf{r}', i\omega + i\omega') W_0(\mathbf{r}, \mathbf{r}', i\omega')$$

We are interested in evaluating the matrix elements of this operator in the molecular orbital basis. Note that both molecular orbitals must have the same crystal momentum in order for it to be conserved in this process. We also apply the identity operator:

$$\langle n\mathbf{k}| \Sigma(i\omega) | n'\mathbf{k} \rangle = -\frac{1}{2\pi} \sum_{m\mathbf{k}'} \int_{-\infty}^{\infty} d\omega' \langle n\mathbf{k}| G_0(i\omega + i\omega') | m\mathbf{k}' \rangle \langle m\mathbf{k}'| W_0(i\omega') | n'\mathbf{k} \rangle \qquad (1.58)$$

The noninteracting Green's function has the form:

$$G_0(\mathbf{r}, \mathbf{r}', i\omega) = \sum_{m\mathbf{k}_m} \frac{\psi_{m\mathbf{k}_m}(\mathbf{r})\psi_{m\mathbf{k}_m}^*(\mathbf{r}')}{i\omega + \epsilon_F - \epsilon_{m\mathbf{k}_m}} \implies G_0(\mathbf{k} - \mathbf{q}, i\omega + i\omega') = \sum_{m\mathbf{k} - \mathbf{q}} \frac{\psi_{m\mathbf{k} - \mathbf{q}}\psi_{m\mathbf{k} - \mathbf{q}}^*}{i(\omega + \omega') + \epsilon_F - \epsilon_{m\mathbf{k} - \mathbf{q}}}$$

so that the above equation simplifies to:

$$\Sigma_{nn'}(\mathbf{k}, i\omega) = -\frac{1}{2\pi N_{\mathbf{k}}} \sum_{m\mathbf{q}} \int_{-\infty}^{\infty} d\omega' \frac{(n\mathbf{k}, m\mathbf{k} - \mathbf{q} \mid W_0(\mathbf{q}, i\omega) \mid m\mathbf{k} - \mathbf{q}, n'\mathbf{k})}{i(\omega + \omega') + \epsilon_F - \epsilon_{m\mathbf{k} - \mathbf{q}}}$$
(1.59)

#### **Screened Coulomb Interaction**

$$(n\mathbf{k}, m\mathbf{k} - \mathbf{q} | W_0(\mathbf{q}, i\omega) | m\mathbf{k} - \mathbf{q}, n'\mathbf{k}) = \int \int d\mathbf{r}_1 d\mathbf{r}_2 \psi_{n\mathbf{k}}^*(\mathbf{r}_1) \psi_{m\mathbf{k} - \mathbf{q}}(\mathbf{r}_1) W_0(\mathbf{q}, \mathbf{r}_1, \mathbf{r}_2, i\omega) \psi_{m\mathbf{k} - \mathbf{q}}^*(\mathbf{r}_2) \psi_{n'\mathbf{k}}(\mathbf{r}_2)$$

We expand the orbital pair product  $\psi_{n\mathbf{k}}^*(\mathbf{r})\psi_{m\mathbf{k}-\mathbf{q}}(\mathbf{r})$  in the auxiliary basis

$$\psi_{n\mathbf{k}}^*(\mathbf{r})\psi_{m\mathbf{k}-\mathbf{q}}(\mathbf{r}) = \sum_{P} b_{P\mathbf{q}}^{n\mathbf{k},m\mathbf{k}-\mathbf{q}}\phi_{P\mathbf{q}}(\mathbf{r})$$

and

$$\psi_{m\mathbf{k}-\mathbf{q}}^*(\mathbf{r})\psi_{n'\mathbf{k}}(\mathbf{r}) = \sum_{Q} b_{Q(-\mathbf{q})}^{m\mathbf{k}-\mathbf{q},n'\mathbf{k}} \phi_{Q(-\mathbf{q})}(\mathbf{r})$$
(1.60)

where we have recognized the fact that in the former there is a momentum transfer of  $\mathbf{q}$ , and in the latter, there is a momentum transfer of  $-\mathbf{q}$ . Substituting in gives

$$(n\mathbf{k}, m\mathbf{k} - \mathbf{q} | W_0(\mathbf{q}, i\omega) | m\mathbf{k} - \mathbf{q}, n'\mathbf{k})$$
 (1.61)

$$= \sum_{PQ} b_{P\mathbf{q}}^{n\mathbf{k},m\mathbf{k}-\mathbf{q}} \left[ \iint d\mathbf{r}_1 d\mathbf{r}_2 \phi_{P\mathbf{q}}(\mathbf{r}_1) W_0(\mathbf{q},\mathbf{r}_1,\mathbf{r}_2,i\omega) \phi_{Q(-\mathbf{q})}(\mathbf{r}_2) \right] b_{Q(-\mathbf{q})}^{m\mathbf{k}-\mathbf{q},n'\mathbf{k}}$$
(1.62)

with

$$b_{P\mathbf{q}}^{n\mathbf{k},m\mathbf{k}-\mathbf{q}} = \sum_{R} (n\mathbf{k}, m\mathbf{k} - \mathbf{q} \mid R\mathbf{q}) \cdot \mathbf{J}_{RP}^{-1}(\mathbf{q})$$
(1.63)

Now is a good place to recall their definition of the density fitting where the ERIs are represented as:

$$\left(p\mathbf{k}_{p}q\mathbf{k}_{q}\mid r\mathbf{k}_{r}s\mathbf{k}_{s}\right)=\sum_{PQ}\left(p\mathbf{k}_{p}q\mathbf{k}_{q}\mid P\mathbf{k}_{pq}\right)\mathbf{J}_{PQ}^{-1}\left(Q\mathbf{k}_{rs}\mid r\mathbf{k}_{r}s\mathbf{r}_{s}\right),$$

with

$$\mathbf{J}_{PQ}(\mathbf{k}) = \iint d\mathbf{r} d\mathbf{r}' \phi_{P(-\mathbf{k})}(\mathbf{r}) \frac{1}{|\mathbf{r} - \mathbf{r}'|} \phi_{Q\mathbf{k}}(\mathbf{r}'),$$

$$(Q\mathbf{k}_{rs} \mid r\mathbf{k}_{r} s \mathbf{k}_{s}) = \iint d\mathbf{r} d\mathbf{r}' \phi_{Q\mathbf{k}_{rs}}(\mathbf{r}) \frac{1}{|\mathbf{r} - \mathbf{r}'|} \phi_{r\mathbf{k}_{r}}^{*}(\mathbf{r}') \phi_{s\mathbf{k}_{s}}(\mathbf{r}').$$

Note that these b are then not yet our Cholesky vectors, since each one contains  $\frac{|\mathbf{r}-\mathbf{r}'|}{|\mathbf{r}-\mathbf{r}'|}$  scaling, i.e., there should be a factor of  $\mathbf{J}^{-\frac{1}{2}}$  instead of  $\mathbf{J}^{-1}$  in 1.63 if we are to apply the Cholesky vectors. At this point, we use the expansion of the screened Coulomb interaction:

$$W_0 = v + v\chi_0 v + v\chi_0 v\chi_0 v + \cdots \tag{1.64}$$

$$= v(1 + \chi_0 v + \chi_0 v \chi_0 v + \cdots) \tag{1.65}$$

$$= v^{1/2} (1 - \chi_0)^{-1} v^{1/2} \tag{1.66}$$

simplifying to

$$(n\mathbf{k}, m\mathbf{k} - \mathbf{q} | W_0(\mathbf{q}, i\omega) | m\mathbf{k} - \mathbf{q}, n'\mathbf{k}) = \sum_{PQ} b_{P\mathbf{q}}^{n\mathbf{k}, m\mathbf{k} - \mathbf{q}} \left[ \mathbf{J}^{\frac{1}{2}} \left( \mathbf{I} - \mathbf{\Pi}(\mathbf{q}, i\omega) \right) \mathbf{J}^{\frac{1}{2}} \right]_{PQ}^{-1} b_{Q(-\mathbf{q})}^{m\mathbf{k} - \mathbf{q}, n'\mathbf{k}}$$

$$(1.67)$$

$$= \sum_{PQ} v_P^{nm} \left[ \mathbf{I} - \mathbf{\Pi} \left( \mathbf{q}, i\omega' \right) \right]_{PQ}^{-1} v_Q^{mn'}$$
 (1.68)

where  $\mathbf{J}_{PQ}$  is the Coulomb interaction projected onto the auxiliary basis, and we have defined

$$v_{P\mathbf{q}}^{n\mathbf{k},m\mathbf{k}-\mathbf{q}} = \sum_{pq} C_{pn}(\mathbf{k}) C_{qm}(\mathbf{k} - \mathbf{q}) v_{P\mathbf{q}}^{p\mathbf{k},q\mathbf{k}-\mathbf{q}}$$
(1.69)

with

$$v_{P\mathbf{q}}^{\mathbf{pk},q\mathbf{k}-\mathbf{q}} = \sum_{R} (p\mathbf{k}, q\mathbf{k} - \mathbf{q} \mid R\mathbf{q}) \mathbf{J}_{RP}^{-1/2}(\mathbf{q})$$
(1.70)

If we first rename  $\mathbf{k}' = \mathbf{k} - \mathbf{q} \implies \mathbf{k} = \mathbf{k}' + \mathbf{q}$ , and then we are free to redefine  $\mathbf{q} \to -\mathbf{q}$ , so that 1.69 becomes

$$v_{P-\mathbf{q}}^{n\mathbf{k}-\mathbf{q},m\mathbf{k}} = \sum_{pq} C_{pn}(\mathbf{k} - \mathbf{q}) C_{qm}(\mathbf{k}) v_{P-\mathbf{q}}^{p\mathbf{k}-\mathbf{q},q\mathbf{k}}$$
(1.71)

but we know that the bare Coulomb potential projected onto the auxiliary basis is given by

$$v_{P\mathbf{q}}^{n\mathbf{k},m\mathbf{k}-\mathbf{q}} = \sum_{pq} C_{pn}(\mathbf{k}) C_{qm}(\mathbf{k} - \mathbf{q}) v_{P\mathbf{q}}^{p\mathbf{k},q\mathbf{k}-\mathbf{q}}$$
(1.72)

To ease notation, some momentum labels are suppressed in the above and following equations (e.g., we will use  $b_P^{nm}$  to denote  $b_{P\mathbf{q}}^{n\mathbf{k},m\mathbf{k}-\mathbf{q}}$ ). Using Eqs. 19-21, the matrix elements of  $W_0$  are computed as

$$(n\mathbf{k}, m\mathbf{k} - \mathbf{q} | W_0 | m\mathbf{k} - \mathbf{q}, n'\mathbf{k})$$

$$= \sum_{PQ} b_P^{nm} \left[ \iint d\mathbf{r} d\mathbf{r}' \phi_{P\mathbf{q}}(\mathbf{r}) W_0(\mathbf{r}, \mathbf{r}', i\omega') \phi_{Q(-\mathbf{q})}(\mathbf{r}') \right] b_Q^{mn'}$$

$$= \sum_{PQ} b_P^{nm} \left[ \mathbf{J}_{PQ}(\mathbf{q}) + \left( \mathbf{J}^{1/2} \mathbf{\Pi} \mathbf{J}^{1/2} \right)_{PQ}(\mathbf{q}) + \dots \right] b_Q^{mn'}$$

$$= \sum_{PQ} v_P^{nm} \left[ \mathbf{I} - \mathbf{\Pi} \left( \mathbf{q}, i\omega' \right) \right]_{PQ}^{-1} v_Q^{mn'}$$

The 3-center 2-electron integral  $v_P^{nm}$  between auxiliary basis function P and molecular orbital pairs nm is obtained from an AO to MO transformation of the GDF AO integrals defined in Eq. 15:

$$v_P^{nm} = \sum_p \sum_q C_{pn}(\mathbf{k}) C_{qm}(\mathbf{k} - \mathbf{q}) v_{P\mathbf{q}}^{p\mathbf{k},q\mathbf{k} - \mathbf{q}}$$

where  $C(\mathbf{k})$  refers to the MO coefficients in the AO basis.  $\Pi(\mathbf{q}, i\omega')$  in Eq. 22 is an auxiliary density response function:

$$\Pi_{PQ}(\mathbf{q}, i\omega') = \frac{2}{N_{\mathbf{k}}} \sum_{\mathbf{k}} \sum_{i}^{\text{occ}} \sum_{a}^{\text{vir}} v_{P}^{ia} \frac{\epsilon_{i\mathbf{k}} - \epsilon_{a\mathbf{k} - \mathbf{q}}}{\omega'^{2} + (\epsilon_{i\mathbf{k}} - \epsilon_{a\mathbf{k} - \mathbf{q}})^{2}} v_{Q}^{ai}$$

### 1.3 UHF formalism

The first thing to do is to solve the Casida equation for the polarizability in the direct formulation of the RPA:

$$\begin{pmatrix} \mathbf{A} & \mathbf{B} \\ \mathbf{B}^* & \mathbf{A}^* \end{pmatrix} \begin{pmatrix} \mathbf{X} \\ \mathbf{Y} \end{pmatrix} = \begin{pmatrix} \Omega & 0 \\ 0 & -\Omega \end{pmatrix} \begin{pmatrix} \mathbf{X} \\ \mathbf{Y} \end{pmatrix} \tag{1.73}$$

with **A** and **B** given by

$$\mathbf{A}_{ia,jb}^{\sigma\sigma'} = \delta_{ij}\delta_{ab}\delta_{\sigma\sigma'}(\varepsilon_a - \varepsilon_i) + (i_{\sigma}a_{\sigma}|b_{\sigma'}j_{\sigma'})$$

$$\mathbf{B}_{ia,jb}^{\sigma\sigma'} = (i_{\sigma}a_{\sigma}|j_{\sigma'}b_{\sigma'})$$
(1.74)

Therefore, with the different spins we form a super matrix:

$$\begin{pmatrix}
\begin{pmatrix}
\mathbf{A}_{\alpha\alpha} & \mathbf{A}_{\alpha\beta} \\
\mathbf{A}_{\beta\alpha} & \mathbf{A}_{\beta\beta}
\end{pmatrix} & \begin{pmatrix}
\mathbf{B}_{\alpha\alpha} & \mathbf{B}_{\alpha\beta} \\
\mathbf{B}_{\beta\alpha} & \mathbf{B}_{\beta\beta}
\end{pmatrix} \\
\begin{pmatrix}
\mathbf{B}_{\alpha\alpha}^* & \mathbf{B}_{\alpha\beta}^* \\
\mathbf{B}_{\alpha\alpha}^* & \mathbf{B}_{\alpha\beta}^* \\
\mathbf{B}_{\beta\alpha}^* & \mathbf{B}_{\beta\beta}^*
\end{pmatrix} & \begin{pmatrix}
\mathbf{A}_{\alpha\alpha}^* & \mathbf{A}_{\alpha\beta}^* \\
\mathbf{A}_{\beta\alpha}^* & \mathbf{A}_{\beta\beta}^* \\
\end{pmatrix} \end{pmatrix} \begin{pmatrix}
\mathbf{X}_{\alpha\alpha} & \mathbf{X}_{\alpha\beta} \\
\mathbf{X}_{\beta\alpha} & \mathbf{X}_{\beta\beta} \\
\mathbf{Y}_{\alpha\alpha} & \mathbf{Y}_{\alpha\beta} \\
\mathbf{Y}_{\beta\alpha} & \mathbf{Y}_{\beta\beta}
\end{pmatrix} = \begin{pmatrix}
\Omega & 0 & 0 & 0 \\
0 & \Omega & 0 & 0 \\
0 & 0 & -\Omega & 0 \\
0 & 0 & 0 & -\Omega
\end{pmatrix} \begin{pmatrix}
\mathbf{X}_{\alpha\alpha} & \mathbf{X}_{\alpha\beta} \\
\mathbf{X}_{\beta\alpha} & \mathbf{X}_{\beta\beta} \\
\mathbf{Y}_{\alpha\alpha} & \mathbf{Y}_{\alpha\beta} \\
\mathbf{Y}_{\beta\alpha} & \mathbf{Y}_{\beta\beta}
\end{pmatrix}$$

$$(1.75)$$

This implies a way to get the excitation energies  $\Omega^{\mu}$  and the eigenvectors  $\mathbf{X}^{\mu}$  and  $\mathbf{Y}^{\mu}$  for a certain spin channel  $\sigma$ . Next, for each spin channel, we need to formulate the matrix  $\mathbf{M}^{\mu}$ , which is used to form the transition densities.

$$M_{iajb}^{\mu} = X_{ia}^{\mu} X_{jb}^{\mu} + X_{ia}^{\mu} Y_{jb}^{\mu} + Y_{ia}^{\mu} X_{jb}^{\mu} + Y_{ia}^{\mu} Y_{jb}^{\mu}$$

$$\tag{1.76}$$

With these quantities, we can then form the self energy for the given spin channel:

$$\Sigma_{pq}^{c}(\omega) = \sum_{jbkc} \sum_{\mu} \left( \sum_{i} \frac{(ip \mid jb)(iq \mid kc)}{\omega - \Omega_{\mu} - \varepsilon_{i}^{\text{MF}} - i\eta} + \sum_{a} \frac{(ap \mid jb)(aq \mid kc)}{\omega + \Omega_{\mu} - \varepsilon_{a}^{\text{MF}} + i\eta} \right) M_{jbkc}^{\mu}$$
(1.77)

But in solving the case partial equation, we are just interested in the real, diagonal part of the self energy, so this reduces to:

$$\Sigma_{pp}^{c}(\omega) = \sum_{jbkc} \sum_{\mu} \left( \sum_{i} \frac{(ip \mid jb)(ip \mid kc)}{\omega - \Omega_{\mu} - \varepsilon_{i}^{\text{MF}}} + \sum_{a} \frac{(ap \mid jb)(ap \mid kc)}{\omega + \Omega_{\mu} - \varepsilon_{a}^{\text{MF}}} \right) M_{jbkc}^{\mu}$$
(1.78)

## 1.4 Deriving linear response: 11/22

### 1.4.1 The Fundamentals

The motivation for this is to be able to understand why the poles of the screened Coulomb interaction are the same as those of the fully interacting polarizability, which are given by the frequencies of the RPA, obtained by diagonalizing the Casida equation. And then we want to be able to connect why  $W_0 = v + v\chi_0 v + \ldots = \frac{v}{1-\chi_0 v}$  where  $W_0$  is the screened Coulomb interaction and  $\chi_0$  is the non-interacting polarizability with Lehmann representation.

$$\chi_0(\mathbf{r}, \mathbf{r}', \omega) = \sum_{ia} \frac{\psi_i(\mathbf{r})\psi_a^*(\mathbf{r}')\psi_i(\mathbf{r}')\psi_a^*(\mathbf{r})}{\omega - (\epsilon_a - \epsilon_i) + i\eta \operatorname{sgn}(\epsilon_a - \epsilon_i - \mu)}$$
(1.79)

to do so, one must understand the reformulation of based on the density matrix as for posed by Furche [?]. Alternatively, let us start from the known Dyson equation that relates the fully interacting Green's function to the non-interacting one. We know the integral form of the Dyson equation is

$$G(1,2) = G_0(1,2) + \int d3d4G_0(1,3)\Sigma(3,4)G(4,2)$$
(1.80)

but we proceed symbolically to get

$$G = G_0 + G_0 \Sigma G \tag{1.81}$$

$$(I - G_0 \Sigma) G = G_0 \tag{1.82}$$

$$G = (I - G_0 \Sigma)^{-1} G_0 \tag{1.83}$$

$$G = (G_0 (G_0^{-1} - \Sigma))^{-1} G_0$$
 (1.84)

$$G = (G_0^{-1} - \Sigma)^{-1} \tag{1.85}$$

Now, we also know the Dyson equation for the polarizability is

$$\chi(\omega, x_1, x_2) = \chi_0(\omega, x_1, x_2) + \int dx dx' \chi_0(\omega, x_1, x)$$

$$\times \left(\frac{1}{|\mathbf{r} - \mathbf{r}'|} + f_{xc}(\omega, x, x')\right) \chi(\omega, x', x_2)$$
(1.86)

In the RPA, we neglect the exchange correlation kernel  $f_{xc}$ , so we have

$$\chi_{RPA} = \chi_0 + \chi_0 v \chi_{RPA} = \frac{\chi_0}{1 - v \chi_0} \tag{1.87}$$

where in the final step we used the symbolic manipulation that was used before. So  $W_0 = \frac{v}{1-v\chi_0} = v\chi_{RPA}\chi_0^{-1} = v\left(\frac{\chi_0\chi_0^{-1}}{1-v\chi_0}\right) = \frac{v}{1-v\chi_0}$ . Therefore, we see why the poles of  $W_0$  are the same as those of  $\chi_{RPA}$ , since they have a linear relationship. Now, we will proceed to derive the poles of  $\chi_{RPA}$ . But first we must introduce the density matrix based linear response theory.

#### 1.4.2 Introduction to TDKS

In time-dependent density functional theory (TDDFT), the **Time-Dependent Kohn-Sham (TDKS)** equations describe a system of N noninteracting fermions that reproduce the same time-dependent density  $\rho(t, \mathbf{r})$  as the interacting system. The TDKS equations are given by:

$$i\frac{\partial}{\partial t}\varphi_j(t,\mathbf{r}) = H[\rho](t,\mathbf{r})\varphi_j(t,\mathbf{r})$$
 (1.88)

where  $j=1,\ldots,N$  indexes the Kohn-Sham orbitals  $\varphi_j(t,\mathbf{r})$ , and  $H[\rho](t,\mathbf{r})$  is the effective one-particle Hamiltonian defined as:

$$H[\rho](t,\mathbf{r}) = \frac{\boldsymbol{\pi}^2(t,\mathbf{r})}{2} + v_{\text{eff}}[\rho](t,\mathbf{r})$$
(1.89)

where  $v_{\text{eff}}[\rho](t, \mathbf{r}) = v_{\text{ext}}(t, \mathbf{r}) + v_{\text{H}}[\rho](t, \mathbf{r}) + v_{\text{xc}}[\rho](t, \mathbf{r}).$ 

The operator  $\pi(t, \mathbf{r})$  is known as the **kinetic momentum operator**. In the presence of an electromagnetic field, the kinetic momentum operator is modified from the canonical momentum operator  $\mathbf{p}$  to include the effects of the vector potential  $\mathbf{A}_{\text{ext}}(t, \mathbf{r})$ :

$$\pi(t, \mathbf{r}) = \mathbf{p} + \frac{1}{c} \mathbf{A}_{\text{ext}}(t, \mathbf{r})$$
 (1.90)

Here,  $\mathbf{p} = -i\hbar\nabla$  is the canonical momentum operator, and c is the speed of light. The vector potential  $\mathbf{A}_{\rm ext}(t,\mathbf{r})$  accounts for the influence of external perturbative electromagnetic fields on the system. Why it does the influence of the vector potential not just all go into the  $v_{\rm eff}$ ?

But we know that under a gauge transformation, the physical observables will be invariant, while the orbits will merely acquire a **phase factor**:

$$\varphi_j(t, \mathbf{r}) \to \varphi'_j(t, \mathbf{r}) = \varphi_j(t, \mathbf{r}) \exp\left(-\frac{i}{c}\psi(t, \mathbf{r})\right)$$
 (1.91)

Therefore observables, like the density or current density will be unaffected by this gauge transformation.

### 1.4.3 Density Matrix Formulation in TDKS Theory

In Time-Dependent Density Functional Theory (TDDFT), the **Time-Dependent Kohn-Sham (TDKS)** equations 1.88 describe a system of N noninteracting fermions that reproduce the same time-dependent electron density  $\rho(t, \mathbf{r})$  as the interacting system. An alternative formulation of TDKS theory utilizes the one-particle density matrix  $\gamma(t, \mathbf{r}, \mathbf{r}')$ , which offers advantage because it introduces a basis that one can exploit computationally. The one-particle density matrix is defined as:

$$\gamma(t, \mathbf{r}, \mathbf{r}') = \sum_{j=1}^{N} \varphi_j(t, \mathbf{r}) \varphi_j^*(t, \mathbf{r}')$$
(1.92)

and it is idempotent, meaning that

$$\gamma^2(t, \mathbf{r}, \mathbf{r}') = \gamma(t, \mathbf{r}, \mathbf{r}') \tag{1.93}$$

See section 5.1 for a proof.

Now, we would like to derive an evolution equation for the density matrix in analogy with the one we already have for the KS orbitals in equation 1.88. The result is

$$i\frac{\partial}{\partial t}\gamma(t) = [H[\rho](t), \gamma(t)]$$
 (1.94)

See section 5.2 for proof. For the purposes of response theory, it is convenient to consider external scalar potentials,

$$v_{\text{ext}}(t,x) = v^{(0)}(x) + \sum_{\alpha} \lambda_{\alpha} \left( v^{(\alpha)}(\omega_{\alpha}, x) e^{i\omega_{\alpha}t} + v^{(\alpha)}(-\omega_{\alpha}, x) e^{-i\omega_{\alpha}t} \right)$$
(1.95)

and longitudinal vector potentials

$$\mathbf{A}_{\text{ext}}(t,x) = \sum_{\alpha} \lambda_{\alpha} \left( \mathbf{A}^{(\alpha)} \left( \omega_{\alpha}, x \right) e^{i\omega_{\alpha}t} + \mathbf{A}^{(\alpha)} \left( -\omega_{\alpha}, x \right) e^{-i\omega_{\alpha}t} \right)$$
(1.96)

Note that we are not considering transverse vector potentials, as we would get if we had a magnetic field. The cemetery of the Fourier component is fate in section 5.3. Next, we need to determine the derivatives of the observables with respect to a perturbation. We have Any time-dependent expectation value  $f_{\lambda}(t)$  is a function of the coupling strength vector  $\lambda$ . Its response to the external perturbation is defined by its derivatives with respect to  $\lambda$  at  $\lambda = 0$ . For monochromatic perturbations, the derivatives exhibit a characteristic time dependence,

$$f_{\lambda}(t)|_{\lambda=0} = f^{(0)},$$
 (1.97)

$$\left. \frac{\partial}{\partial \lambda_{\alpha}} f_{\lambda}(t) \right|_{\lambda=0} = f^{(\alpha)} \left( \omega_{\alpha} \right) e^{i\omega_{\alpha}t} + f^{(\alpha)} \left( -\omega_{\alpha} \right) e^{-i\omega_{\alpha}t}, \tag{1.98}$$

(1.99)

Proof of the second expression is given in section 5.4. The third expression follows by taking the derivative of the second expression.

$$\frac{\partial^2}{\partial \lambda_{\alpha} \partial \lambda_{\beta}} f_{\lambda}(t) \bigg|_{\lambda=0} = f^{(\alpha\beta)} (\omega_{\alpha}, \omega_{\beta}) e^{i(\omega_{\alpha} + \omega_{\beta})t} + f^{(\alpha\beta)} (\omega_{\alpha}, -\omega_{\beta}) e^{i(\omega_{\alpha} - \omega_{\beta})t}$$
(1.100)

$$+ f^{(\alpha\beta)} \left( -\omega_{\alpha}, \omega_{\beta} \right) e^{i\left( -\omega_{\alpha} + \omega_{\beta} \right)t} + f^{(\alpha\beta)} \left( -\omega_{\alpha}, -\omega_{\beta} \right) e^{-i\left( \omega_{\alpha} + \omega_{\beta} \right)t}$$
 (1.101)

These expressions define the frequency dependent response of  $f_{\lambda}$  up to second order. The key step is to realize that when evaluating the expectation value of some observable O, we must consider the coupling to the density matrix, i.e.  $f_{\lambda}(t) = \operatorname{tr}(O(t)\gamma_{\lambda}(t))$ .

The route to frequency-dependent response properties is then obvious: (1) Calculate the frequency-dependent KS density matrix response by differentiation of Eqs. (8) and (10); (2) Take the trace with O. The interacting response can be calculated from the noninteracting KS system because the TDKS density matrix yields the interacting density and current density as it follows from Eq. (9). **Better understanding needed.** 

For the idempotency constraint, expansion up to second order yields, in shorthand notation,

$$\gamma^{(0)} = \gamma^{(0)}\gamma^{(0)},\tag{16}$$

$$\gamma^{(\alpha)} = \gamma^{(0)}\gamma^{(\alpha)} + \gamma^{(\alpha)}\gamma^{(0)},\tag{17}$$

$$\gamma^{(\alpha\beta)} = \gamma^{(0)}\gamma^{(\alpha\beta)} + \gamma^{(\alpha)}\gamma^{(\beta)} + \gamma^{(\beta)}\gamma^{(\alpha)} + \gamma^{(\alpha\beta)}\gamma^{(0)}.$$
 (18)

The equations of motion up to second order read

$$0 = \left[H^{(0)}, \gamma^{(0)}\right] \tag{1.102}$$

$$\omega_{\alpha} \gamma^{(\alpha)} = [H^{(0)}, \gamma^{(\alpha)}] + [H^{(\alpha)}, \gamma^{(0)}]$$
(1.103)

$$(\omega_{\alpha} + \omega_{\beta}) \gamma^{(\alpha\beta)} = [H^{(0)}, \gamma^{(\alpha\beta)}] + [H^{(\alpha)}, \gamma^{(\beta)}] + [H^{(\beta)}, \gamma^{(\alpha)}] + [H^{(\alpha\beta)}, \gamma^{(0)}]$$
(1.104)

Proof of this series is given in section 5.5.

## 1.4.4 Proving $\chi_{RPA} = \frac{\chi_0}{1 - v\chi_0}$ : 11/29

### Trying direct evaluation

We know

$$\chi_{RPA}^{-1}(\omega) = \frac{1 - v\chi_0}{\chi_0} = \chi_0^{-1} - \mathbf{v}$$
 (1.105)

The Lehmann representation for  $\chi_0$  is

$$\chi_0(\mathbf{r}, \mathbf{r}', \omega) = \sum_{ia} \frac{\psi_i(\mathbf{r}) \psi_a^*(\mathbf{r}') \psi_i(\mathbf{r}') \psi_a^*(\mathbf{r})}{\omega \operatorname{sgn}(\epsilon_a - \epsilon_i - \mu) + \underbrace{(\epsilon_a - \epsilon_i)}_{KS \operatorname{bare} \Omega_0} + i\eta \operatorname{sgn}(\epsilon_a - \epsilon_i - \mu)}$$
(1.106)

Let's start by considering the right-hand side of equation 1.105. We know that in the particle-hole basis  $\chi_0(\omega) = \chi_0^+(\omega) + \chi_0^-(\omega) = \begin{pmatrix} \chi_0^+(\omega) & 0 \\ 0 & \chi_0^-(\omega) \end{pmatrix}$  is diagonal, where we define  $\chi_0^{\pm}(\omega) = \frac{1}{\pm \omega + [\epsilon_a - \epsilon_i]}$  as the KS excitation/de-excitations polarizabilities.

$$\chi_0^{-1}(\omega) = \begin{pmatrix} \frac{1}{\chi_0^+(\omega)} & 0\\ 0 & \frac{1}{\chi_0^-(\omega)} \end{pmatrix} = \begin{pmatrix} \omega + [\epsilon_a - \epsilon_i] & 0\\ 0 & -\omega + [\epsilon_a - \epsilon_i] \end{pmatrix}$$
(1.107)

The Coulomb interaction in the particle-hole basis is

$$\mathbf{v} = \begin{pmatrix} \mathbf{v}^{++} & \mathbf{v}^{+-} \\ \mathbf{v}^{-+} & \mathbf{v}^{--} \end{pmatrix} \tag{1.108}$$

Note the permutational symmetries, so  $v_{pq,rs}^{++} \equiv (ia|jb) = (ai|bj) \equiv v_{pq,rs}^{--}$  and  $v_{pq,rs}^{+-} \equiv (ia|bj) = (ai|jb) \equiv v_{pq,rs}^{-+}$ . So the RHS of equation 1.105 is

$$\chi_{RPA}^{-1}(\omega) = \chi_0^{-1}(\omega) - \mathbf{v} = \begin{pmatrix} (\omega + [\epsilon_a - \epsilon_i]) - \mathbf{v}^{++} & -\mathbf{v}^{+-} \\ -\mathbf{v}^{-+} & (-\omega + [\epsilon_a - \epsilon_i]) - \mathbf{v}^{--} \end{pmatrix} = \omega \mathbf{\Sigma}_{\mathbf{z}} + \mathbf{M}$$
(1.109)

where

$$\Sigma_{\mathbf{z}} = \begin{pmatrix} \mathbf{I} & 0 \\ 0 & -\mathbf{I} \end{pmatrix} \quad \text{and} \quad \mathbf{M} = \begin{pmatrix} \mathbf{A} & \mathbf{B} \\ \mathbf{B} & \mathbf{A} \end{pmatrix}$$
 (1.110)

where  $A_{ij,ab} = \delta_{ij}\delta_{ab}\left(\epsilon_a - \epsilon_i\right) - (ia|jb)$  and  $B_{ij,ab} = -(ia|bj)$ . So we have found that  $\chi_{RPA}(\omega) = \left[\omega \Sigma_{\mathbf{z}} + \mathbf{M}\right]^{-1}$ . And so we recover

$$\chi_{RPA}(\omega) = \left[ \begin{pmatrix} \mathbf{A} & \mathbf{B} \\ \mathbf{B} & \mathbf{A} \end{pmatrix} + \omega \begin{pmatrix} \mathbf{I} & 0 \\ 0 & -\mathbf{I} \end{pmatrix} \right]^{-1}$$
(1.111)

To forced further, recognize that the matrix  $\omega \Sigma_z + M$  is diagonal in the RPA eigenbasis, so we can write

$$\omega \Sigma_{\mathbf{z}} + \mathbf{M} = \begin{pmatrix} \mathbf{X} \\ \mathbf{Y} \end{pmatrix} \begin{pmatrix} \mathbf{\Omega} - \omega & 0 \\ 0 & \mathbf{\Omega} + \omega \end{pmatrix} \begin{pmatrix} \mathbf{X} \\ \mathbf{Y} \end{pmatrix}^{\dagger}$$
(1.112)

$$\chi_{RPA}(\omega) = (\omega \Sigma_{\mathbf{z}} + \mathbf{M})^{-1} = \begin{pmatrix} \mathbf{X} \\ \mathbf{Y} \end{pmatrix} \begin{pmatrix} \frac{1}{\Omega - \omega} & 0 \\ 0 & \frac{1}{\Omega + \omega} \end{pmatrix} \begin{pmatrix} \mathbf{X} \\ \mathbf{Y} \end{pmatrix}^{\mathsf{T}}$$
(1.113)

$$\chi_{RPA}(\omega) = \sum_{\mathbf{I}} \left[ \frac{1}{\Omega_{\mathbf{I}} + \omega} \begin{pmatrix} X^{\mathbf{I}} \\ Y^{\mathbf{I}} \end{pmatrix} \begin{pmatrix} X^{\mathbf{I}} & Y^{\mathbf{I}} \end{pmatrix} + \frac{1}{\Omega_{\mathbf{I}} - \omega} \begin{pmatrix} Y^{\mathbf{I}} \\ X^{\mathbf{I}} \end{pmatrix} \begin{pmatrix} Y^{\mathbf{I}} & X^{\mathbf{I}} \end{pmatrix} \right]$$
(1.114)

#### How to determine the excitations?

To just determine the locations of the poles, we take a different route here. We have the form:

 $\chi_{RPA} = \chi_0 + \chi_0 v \chi_{RPA} = \frac{\chi_0}{1 - v \gamma_0}$ (1.115)

This implies that we are faced with a matrix inversion problem. The condition for the matrix  $[\mathbf{I} - \mathbf{v}\chi_0(\omega)]$  to be invertible is one this matrix is non-singular; therefore, the poles of the RPA occur where  $[\mathbf{I} - \mathbf{v}\chi_{\mathbf{0}}(\omega)]$  is singular, i.e. where det  $[\mathbf{I} - \mathbf{v}\chi_{\mathbf{0}}(\omega)] = 0$ . This condition implies that we must have a nonzero eigenvector  $\mathbf{F}$  such that

$$[\mathbf{I} - \mathbf{v}\chi_{\mathbf{0}}(\omega)]\mathbf{F} = 0 \implies \mathbf{v}\chi_{\mathbf{0}}(\omega)\mathbf{F} = \mathbf{F}$$
(1.116)

We need to determine what the matrix element of the operator  $\mathbf{v}\chi_0(\omega)$  is in a basis that we will specify later. With the resolution of the identity, we have

$$\langle pq | \mathbf{v} \chi_0(\omega) | rs \rangle = \sum_{tu} \langle pq | \mathbf{v} | tu \rangle \langle tu | \chi_0(\omega) | rs \rangle$$
 (1.117)

But we know that the  $\chi_0$  is diagonal in a particle-hole basis, so we will have

$$\langle \tilde{pq} | \mathbf{v} \chi_0(\omega) | \tilde{rs} \rangle = v_{\tilde{pq}\tilde{rs}} \chi_{0,\tilde{rs}}(\omega)$$
 (1.118)

where it is understood that  $\tilde{pq}$ ,  $\tilde{rs}$  form an occupied-virtual pair. Now consider partitioning  $\chi_0$  into two pieces,  $\chi_0^+$  for OV excitations and  $\chi_0^-$  for VO de-excitations:

$$\chi_0(\omega) = \chi_0^+(\omega) + \chi_0^-(\omega)$$
(1.119)

We know that since  $\chi_0^{\pm}(\omega) = \frac{1}{\pm(\omega - [\epsilon_a - \epsilon_i])} \implies \chi_0^{+}(\omega) = \tilde{\chi}_0(\omega)$  and  $\chi_0^{-}(\omega) = -\tilde{\chi}_0(\omega)$ , where  $\tilde{\chi}_0(\omega) = \frac{1}{\omega - [\epsilon_a - \epsilon_i]}$ . Notice from equation 1.118 that the occupied virtual combination of  $\chi_0$ constrains the second index of  $\mathbf{v}$ , so we can formulate a matrix for equation 1.116 in this pair basis:

$$\begin{pmatrix} \mathbf{v}^{++} \tilde{\chi}_{0} (\omega) & -\mathbf{v}^{+-} \tilde{\chi}_{0} (\omega) \\ \mathbf{v}^{-+} \tilde{\chi}_{0} (\omega) & -\mathbf{v}^{--} \tilde{\chi}_{0} (\omega) \end{pmatrix} \begin{pmatrix} \mathbf{F}^{+} \\ \mathbf{F}^{-} \end{pmatrix} = \begin{pmatrix} \mathbf{F}^{+} \\ \mathbf{F}^{-} \end{pmatrix}$$
(1.120)

This implies the system of equations

$$\left(\mathbf{v}^{++}\tilde{\chi}_{0}\left(\omega\right)-\mathbf{I}\right)\mathbf{F}^{+}-\mathbf{v}^{+-}\tilde{\chi}_{0}\left(\omega\right)\mathbf{F}^{-}=0$$
(1.121)

$$\mathbf{v}^{-+}\tilde{\chi}_{0}\left(\omega\right)\mathbf{F}^{+}-\left(\mathbf{v}^{--}\tilde{\chi}_{0}\left(\omega\right)+\mathbf{I}\right)\mathbf{F}^{-}=0$$
(1.122)

but we can multiply through by  $\tilde{\chi}_0(\omega)^{-1} = \omega - [\epsilon_a - \epsilon_i]$  to yield

$$\mathbf{v}^{++}\mathbf{F}^{+} - \mathbf{v}^{+-}\mathbf{F}^{-} = (\omega - [\epsilon_a - \epsilon_i])\mathbf{F}^{+}$$
(1.123)

$$\mathbf{v}^{-+}\mathbf{F}^{+} - \mathbf{v}^{--}\mathbf{F}^{-} = -\left(\omega - \left[\epsilon_{a} - \epsilon_{i}\right]\right)\mathbf{F}^{-} \tag{1.124}$$

$$\begin{pmatrix} \mathbf{v}^{++} & \mathbf{v}^{+-} \\ \mathbf{v}^{-+} & \mathbf{v}^{--} \end{pmatrix} \begin{pmatrix} \mathbf{F}^{+} \\ -\mathbf{F}^{-} \end{pmatrix} = (\omega - [\epsilon_a - \epsilon_i]) \begin{pmatrix} \mathbf{F}^{+} \\ -\mathbf{F}^{-} \end{pmatrix}$$
(1.125)

$$\begin{pmatrix} \mathbf{v}^{++} & \mathbf{v}^{+-} \\ \mathbf{v}^{-+} & \mathbf{v}^{--} \end{pmatrix} \begin{pmatrix} \mathbf{F}^{+} \\ -\mathbf{F}^{-} \end{pmatrix} = \omega \begin{pmatrix} \mathbf{F}^{+} \\ -\mathbf{F}^{-} \end{pmatrix} - [\epsilon_{a} - \epsilon_{i}] \begin{pmatrix} \mathbf{F}^{+} \\ -\mathbf{F}^{-} \end{pmatrix}$$

$$\begin{pmatrix} [\epsilon_{a} - \epsilon_{i}] + \mathbf{v}^{++} & \mathbf{v}^{+-} \\ \mathbf{v}^{-+} & [\epsilon_{a} - \epsilon_{i}] + \mathbf{v}^{--} \end{pmatrix} \begin{pmatrix} \mathbf{F}^{+} \\ -\mathbf{F}^{-} \end{pmatrix} = \omega \begin{pmatrix} \mathbf{F}^{+} \\ -\mathbf{F}^{-} \end{pmatrix}$$

$$(1.126)$$

$$\begin{pmatrix} [\epsilon_a - \epsilon_i] + \mathbf{v}^{++} & \mathbf{v}^{+-} \\ \mathbf{v}^{-+} & [\epsilon_a - \epsilon_i] + \mathbf{v}^{--} \end{pmatrix} \begin{pmatrix} \mathbf{F}^+ \\ -\mathbf{F}^- \end{pmatrix} = \omega \begin{pmatrix} \mathbf{F}^+ \\ -\mathbf{F}^- \end{pmatrix}$$
(1.127)

where now we recognize A and B with elements in the particle-hole basis as

$$A_{iajb} = \delta_{ij}\delta_{ab} \left(\epsilon_a - \epsilon_i\right) - \left(ia|jb\right) \tag{1.128}$$

$$B_{iajb} = (ia|bj) (1.129)$$

and note that we can make this work because (ia|jb) = (ai|bj) and (ia|bj) = (ai|jb) by permutation symmetry, so we can rewrite the matrix problem as

$$\begin{pmatrix} \mathbf{A} & \mathbf{B} \\ -\mathbf{B} & -\mathbf{A} \end{pmatrix} \begin{pmatrix} \mathbf{F}^{+} \\ \mathbf{F}^{-} \end{pmatrix} = \omega \sigma_{\mathbf{z}} \begin{pmatrix} \mathbf{F}^{+} \\ \mathbf{F}^{-} \end{pmatrix}$$
(1.130)

with Pauli matrix  $\sigma_{\mathbf{z}} = \begin{pmatrix} \mathbf{I} & 0 \\ 0 & -\mathbf{I} \end{pmatrix}$ . So, by solving this eigenvalue problem, we determine the poles of the RPA  $\omega \equiv \mathbf{\Omega}^I$  with excitation vectors  $\begin{pmatrix} \mathbf{F}^+ \\ \mathbf{F}^- \end{pmatrix} \equiv \begin{pmatrix} \mathbf{X}^\mathrm{I} \\ \mathbf{Y}^\mathrm{I} \end{pmatrix}$ .

## Chapter 2

# Dyson to GQME: 12/13

The Generalized Quantum Master Equation (GQME) is given by

$$\dot{\mathcal{C}}(t) = \mathcal{C}(t)\Omega_1 - \int_0^t d\tau \, \mathcal{C}(t-\tau)\mathcal{K}_1(\tau) + D(t)$$
(2.1)

where the correlation function is defined as

$$C(t) = (\hat{\mu} \mid \hat{\mu}(t)), \tag{2.2}$$

the higher-order moments are

$$\Omega_n \equiv \left( (i\mathcal{L})^n \hat{\mu}, \hat{\mu} \right) / (\hat{\mu}, \hat{\mu}) \tag{2.3}$$

with the auxiliary kernels

$$K_n(t) \equiv \left( (i\mathcal{L})^n \hat{f}(t), \hat{\mu} \right) / (\hat{\mu}, \hat{\mu})$$
(2.4)

 $\hat{f}(t)$  in the above equation is referred to as the random force operator

$$\hat{f}(t) \equiv e^{it\mathcal{L}\mathcal{L}} \mathcal{Q}i\mathcal{L}\hat{\mu} \tag{2.5}$$

with  $Q = \mathcal{I} - \mathcal{P}$  being the complementary projection operator. But this becomes complicated, so Wenjie found that we can express  $\mathcal{K}_1(t)$  without time evolution using  $\hat{f}(0) = Qi\mathcal{L}\hat{\mu}$  and, we get

$$K_n(0) = \Omega_{n+1} - \Omega_n \Omega_1 \tag{2.6}$$

Therefore, we only need to consider  $\dot{K}_1(t)$ , which can be obtained directly:

$$\dot{K}_1(t) = \frac{(i\mathcal{L}\dot{\hat{f}}(t), \hat{\mu})}{(\hat{\mu}, \hat{\mu})} = K_2(t) - \Omega_1 K_1(t)$$
(2.7)

Similarly, we can show that the auxiliary kernels are coupled through

$$\dot{K}_n(t) = K_{n+1}(t) - \Omega_n K_1(t). \tag{2.8}$$

We expect that the higher order auxiliary kernels will decay quickly, so we can truncate the series at some finite n. The moments of the memory kernel are

$$\Omega_n = \frac{((i\mathcal{L})^n \hat{\mu}, \hat{\mu})}{(\hat{\mu}, \hat{\mu})}, \tag{2.9}$$

with  $\mathcal{L}$  being the Liouville superoperator with  $\mathcal{L}\hat{\mu} = \left[\hat{H}, \hat{\mu}\right]$ . The construction of the numerator in equation 2.9 can be thought of as the generation of a Krylov subspace up to level n, i.e. we need to build up  $\mathcal{K}_n(\mathcal{L}, \hat{\mu}) = \operatorname{span}\{\hat{\mu}, (i\mathcal{L})\hat{\mu}, (i\mathcal{L})^2\hat{\mu}, \dots, (i\mathcal{L})^{n-1}\hat{\mu}\}$ , where  $\hat{\mu} = \hat{c}$  or  $\hat{c}^{\dagger}$ . In the case if we choose  $\hat{\mu} = \hat{c}$ , we get the lesser Green's function

$$C(t) = (\hat{c}, \hat{c}(t)) \equiv \langle \hat{c}^{\dagger}(0)\hat{c}(t)\rangle = \frac{G^{<}(t)}{i}$$
(2.10)

whereas if we chose  $\hat{\mu} = \hat{c}^{\dagger}$ , we get the greater Green's function

$$C(t) = (\hat{c}^{\dagger}, \hat{c}^{\dagger}(t)) \equiv \langle \hat{c}(0)\hat{c}^{\dagger}(t)\rangle = -\frac{G^{\gt}(t)}{i}$$
(2.11)

Then we can construct the retarded Green's function as

$$G_R(t) = \Theta(t) \left( G^{<}(t) - G^{>}(t) \right) \tag{2.12}$$

Using Krylov subspace methods, one never has to construct the Liouvillian matrix, but instead can directly compute the extremal eigenvalues and eigenvectors of  $\mathcal{L}$  by considering the action of  $\mathcal{L}$  on the Krylov subspace.

# 2.1 Explicit Construction of the Liouville Superoperator

Consider that we are working with the upfolded Hamiltonian

$$\mathbf{H} = \begin{pmatrix} \mathbf{f} & \mathbf{W} \\ \mathbf{W}^{\dagger} & \mathbf{d} \end{pmatrix} \tag{2.13}$$

where we again have a physical space  $\mathbf{f}$  and a bath space  $\mathbf{d}$ , whose coupling is given by  $\mathbf{W}$ . Tell me what would happened if we considered the action of this on the composite operator vector defined lower? Lets consider making a Krylov subspace, corresponding to repeated applications of the Liouville superoperator to the initial operator  $\hat{\boldsymbol{\mu}}$ . Now, The idea is to define a composite operator vector

$$\hat{\boldsymbol{\mu}} \equiv \begin{pmatrix} \hat{\mu}_1 \\ \hat{\mu}_2 \end{pmatrix} = \begin{pmatrix} \hat{c} \\ \hat{c}^{\dagger} \end{pmatrix}, \tag{2.14}$$

where  $\hat{c}$  is the annihilation operator and  $\hat{c}^{\dagger}$  is the creation operator. Notice that the equation of motion for the Green's function is

$$(i\partial_t - \hat{H}_0)G(t, t') = \delta(t - t') + \int_{-\infty}^{\infty} d\tau \, \Sigma(t, \tau) \, G(\tau, t')$$
 (2.15)

$$\rightarrow G(\dot{t}, t') = -i\hat{H}_0 G(t, t') - i\delta(t - t') + \int_{-\infty}^{\infty} d\tau' \, \Sigma(t, \tau') \, G(\tau', t') \tag{2.16}$$

So our task becomes to figure out how

$$-i\hat{H}_0G(t,t') - i\delta(t-t') + \int_{-\infty}^{\infty} d\tau' \,\Sigma(t,\tau') \,G(\tau',t') = \mathcal{C}(t)\Omega_1 - \int_0^t d\tau \,\mathcal{C}(t-\tau)\mathcal{K}(\tau) + D(t) \quad (2.17)$$

I feel like it should be the case that  $-i\hat{H}_0G(t,t') = \mathcal{C}(t)\Omega_1$ . Do you think that this should be the case or no? Because I think we can agree that the equation that comes first should be the same as the equation of motion for the greens function. And then try to apply a decomposed Hamiltonia like  $H = H_0 + V$  to the first equation, so that we can see what happens.

To simple by things as much as possible initially consider that we only use the noninteracting Hamiltonian  $\hat{H}_0 = \epsilon \hat{c}^{\dagger} \hat{c}$  in the action of the Liouvillian.

$$\Omega_1 = \frac{((i\mathcal{L})\hat{\boldsymbol{\mu}}, \hat{\boldsymbol{\mu}})}{(\hat{\boldsymbol{\mu}}, \hat{\boldsymbol{\mu}})} = \frac{\left((i[\hat{H}_0, \hat{\boldsymbol{\mu}}], \hat{\boldsymbol{\mu}}\right)}{(\hat{\boldsymbol{\mu}}, \hat{\boldsymbol{\mu}})}$$
(2.18)

If we just consider the numerator, we see that

$$\left( (i[\hat{H}_0, \hat{\boldsymbol{\mu}}], \hat{\boldsymbol{\mu}}) = \left( i[\hat{H}_0, \hat{c}], \hat{c} \right) + \left( i[\hat{H}_0, \hat{c}^{\dagger}], \hat{c}^{\dagger} \right)$$
(2.19)

Considering just the first term

$$\left(i[\hat{H}_0,\hat{c}],\hat{c}\right) = -i\epsilon([\hat{c}^{\dagger}\hat{c},\hat{c}],\hat{c}) = -i\epsilon(\hat{c},\hat{c}) = -i\epsilon(1 - f(\epsilon))$$
(2.20)

and the second term

$$\left(i[\hat{H}_0,\hat{c}^{\dagger}],\hat{c}^{\dagger}\right) = -i\epsilon([\hat{c}^{\dagger}\hat{c},\hat{c}^{\dagger}],\hat{c}^{\dagger}) = -i\epsilon(\hat{c}^{\dagger},\hat{c}^{\dagger}) = -i\epsilon f(\epsilon)$$
(2.21)

which can be summarized as

$$\left(i[\hat{H}_0,\hat{\boldsymbol{\mu}}],\hat{\boldsymbol{\mu}}\right) = -i\epsilon \implies \Omega_1 = -i\epsilon$$
 (2.22)

Now, the equation of motion for the interacting Green's function is given by

$$\left(i\frac{\partial}{\partial t} - h_0\right)G(t, t') = \delta(t - t') + \int dt'' \Sigma(t, t'')G(t'', t') \tag{2.23}$$

$$\frac{\partial}{\partial t}G(t,t') = \underbrace{-ih_0G(t,t')}_{\Omega_1C(t)} -i\delta(t-t') + \int dt'' \Sigma(t,t'')G(t'',t')$$
(2.24)

Now come if we consider the higher-order moments

$$\Omega_n \equiv \frac{((i\mathcal{L})^n \hat{\boldsymbol{\mu}}, \hat{\boldsymbol{\mu}})}{(\hat{\boldsymbol{\mu}}, \hat{\boldsymbol{\mu}})} = \frac{(i)^n (\mathcal{L}^n \hat{\boldsymbol{\mu}}, \hat{\boldsymbol{\mu}})}{(\hat{\boldsymbol{\mu}}, \hat{\boldsymbol{\mu}})} = (i)^n \left( [\hat{H}, [\hat{H}, [\hat{H}, \dots, \hat{\boldsymbol{\mu}}]]] \dots \right)$$
(2.25)

where it is implied that we are applying the commutator n times. We want to answer the form for the  $\hat{\mathbf{H}}^{G_0W_0}$  Hamiltonian, which has the super matrix form of

$$\begin{bmatrix} \mathbf{f} + \mathbf{\Sigma}_{\infty} & \mathbf{W} \\ \mathbf{W}^{\dagger} & \mathbf{d} \end{bmatrix}$$
 (2.26)

and the memory kernel

$$\mathcal{K}(t) = \left( \mathbf{A} \left| \mathcal{L} \mathcal{Q} e^{i\mathcal{Q}\mathcal{L}t} \mathcal{Q} \mathcal{L} \right| \mathbf{A} \right), \tag{2.27}$$

where  $Q = \mathcal{I} - \mathcal{P}$  is the complementary projection operator.

## Chapter 3

# Moment-Conserving GW

## 3.1 Löwdin Downfolding: 2/2/2025

We know that the definition of a Green's function associated with some Hamiltonian  $\boldsymbol{H}$  is given by:

$$(\omega - \boldsymbol{H}) \boldsymbol{G} = \boldsymbol{I} \tag{3.1}$$

where we can consider both cases to be fully interacting. The downfolding tells us to separate into a system S and auxiliary space L, so we have:

$$\begin{pmatrix} \omega - \begin{pmatrix} \boldsymbol{H}_{\mathcal{S}\mathcal{S}} & \boldsymbol{H}_{\mathcal{S}\mathcal{L}} \\ \boldsymbol{H}_{\mathcal{L}\mathcal{S}} & \boldsymbol{H}_{\mathcal{L}\mathcal{L}} \end{pmatrix} \end{pmatrix} \begin{pmatrix} \boldsymbol{G}_{\mathcal{S}\mathcal{S}} & \boldsymbol{G}_{\mathcal{S}\mathcal{L}} \\ \boldsymbol{G}_{\mathcal{L}\mathcal{S}} & \boldsymbol{G}_{\mathcal{L}\mathcal{L}} \end{pmatrix} = \begin{pmatrix} \boldsymbol{I}_{\mathcal{S}\mathcal{S}} & \boldsymbol{0} \\ \boldsymbol{0} & \boldsymbol{I}_{\mathcal{L}\mathcal{L}} \end{pmatrix} .$$

or:

$$egin{pmatrix} \omega - oldsymbol{H}_{\mathcal{SS}} & -oldsymbol{H}_{\mathcal{SL}} \ -oldsymbol{H}_{\mathcal{LS}} & \omega - oldsymbol{H}_{\mathcal{LL}} \end{pmatrix} egin{pmatrix} oldsymbol{G}_{\mathcal{SS}} & oldsymbol{G}_{\mathcal{SL}} \ oldsymbol{G}_{\mathcal{LL}} \end{pmatrix} = egin{pmatrix} oldsymbol{I}_{\mathcal{SS}} & oldsymbol{0} \ oldsymbol{0} & oldsymbol{I}_{\mathcal{LL}} \end{pmatrix} \,.$$

Multiplying out the matrices, the S block gives:

$$(\omega - \mathbf{H}_{SS}) \mathbf{G}_{SS} - \mathbf{H}_{SL} \mathbf{G}_{LS} = \mathbf{I}_{SS}. \tag{3.2}$$

Similarly, the  $\mathcal{L}$  block gives:

$$-\boldsymbol{H}_{\mathcal{LS}}\boldsymbol{G}_{\mathcal{SS}} + (\omega - \boldsymbol{H}_{\mathcal{LL}})\boldsymbol{G}_{\mathcal{LS}} = \boldsymbol{0}.$$
(3.3)

Assuming  $(\omega - \mathbf{H}_{\mathcal{LL}})$  is invertible, from the second equation we obtain:

$$G_{\mathcal{LS}} = (\omega - H_{\mathcal{LL}})^{-1} H_{\mathcal{LS}} G_{\mathcal{SS}}$$
.

Substituting this into the first equation:

$$\left(\omega-oldsymbol{H}_{\mathcal{SS}}
ight)oldsymbol{G}_{\mathcal{SS}}-oldsymbol{H}_{\mathcal{SL}}\left(\omega-oldsymbol{H}_{\mathcal{LL}}
ight)^{-1}oldsymbol{H}_{\mathcal{LS}}oldsymbol{G}_{\mathcal{SS}}=oldsymbol{I}_{\mathcal{SS}}\,.$$

Factorizing  $G_{SS}$ :

$$\left[ (\omega - \boldsymbol{H}_{\mathcal{S}\mathcal{S}}) - \boldsymbol{H}_{\mathcal{S}\mathcal{L}} (\omega - \boldsymbol{H}_{\mathcal{L}\mathcal{L}})^{-1} \, \boldsymbol{H}_{\mathcal{L}\mathcal{S}} \right] \, \boldsymbol{G}_{\mathcal{S}\mathcal{S}} = \boldsymbol{I}_{\mathcal{S}\mathcal{S}} \, .$$

Thus:

$$G_{SS} = [(\omega - H_{SS}) - H_{SL} (\omega - H_{LL})^{-1} H_{LS}]^{-1}$$

Now, notice that:

$$[\boldsymbol{G}_{\mathcal{SS}}^{0}(\omega)]^{-1} \equiv \omega - \boldsymbol{H}_{\mathcal{SS}} = \omega - (\boldsymbol{F} + \boldsymbol{\Sigma}(\infty)), \implies \boldsymbol{H}_{\mathcal{SS}} = \boldsymbol{F} + \boldsymbol{\Sigma}(\infty).$$

where  $\Sigma(\infty)$  is the static self-energy (0 for a HF mean-field reference), and F is the Fock matrix. Identifying the coupling matrices:

$$oldsymbol{H}_{\mathcal{SL}} = oldsymbol{W}, \quad oldsymbol{H}_{\mathcal{LS}} = oldsymbol{W}^\dagger$$
 .

and

$$H_{\mathcal{LL}} = d$$
.

Now let us make an ansatz for the upfolded Hamiltonian:

$$m{H}_{ ext{Upfolded}} = egin{pmatrix} m{F} + m{\Sigma}(\infty) & m{W} \ m{W}^\dagger & m{d} \end{pmatrix}.$$

Then, the resolvent is given by:

$$(\omega - \mathbf{H}_{\text{Upfolded}}) = \begin{pmatrix} \omega - \mathbf{F} - \mathbf{\Sigma}(\infty) & -\mathbf{W} \\ -\mathbf{W}^{\dagger} & \omega - \mathbf{d} \end{pmatrix}$$
(3.4)

Because we are interested in  $G_{SS}(\omega)$ , we care about  $(\omega - \mathbf{H}_{Upfolded})^{-1}$  in the upper left block, which is the Schur complement of  $(\omega - \mathbf{H}_{Upfolded})$  with respect to  $\omega - \mathbf{d}$ , defined as:

$$G_{SS}(\omega) = \left(\frac{(\omega - \mathbf{H}_{Upfolded})}{\omega - \mathbf{d}}\right)_{SS}^{-1} = \left(\omega - (\mathbf{F} + \mathbf{\Sigma}(\infty)) - \mathbf{W}\left[\omega - \mathbf{d}\right]^{-1}\mathbf{W}^{\dagger}\right)^{-1}$$
(3.5)

so the ansatz is correct.

## 3.2 Cumulant Idea

The definition of the cumulant ansatz for the Green's function is given by:

$$G_{SS}(t) = G_{SS}^{0}(t)e^{C(t)}$$
(3.6)

where C(t) is the cumulant and  $G^0_{SS}(t)$  is the HF Green's function. By relating the Dyson equation to the Taylor series expansion of the exponential, we can write:

$$\boldsymbol{G}_{SS}^{0}(t)\boldsymbol{C}(t) = \iint dt_{1}dt_{2}\boldsymbol{G}_{SS}^{0}(t-t_{1})\boldsymbol{\Sigma}^{c}(t_{1}-t_{2})\boldsymbol{G}_{SS}^{0}(t_{2})$$
(3.7)

Projecting to the spin-orbital basis and inserting the resolution of the identity, we get:

$$\sum_{r} \langle p | \mathbf{G}_{SS}^{0}(t) | r \rangle \langle r | \mathbf{C}(t) | q \rangle = \sum_{rs} \iint dt_{1} dt_{2} \langle p | \mathbf{G}_{SS}^{0}(t - t_{1}) | r \rangle \langle r | \mathbf{\Sigma}^{c}(t_{1} - t_{2}) | s \rangle \langle s | \mathbf{G}_{SS}^{0}(t_{2}) | q \rangle$$

(3.8)

$$\sum_{r} \boldsymbol{G}_{pr}^{0}(t) \boldsymbol{C}_{rq}(t) = \sum_{rs} \iint dt_{1} dt_{2} \boldsymbol{G}_{ps}^{0}(t - t_{1}) \boldsymbol{\Sigma}_{sr}^{c}(t_{1} - t_{2}) \boldsymbol{G}_{rq}^{0}(t_{2})$$
(3.9)

$$\boldsymbol{G}_{pp}^{0}(t)\boldsymbol{C}_{pq}(t) = \underbrace{\iint dt_{1}dt_{2}\boldsymbol{G}_{pp}^{0}(t-t_{1})\boldsymbol{\Sigma}_{pq}^{c}(t_{1}-t_{2})\boldsymbol{G}_{qq}^{0}(t_{2})}_{(3.10)}$$

where  $G_{SS}^0(t) \equiv G^0(t)$  and in the last step we used the fact that the HF Green's function is diagonal in the spin-orbital basis, specifically  $G_{pp}^0(t) = -i\Theta(t)e^{-i\epsilon_p t}$ , where  $\epsilon_p$  is the HF energy of the p-th spin-orbital. The formula for the inverse Fourier transform is given by:

$$f(t) = \int \frac{\mathrm{d}\omega}{2\pi} e^{-i\omega t} f(\omega) \tag{3.11}$$

which implies that

$$\boldsymbol{G}_{pp}^{0}(t-t_{1}) = \int \frac{\mathrm{d}\omega}{2\pi} e^{-i\omega(t-t_{1})} \boldsymbol{G}_{pp}^{0}(\omega)$$
(3.12)

$$\Sigma_{pq}^{c}(t_1 - t_2) = \int \frac{\mathrm{d}\omega'}{2\pi} e^{-i\omega'(t_1 - t_2)} \Sigma_{pq}^{c}(\omega')$$
(3.13)

$$\boldsymbol{G}_{qq}^{0}(t_{2}) = \int \frac{\mathrm{d}\omega''}{2\pi} e^{-i\omega''t_{2}} \boldsymbol{G}_{qq}^{0}(\omega'')$$
(3.14)

and plugging into the double time integral \*, we get:

$$* = \iint dt_1 dt_2 \left[ \int \frac{d\omega}{2\pi} e^{-i\omega(t-t_1)} \boldsymbol{G}_{pp}^0(\omega) \right] \left[ \int \frac{d\omega'}{2\pi} e^{-i\omega'(t_1-t_2)} \boldsymbol{\Sigma}_{pq}^c(\omega') \right] \left[ \int \frac{d\omega''}{2\pi} e^{-i\omega''t_2} \boldsymbol{G}_{qq}^0(\omega'') \right]$$

(3.15)

$$= \underbrace{\int dt_1 e^{-i(\omega' - \omega)t_1} \int dt_2 e^{-i(\omega'' - \omega')t_2}}_{4\pi^2 \delta(\omega' - \omega)\delta(\omega'' - \omega')} \iiint d\omega d\omega' d\omega'' \frac{e^{-i\omega t}}{8\pi^3} \boldsymbol{G}_{pp}^0(\omega) \boldsymbol{\Sigma}_{pq}^c(\omega') \boldsymbol{G}_{qq}^0(\omega'')$$
(3.16)

$$= \int \frac{\mathrm{d}\omega}{2\pi} e^{-i\omega t} \boldsymbol{G}_{pp}^{0}(\omega) \boldsymbol{\Sigma}_{pq}^{c}(\omega) \boldsymbol{G}_{qq}^{0}(\omega)$$
(3.17)

But now note that from the left hand side of eqn. 3.10, we can divide out the HF Green's function to get:

$$\boldsymbol{C}_{pq}(t) = i \int \frac{\mathrm{d}\omega}{2\pi} e^{-i(\omega - \epsilon_p)t} \boldsymbol{G}_{pp}^0(\omega) \boldsymbol{\Sigma}_{pq}^c(\omega) \boldsymbol{G}_{qq}^0(\omega)$$
(3.18)

Now, we insert the upfolded form for the self-energy, which is frequency independent as

$$oldsymbol{\Sigma}^c_{pq}(\omega) \equiv oldsymbol{\Sigma}^c_{pq} = egin{pmatrix} oldsymbol{\Sigma}(\infty) & oldsymbol{W}^< & oldsymbol{W}^> \ oldsymbol{W}^{\dagger <} & oldsymbol{d}^< & oldsymbol{0} \ oldsymbol{W}^{\dagger >} & oldsymbol{0} & oldsymbol{d}^> \end{pmatrix}_{pq}$$
. Customarily this is the point where the diagonal

approximation for the self-energy is introduced instead. We also know that  $G_{pp}^0(\omega) = \frac{I}{\omega - \epsilon_p}$ . We can then write:

$$C_{pq}(t) = i\Sigma_{pq}^{c} \int \frac{d\omega}{2\pi} \frac{e^{-i(\omega - \epsilon_{p})t}}{(\omega - \epsilon_{p})(\omega - \epsilon_{q})}$$
(3.19)

$$= i \frac{\Sigma_{pq}^{c}}{\epsilon_{q} - \epsilon_{p}} \left[ \underbrace{\int \frac{d\omega}{2\pi} \frac{e^{-i(\omega - \epsilon_{p})t}}{\omega - \epsilon_{p}}}_{-i\Theta(t)} - \underbrace{\int \frac{d\omega}{2\pi} \frac{e^{-i(\omega - \epsilon_{q})t}}{\omega - \epsilon_{q}}}_{e^{-i(\epsilon_{q} - \epsilon_{p})t}(-i\Theta(t))} \right]$$
(3.20)

$$= \frac{\Sigma_{pq}^{c}}{\epsilon_{q} - \epsilon_{p}} \Theta(t) \left[ 1 - e^{-i(\epsilon_{q} - \epsilon_{p})t} \right]$$
(3.21)

(3.22)

So now we insert this expression into our original ansatz for the cumulant, and we get:

$$G(t) = G^{0}(t)e^{\frac{\sum_{pq}^{c}}{\epsilon_{q} - \epsilon_{p}}\Theta(t)\left[1 - e^{-i(\epsilon_{q} - \epsilon_{p})t}\right]}$$
(3.23)

### 3.3 Lanczos Iteration

The block tridiagonal form can be expressed as:

$$\tilde{\mathbf{H}}_{\text{tri}} = \tilde{\mathbf{q}}^{(j),\dagger} \begin{bmatrix} \mathbf{f} + \mathbf{\Sigma}_{\infty} & \mathbf{W} \\ \mathbf{W}^{\dagger} & \mathbf{d} \end{bmatrix} \tilde{\mathbf{q}}^{(j)}$$

$$= \begin{bmatrix} \mathbf{f} + \mathbf{\Sigma}_{\infty} & \mathbf{L} & \mathbf{0} \\ \mathbf{L}^{\dagger} & \mathbf{H}_{1} & \mathbf{C}_{1} & \\ & \mathbf{C}_{1}^{\dagger} & \mathbf{H}_{2} & \mathbf{C}_{2} \\ & & \mathbf{C}_{2}^{\dagger} & \mathbf{H}_{3} & \ddots \\ & & \ddots & \ddots & \mathbf{C}_{j-1} \\ \mathbf{0} & & & & \mathbf{C}_{j+1}^{\dagger} & \mathbf{H}_{k} \end{bmatrix}$$
(3.24)

where we define  $\tilde{\mathbf{q}}^{(j)}$  as

$$\tilde{\mathbf{q}}^{(j)} = \begin{bmatrix} \mathbf{I} & \mathbf{0} \\ \mathbf{0} & \mathbf{q}^{(j)} \end{bmatrix} \tag{3.26}$$

This formulation becomes exact when the level j equals N, the dimension of the original hamiltonian; this corresponds to considering up to the highest moment of the self-energy, i.e. n france from  $1, \ldots, N$ . in practice, however, we always truncate the Krylov subspace to some j < N. Note that the tridiagonal form never actually forces us to compute  $\mathbf{W}$  or  $\mathbf{d}$ , as desired.

### 3.3.1 Creation of Krylov Subspace

Formally, the Krylov subspace of level j is given as  $\mathbf{q}^{(j)} = [\mathbf{q}_1, \mathbf{q}_2, \cdots, \mathbf{q}_j]$ , where the projection of the full Hamiltonian onto this subspace gives the tridiagonal form of equation 3.25. To start building up the subspace, we need to determine  $\mathbf{q}_1$  via QR decomposition of the exact GW couplings as  $\mathbf{W}^{\dagger} = \mathbf{q}_1 \mathbf{L}^{\dagger} \to \mathbf{q}_1 = \mathbf{W}^{\dagger} \mathbf{L}^{\dagger,-1}$ .  $\mathbf{L}$  is defined in terms of just the 0th order self-energy moment as  $\mathbf{L}^{\dagger} = (\mathbf{\Sigma}^{(0)})^{\frac{1}{2}}$ . We build up the subsequent  $q_i$  vectors through a three-term recurrence

$$\mathbf{q}_{i+1}\mathbf{C}_{i}^{\dagger} = \left[\mathbf{d}\mathbf{q}_{i} - \mathbf{q}_{i}\mathbf{H}_{i} - \mathbf{q}_{i-1}\mathbf{C}_{i-1}\right], \tag{3.27}$$

where the on-diagonal blocks are defined as

$$\mathbf{H}_i = \mathbf{q}_i^{\dagger} \mathbf{d} \mathbf{q}_i \tag{3.28}$$

Notice that to form the initial vector  $\mathbf{q}_1$  we would need  $\mathbf{W}^{\dagger}$  and to continue building the subspace, we would need  $\mathbf{d}$ , so to avoid this, we introduce the self-energy moments.

### 3.3.2 A sketch of the implicit Lanczos method

Due to Garnet's paper on the quasi-boson  $G_0W_0$  method, we know that we have a form for an upfolded  $G_0W_0$  Hamiltonian as

$$\boldsymbol{H}_{\text{Upfolded}}^{G_0W_0} = \begin{pmatrix} \boldsymbol{F} + \boldsymbol{\Sigma}(\infty) & \boldsymbol{W}^{<} & \boldsymbol{W}^{>} \\ \boldsymbol{W}^{<,\dagger} & \boldsymbol{d}^{<} & 0 \\ \boldsymbol{W}^{>,\dagger} & 0 & \boldsymbol{d}^{>} \end{pmatrix}$$
(3.29)

with quantities defined separately for lesser and greater parts. But we will just focus on the lesser part for now, where matrix elements of the screened interaction  $W^{<}$  are given by

$$W_{pkv}^{\leq} = \sum_{ia} (pk|ia) \left( X_{ia}^{v} + Y_{ia}^{v} \right) \tag{3.30}$$

and

$$d_{kv,lv'}^{\leq} = (\epsilon_k - \Omega_v) \,\delta_{k,l} \delta_{v,v'} \tag{3.31}$$

Note that we get the factor of  $\sqrt{2}$  accompanying all ERIs in RHF because we are considering the expectation value of the form

$$\frac{1}{2} \sum_{pqrs} \langle pq | | rs \rangle \langle \Psi_0 | \left( \hat{T}_i^{a,\alpha} + \hat{T}_i^{a,\beta} \right)^{\dagger} \left( a_p^{\dagger} a_q^{\dagger} a_s a_r \left( \hat{T}_i^{a,\alpha} + \hat{T}_i^{a,\beta} \right) \right) | \Psi_0 \rangle \tag{3.32}$$

where we have the excitation operator for a given spin channel as  $\hat{T}_i^{a,\sigma} = a_a^{\dagger,\sigma} a_i^{\sigma}$  and a singlet state for the RHF ground state carrying a factor of  $\frac{1}{\sqrt{2}}$ . If we apply Wick's theorem to this string, we get a contribution from both the  $\alpha$  and  $\beta$  channels, so  $\frac{1}{\sqrt{2}} + \frac{1}{\sqrt{2}} = \sqrt{2}$ . Through the introduction of a Krylov subspace as

$$\tilde{\boldsymbol{H}}_{\text{Upfolded}}^{G_0W_0} = \tilde{\boldsymbol{Q}}^{(n,\dagger)} \boldsymbol{H}_{\text{Upfolded}}^{G_0W_0} \tilde{\boldsymbol{Q}}^{(n)}$$
(3.33)

where  $Q^{(n)} \equiv (q_1 \ q_2 \ \cdots \ q_n)$  is the block Krylov subspace spanned by the Lanczos vectors, but we want to preserve the physical space of  $F + \Sigma(\infty)$  so we are really interested in the projection matrix

$$\tilde{\boldsymbol{Q}}^{(n)} = \begin{pmatrix} \boldsymbol{I} & \boldsymbol{0} \\ \boldsymbol{0} & \boldsymbol{Q}^{(n)} \end{pmatrix} \tag{3.34}$$

To get a gist of what the block Lanczos will do, let's just consider the case where we have two Lanczos vectors, so that we have a projection matrix of the form

$$\tilde{\boldsymbol{Q}}^{(2)} = \begin{pmatrix} \boldsymbol{I} & \boldsymbol{0} \\ \boldsymbol{0} & \boldsymbol{Q}^{(2)} \end{pmatrix} \tag{3.35}$$

So

$$\tilde{\boldsymbol{H}}_{\text{upfolded}}^{\text{Lanczos Iter 2}} = \tilde{\boldsymbol{Q}}^{(2,\dagger)} \boldsymbol{H}_{\text{Upfolded}}^{G_0 W_0} \tilde{\boldsymbol{Q}}^{(2)}$$
(3.36)

$$= \begin{pmatrix} I & 0 \\ 0 & Q^{(2)\dagger} \end{pmatrix} \begin{pmatrix} F + \Sigma(\infty) & W \\ W^{\dagger} & d \end{pmatrix} \begin{pmatrix} I & 0 \\ 0 & Q^{(2)} \end{pmatrix}$$
(3.37)

$$= \begin{pmatrix} I & \mathbf{0} \\ \mathbf{0} & [\mathbf{q}_1^{\dagger} & \mathbf{q}_2^{\dagger}] \end{pmatrix} \begin{pmatrix} F + \mathbf{\Sigma}(\infty) & \mathbf{W} \\ \mathbf{W}^{\dagger} & \mathbf{d} \end{pmatrix} \begin{pmatrix} I & \mathbf{0} \\ \mathbf{0} & [\mathbf{q}_1 & \mathbf{q}_2] \end{pmatrix}$$
(3.38)

$$= \begin{pmatrix} \mathbf{F} + \mathbf{\Sigma}(\infty) & \mathbf{W} \\ [\mathbf{q}_1^{\dagger} & \mathbf{q}_2^{\dagger}] \mathbf{W}^{\dagger} & [\mathbf{q}_1^{\dagger} & \mathbf{q}_2^{\dagger}] \mathbf{d} \end{pmatrix} \begin{pmatrix} \mathbf{I} & \mathbf{0} \\ \mathbf{0} & [\mathbf{q}_1 & \mathbf{q}_2] \end{pmatrix}$$
(3.39)

$$= \begin{pmatrix} F + \Sigma(\infty) & [\underbrace{Wq_1}_{\mathcal{Q}^{\dagger}\mathcal{Q}=\mathcal{R}} & \underbrace{Wq_2}_{0}] \\ & \underbrace{Q^{\dagger}\mathcal{Q}=\mathcal{R}}_{\mathcal{Q}^{\dagger}\mathcal{Q}=\mathcal{R}} & \underbrace{Q^{\dagger}_{1}dq_{2}}_{\mathcal{Q}_{1}} \\ & \underbrace{Q^{\dagger}_{1}dq_{1}}_{\mathcal{Q}^{\dagger}\mathcal{Q}} & \underbrace{Q^{\dagger}_{2}dq_{2}}_{\mathcal{Q}_{2}} \\ & \underbrace{Q^{\dagger}_{2}dq_{1}}_{\mathcal{Q}^{\dagger}} & \underbrace{Q^{\dagger}_{2}dq_{2}}_{\mathcal{M}_{2}} \end{pmatrix}$$

$$(3.40)$$

where in the last line we used the fact that we are free to choose our  $q_1 \equiv \mathcal{Q}$  as the same orthogonal matrix employed in constructing the QR decomposition of  $W^{\dagger}$  as

$$W^{\dagger} = \mathcal{Q} \mathcal{R}^{\dagger} \implies q_1 = W^{\dagger} \mathcal{R}^{-1,\dagger}$$
 (3.41)

### 3.3.3 Physical-auxiliary coupling W

Now consider  $WW^{\dagger} = \mathcal{RQQ}^{\dagger}\mathcal{R}^{\dagger} = \mathcal{RR}^{\dagger}$ . So if we can do a Cholesky decomposition of  $WW^{\dagger}$  we would get access to  $\mathcal{R}$ . But consider that

$$\Sigma_{pq}^{(n,<)} = \sum_{ia,jb,k} \sum_{\mu} (pk|ia) \left( X_{ia}^{\mu} + Y_{ia}^{\mu} \right) \left( X_{jb}^{\mu} + Y_{jb}^{\mu} \right) (qk|jb) [(\epsilon_k - \Omega_{\mu})]^n$$
 (3.42)

so it becomes clear that  $WW^{\dagger} = \Sigma^0 \implies \mathcal{R} = (WW^{\dagger})^{1/2} = (\Sigma^0)^{1/2}$ , which we hope to be able to accomplish with the Cholesky decomposition.

### 3.3.4 Auxiliary-auxiliary space d: the working equations

From above, notice that

$$M_i = q_i^{\dagger} dq_i$$
 and  $C_i = q_i^{\dagger} dq_{i+1}$  (3.43)

To get these, define

$$\boldsymbol{S}_{i,j}^{(n)} = \boldsymbol{q}_i^{\dagger} \boldsymbol{d}^n \boldsymbol{q}_j \tag{3.44}$$

So it must be that  $S_{0,j}^{(n)} = S_{i,0}^{(n)} = 0$  for all  $i, j, S_{i,j}^{(0)} = \delta_{ij} \mathbf{I}$ , and we can demand Hermiticity, so that  $S_{i,j}^{(n)} = S_{j,i}^{(n)\dagger}$ . We are able to initialize S using

$$S_{1,1}^{(n)} = q_1^{\dagger} d^n q_1 = R^{-1} W d^n W^{\dagger} R^{-1,\dagger} = R^{-1} \Sigma^{(n)} R^{-1,\dagger}$$
(3.45)

So looking at eqn. 3.117, we see that the on-diagonal elements are  $M_i = S_{i,i}^{(1)}$  while the off-diagonal elements are  $C_i = S_{i,i+1}^{(1)}$ . The familiar three-term Lanczos recurrence is

$$\mathbf{q}_{i+1}\mathbf{C}_{i}^{\dagger} = \left[\mathbf{d}\mathbf{q}_{i} - \mathbf{q}_{i}\mathbf{M}_{i} - \mathbf{q}_{i-1}\mathbf{C}_{i-1}\right] \implies \mathbf{q}_{i+1} = \left[\mathbf{d}\mathbf{q}_{i} - \mathbf{q}_{i}\mathbf{M}_{i} - \mathbf{q}_{i-1}\mathbf{C}_{i-1}\right]\mathbf{C}_{i}^{\dagger,-1} \qquad (3.46)$$

where the participants are block vectors and we have assumed that  $C_i$  is invertible. Let us start by considering the form of

$$S_{i+1,i}^{n} \equiv q_{i+1}^{\dagger} d^{n} q_{i} = \left[ \left[ \mathbf{d} \mathbf{q}_{i} - \mathbf{q}_{i} \mathbf{M}_{i} - \mathbf{q}_{i-1} \mathbf{C}_{i-1} \right] C_{i}^{\dagger,-1} \right]^{\dagger} d^{n} q_{i}$$
(3.47)

$$= \boldsymbol{C}_{i}^{-1} \left[ \underbrace{\boldsymbol{q}_{i}^{\dagger} \boldsymbol{d}^{n+1} \boldsymbol{q}_{i}}_{\boldsymbol{S}_{i,i}^{n+1}} - \boldsymbol{M}_{i} \underbrace{\boldsymbol{q}_{i}^{\dagger} \boldsymbol{d}^{n} \boldsymbol{q}_{i}}_{\boldsymbol{S}_{i,i}^{n}} - \boldsymbol{C}_{i-1}^{\dagger} \underbrace{\boldsymbol{q}_{i-1}^{\dagger} \boldsymbol{d}^{n} \boldsymbol{q}_{i}}_{\boldsymbol{S}_{i-1,i}^{n}} \right]$$
(3.48)

$$= C_i^{-1} \left[ S_{i,i}^{n+1} - M_i S_{i,i}^n - C_{i-1}^{\dagger} S_{i-1,i}^n \right]$$
(3.49)

Similarly

$$S_{i+1,i+1}^{n} \equiv \boldsymbol{q}_{i+1}^{\dagger} \boldsymbol{d}^{n} \boldsymbol{q}_{i+1} = \left[ \left[ \mathbf{d} \mathbf{q}_{i} - \mathbf{q}_{i} \mathbf{M}_{i} - \mathbf{q}_{i-1} \mathbf{C}_{i-1} \right] \boldsymbol{C}_{i}^{\dagger,-1} \right]^{\dagger} \boldsymbol{d}^{n} \left[ \mathbf{d} \mathbf{q}_{i} - \mathbf{q}_{i} \mathbf{M}_{i} - \mathbf{q}_{i-1} \mathbf{C}_{i-1} \right] \boldsymbol{C}_{i}^{\dagger,-1}$$

$$(3.50)$$

$$(3.51)$$

$$S_{i+1,i+1}^{n} = C_{i}^{-1} \left[ S_{i,i}^{n+2} + M_{i} S_{i,i}^{n} M_{i} + C_{i-1}^{\dagger} S_{i-1,i-1}^{n} C_{i-1} \right.$$

$$\left. - \underbrace{q_{i}^{\dagger} d^{n+1} q_{i}}_{S_{i,i}^{n+1}} M_{i} - M_{i} \underbrace{q_{i}^{\dagger} d^{n+1} q_{i}}_{S_{i,i}^{n+1}} \right.$$

$$\left. - \underbrace{q_{i}^{\dagger} d^{n+1} q_{i-1}}_{S_{i,i-1}^{n+1}} C_{i-1} - C_{i-1}^{\dagger} \underbrace{q_{i-1}^{\dagger} d^{n+1} q_{i}}_{S_{i-1,i}^{n+1}} \right.$$

$$\left. + M_{i} \underbrace{q_{i}^{\dagger} d^{n} q_{i-1}}_{S_{i,i-1}^{n}} C_{i-1} + C_{i-1}^{\dagger} \underbrace{q_{i-1}^{\dagger} d^{n} q_{i}}_{S_{i-1,i}^{n}} M_{i} \right] C_{i}^{\dagger,-1}$$

$$= C_{i}^{-1} \left[ S_{i,i}^{n+2} + M_{i} S_{i,i}^{n} M_{i} + C_{i-1}^{\dagger} S_{i-1,i-1}^{n} C_{i-1} - C_{i-1} - C_{i-1} \right.$$

$$\left. - P(S_{i,i}^{n+1} M_{i}) - P(S_{i,i-1}^{n+1} C_{i-1}) + P(M_{i} S_{i,i-1}^{n} C_{i-1}) \right] C_{i}^{\dagger,-1}$$

$$(3.52)$$

where we introduced the permutation operator as  $P(\mathbf{A}) = \mathbf{A} + \mathbf{A}^{\dagger}$ . Setting n = 0 we get

$$\boldsymbol{I} = \boldsymbol{C}_{i}^{-1} \left[ \boldsymbol{S}_{i,i}^{2} + \boldsymbol{M}_{i}^{2} + \boldsymbol{C}_{i-1}^{\dagger} \boldsymbol{C}_{i-1} - P\left(\boldsymbol{M}_{i}^{2}\right) - P\left(\boldsymbol{S}_{i,i-1}^{1} \boldsymbol{C}_{i-1}\right) + 0 \right] \boldsymbol{C}_{i}^{\dagger,-1}$$

$$(3.54)$$

$$\implies \boldsymbol{C}_{i}^{2} = \boldsymbol{C}_{i} \boldsymbol{C}_{i}^{\dagger} = \left[ \boldsymbol{S}_{i,i}^{2} + \boldsymbol{M}_{i}^{2} + \boldsymbol{C}_{i-1}^{\dagger} \boldsymbol{C}_{i-1} - P\left(\boldsymbol{S}_{i,i}^{1} \boldsymbol{M}_{i}\right) - P\left(\boldsymbol{S}_{i,i-1}^{1} \boldsymbol{C}_{i-1}\right) \right]$$

$$(3.55)$$

$$\implies C_i C_i^{\dagger} \equiv q_i^{\dagger} dq_{i+1} q_{i+1}^{\dagger} dq_i = q_i^{\dagger} d^2 q_i \tag{3.56}$$

The hope is that we can use the Cholesky QR algorithm to take the effective matrix square root to get the  $C_i$ .

## 3.4 Efficient generation of the self-energy moments

Since the self-energy is defined as a convolution between the interacting Green's function and the screened Coulomb potential in GW we can use the known formula for the moment distribution of a convolution of two quantities to get

$$\Sigma_{pq}^{(n,<)} = \sum_{ia,jb,k} \sum_{t=0}^{n} \binom{n}{t} (-1)^t \epsilon_k^{n-t} (pk \mid ia) \eta_{ia,jb}^{(t)} (qk \mid jb), \tag{3.57}$$

where we have defined the density-density response moment as

$$\eta_{ia,jb}^{(n)} = \sum_{v} \left( X_{ia}^{v} + Y_{ia}^{v} \right) \Omega_{v}^{n} \left( X_{jb}^{v} + Y_{jb}^{v} \right) \tag{3.58}$$

# 3.4.1 An alternative formulation of the RPA polarizability with the density response moments

The Casida equation is given by

$$\begin{pmatrix} A & B \\ -B & -A \end{pmatrix} \begin{pmatrix} X & Y \\ Y & X \end{pmatrix} \begin{pmatrix} \Omega & 0 \\ 0 & -\Omega \end{pmatrix} = \begin{pmatrix} X & Y \\ Y & X \end{pmatrix}$$
(3.59)

where the excitation X and de-excitation Y eigenvectors form the biorthogonal set as

$$(\boldsymbol{X} + \boldsymbol{Y})(\boldsymbol{X} - \boldsymbol{Y})^{T} = (\boldsymbol{X} + \boldsymbol{Y})^{T}(\boldsymbol{X} - \boldsymbol{Y}) = \boldsymbol{I}$$
(3.60)

The  $\boldsymbol{A}$  and  $\boldsymbol{B}$  matrices are defined as

$$A_{ia,jb} = (\epsilon_a - \epsilon_i) \,\delta_{ij} \delta_{ab} + \mathcal{K}_{ia,bj} \tag{3.61}$$

$$B_{ia,jb} = \mathcal{K}_{ia,jb} \tag{3.62}$$

and I take the direct approximation, so  $\mathcal{K}_{ia,jb} = (ia \mid jb) = \mathcal{K}_{ia,bj}$ . As I showed earlier this year, these neutral excitation energies  $\Omega$  define the poles of the polarizability  $\chi_{\text{RPA}}(\omega)$  as

$$\chi_{\text{RPA}}(\omega) = \begin{bmatrix} \mathbf{X} & \mathbf{Y} \\ \mathbf{Y} & \mathbf{X} \end{bmatrix} \begin{bmatrix} \omega \mathbf{I} - \mathbf{\Omega} & \mathbf{0} \\ \mathbf{0} & -\omega \mathbf{I} - \mathbf{\Omega} \end{bmatrix}^{-1} \begin{bmatrix} \mathbf{X} & \mathbf{Y} \\ \mathbf{Y} & \mathbf{X} \end{bmatrix}^{T}.$$
 (3.63)

where  $\chi_{\text{RPA}}(\omega) = (\chi_0(\omega)^{-1} - \mathcal{K})^{-1}$ , where  $\chi_0(\omega)$  is the irreducible polarizability of the reference state. Because of the bosonic-like symmetry in equation 3.63, we can write a more compact form of the RPA polarizability as

$$\eta(\omega) = (\mathbf{X} + \mathbf{Y})(\omega \mathbf{I} - \mathbf{\Omega})^{-1}(\mathbf{X} + \mathbf{Y})^{T}.$$
(3.64)

However, here we are interested in the spectral moments of the compactified RPA polarizability in equation 3.64 over all RPA excitation energies, which is given as

$$\eta_{ia,jb}^{(n)} = -\frac{1}{\pi} \int_0^\infty \operatorname{Im} \left[ \eta_{ia,jb}(\omega) \right] \omega^n d\omega = -\frac{1}{\pi} \left( \boldsymbol{X} + \boldsymbol{Y} \right) \left[ \int_0^\infty \operatorname{Im} \left[ \left( \omega \mathbf{I} - \boldsymbol{\Omega} \right)^{-1} \right] \omega^n d\omega \right] \left( \boldsymbol{X} + \boldsymbol{Y} \right)^T.$$
(3.65)

where in the final equality we have just factored out the frequency independent piece. Now, we can use the identity from complex analysis that  $\frac{1}{-\Omega+I(\omega+i\eta)}=P\left(\frac{1}{\omega I-\Omega}\right)-i\pi\delta(\omega-\Omega)\Longrightarrow \operatorname{Im}\left[\frac{1}{-\Omega+I(\omega+i\eta)}\right]=-\pi\delta(\omega-\Omega)$ , where P is the Cauchy principal value, so

$$\int_{0}^{\infty} \operatorname{Im}\left[\left(\omega \mathbf{I} - \mathbf{\Omega}\right)^{-1}\right] \omega^{n} d\omega = -\pi \int_{0}^{\infty} \delta(\omega - \mathbf{\Omega}) \omega^{n} d\omega = -\pi \mathbf{\Omega}^{n}.$$
 (3.66)

and plugging back and, we get  $\boldsymbol{\eta}^{(n)} = (\boldsymbol{X} + \boldsymbol{Y}) (\Omega^n) (\boldsymbol{X} + \boldsymbol{Y})^T$ .

### 3.4.2 Getting $\eta$ with lower scaling

Now we will make use of the symmetric formulation of the RPA problem by Fillip Furche. The RPA eigenvalue problem is given by

$$\begin{pmatrix} A & B \\ -B & -A \end{pmatrix} \begin{pmatrix} X \\ Y \end{pmatrix} = \Omega \begin{pmatrix} X \\ Y \end{pmatrix}$$
(3.67)

so we can get the coupled equations

$$AX + BY = \Omega X \tag{3.68}$$

$$-BX - AY = \Omega Y \tag{3.69}$$

Adding and subtracting these two equations, we get

$$(\boldsymbol{A} - \boldsymbol{B})(\boldsymbol{X} - \boldsymbol{Y}) = (\boldsymbol{X} + \boldsymbol{Y})\Omega \implies (\boldsymbol{A} - \boldsymbol{B}) = (\boldsymbol{X} + \boldsymbol{Y})\Omega(\boldsymbol{X} - \boldsymbol{Y})^{\dagger} \implies \boxed{\boldsymbol{\eta}^{(1)} = \boldsymbol{A} - \boldsymbol{B}}$$
(3.70)

$$(\mathbf{A} + \mathbf{B})(\mathbf{X} + \mathbf{Y}) = (\mathbf{X} - \mathbf{Y})\Omega \implies (\mathbf{A} + \mathbf{B}) = (\mathbf{X} - \mathbf{Y})\Omega(\mathbf{X} + \mathbf{Y})^{\dagger}$$
(3.71)

But then also notice that since  $\boldsymbol{\eta}^{(0)} = (\boldsymbol{X} + \boldsymbol{Y}) (\boldsymbol{X} + \boldsymbol{Y})^{\dagger}$ , we can write  $\boldsymbol{\eta}^{(1)} = \boldsymbol{\eta}^{(0)} (\boldsymbol{A} + \boldsymbol{B}) \boldsymbol{\eta}^{(0)}$ . Continuing on, we arrive at Furche's symmetric formulation of the RPA problem as

$$(\mathbf{A} - \mathbf{B})(\mathbf{A} + \mathbf{B})(\mathbf{X} + \mathbf{Y}) = (\mathbf{X} + \mathbf{Y})\Omega^{2}$$
(3.72)

By right multiplying by  $(\boldsymbol{X} + \boldsymbol{Y})^{\dagger}$ , this leads to

$$(\mathbf{A} - \mathbf{B}) (\mathbf{A} + \mathbf{B}) \boldsymbol{\eta}^{(0)} = \boldsymbol{\eta}^{(2)}$$
(3.73)

This is suggestive of a recursive relation with the form

$$\boldsymbol{\eta}^{(m)} = (\boldsymbol{A} - \boldsymbol{B}) (\boldsymbol{A} + \boldsymbol{B}) \boldsymbol{\eta}^{(m-2)}$$
(3.74)

(3.75)

But we have already found that  $\boldsymbol{\eta}^{(1)} = \boldsymbol{A} - \boldsymbol{B} \implies \boldsymbol{\eta}^{(2)} = \boldsymbol{\eta}^{(1)} (\boldsymbol{A} + \boldsymbol{B}) \boldsymbol{\eta}^{(0)}$ . Now plug in  $\boldsymbol{\eta}^{(1)} = \boldsymbol{\eta}^{(0)} (\boldsymbol{A} + \boldsymbol{B}) \boldsymbol{\eta}^{(0)}$  to get

$$\eta^{(2)} = \eta^{(0)} (\mathbf{A} + \mathbf{B}) \eta^{(0)} (\mathbf{A} + \mathbf{B}) \eta^{(0)} = \left[ \eta^{(0)} (\mathbf{A} + \mathbf{B}) \right]^2 \eta^{(0)}$$
(3.76)

Repeating gives

$$\boldsymbol{\eta}^{(m)} = \left[\boldsymbol{\eta}^{(0)} \left(\boldsymbol{A} + \boldsymbol{B}\right)\right]^{m} \boldsymbol{\eta}^{(0)}$$
(3.77)

and then to initialize, consider

$$\mathbf{A} - \mathbf{B} = \boldsymbol{\eta}^{(0)} \left( \mathbf{A} + \mathbf{B} \right) \boldsymbol{\eta}^{(0)} \tag{3.78}$$

$$(\mathbf{A} - \mathbf{B})(\mathbf{A} + \mathbf{B}) = \boldsymbol{\eta}^{(0)}(\mathbf{A} + \mathbf{B})\boldsymbol{\eta}^{(0)}(\mathbf{A} + \mathbf{B}) = \left[\boldsymbol{\eta}^{(0)}(\mathbf{A} + \mathbf{B})\right]^{2}$$
(3.79)

$$\implies \eta^{(0)} = [(A - B)(A + B)]^{1/2}(A + B)^{-1}$$
(3.80)

We do assume a lot of things in going to equation 3.80, but I have

assert np.all(np.linalg.eigvals(ApB) > 0)

# 3.4.3 Bringing it back to outline an efficient procedure for Computing Self-Energy Moments

The moments of the lesser and greater parts of the self-energy have the form

$$\Sigma_{pq}^{(n,<)} = \sum_{ia,jb,k} \sum_{t=0}^{n} \binom{n}{t} (-1)^{t} \epsilon_{k}^{n-t} (pk \mid ia) \eta_{ia,jb}^{(t)} (qk \mid jb),$$

$$\Sigma_{pq}^{(n,>)} = \sum_{ia,jb,c} \sum_{t=0}^{n} \binom{n}{t} \epsilon_{c}^{n-t} (pc \mid ia) \eta_{ia,jb}^{(t)} (qc \mid jb),$$
(3.81)

respectively. The energies  $\epsilon$  and integrals  $(pq \mid rs)$  are known to us from the prior mean-field calculation, but the density-density moments  $\eta^{(t)}$  are not, so we need to compute them. We find them to be defined as

$$\boldsymbol{\eta}^{(m)} = \left[\boldsymbol{\eta}^{(0)}(\mathbf{A} + \mathbf{B})\right]^m \boldsymbol{\eta}^{(0)}, \tag{3.82}$$

with the initial

$$\eta^{(0)} = [(\mathbf{A} - \mathbf{B})(\mathbf{A} + \mathbf{B})]^{\frac{1}{2}}(\mathbf{A} + \mathbf{B})^{-1}$$
(3.83)

where we have **A** and **B** with components  $A_{ia,jb} = \delta_{ij}\delta_{ab} (\epsilon_a - \epsilon_i) + (ia \mid jb)$  and  $B_{ia,jb} = (ia \mid jb)$ . Note that the scaling of the above would be improved by introducing low-rank approximations to the electron repulsion integrals (ERIs) via the Cholesky decomposition or tensor hyper-contraction, which express these 4-index quantities as a sum of two 3- or five 2-index tensors, respectively.

## 3.5 Avoiding auxiliary space

To avoid using the large auxiliary quantities, we introduce an operator which will project onto the block Lanczos space as  $\mathbf{S}_{i,j}^{(n)} = \mathbf{q}_i^{\dagger} \mathbf{d}^n \mathbf{q}_j$ . But due to the orthogonality of the Krylov subspace, we have  $q_i^{\dagger} q_j = 0$  for  $i \neq j$ , so  $\mathbf{S}_{i,j}^{(n)} = 0$  for |i - j| > 1. Furthermore, due to Hermiticity,  $\mathbf{S}_{i,j}^{(n)} = \mathbf{S}_{j,i}^{(n)\dagger}$ . We can start with the initial definition  $\mathbf{S}_{1,1}^{(n)} = \mathbf{q}_1^{\dagger} \mathbf{d}^n \mathbf{q}_1 = \mathbf{L}^{-1} \mathbf{\Sigma}_{i,j}^{(n)} \mathbf{L}^{-1,\dagger}$ .

Then we know that the recurrence relation is given by

$$\mathbf{q}_{i+1}\mathbf{C}_{i}^{\dagger} = [\mathbf{d}\mathbf{q}_{i} - \mathbf{q}_{i}\mathbf{H}_{i} - \mathbf{q}_{i-1}\mathbf{C}_{i-1}] \tag{3.84}$$

I will not repeat the algebraic manipulations, but Backhouse shows in his thesis that the following relations follow

$$\mathbf{S}_{i+1,i}^{(n)} = \mathbf{q}_{i+1}^{\dagger} \mathbf{d}^{n} \mathbf{q}_{i} = \mathbf{C}_{i}^{-1} \left[ \mathbf{S}_{i,i}^{(n+1)} - \mathbf{H}_{i} \mathbf{S}_{i,i}^{(n)} - \mathbf{C}_{i-1}^{\dagger} \mathbf{S}_{i-1,i}^{(n)} \right]$$

$$\mathbf{S}_{i+1,i+1}^{(n)} = \mathbf{q}_{i+1}^{\dagger} \mathbf{d}^{n} \mathbf{q}_{i+1} = \mathbf{C}_{i}^{-1} \left[ \mathbf{S}_{i,i}^{(n+2)} + \mathbf{H}_{i} \mathbf{S}_{i,i}^{(n)} \mathbf{H}_{i} + \mathbf{C}_{i-1}^{\dagger} \mathbf{S}_{i-1,i-1}^{(n)} \mathbf{C}_{i-1} \right]$$

$$-P \left( \mathbf{S}_{i,i}^{(n+1)} \mathbf{H}_{i} \right) + P \left( \mathbf{H}_{i} \mathbf{S}_{i,i-1}^{(n)} \mathbf{C}_{i-1} \right) - P \left( \mathbf{S}_{i,i-1}^{(n+1)} \mathbf{C}_{i-1} \right) \right] \mathbf{C}_{i}^{-1,\dagger},$$

$$(3.85)$$

and then solving for  $C_i^2$ 

$$\mathbf{C}_{i}^{2} = \left[\mathbf{S}_{i,i}^{(2)} + \mathbf{H}_{i}^{2} + \mathbf{C}_{i-1}^{\dagger} \mathbf{C}_{i-1} - P\left(\mathbf{S}_{i,i}^{(1)} \mathbf{H}_{i}\right) - P\left(\mathbf{S}_{i,i-1}^{(1)} \mathbf{C}_{i-1}\right)\right]$$
(3.86)

and we can also find the on-diagonal H matrices using

$$\mathbf{H}_i = \mathbf{q}_i^{\dagger} \mathbf{d} \mathbf{q}_i = \mathbf{S}_{i,i}^{(1)}. \tag{3.87}$$

Now, we know how to compute every quantity of equation 3.25.

## 3.6 Spectral Function

When we diagonalize the upfolded Hamiltonian, the eigenpairs are composed of charged excitation energies  $E_k$  (analogous to QP energies from the QP equation, but without a diagonal approximation to the self-energy) and eigenvectors  $\mathbf{u}_k$ , which can be transformed into the Dyson orbitals  $\boldsymbol{\psi}_k$  via

$$\psi_k = LPu_k \tag{3.88}$$

where P is the projection operator into the physical space

```
eigvec_phys = extracted_eigvecs[j][:mf.nbsf]
```

and then L, depending on whether we have an occupied or virtual state, is

```
moments_less = [form_moment(i, "lesser", mf) for i in range(2*
n_initial+2)]
moments_great = [form_moment(i, "greater", mf) for i in range(2*
n_initial+2)]

L_less = np.linalg.cholesky(moments_less[0], upper=False)
L_great = np.linalg.cholesky(moments_great[0], upper=False)
L = (L_less, L_great)
```

Then, we can construct the spectral function via

$$A(\omega) = \sum_{k} ||\psi_{k}||^{2} \underbrace{\frac{1}{\pi} \frac{\eta}{(\omega - E_{k})^{2} + \eta^{2}}}_{\text{Lorentzian}}$$
(3.89)

where  $\eta$  is a broadening parameter. In general, the spectral function is known to have the form

$$A(\omega) = -\frac{1}{\pi} \text{Im} G(\omega) \tag{3.90}$$

and the Dyson equation is

$$G(\omega) = G_0(\omega) + G_0(\omega)\Sigma(\omega)G(\omega) = \underbrace{\frac{1}{\omega - \epsilon_0} - \Sigma(\omega)}_{G_0(\omega)^{-1}}$$
(3.91)

but then we know that  $\Sigma(\omega) = \Sigma_R(\omega) + i\Sigma_I(\omega)$ , so we can write

$$G(\omega) = \frac{1}{\omega - \epsilon_0 - \Sigma_R(\omega) - i\Sigma_I(\omega)}$$
(3.92)

Defining  $x(\omega) = \omega - \epsilon_0 - \Sigma_R(\omega)$  and  $y(\omega) = \Sigma_I(\omega)$ , we can rewrite the Dyson equation as

$$G(\omega) = \frac{1}{x(\omega) - iy(\omega)} \times \frac{x(\omega) + iy(\omega)}{x(\omega) + iy(\omega)} = \frac{x(\omega) + iy(\omega)}{(x(\omega))^2 + (y(\omega))^2} \implies \operatorname{Im}G(\omega) = \frac{y(\omega)}{(x(\omega))^2 + (y(\omega))^2}$$
(3.93)

We have the following fully analytic expression for the self-energy

$$\Sigma_{pp}^{\text{corr}}(\omega) = \sum_{\mu}^{\text{RPA}} \left( \sum_{i}^{\text{occupied}} \frac{w_{pi}^{\mu} w_{ip}^{\mu}}{\omega - (\epsilon_{i} - \Omega_{\mu}) + i\eta} + \sum_{a}^{\text{virtual}} \frac{w_{pa}^{\mu} w_{ap}^{\mu}}{\omega - (\epsilon_{a} + \Omega_{\mu}) - i\eta} \right)$$
(3.94)

with

$$w_{pq}^{\mu} = \sum_{jb} (pq|jb) \left( X_{jb}^{\mu} + Y_{jb}^{\mu} \right) \tag{3.95}$$

So

$$\Sigma_{pp}^{R}(\omega) = \sum_{\mu}^{\text{RPA}} \left( \sum_{i}^{\text{occupied}} \frac{w_{pi}^{\mu} w_{ip}^{\mu}}{\omega - (\epsilon_{i} - \Omega_{\mu})} + \sum_{a}^{\text{virtual}} \frac{w_{pa}^{\mu} w_{ap}^{\mu}}{\omega - (\epsilon_{a} + \Omega_{\mu})} \right)$$
(3.96)

and because  $\frac{1}{\omega \pm i\eta} = \mathcal{P} \frac{1}{\omega} \mp i\pi \delta(\omega)$ , we can write

$$\Sigma_{pp}^{I}(\omega) = \pi \sum_{\mu}^{\text{RPA}} \left( \sum_{a}^{\text{virtual}} w_{pa}^{\mu} w_{ap}^{\mu} \delta(\omega - (\epsilon_{a} + \Omega_{\mu}) - \sum_{i}^{\text{occupied}} w_{pi}^{\mu} w_{ip}^{\mu} \delta(\omega - (\epsilon_{i} - \Omega_{\mu})) \right)$$
(3.97)

This brings us back to a Lorentzian

$$A(\omega) = -\frac{1}{\pi} \sum_{p} \frac{\eta(\omega)}{\left(\omega - \epsilon_0 - \sum_{pp}^{R}(\omega)\right)^2 + \eta(\omega)^2}$$
(3.98)

with  $\eta(\omega) \equiv \Sigma_{pp}^{I}(\omega)$ . We need to compute

$$\eta(\omega) = \left(\sum_{\mu}^{\text{RPA}} \left(\sum_{a}^{\text{virtual}} w_{pa}^{\mu} w_{ap}^{\mu} \left(\frac{\gamma}{\left(\omega - (\epsilon_{a} + \Omega_{\mu})\right)^{2} + \gamma^{2}}\right) - \sum_{i}^{\text{occupied}} w_{pi}^{\mu} w_{ip}^{\mu} \left(\frac{\gamma}{\left(\omega - (\epsilon_{i} - \Omega_{\mu})\right)^{2} + \gamma^{2}}\right)\right)\right)_{p}$$
(3.99)

at each  $\omega = E_{QP}$  that we found in the QP equation. Meanwhile, we know that the QP renormalization factor is given by

$$Z_p = \left(1 - \frac{\partial \Sigma_{pp}^R(\omega)}{\partial \omega} \Big|_{\omega = E_{QP}}\right)^{-1} \tag{3.100}$$

So I need to prove the equivalence between this form and the one I have above using  $\eta(\omega)$ .

## $3.7 \quad 6/2$ Conclusions

## 3.7.1 Hypothesis

If one maintains orthogonality in the Krylov subspace, the Ritz eigenpairs will converge to exact diagonalization within numerical precision. However, it is a fact that if the residual of the recurrence relation 3.113 loses full rank, the off-diagonal elements get a high condition number and are no longer invertible. So we have seen that MC-GW can replicate the block Lanczos up until this condition is met, but no longer afterwards. So even though MC-GW enforces orthogonality of the Krylov subspace even after the residual does lose its full rank, the off-diagonal elements are then no longer invertible, which explains the numerical issues that I observe after this point.

## 3.7.2 Systems of Study

Given my physical dimension  $n_{\text{nocc}}$ , the auxiliary dimension is of size  $n_{\text{channel}} \times n_{\text{virt}} \times n_{\text{occ}}$ , where  $n_{\text{channel}}$  is either  $n_{\text{occ}}$  or  $n_{\text{virt}}$ , depending on whether we are dealing with the lesser or greater channel; it becomes clear that the block size, which is determined by the physical dimension, will be  $O(n_{\text{occ}})$ , whereas the size of the auxiliary dimension will be  $O(n_{\text{channel}} \times n_{\text{virt}} \times n_{\text{occ}})$ . I have been doing my tests up until now with a minimal basis, where at least one of these  $n_{\text{s}}$  is small, so I wonder if when I move to a larger basis there will be a noticeable difference with the large discrepancy between n and  $n^3$ . Perhaps we will be able to exactly converge the low-lying orbital MO QPEs, but not the higher ones, but this is typically all that is desired of the method.

# $3.8 \quad 6/10 \text{ plots}$

As you probably have seen, in the theory of MC-GW, the equations come in two forms; the first gives the explicit definition of the quantity in terms of a projection on the Krylov subspace and the second gives the recurrence relation that is used in MC-GW.

$$S_{i+1,i}^{n} \equiv q_{i+1}^{\dagger} d^{n} q_{i} = C_{i}^{-1} \left[ S_{i,i}^{n+1} - M_{i} S_{i,i}^{n} - C_{i-1}^{\dagger} S_{i-1,i}^{n} \right]$$
(3.101)

or

$$S_{i+1,i+1}^{n} \equiv q_{i+1}^{\dagger} d^{n} q_{i+1} = C_{i}^{-1} \left[ S_{i,i}^{n+2} + M_{i} S_{i,i}^{n} M_{i} + C_{i-1}^{\dagger} S_{i-1,i-1}^{n} C_{i-1} - P(S_{i,i}^{n+1} M_{i}) - P(S_{i,i-1}^{n+1} C_{i-1}) + P(M_{i} S_{i,i-1}^{n} C_{i-1}) \right] C_{i}^{\dagger,-1}$$
(3.102)

or

$$C_{i}^{2} = C_{i}C_{i}^{\dagger} = S_{i,i}^{2} + M_{i}^{2} + C_{i-1}^{\dagger}C_{i-1} - P\left(S_{i,i}^{1}M_{i}\right) - P\left(S_{i,i-1}^{1}C_{i-1}\right)$$
(3.103)

with

$$C_i = q_i^{\dagger} dq_{i+1} \tag{3.105}$$

The curve that I label as MCGW dense corresponds to the middle equality, MCGW sparse corresponds to the right equality, and the black curve is the standard block Lanczos with matrix vector products. Then, I also display the solution that Booth code gives. There is a reason why I am not able to get the MCGW sparse curve to match the dense curve. Recall that to solve 3.107, we want to perform a Cholesky decomposition of the matrix  $C_i^2$  to get  $C_i$ . So the matrix should be symmetric and also positive semidefinite. I always begin by trying to use the modified Cholesky algorithm to perform this matrix square root. However, even in the exact dense theory, the matrix  $C_i^2$  is not always even positive semidefinite (it has negative eigenvalues), so I cannot use the modified Cholesky algorithm to solve in that case. In addition, due to the rise in numerical imprecision, the matrix also becomes non-symmetric, so in these cases I just symmetrize it, get the eigendecomposition, and do the matrix square root that way. But in exact arithmetic, this is a Hermitian theory, so I shouldn't have to symmetrize like this and in doing so I introduce some error. As can be seen, this gives a fairly good match with what Booth codes predict; we are probably doing similar things.

# 3.9 Canonical orthogonalization

The numerical issue with the theory of MCGW seems to lie with then obtaining the appropriate off-diagonals from

$$\boldsymbol{C}_{i}^{2} = \boldsymbol{C}_{i} \boldsymbol{C}_{i}^{\dagger} = \boldsymbol{S}_{i,i}^{2} + \boldsymbol{M}_{i}^{2} + \boldsymbol{C}_{i-1}^{\dagger} \boldsymbol{C}_{i-1} - P\left(\boldsymbol{S}_{i,i}^{1} \boldsymbol{M}_{i}\right) - P\left(\boldsymbol{S}_{i,i-1}^{1} \boldsymbol{C}_{i-1}\right)$$
(3.106)

(3.107)

Let us now try to achieve the same result using canonical orthogonalization. Once we build  $C_i^2$  we can take the eigendecomposition of the matrix to get

$$C_i^2 = U_i \Lambda_i U_i^{\dagger} \tag{3.108}$$

# 3.10 Krylov subspace orthogonality contradiction

## 3.10.1 Theoretical background

There seems to be a contradiction regarding the orthogonality of the Krylov vectors in MCGW. We know that I am able to see convergence to ED to within machine precision when I explicitly build up a Krylov subspace, taking care to ensure orthogonality between Krylov vectors, by both doing Gram-Schmidt on each new vector with all previous ones and then throwing away the small singular values if there is a lack of new directions. This is the exact block Lanczos with matrix vector products curve that I sent Tuesday. Now, MCGW claims to enforce orthogonality in the Krylov subspace by doing  $S_{i,j}^{(0)} = \delta_{ij} I$  with

$$\boldsymbol{S}_{i,j}^{(n)} = \boldsymbol{q}_i^{\dagger} \boldsymbol{d}^n \boldsymbol{q}_j \tag{3.109}$$

This is what rationalizes the simplification in the rewriting of the coupling block W as R padded with 0s. For a brief refresher, see the flow in the equations below in the example of a Krylov subspace which is containing only two block vectors.

$$\tilde{\boldsymbol{H}}_{\text{upfolded}}^{\text{Lanczos Iter 2}} = \tilde{\boldsymbol{Q}}^{(2,\dagger)} \boldsymbol{H}_{\text{Upfolded}}^{G_0 W_0} \tilde{\boldsymbol{Q}}^{(2)} \tag{3.110}$$

$$= \begin{pmatrix} \boldsymbol{I} & \boldsymbol{0} \\ \boldsymbol{0} & [\boldsymbol{q}_1^{\dagger} & \boldsymbol{q}_2^{\dagger}] \end{pmatrix} \begin{pmatrix} \boldsymbol{F} + \boldsymbol{\Sigma}(\infty) & \boldsymbol{W} \\ \boldsymbol{W}^{\dagger} & \boldsymbol{d} \end{pmatrix} \begin{pmatrix} \boldsymbol{I} & \boldsymbol{0} \\ \boldsymbol{0} & [\boldsymbol{q}_1 & \boldsymbol{q}_2] \end{pmatrix}$$

$$= \begin{pmatrix} \boldsymbol{F} + \boldsymbol{\Sigma}(\infty) & [\boldsymbol{W}\boldsymbol{q}_1 & \boldsymbol{W}\boldsymbol{q}_2] \\ \boldsymbol{\mathcal{R}}\boldsymbol{\mathcal{Q}}^{\dagger}\boldsymbol{\mathcal{Q}} = \boldsymbol{\mathcal{R}} & \boldsymbol{0} \\ (\boldsymbol{q}_1^{\dagger}\boldsymbol{d}\boldsymbol{q}_1 & \boldsymbol{q}_1^{\dagger}\boldsymbol{d}\boldsymbol{q}_2) \\ \boldsymbol{\mathcal{Q}}^{\dagger}\boldsymbol{\mathcal{Q}} + \boldsymbol{\mathcal{R}}^{\dagger} = \boldsymbol{\mathcal{R}}^{\dagger} & \boldsymbol{q}_2^{\dagger}\boldsymbol{W}^{\dagger} \end{bmatrix} \begin{pmatrix} \boldsymbol{q}_1^{\dagger}\boldsymbol{d}\boldsymbol{q}_1 & \boldsymbol{q}_1^{\dagger}\boldsymbol{d}\boldsymbol{q}_2 \\ \boldsymbol{M}_1 & \boldsymbol{C}_1 \\ \boldsymbol{q}_2^{\dagger}\boldsymbol{d}\boldsymbol{q}_1 & \boldsymbol{q}_2^{\dagger}\boldsymbol{d}\boldsymbol{q}_2 \end{pmatrix} \end{pmatrix} \tag{3.112}$$

But now recall that MCGW relies on the three-term Lanczos recurrence of

$$\mathbf{q}_{i+1}\mathbf{C}_{i}^{\dagger} = \left[\mathbf{d}\mathbf{q}_{i} - \mathbf{q}_{i}\mathbf{M}_{i} - \mathbf{q}_{i-1}\mathbf{C}_{i-1}\right] \implies \mathbf{q}_{i+1} = \left[\mathbf{d}\mathbf{q}_{i} - \mathbf{q}_{i}\mathbf{M}_{i} - \mathbf{q}_{i-1}\mathbf{C}_{i-1}\right]\mathbf{C}_{i}^{\dagger,-1} \quad (3.113)$$

which is used in the definition of the Ses. Crucially, orthogonality is never enforced in these relations, so after sufficient iterations we get  $\mathbf{q}_i^{\dagger}\mathbf{q}_j \neq 0$  for  $i \neq j$ . This is in contradiction with our initial assumption that the coupling block  $\mathbf{W}$  can be written as  $\mathbf{R}$  padded with zeros. So it is perfectly natural to not see MCGW converge to ED to within machine precision as is seen in the MCGW sparse curve of my plots. Recall that the MCGW dense curve was obtained by explicitly plugging in the Krylov vector from the exact block Lanczos with matrix vector products into 3.109; this was done with genuinely orthogonalized Krylov vectors and hence we saw an exact convergence to ED similar to that of the matrix vector product implementation.

### 3.10.2 Comment about the effect of minimal basis

It is true that the fact that I was experiencing a lack of new directions was an artifact of me using a minimal basis in my calculations. This put the size of the physical space, which scales as O(O+V), on par with the auxiliary space, which scales as  $O(O^2V)$  or  $O(V^2O)$ , depending on whether we are dealing with the lesser or greater channel, respectively. The dimension of the auxiliary space will grow faster than that of the physical space as we move to larger system sizes, so it is expected that this effect would be negligible by using a larger basis in my calculation. However, this does not solve the issue of not being able to maintain orthogonality by doing Gram-Schmidt of a new vector on all previous vectors. But, I did a numerical tests of the effect of a lack of reorthogonalization on the Ritz values a few months ago, and the effect was negligible.

# 3.10.3 Next steps

Should I run this with double zeta basis in C++? At the end of the day, I have just speculated, and I haven't proven anything, so it would be nice to see if what I hypothesis plays out in practice.

# 3.11 A hierarchy of approximations to block Lanczos with reorthogonalization

To simplify the discussion, I will just consider a single channel here.

### 3.11.1 Exact solution

The QP energies are given by diagonalizing the upfolded Hamiltonian

$$\boldsymbol{H}_{\text{upfolded}} = \begin{pmatrix} \boldsymbol{F} & \boldsymbol{W} \\ \boldsymbol{W}^{\dagger} & \boldsymbol{d} \end{pmatrix} \tag{3.114}$$

We can accomplish the same thing by considering the Ritz values as obtained by projecting the upfolded Hamiltonian, which is called exact d, onto an orthonormal Krylov subspace as

This is what the matrix-vector product and exact d both do when we apply reorthogonalization.

## 3.11.2 Matrix-vector product without reorthogonalization

This is one step down from the exact answer. The only reason that we are able to set  $q_i dq_j = 0$  for i > j + 1 in 3.117 is because we are assuming an orthogonal basis where

$$\boldsymbol{dq}_{j} = \boldsymbol{C}_{j-1}\boldsymbol{q}_{j-1} + \boldsymbol{M}_{j}\boldsymbol{q}_{j} + \boldsymbol{C}_{j+1}\boldsymbol{q}_{j+1} \implies \boldsymbol{q}_{j+2}^{\dagger}\boldsymbol{dq}_{j} = \boldsymbol{q}_{j+2}^{\dagger}\left(\boldsymbol{C}_{j-1}^{\dagger}\boldsymbol{q}_{j-1} + \boldsymbol{M}_{j}^{\dagger}\boldsymbol{q}_{j} + \boldsymbol{C}_{j+1}^{\dagger}\boldsymbol{q}_{j+1}\right) = \boldsymbol{0}$$
(3.118)

We will call this assumption 1. But in general, with a non-orthogonal basis, we have

$$dq_j = \sum_{\ell=1}^{j+1} \boldsymbol{H}_{j,\ell} q_{\ell}. \tag{3.119}$$

This becomes the case after a few Lanczos iterations, which I have shown before. However, the fact that we have written the coupling block as  $\mathbf{R}$  padded with zeros (I will refer to this as assumption 2) is consistent with the way that we have constructed the auxiliary block, so even though we have violated assumption 1, assumption 2 still holds, and we get results one level down from the exact solution.

# 3.11.3 Exact d without reorthogonalization

Because this is just projecting d onto an increasingly non-orthogonal Krylov subspace, we are no longer making assumption 1. To make it more clear what I mean here, consider that we are no longer assuming a tridiagonal form for the auxiliary block in 3.117, as we are allowing for nonzero  $q_i dq_j$  with i > j + 1. But now we are violating assumption 2, which

is why this gives different results than 3.11.2. Do we want to understand under what conditions approximation 1 seems to do better than 2, and if so, then how? It doesn't seem to matter in the context of MC-GW, but maybe you have a broader vision.

### 3.11.4 MCGW with recurrence relation

In our derivation of the extended recurrence relation, there were 3 instances where we assumed that we could substitute in our expression for  $q_{i+1}$  from 3.118 as

$$\mathbf{q}_{i+1} = \left[ \mathbf{d}\mathbf{q}_i - \mathbf{q}_i \mathbf{M}_i - \mathbf{q}_{i-1} \mathbf{C}_{i-1} \right] \mathbf{C}_i^{\dagger,-1}$$
(3.120)

2 were when defined a  $S_{i+1,i+1}^n$  and 1 was when we defined  $S_{i+1,i}^n$ . This uses more approximations than 3.11.3, so it makes sense that we get worse results.

# 3.12 Next steps

I have a C++ code that I can run now for the dzvp basis, but it is still to slow, which was why I was only able to show you a limited amount of results last time. I will talk to people in the group about how to make it faster. I guess you want me to generate plots like this one for the larger system sizes.

# Chapter 4

# Berkelbach's GW

#### Formulation for the dTDA 4.1

#### 4.1.1Comparing to Booth's ED

For simplicity, we will just work with a single candle, the lesser one. In the dTDA case, Booth's formulation for the upfolded Hamiltonian is

$$\boldsymbol{H} = \begin{pmatrix} \boldsymbol{F} & \boldsymbol{W} \\ \boldsymbol{W}^{\dagger} & \boldsymbol{d} \end{pmatrix} \tag{4.1}$$

where we have the definitions

$$W_{pkv} = \sum_{ia} (pk|ia) X_{ia}^{v} \quad \text{and} \quad d_{kv,lv'} = (\epsilon_k - \Omega_v) \, \delta_{k,l} \delta_{v,v'}$$

$$\tag{4.2}$$

Now, Tim's version of the Hamiltonian is given by

$$\boldsymbol{H} = \begin{pmatrix} \boldsymbol{F} & \boldsymbol{V}^{2\text{h1p}} \\ (\boldsymbol{V}^{2\text{h1p}})^{\dagger} & \boldsymbol{C}^{2\text{hlp}} \end{pmatrix}$$
(4.3)

where the definitions of the matrix elements are

$$V_{p,k[lc]}^{2 \text{ h1p}} = \langle pc|kl \rangle \equiv (pk|lc) \tag{4.4}$$

$$V_{p,k[lc]}^{2 \text{ h1p}} = \langle pc|kl \rangle \equiv (pk|lc)$$

$$C_{i[ja],k[lc]}^{2 \text{ h1p}} = [(\epsilon_i + \epsilon_j - \epsilon_a) \, \delta_{jl} \delta_{ac} - \langle jc|al \rangle] \, \delta_{ik}$$

$$(4.4)$$

and in particular, we have a definition

$$C^{2\text{hlp}} = \epsilon^{1 \text{ h}} \oplus (-A) = \epsilon^{1 \text{ h}} \otimes 1 + 1 \otimes (-A)$$

$$(4.6)$$

Let us define a unitary rotation  $U = 1 \oplus_{\text{diag}} X$ . Application of this unitary to the Hamiltonian will not change the spectrum and actually transforms the problem into

$$\boldsymbol{H}' = \boldsymbol{U}^{\dagger} \boldsymbol{H} \boldsymbol{U} = \begin{pmatrix} 1 & 0 \\ 0 & \boldsymbol{X}^{\dagger} \end{pmatrix} \begin{pmatrix} \boldsymbol{F} & \boldsymbol{V}^{2\text{h1p}} \\ (\boldsymbol{V}^{2 \text{ h1p}})^{\dagger} & \boldsymbol{C}^{2\text{hlp}} \end{pmatrix} \begin{pmatrix} 1 & 0 \\ 0 & \boldsymbol{X} \end{pmatrix} = \begin{pmatrix} \boldsymbol{F} & \boldsymbol{V}^{2\text{h1p}} \boldsymbol{X} \\ \boldsymbol{X}^{\dagger} (\boldsymbol{V}^{2 \text{ h1p}})^{\dagger} & \boldsymbol{X}^{\dagger} \boldsymbol{C}^{2\text{hlp}} \boldsymbol{X} \end{pmatrix}. \tag{4.7}$$

Now let us evaluate  $X^{\dagger}C^{2\mathrm{hlp}}X$  in the dTDA case. We have

$$X^{\dagger}C^{2\mathrm{hlp}}X = X^{\dagger} \left( \epsilon^{1 \mathrm{h}} \otimes 1 + 1 \otimes (-A) \right) X \tag{4.8}$$

$$= X^{\dagger} \left( \epsilon^{1 \text{ h}} \otimes 1 \right) X + X^{\dagger} \left( 1 \otimes (-A) \right) X \tag{4.9}$$

$$= \left(\epsilon^{1 \text{ h}} \underbrace{\mathbf{X}^{\dagger} \mathbf{X}}_{\mathbf{I}}\right) \otimes \left(\underbrace{\mathbf{X}^{\dagger} \mathbf{X}}_{\mathbf{I}}\right) + \left(\underbrace{\mathbf{X}^{\dagger} \mathbf{X}}_{\mathbf{I}}\right) \otimes \left(-\underbrace{\mathbf{X}^{\dagger} \mathbf{A} \mathbf{X}}_{\mathbf{Q}}\right) \tag{4.10}$$

$$= \epsilon^{1 \text{ h}} \oplus (-\Omega) \equiv \mathbf{d}^{<} \tag{4.11}$$

where we have used the fact that  $X^{\dagger}X = 1$ , since X is unitary. Similarly, we can evaluate the other term

$$\boldsymbol{V}^{2 \text{ hlp}} \boldsymbol{X} = \boldsymbol{W}^{<} \equiv \sum_{lc} (pk|lc) X_{lc}^{v}$$
(4.12)

(4.13)

where we have used the definition of the W matrix in Booth's formulation, as given in (4.2).

## 4.1.2 Deriving the matrix vector products

Now, we can define a vector  $\mathbf{R} = (r_i, r_a, r_{i[jb]}, r_{[jb]a})$ . Application of the Hamiltonian to this vector gives us the matrix-vector product  $\mathbf{H}\mathbf{R} = \boldsymbol{\sigma}$ , where  $\boldsymbol{\sigma} = (\sigma_i, \sigma_a, \sigma_{i[jb]}, \sigma_{[jb]a})$ . Consider

$$\boldsymbol{H}\boldsymbol{R} = \begin{pmatrix} \boldsymbol{F} & \boldsymbol{V}^{2\text{h1p}} & \boldsymbol{V}^{2\text{p1 h}} \\ (\boldsymbol{V}^{2\text{h1p}})^{\dagger} & \boldsymbol{C}^{2\text{hlp}} & \boldsymbol{0} \\ (\boldsymbol{V}^{2\text{plh}})^{\dagger} & \boldsymbol{0} & \boldsymbol{C}^{2\text{plh}} \end{pmatrix} \begin{pmatrix} r_i \\ r_a \\ r_{i[jb]} \\ r_{[jb]a} \end{pmatrix} = \begin{pmatrix} \sigma_i \\ \sigma_a \\ \sigma_{i[jb]} \\ \sigma_{[jb]a} \end{pmatrix}$$
(4.14)

(4.15)

Let us enumerate now what we actually will get:

$$\sigma_i = \sum_{j} f_{ij} r_j + \sum_{b} f_{ib} r_b + \sum_{klc} \langle i c \mid k l \rangle r_{k[lc]} + \sum_{kcd} \langle i k \mid d c \rangle r_{[kc]d}, \tag{4.16}$$

$$\sigma_a = \sum_{j} f_{aj} r_j + \sum_{b} f_{ab} r_b + \sum_{klc} \langle a c \mid k l \rangle r_{k[lc]} + \sum_{kcd} \langle a k \mid d c \rangle r_{[kc]d}, \tag{4.17}$$

$$\sigma_{i[ja]} \sum_{k} \langle k \, a \mid i \, j \rangle \, r_k + \sum_{b} \langle b \, a \mid i \, j \rangle \, r_b + (\epsilon_i + \epsilon_j - \epsilon_a) \, r_{i[ja]} - \sum_{lc} \langle j \, c \mid a \, l \rangle \, r_{i[lc]}$$
(4.18)

$$\sigma_{[ia]b} = \sum_{j} \langle j i \mid b a \rangle r_{j} + \sum_{c} \langle c i \mid b a \rangle r_{c} + (\epsilon_{a} + \epsilon_{b} - \epsilon_{i}) r_{[ia]b} + \sum_{kc} \langle a k \mid i c \rangle r_{[kc]b}. \quad (4.19)$$

# 4.2 Formulation for the dRPA

# 4.2.1 Deriving the dRPA approach as they propose

So we start with this generalized eigenvalue equation

$$\mathbf{M}oldsymbol{Z} = \mathbf{N}oldsymbol{Z} \left(egin{array}{cc} oldsymbol{\Omega}_+ & 0 \ 0 & -oldsymbol{\Omega}_+ \end{array}
ight)$$

where

$$\mathbf{M} = \left(egin{array}{cc} \mathbf{A} & \mathbf{B} \\ \mathbf{B} & \mathbf{A} \end{array}
ight) \quad \mathbf{N} = \left(egin{array}{cc} 1 & \mathbf{0} \\ \mathbf{0} & -1 \end{array}
ight) \quad oldsymbol{Z} = \left(egin{array}{cc} oldsymbol{X} & oldsymbol{Y} \\ oldsymbol{Y} & oldsymbol{X} \end{array}
ight)$$

and  $\Omega_+$  is a diagonal matrix of positive excitation energies. Left multiplying both sides by N and right multiplying by  $Z^{-1}$  gives us

$$NM\underbrace{ZZ^{-1}}_{1} = \underbrace{NN}_{1}Z\begin{pmatrix} \Omega_{+} & 0 \\ 0 & -\Omega_{+} \end{pmatrix}Z^{-1} \implies -NM = -Z\begin{pmatrix} \Omega_{+} & 0 \\ 0 & -\Omega_{+} \end{pmatrix}Z^{-1}$$
(4.20)

Now we can use the fact that the action of a scalar function f, such as the step function, on a diagonalizable matrix  $X \equiv Y\Lambda Y^{-1}$  can be expressed as

$$f(\boldsymbol{X}) = \boldsymbol{Y} f(\boldsymbol{\Lambda}) \boldsymbol{Y}^{-1}$$

so we can write

$$\Theta(-NM) = Z \begin{pmatrix} \Theta(-\Omega_{+}) & 0 \\ 0 & \Theta(\Omega_{+}) \end{pmatrix} Z^{-1} = Z \begin{pmatrix} 0 & 0 \\ 0 & 1 \end{pmatrix} Z^{-1}$$
(4.21)

and so it becomes clear that if we define  $\tilde{M} = M + \eta N\Theta(-NM)$ , we can write

$$\tilde{\boldsymbol{M}}\boldsymbol{Z} = \boldsymbol{M}\boldsymbol{Z} + \eta \boldsymbol{N}\boldsymbol{\Theta}(-\boldsymbol{N}\boldsymbol{M})\boldsymbol{Z} = \boldsymbol{N}\boldsymbol{Z}\begin{pmatrix} \boldsymbol{\Omega}_{+} & \boldsymbol{0} \\ \boldsymbol{0} & -\boldsymbol{\Omega}_{+} \end{pmatrix} + \boldsymbol{N}\boldsymbol{Z}\begin{pmatrix} \boldsymbol{0} & \boldsymbol{0} \\ \boldsymbol{0} & \boldsymbol{\eta} \end{pmatrix} \underbrace{\boldsymbol{Z}^{-1}\boldsymbol{Z}}_{1} = \boldsymbol{N}\boldsymbol{Z}\begin{pmatrix} \boldsymbol{\Omega}_{+} & \boldsymbol{0} \\ \boldsymbol{0} & -\boldsymbol{\Omega}_{+} + \boldsymbol{\eta} \end{pmatrix}$$

$$(4.22)$$

Now they define a super-matrix,

$$\boldsymbol{H} = \begin{pmatrix} \boldsymbol{F} & \boldsymbol{V}^{2\text{h1p}} & \boldsymbol{V}^{2\text{h1p}} & \boldsymbol{V}^{2\text{plh}} & \boldsymbol{V}^{2\text{plh}} \\ (\boldsymbol{V}^{2\text{h1p}})^{\dagger} & & & & & \\ (\boldsymbol{V}^{2\text{h1p}})^{\dagger} & & \boldsymbol{C}^{2\text{hlp}} & & \boldsymbol{0} \\ (\boldsymbol{V}^{2\text{plh}})^{\dagger} & & & & & \\ (\boldsymbol{V}^{2\text{plh}})^{\dagger} & & & & & & \\ (\boldsymbol{V}^{2\text{plh}})^{\dagger} & & & & & & \\ \end{pmatrix}$$
(4.23)

where  $\mathbf{C}^{2\mathrm{hlp}} = \varepsilon^{1\ \mathrm{h}} \oplus (-\tilde{\mathbf{M}})$  and  $\mathbf{C}^{2\mathrm{plh}} = \varepsilon^{1\mathrm{p}} \oplus \tilde{\mathbf{M}}$ , which for which we use the super-metric

$$\mathcal{N} = \left( \begin{array}{ccc} \mathbf{1} & 0 & 0 \\ 0 & \mathbf{1} \oplus \mathbf{N} & 0 \\ 0 & 0 & \mathbf{1} \oplus \mathbf{N} \end{array} \right)$$

and the excitation vector is  $\mathbf{R}=(\mathbf{R}^{1\mathrm{h}+1\mathrm{p}},\mathbf{R}^{2\mathrm{h}1\mathrm{p}},\mathbf{R}^{2\mathrm{p}1\mathrm{h}})$  with  $\mathbf{R}^{2\mathrm{h}1\mathrm{p}}=(\mathbf{r}^{2\mathrm{h}1\mathrm{p}},\bar{\mathbf{r}}^{2\mathrm{h}1\mathrm{p}})$  and  $\mathbf{R}^{2\mathrm{p}1\mathrm{h}}=(\mathbf{r}^{2\mathrm{p}1\mathrm{h}},\bar{\mathbf{r}}^{2\mathrm{p}1\mathrm{h}})$ . So  $\mathcal{N}H\mathbf{R}=\mathbf{R}\mathbf{E}$  leads to:

$$\begin{pmatrix} \mathbf{1} & 0 & 0 \\ 0 & \mathbf{1} \oplus \mathbf{N} & 0 \\ 0 & 0 & \mathbf{1} \oplus \mathbf{N} \end{pmatrix} \begin{pmatrix} \mathbf{F} & \mathbf{V}^{2\text{h1p}} & \mathbf{V}^{2\text{h1p}} & \mathbf{V}^{2\text{plh}} & \mathbf{V}^{2\text{plh}} \\ (\mathbf{V}^{2\text{h1p}})^{\dagger} & & & & & \\ (\mathbf{V}^{2\text{h1p}})^{\dagger} & & \mathbf{C}^{2\text{hlp}} & & \mathbf{0} \\ (\mathbf{V}^{2\text{plh}})^{\dagger} & & & & & \\ (\mathbf{V}^{2\text{plh}})^{\dagger} & & & & & \\ (\mathbf{V}^{2\text{plh}})^{\dagger} & & & & & \\ & & & & & & \\ \end{pmatrix} \begin{pmatrix} \mathbf{r}^{1\text{h+1p}} \\ \mathbf{r}^{2\text{h1p}} \\ \mathbf{r}^{2\text{h1p}} \\ \mathbf{r}^{2\text{p1h}} \\ \mathbf{r}^{2\text{p1h}} \end{pmatrix} = E \begin{pmatrix} \mathbf{r}^{1\text{h+1p}} \\ \mathbf{r}^{2\text{h1p}} \\ \mathbf{r}^{2\text{p1h}} \\ \mathbf{r}^{2\text{p1h}} \end{pmatrix}$$

$$(4.24)$$

The 5 coupled equations can be written as

$$Fr^{1h+1p} + V^{2h1p} (r^{2h1p} + \bar{r}^{2h1p}) + V^{2plh} (r^{2p1h} + \bar{r}^{2p1h}) = Er^{1h+1p}$$
 (4.25)

$$(\boldsymbol{V}^{2\text{h1p}})^{\dagger} \boldsymbol{r}^{1\text{h+1p}} + [\tilde{\boldsymbol{C}}^{2\text{hlp}}]^{xx} \boldsymbol{r}^{2\text{h1p}} + [\tilde{\boldsymbol{C}}^{2\text{hlp}}]^{xd} \bar{\boldsymbol{r}}^{2\text{h1p}} = E \boldsymbol{r}^{2\text{h1p}}$$
(4.26)

$$(\boldsymbol{V}^{2\text{h1p}})^{\dagger} \boldsymbol{r}^{1\text{h+1p}} + [\tilde{\boldsymbol{C}}^{2\text{hlp}}]^{dx} \boldsymbol{r}^{2\text{h1p}} + [\tilde{\boldsymbol{C}}^{2\text{hlp}}]^{dd} \bar{\boldsymbol{r}}^{2\text{h1p}} = E \bar{\boldsymbol{r}}^{2\text{h1p}}$$
(4.27)

$$(\boldsymbol{V}^{2\text{plh}})^{\dagger} \boldsymbol{r}^{1\text{h}+1\text{p}} + [\tilde{\boldsymbol{C}}^{2\text{plh}}]^{xx} \boldsymbol{r}^{2\text{p1h}} + [\tilde{\boldsymbol{C}}^{2\text{plh}}]^{xd} \bar{\boldsymbol{r}}^{2\text{p1h}} = E \boldsymbol{r}^{2\text{p1h}}$$
(4.28)

$$(\boldsymbol{V}^{2\text{plh}})^{\dagger} \boldsymbol{r}^{1\text{h}+1\text{p}} + [\tilde{\boldsymbol{C}}^{2\text{plh}}]^{dx} \boldsymbol{r}^{2\text{p1h}} + [\tilde{\boldsymbol{C}}^{2\text{plh}}]^{dd} \bar{\boldsymbol{r}}^{2\text{p1h}} = E\bar{\boldsymbol{r}}^{2\text{p1h}}$$
(4.29)

where now we have  $\tilde{C}^{2\text{hlp}} = \varepsilon^{1 \text{ h}} \oplus (-N\tilde{M})$  and  $\tilde{C}^{2\text{plh}} = \varepsilon^{1\text{p}} \oplus N\tilde{M}$ . Solving the last two equations gives:

$$\boldsymbol{r}^{2\text{h1p}} + \bar{\boldsymbol{r}}^{2\text{h1p}} = \left(E\boldsymbol{1} - \tilde{\boldsymbol{C}}^{2\text{hlp}}\right)^{-1} \left( \left(\boldsymbol{V}^{2\text{h1p}}\right)^{\dagger} \boldsymbol{r}^{1\text{h+1p}} - \tilde{\boldsymbol{C}}^{2\text{plh}} \left(\boldsymbol{r}^{2\text{p1h}} + \bar{\boldsymbol{r}}^{2\text{p1h}}\right) \right)$$
(4.30)

$$\boldsymbol{r}^{2\text{plh}} + \bar{\boldsymbol{r}}^{2\text{plh}} = \left(\tilde{\boldsymbol{C}}^{2\text{plh}} - E\boldsymbol{1}\right)^{-1} \left( \left(\boldsymbol{V}^{2\text{plh}}\right)^{\dagger} \boldsymbol{r}^{1\text{h}+1\text{p}} + \tilde{\boldsymbol{C}}^{2\text{hlp}} \left(\boldsymbol{r}^{2\text{hlp}} + \bar{\boldsymbol{r}}^{2\text{hlp}}\right) \right)$$
(4.31)

The eigenvalue equation is then given by what they claim gives  $HR = \mathcal{N}RE$  and then leading to the non-Hermitian eigenvalue equation  $\mathcal{N}HR = RE$ .

### Downfolding equivalence

Let us carry this matrix multiplication out, just for the 2h1p sector. Collect the relevant components of the excitation vector into a block vector:

$$oldsymbol{R} = egin{pmatrix} oldsymbol{r}_i & oldsymbol{r}_a & oldsymbol{r}_{i[jb]} & ar{oldsymbol{r}}_{i[jb]} \end{pmatrix}^\dagger$$

We will define the total dimension of this excitation vector as  $T = P + A_O$ , where P = O + V and  $A_O = 2O^2V$ . The super-Hamiltonian decomposes as:

$$oldsymbol{H} = egin{pmatrix} \left(oldsymbol{F}_{oo} & oldsymbol{F}_{ov} \ oldsymbol{F}_{vo} & oldsymbol{F}_{vv} 
ight)_{PP} & \left(oldsymbol{V}_{p,i[jb]}^{2 ext{hlp}} & oldsymbol{V}_{p,i[jb]}^{2 ext{hlp}} 
ight)_{PA_O} \ \left(oldsymbol{V}_{p,k[lc]}^{2 ext{hlp}} & oldsymbol{V}_{p,k[lc]}^{2 ext{hlp}} 
ight)_{A_OP} & \left(\left(arepsilon^{1h}
ight) \oplus_{ ext{kron}} - ilde{oldsymbol{M}}_{jb,jb;lc,lc} 
ight)_{A_OA_O} \end{pmatrix}$$

### Equivalence to Booth form

We would need to perform a rotation by The matrix-vector product becomes:

$$oldsymbol{\sigma} = oldsymbol{\mathcal{N}} egin{pmatrix} \left(egin{array}{ccc} oldsymbol{F}_{oo} & oldsymbol{F}_{ov} \\ oldsymbol{F}_{vo} & oldsymbol{F}_{vv} \end{array}
ight)_{PP} \left(oldsymbol{r}_{i} \\ oldsymbol{r}_{p,i[jb]} & oldsymbol{V}_{p,i[jb]}^{2\mathrm{h1p}} & oldsymbol{V}_{p,i[jb]}^{2\mathrm{h1p}} 
ight)_{A_{O}P}^{\dagger} \left(oldsymbol{r}_{i} \\ oldsymbol{r}_{a} 
ight)_{PT} + \left(\left(arepsilon^{1h}
ight) \oplus_{\mathrm{kron}} - oldsymbol{ ilde{M}}_{jb,jb;lc,lc} 
ight)_{A_{O}A_{O}} \left(oldsymbol{r}_{i[jb]} \\ oldsymbol{ar{r}}_{i[jb]} \end{array}
ight)_{A_{O}T} \end{pmatrix}$$

Now we are left to decide what the super-metric  $\mathcal N$  is. We can define it as

$$\mathcal{N} = \begin{pmatrix} I_{PP} & 0 \\ 0 & N_{A_O A_O} \end{pmatrix}$$

with  $N_{A_OA_O} \equiv \mathbf{1}_O \oplus_{\text{kron}} N_{2\text{OV}} = \mathbf{1}_O \otimes \mathbf{1}_{2OV} + \mathbf{1}_O \otimes N_{2\text{OV}}$ . But if we take this definiton, then  $\mathcal{NN} \neq \mathbf{1}$ , so we can't do  $HR = \mathcal{N}RE \implies \mathcal{N}HR = RE$ , as they claim. Next, assume they made a typo, so do  $N_{A_OA_O} = \begin{pmatrix} \mathbf{1}_{O^2V} & 0 \\ 0 & -\mathbf{1}_{O^2V} \end{pmatrix}$ . Now, we have  $N_{A_OA_O}N_{A_OA_O} = \mathbf{1}_{A_OA_O}$  and

$$\mathcal{N}HR = \begin{pmatrix}
\mathbf{1}_{PP} \begin{pmatrix} \mathbf{F}_{oo} & \mathbf{F}_{ov} \\ \mathbf{F}_{vo} & \mathbf{F}_{vv} \end{pmatrix}_{PP} \begin{pmatrix} \mathbf{r}_{i} \\ \mathbf{r}_{a} \end{pmatrix}_{PT} + \mathbf{1}_{PP} \begin{pmatrix} \mathbf{V}_{p,i[jb]}^{2\text{h1p}} & \mathbf{V}_{p,i[jb]}^{2\text{h1p}} \end{pmatrix}_{PA_{O}} \begin{pmatrix} \mathbf{r}_{i[jb]} \\ \bar{\mathbf{r}}_{i[\bar{j}b]} \end{pmatrix}_{A_{O}T} \\
N_{A_{O}A_{O}} \begin{pmatrix} \mathbf{V}_{p,i[jb]}^{2\text{h1p}} & \mathbf{V}_{p,i[jb]}^{2\text{h1p}} \end{pmatrix}_{A_{O}P}^{\dagger} \begin{pmatrix} \mathbf{r}_{i} \\ \mathbf{r}_{a} \end{pmatrix}_{PT} + \mathbf{N}_{A_{O}A_{O}} \begin{pmatrix} (\varepsilon^{1h}) \oplus_{\text{kron}} - \tilde{\mathbf{M}}_{jb,jb;lc,lc} \end{pmatrix}_{A_{O}A_{O}} \begin{pmatrix} \mathbf{r}_{i[jb]} \\ \bar{\mathbf{r}}_{i[\bar{j}b]} \end{pmatrix}_{A_{O}T} \end{pmatrix}$$

$$(4.32)$$

$$=\begin{pmatrix}\begin{pmatrix} \left(\boldsymbol{F}_{oo}\boldsymbol{r}_{i}+\boldsymbol{F}_{ov}\boldsymbol{r}_{a}\right)_{PT}+\left(\boldsymbol{V}_{k,j[lc]}^{2\mathrm{h1p}}\boldsymbol{r}_{i[jb]}+\boldsymbol{V}_{k,i[jb]}^{2\mathrm{h1p}}\bar{\boldsymbol{r}}_{i[\bar{b}]}\right)_{PT}\\\boldsymbol{V}_{c,a[jb]}^{2\mathrm{h1p}}\boldsymbol{r}_{a[jb]}+\boldsymbol{V}_{c,a[jb]}^{2\mathrm{h1p}}\bar{\boldsymbol{r}}_{a[jb]}\end{pmatrix}_{PT}\\\boldsymbol{N}_{A_{O}A_{O}}\begin{pmatrix}\left(\boldsymbol{V}_{k,i[jb]}^{2\mathrm{h1p}}\boldsymbol{r}_{k}+\boldsymbol{V}_{a,i[jb]}^{2\mathrm{h1p}}\boldsymbol{r}_{a}\right)_{A_{O}T}^{\dagger}+\boldsymbol{N}_{A_{O}A_{O}}\begin{pmatrix}\left(\boldsymbol{\varepsilon}^{1h}\right)\oplus_{\mathrm{kron}}-\left(\boldsymbol{\tilde{M}}_{jb,lc}^{xx}&\boldsymbol{\tilde{M}}_{jb,lc}^{xd}\\\boldsymbol{\tilde{M}}_{jb,lc}^{dx}&\boldsymbol{\tilde{M}}_{jb,lc}^{dd}\end{pmatrix}\right)_{A_{O}A_{O}}\begin{pmatrix}\boldsymbol{r}_{i[jb]}\\\bar{\boldsymbol{r}}_{i[\bar{j}b]}\end{pmatrix}_{A_{O}T}\end{pmatrix}$$

$$(4.33)$$

$$=\begin{pmatrix} \begin{pmatrix} \sum_{j} f_{ij}r_{j} + \sum_{b} f_{ib}r_{b} + \sum_{jlc}\langle ic|jl\rangle r_{j[lc]} + \sum_{jlc}\langle ic|jl\rangle \bar{r}_{j[lc]} \\ \sum_{j} f_{aj}r_{j} + \sum_{b} f_{ab}r_{b} + \sum_{jlc}\langle ac|jl\rangle r_{j[lc]} + \sum_{jlc}\langle ac|jl\rangle \bar{r}_{j[lc]} \end{pmatrix}_{PT} \\ \begin{pmatrix} \sum_{k}\langle ka|ij\rangle r_{k} + \sum_{b}\langle ba|ij\rangle r_{b} \\ -\sum_{k}\langle ka|ij\rangle r_{k} - \sum_{b}\langle ba|ij\rangle r_{b} \end{pmatrix}_{A_{O}T} + \mathbf{N}_{A_{O}A_{O}} \begin{pmatrix} (\varepsilon^{1h}) \oplus_{\mathrm{kron}} - \begin{pmatrix} \tilde{\mathbf{M}}_{jx}^{xx} & \tilde{\mathbf{M}}_{jb,lc}^{xd} \\ \tilde{\mathbf{M}}_{jb,lc}^{dx} & \tilde{\mathbf{M}}_{jb,lc}^{dd} \end{pmatrix}_{A_{O}A_{O}} \begin{pmatrix} \mathbf{r}_{i[jb]} \\ \bar{\mathbf{r}}_{i[\bar{j}\bar{b}]} \end{pmatrix}_{A_{O}T} \end{pmatrix}$$

$$(4.34)$$

$$(4.35)$$

# 4.2.2 Proving ED equivalence to Booth for the dRPA using Z-vector trick

For dRPA, we will choose

$$\boldsymbol{C}^{2\text{hlp}} = \epsilon^{1 \text{ h}} \oplus (-\boldsymbol{M}^{1/2}) \implies C_{i[ja],k[lc]}^{2 \text{ hlp}} = \left[\epsilon_i - \boldsymbol{M}_{ja,lc}^{1/2}\right] \delta_{ik}$$
(4.36)

where M is the matrix defined in (4.41). If we choose our unitary transformation to be  $U' = 1 \otimes T$ , where T is the matrix of eigenvectors of the symmetric formulation as

$$MT = \lambda T \tag{4.37}$$

with the matrix M defined as

$$M = (A - B)^{1/2} (A + B)(A - B)^{1/2}$$
(4.38)

with the identities  $\lambda = \Omega^2$  and  $T = \Omega^{1/2} [A - B]^{-1/2} (X + Y)$ , where X and Y are the excitation and de-excitation vectors, respectively.

Then, we can apply the unitary transformation to the Hamiltonian in Tim's formulation as

$$\boldsymbol{H'} = \boldsymbol{U'}^{\dagger} \boldsymbol{H} \boldsymbol{U'} = \begin{pmatrix} \mathbf{1} & \mathbf{0} \\ \mathbf{0} & \boldsymbol{T}^{\dagger} \end{pmatrix} \begin{pmatrix} \boldsymbol{F} & \boldsymbol{V}^{2\text{hlp}} \\ (\boldsymbol{V}^{2 \text{ hlp}})^{\dagger} & \boldsymbol{C}^{2\text{hlp}} \end{pmatrix} \begin{pmatrix} \mathbf{1} & \mathbf{0} \\ \mathbf{0} & \boldsymbol{T} \end{pmatrix} = \begin{pmatrix} \boldsymbol{F} & \boldsymbol{V}^{2\text{hlp}} \boldsymbol{T} \\ \boldsymbol{T}^{\dagger} \begin{pmatrix} \boldsymbol{V}^{2 \text{ hlp}} \end{pmatrix}^{\dagger} & \boldsymbol{T}^{\dagger} \boldsymbol{C}^{2\text{hlp}} \boldsymbol{T} \end{pmatrix}. \tag{4.39}$$

But now let us try evaluating the term  $T^{\dagger}C^{2\text{hlp}}T$  in the dRPA case. We have

$$T^{\dagger}C^{2\mathrm{hlp}}T = T^{\dagger}\left(\epsilon^{1} \otimes 1 + 1 \otimes (-M^{1/2})\right)T$$
 (4.40)

$$= \mathbf{T}^{\dagger} \left( \epsilon^{1 \text{ h}} \otimes \mathbf{1} \right) \mathbf{T} + \mathbf{T}^{\dagger} \left( \mathbf{1} \otimes (-\mathbf{M}^{1/2}) \right) \mathbf{T}$$

$$(4.41)$$

$$= \left( \epsilon^{1 \text{ h}} \underbrace{\mathbf{T}^{\dagger} \mathbf{T}}_{I} \right) \otimes \left( \underbrace{\mathbf{T}^{\dagger} \mathbf{T}}_{I} \right) + \left( \underbrace{\mathbf{T}^{\dagger} \mathbf{T}}_{I} \right) \otimes \left( - \underbrace{\mathbf{T}^{\dagger} \mathbf{M}^{1/2} \mathbf{T}}_{\Omega} \right)$$
(4.42)

(4.43)

Next, let us consider the product  $V^{2 \text{ hlp}} T$ . We have

$$V^{2 \text{ h1p}} T = \sum_{ia} (pk|ia) \left( X_{ia}^v + Y_{ia}^v \right) \Omega_{ia;v}^{1/2} \underbrace{\sum_{jb} \left( A_{ia,jb} - B_{ia,jb} \right)^{-1/2}}_{\Delta_{ia;v}^{-1/2}}$$
(4.44)

(4.45)

Now we can approximate  $\frac{\Omega_{ia;v}}{\Delta_{ia;v}} \approx \sqrt{1 + 2\frac{K_{ia,jb;v}}{\Delta_{ia;v}}} \Longrightarrow \left(\frac{\Omega_{ia;v}}{\Delta_{ia;v}}\right)^{1/2} \approx \sqrt[4]{1 + 2\frac{K_{ia,jb;v}}{\Delta_{ia;v}}}$ . For weak to moderately correlated systems, we can assume that the ratio of Coulomb coupling to orbital energy gaps  $\frac{K_{ia,jb;v}}{\Delta_{ia;v}}$  is negligible, so we can just assume that  $\left(\frac{\Omega_{ia;v}}{\Delta_{ia;v}}\right)^{1/2} \approx 1$ .

# 4.2.3 Evaluating the matrix-vector products

Now, we can define a vector  $\mathbf{R} = (r_i, r_a, r_{i[jb]}, r_{[jb]a})$ . Application of the Hamiltonian to this vector gives us the matrix-vector product  $\mathbf{H'R} = \boldsymbol{\sigma}$ , where  $\boldsymbol{\sigma} = (\sigma_i, \sigma_a, \sigma_{i[jb]}, \sigma_{[jb]a})$ .

So the one body matrix vector products are the same as in the dTDA case

$$\sigma_{i} = \sum_{j} f_{ij} r_{j} + \sum_{b} f_{ib} r_{b} + \sum_{klc} \langle i c \mid k l \rangle r_{k[lc]} + \sum_{kcd} \langle i k \mid d c \rangle r_{[kc]d}, \tag{4.46}$$

$$\sigma_a = \sum_{i} f_{aj} r_j + \sum_{b} f_{ab} r_b + \sum_{klc} \langle a c \mid k l \rangle r_{k[lc]} + \sum_{kcd} \langle a k \mid d c \rangle r_{[kc]d}, \qquad (4.47)$$

(4.48)

For the two-body matrix vector products, we have

$$\sigma_{i[ja]} = \sum_{k} \langle k \, a \mid i \, j \rangle \, r_k + \sum_{b} \langle b \, a \mid i \, j \rangle \, r_b + \sum_{klc} C_{i[ja],k[lc]}^{2 \text{ hlp}} r_{i[lc]}$$

$$(4.49)$$

Now, consider just

$$\sum_{klc} C_{i[ja],k[lc]}^{2 \text{ hlp}} r_{i[lc]} = \sum_{klc} \left[ \left( \epsilon_i - \boldsymbol{M}_{ja,lc}^{1/2} \right) \delta_{ik} r_{i[lc]} \right] = \sum_{lc} \left( \epsilon_i - \boldsymbol{M}_{ja,lc}^{1/2} \right) r_{i[lc]}$$
(4.50)

Consider now just the term

$$\sum_{lc} \mathbf{M}_{ja,lc}^{1/2} r_{i[lc]} = \sum_{lc} \left[ \underbrace{\Delta_{ja}^{1/2} (\Delta_{ja} + 2(ja|lc))_{ja,lc} \Delta_{ja}^{1/2}}_{\mathbf{D}} \right]^{1/2} r_{i[lc]}. \tag{4.51}$$

Let us consider instead the similarity transformed  $\tilde{\boldsymbol{D}} = \Delta^{-1/2}\boldsymbol{D}\Delta^{-1/2} = (\Delta+2K)$ . This is useful because we have  $D^{1/2} = \Delta^{1/2}\tilde{D}^{1/2}\Delta^{1/2} = (\Delta+2K)^{1/2}$  for our SPD D.Now make the change of variables  $u = \Delta^{1/2}r \implies r = \Delta^{-1/2}u$ . Then we have  $(\Delta+2K)\Delta^{-1/2}u = \Delta^{1/2}\left[I + 2\Delta^{-1/2}K\Delta^{-1/2}\right]u$ . Again, because we are working with weak to moderately correlated systems, we can assume that L is small, and neglect it for now. So we have

$$\sum_{lc} \mathbf{M}_{ja,lc}^{1/2} r_{i[lc]} \approx \sum_{lc} \Delta^{1/2} I \Delta^{1/2} r_{i[lc]} \approx \sum_{lc} \Delta r_{i[lc]}$$

$$\tag{4.52}$$

We can rewrite this as So really we have reduced the scaling to  $O(O^3V^2)$ . If we do a similar thing for the other term, we have

$$\sigma_{[ia]b} = \sum_{i} \langle j i \mid b a \rangle r_j + \sum_{c} \langle c i \mid b a \rangle r_c + \sum_{kc} \left[ \epsilon_b + \boldsymbol{M}_{ia,kc}^{1/2} \right] r_{[kc]b}$$
(4.53)

which gives the scaling of  $O(O^2V^3)$ .

# 4.3 Extra

Recall that we are trying to recover the form

$$\boldsymbol{H} = \begin{pmatrix} \boldsymbol{F} & \boldsymbol{W}^{<} \\ \boldsymbol{W}^{<,\dagger} & \boldsymbol{d}^{<} \end{pmatrix} \tag{4.54}$$

where now we have the definitions

$$W_{pkv}^{\leq} = \sum_{ia} (pk|ia) \left( X_{ia}^{v} + Y_{ia}^{v} \right) \tag{4.55}$$

$$d_{kv,lv'}^{\leq} = (\epsilon_k - \Omega_v) \, \delta_{k,l} \delta_{v,v'} \tag{4.56}$$

where now  $\Omega$  is taken from the eigenproblem  $MT = \Omega^2 T$ , which is the symmetric formulation of the RPA. Application of downfolding on this upfolded form then gave us the expression

$$\Sigma(\omega) = \mathbf{W}^{\langle}(\omega \mathbf{I} - \mathbf{d}^{\langle})^{-1} \mathbf{W}^{\langle,\dagger} + \mathbf{W}^{\rangle}(\omega \mathbf{I} - \mathbf{d}^{\rangle})^{-1} \mathbf{W}^{\rangle,\dagger}$$
(4.57)

Meanwhile, downfolding Tim's formulation gives us the expression for the correlation self energy

$$\Sigma(\omega) = \mathbf{V}^{2\text{hlp}} \left[ \omega \mathbf{1} - \mathbf{C}^{2\text{hlp}} \right]^{-1} \left[ \mathbf{V}^{2\text{hlp}} \right]^{\dagger} + \mathbf{V}^{2\text{plh}} \left[ \omega \mathbf{1} - \mathbf{C}^{2\text{plh}} \right]^{-1} \left[ \mathbf{V}^{2\text{plh}} \right]^{\dagger}$$
(4.58)

where in the dTDA case the supermatrix is

$$\mathbf{H} = \begin{pmatrix} \mathbf{f} & \mathbf{V}^{2\text{hlp}} & \mathbf{V}^{2\text{pl h}} \\ (\mathbf{V}^{2\text{ hlp}})^{\dagger} & \mathbf{C}^{2\text{hlp}} & \mathbf{0} \\ (\mathbf{V}^{2\text{plh}})^{\dagger} & \mathbf{0} & \mathbf{C}^{2\text{plh}} \end{pmatrix}$$
(4.59)

with the definitions

$$V_{p,k|lc|}^{2 \text{ hlp}} = \langle pc|kl \rangle \tag{4.60}$$

$$V_{p,|kc|d}^{2p1\ h} = \langle pk|dc\rangle \tag{4.61}$$

$$C_{i[ja],k[lc]}^{2 \text{ hlp}} = \left[ \left( \epsilon_i + \epsilon_j - \epsilon_a \right) \delta_{jl} \delta_{ac} - \left\langle jc | al \right\rangle \right] \delta_{ik}$$
(4.62)

$$V_{p,k[lc]}^{2 \text{ h1p}} = \langle pc|kl \rangle$$

$$V_{p,k[lc]}^{2 \text{ p1 h}} = \langle pk|dc \rangle$$

$$C_{i[ja],k[lc]}^{2 \text{ h1p}} = [(\epsilon_i + \epsilon_j - \epsilon_a) \, \delta_{jl} \delta_{ac} - \langle jc|al \rangle] \, \delta_{ik}$$

$$C_{[ia]b,[kc]d}^{2 \text{ p1 h}} = [(\epsilon_a + \epsilon_b - \epsilon_i) \, \delta_{ik} \delta_{ac} + \langle ak|ic \rangle] \, \delta_{bd}$$

$$(4.63)$$

Presumably they should be equivalent but I am a little bet confused. In both kisses we have physical and auxiliary spaces. However, Tim uses a bare Coulomb interaction for the coupling matrix, while Booth uses the screened interaction. Then also the definitions of the auxiliary spaces d versus C are different, where the former makes use of the  $\Omega$ , assuming that the RPA calculation has already been performed, while the latter is just defined by the A matrix in the dTDA case. Please show how to get from Tim's formulation to Booth's formulation, or vice versa, and how the two are equivalent. let's just focus on a single channel, so just the lesser channel. Let us start by proving for the dTDA case, where we have the definitions for Booth's formulation

$$W_{pkv}^{<} = \sum_{ia} (pk|ia) (X_{ia}^{v})$$
(4.64)

$$d_{kv,lv'}^{\leq} = (\epsilon_k - \Omega_v) \,\delta_{k,l} \delta_{v,v'} \tag{4.65}$$

where in the dtda case, we are working with the eigenproblem  $AX = \Omega X$ . while in Tim's formulation we have

$$\mathbf{C}^{2\text{hlp}} = \epsilon^{1 \text{ h}} \oplus (-\mathbf{A}) \tag{4.66}$$

(4.67)

with the metrics aliments defined as

$$V_{p,k|lc|}^{2 \text{ hlp}} = \langle pc|kl \rangle \tag{4.68}$$

$$V_{p,k[lc]}^{2 \text{ hlp}} = \langle pc|kl \rangle$$

$$C_{i[ja],k[lc]}^{2 \text{ hlp}} = [(\epsilon_i + \epsilon_j - \epsilon_a) \, \delta_{jl} \delta_{ac} - \langle jc|al \rangle] \, \delta_{ik}$$

$$(4.68)$$

## Derivation with RPA screening using Furche's sym-4.4 metric formulation

In the symmetric formulation, we have the hermitian eigenproblem:

$$MT = \lambda T \tag{4.70}$$

where we define

$$M = (A - B)^{1/2} (A + B) (A - B)^{1/2}$$
(4.71)

with the  $\boldsymbol{A}$  and  $\boldsymbol{B}$  matrices defined as

$$A_{ia,jb} = (\epsilon_a - \epsilon_i) \, \delta_{ij} \delta_{ab} + \mathcal{K}_{ia,bj} \tag{4.72}$$

$$B_{ia,jb} = \mathcal{K}_{ia,jb}. \tag{4.73}$$

From this eigenproblem, we can get the quantities that we need for GW as

$$\Omega_n = \sqrt{\lambda_n}$$

$$(X+Y)_n = \frac{1}{\sqrt{\Omega_n}} (A-B)^{1/2} T_n$$
(4.74)

In direct RPA, one defines a frequency-independent "super-matrix" H,

$$\mathbf{H} = \left(egin{array}{ccc} \mathbf{f} & \mathbf{V}^{2 ext{hlp}} & \mathbf{V}^{2 ext{pl h}} & \mathbf{V}^{2 ext{pl h}} \ \left(\mathbf{V}^{2 ext{ hlp}}
ight)^{\dagger} & \mathbf{C}^{2 ext{hlp}} & \mathbf{0} \ \left(\mathbf{V}^{2 ext{plh}}
ight)^{\dagger} & \mathbf{0} & \mathbf{C}^{2 ext{plh}} \end{array}
ight)$$

If the Tamm-Dancoff approximation is used, we have  $C^{2 \text{ hlp}} = \varepsilon^{1 \text{ h}} \oplus (-A)$  and  $C^{2 \text{plh}} =$  $\varepsilon^{1p} \oplus \mathbf{A}$ .

with matrix elements

$$V_{p,k[lc]}^{2 \text{ h1p}} = \langle pc \mid kl \rangle$$

$$V_{p,k[lc]}^{2\text{p1 h}} = \langle pk \mid dc \rangle$$

$$C_{i[ja],k[lc]}^{2 \text{ h1p}} = [(\varepsilon_i + \varepsilon_j - \varepsilon_a) \, \delta_{jl} \delta_{ac} - \langle jc \mid al \rangle] \, \delta_{ik}$$

$$C_{[ia]b,[kc]d}^{2\text{p1 h}} = [(\varepsilon_a + \varepsilon_b - \varepsilon_i) \, \delta_{ik} \delta_{ac} + \langle ak \mid ic \rangle] \, \delta_{bd}$$

But now, to incorporate the RPA screening, we can instead define  $C^{2\text{h1p}} \equiv \varepsilon^{1\text{ h}} \oplus (-M)$  and  $C^{2\text{p1 h}} \equiv \varepsilon^{1\text{p}} \oplus M$ , where M is the matrix defined above. We can get the matrix elements of M as

$$M_{ia,jb} = (\mathbf{A} - \mathbf{B})_{ia,jb}^{1/2} (\mathbf{A} + \mathbf{B})_{ia,jb} (\mathbf{A} - \mathbf{B})_{ia,jb}^{1/2}$$
(4.75)

$$= \sqrt{\Delta_{ia}} \left[ (\epsilon_a - \epsilon_i) \, \delta_{ij} \delta_{ab} + 2 \mathcal{K}_{ia,bj} \right] \sqrt{\Delta_{jb}}. \tag{4.77}$$

(4.78)

In the second line we have used the fact that in the direct formulation  $\mathbf{A} - \mathbf{B}$  is purely diagonal, so in fact  $(\mathbf{A} - \mathbf{B})^{1/2}$  is just the square root of the diagonal elements, which we denote as  $\sqrt{\Delta_{ia}}$ . Furthermore, we have that the interaction kernels  $\mathcal{K}_{ia,bj} \equiv \mathcal{K}_{ia,jb}$  for the HF reference, where in the physicist's notation we have  $\mathcal{K}_{ia,jb} = \langle ij|ab\rangle$ . So the expressions become

$$\boldsymbol{C}^{2\text{h1p}} = \epsilon^{1 \text{ h}} \oplus (-\boldsymbol{M}) \tag{4.79}$$

$$= \begin{pmatrix} \epsilon^{1h} & 0 \\ 0 & -\mathbf{M} \end{pmatrix} \tag{4.80}$$

and

$$\mathbf{C}^{2\text{p1 h}} = \epsilon^{1\text{p}} \oplus \mathbf{M} \tag{4.81}$$

$$= \begin{pmatrix} \epsilon^{1p} & 0 \\ 0 & \mathbf{M} \end{pmatrix} \tag{4.82}$$

where  $\epsilon^{1 \text{ h}}$  and  $\epsilon^{1 \text{p}}$  are the diagonal matrices of the one-electron energies for the hole and particle states, respectively.

# Chapter 5

# Extraneous derivations

# 5.1 Density matrix idempotency

Given the orthonormality of the Kohn-Sham orbitals:

$$\int d\mathbf{r}\,\varphi_j^*(t,\mathbf{r})\varphi_k(t,\mathbf{r}) = \delta_{jk},$$

we can expand  $\gamma^2(t)$  as follows:

$$\gamma^{2}(t, \mathbf{r}, \mathbf{r}') = \int d\mathbf{r}_{1} \, \gamma(t, \mathbf{r}, \mathbf{r}_{1}) \gamma(t, \mathbf{r}_{1}, \mathbf{r}')$$

$$= \int d\mathbf{r}_{1} \left( \sum_{j} \varphi_{j}(t, \mathbf{r}) \varphi_{j}^{*}(t, \mathbf{r}_{1}) \right) \left( \sum_{k} \varphi_{k}(t, \mathbf{r}_{1}) \varphi_{k}^{*}(t, \mathbf{r}') \right)$$

$$= \sum_{j,k} \varphi_{j}(t, \mathbf{r}) \left( \int d\mathbf{r}_{1} \, \varphi_{j}^{*}(t, \mathbf{r}_{1}) \varphi_{k}(t, \mathbf{r}_{1}) \right) \varphi_{k}^{*}(t, \mathbf{r}')$$

$$= \sum_{j,k} \varphi_{j}(t, \mathbf{r}) \delta_{jk} \varphi_{k}^{*}(t, \mathbf{r}')$$

$$= \sum_{j} \varphi_{j}(t, \mathbf{r}) \varphi_{j}^{*}(t, \mathbf{r}')$$

$$= \gamma(t, \mathbf{r}, \mathbf{r}')$$

Thus,  $\gamma^2(t) = \gamma(t)$ , confirming idempotency.

# 5.2 Density matrix time evolution

First, we take the time derivative of both sides of equation 1.92:

$$i\frac{\partial}{\partial t}\gamma(t,\mathbf{r},\mathbf{r}') = i\frac{\partial}{\partial t} \left( \sum_{j=1}^{N} \varphi_j(t,\mathbf{r})\varphi_j^*(t,\mathbf{r}') \right)$$
(5.1)

$$= \sum_{j=1}^{N} \left[ i \frac{\partial \varphi_j(t, \mathbf{r})}{\partial t} \varphi_j^*(t, \mathbf{r}') + i \varphi_j(t, \mathbf{r}) \frac{\partial \varphi_j^*(t, \mathbf{r}')}{\partial t} \right]$$
 (5.2)

We know that the TDKS equations for the orbitals  $\varphi_j(t, \mathbf{r})$  and their complex conjugates  $\varphi_j^*(t, \mathbf{r}')$  are:

$$i\frac{\partial}{\partial t}\varphi_j(t,\mathbf{r}) = H[\rho](t,\mathbf{r})\varphi_j(t,\mathbf{r})$$
 (5.3)

$$-i\frac{\partial}{\partial t}\varphi_j^*(t, \mathbf{r}') = \varphi_j^*(t, \mathbf{r}')H[\rho](t, \mathbf{r}')$$
(5.4)

where  $H[\rho](t, \mathbf{r})$  is the effective one-particle Hamiltonian.

Substituting these into the time derivative of  $\gamma(t)$ :

$$i\frac{\partial}{\partial t}\gamma(t,\mathbf{r},\mathbf{r}') = \sum_{j=1}^{N} \left[ H[\rho](t,\mathbf{r})\varphi_j(t,\mathbf{r})\varphi_j^*(t,\mathbf{r}') - \varphi_j(t,\mathbf{r})\varphi_j^*(t,\mathbf{r}')H[\rho](t,\mathbf{r}') \right]$$

Notice that the right-hand side can be written in terms of the density matrix  $\gamma(t)$ :

$$\sum_{j=1}^{N} H[\rho](t, \mathbf{r})\varphi_j(t, \mathbf{r})\varphi_j^*(t, \mathbf{r}') = H[\rho](t, \mathbf{r})\gamma(t, \mathbf{r}, \mathbf{r}')$$

$$\sum_{j=1}^{N} \varphi_j(t, \mathbf{r}) \varphi_j^*(t, \mathbf{r}') H[\rho](t, \mathbf{r}') = \gamma(t, \mathbf{r}, \mathbf{r}') H[\rho](t, \mathbf{r}')$$

Therefore, the equation becomes:

$$i\frac{\partial}{\partial t}\gamma(t,\mathbf{r},\mathbf{r}') = H[\rho](t,\mathbf{r})\gamma(t,\mathbf{r},\mathbf{r}') - \gamma(t,\mathbf{r},\mathbf{r}')H[\rho](t,\mathbf{r}')$$

# 5.3 Fourier Transform Symmetry for Real Functions

In the context of response theory, it is essential to ensure that external perturbations are represented by real-valued functions. This requirement imposes a specific symmetry on their Fourier transforms. Specifically, for any real-valued function f(t), its Fourier transform  $F(\omega)$  must satisfy the condition:

$$F(-\omega) = [F(\omega)]^* \tag{5.5}$$

This subsection provides a four-step derivation of this important property.

### Step 1: Definition of the Fourier Transform

The Fourier transform of a real-valued function f(t) is defined as:

$$F(\omega) = \int_{-\infty}^{\infty} f(t)e^{i\omega t} dt$$
 (5.6)

### Step 2: Taking the Complex Conjugate

Since f(t) is real, we have  $f(t) = f^*(t)$ . Taking the complex conjugate of both sides of the Fourier transform yields:

$$[F(\omega)]^* = \left(\int_{-\infty}^{\infty} f(t)e^{i\omega t} dt\right)^*$$

$$= \int_{-\infty}^{\infty} f^*(t)e^{-i\omega t} dt$$

$$= \int_{-\infty}^{\infty} f(t)e^{-i\omega t} dt$$
(5.7)

### Step 3: Evaluating $F(-\omega)$

Substituting  $-\omega$  into the Fourier transform definition:

$$F(-\omega) = \int_{-\infty}^{\infty} f(t)e^{i(-\omega)t} dt$$
$$= \int_{-\infty}^{\infty} f(t)e^{-i\omega t} dt$$
 (5.8)

### Step 4: Establishing the Symmetry

Comparing equations (5.7) and (5.8), we observe that:

$$F(-\omega) = [F(\omega)]^*$$

Thus, for any real-valued function f(t), its Fourier transform satisfies the Hermitian symmetry condition as stated in equation (5.5).

### Conclusion

This symmetry ensures that when constructing real-valued external potentials from their Fourier components, the contributions from positive and negative frequencies are related through complex conjugation. Specifically, in equations (11) and (12) of the response theory framework, the inclusion of both  $e^{i\omega_{\alpha}t}$  and  $e^{-i\omega_{\alpha}t}$  terms with appropriately related spatial components guarantees that the external potentials  $v_{\text{ext}}(t,x)$  and  $\mathbf{A}_{\text{ext}}(t,x)$  remain real-valued functions of time and space.

$$v^{(\alpha)}(-\omega_{\alpha}, x) = [v^{(\alpha)}(\omega_{\alpha}, x)]^*, \quad \mathbf{A}^{(\alpha)}(-\omega_{\alpha}, x) = [\mathbf{A}^{(\alpha)}(\omega_{\alpha}, x)]^*$$

This property is fundamental in ensuring the physical relevance and mathematical consistency of the perturbations applied in density matrix-based response theory.

# 5.4 First-Order Response of an Observable to External Perturbations

In the context of density matrix-based response theory, we are interested in how an observable  $f_{\lambda}(t)$  responds to external perturbations characterized by coupling constants  $\lambda_{\alpha}$ . Specifically, we aim to derive the first-order response given by:

$$\left. \frac{\partial}{\partial \lambda_{\alpha}} f_{\lambda}(t) \right|_{\lambda=0} = f^{(\alpha)}(\omega_{\alpha}) e^{i\omega_{\alpha}t} + f^{(\alpha)}(-\omega_{\alpha}) e^{-i\omega_{\alpha}t}$$
(5.9)

Here,  $f^{(\alpha)}(\omega_{\alpha})$  represents the response of the observable at frequency  $\omega_{\alpha}$ , and  $f^{(\alpha)}(-\omega_{\alpha})$  is its counterpart at the negative frequency.

### Step 1: Expansion of the Observable in Powers of Coupling Constants

Assume that the observable  $f_{\lambda}(t)$  depends on the coupling strengths  $\lambda = \{\lambda_{\alpha}\}$ . We can expand  $f_{\lambda}(t)$  in a Taylor series around  $\lambda = 0$  as follows:

$$f_{\lambda}(t) = f^{(0)}(t) + \sum_{\alpha} \lambda_{\alpha} f^{(\alpha)}(t) + \sum_{\alpha,\beta} \lambda_{\alpha} \lambda_{\beta} f^{(\alpha\beta)}(t) + \cdots$$

Where:

- $\bullet$   $f^{(0)}(t)$  is the unperturbed (zeroth-order) value of the observable.
- $f^{(\alpha)}(t)$  is the first-order response to the perturbation  $\lambda_{\alpha}$ .
- Higher-order terms represent responses to multiple perturbations and their interactions.

### Step 2: Representation of External Perturbations

The external scalar potential is given by Equation (11):

$$v_{\rm ext}(t,x) = v^{(0)}(x) + \sum_{\alpha} \lambda_{\alpha} \left( v^{(\alpha)}(\omega_{\alpha}, x) e^{i\omega_{\alpha}t} + v^{(\alpha)}(-\omega_{\alpha}, x) e^{-i\omega_{\alpha}t} \right)$$

Similarly, the external vector potential is given by Equation (12):

$$\mathbf{A}_{\mathrm{ext}}(t,x) = \sum_{\alpha} \lambda_{\alpha} \left( \mathbf{A}^{(\alpha)}(\omega_{\alpha}, x) e^{i\omega_{\alpha}t} + \mathbf{A}^{(\alpha)}(-\omega_{\alpha}, x) e^{-i\omega_{\alpha}t} \right)$$

These perturbations are \*\*monochromatic\*\* and ensure that the external potentials remain real by including both positive and negative frequency components.

### Step 3: Taking the First Derivative with Respect to $\lambda_{\alpha}$

To obtain the first-order response of the observable  $f_{\lambda}(t)$  to the perturbation  $\lambda_{\alpha}$ , we differentiate the Taylor expansion with respect to  $\lambda_{\alpha}$  and evaluate at  $\lambda = 0$ :

$$\left. \frac{\partial}{\partial \lambda_{\alpha}} f_{\lambda}(t) \right|_{\lambda=0} = f^{(\alpha)}(t)$$

However, due to the form of the external perturbations in Equations (11) and (12), the response  $f^{(\alpha)}(t)$  inherits the time dependence from the perturbations. Specifically, each perturbation  $\lambda_{\alpha}$  introduces oscillations at frequencies  $\omega_{\alpha}$  and  $-\omega_{\alpha}$ . Therefore, the first-order response can be expressed as a sum of contributions from these frequencies:

$$f^{(\alpha)}(t) = f^{(\alpha)}(\omega_{\alpha})e^{i\omega_{\alpha}t} + f^{(\alpha)}(-\omega_{\alpha})e^{-i\omega_{\alpha}t}$$

Substituting this into the derivative, we obtain:

$$\left. \frac{\partial}{\partial \lambda_{\alpha}} f_{\lambda}(t) \right|_{\lambda=0} = f^{(\alpha)}(\omega_{\alpha}) e^{i\omega_{\alpha}t} + f^{(\alpha)}(-\omega_{\alpha}) e^{-i\omega_{\alpha}t}$$

# 5.5 Taylor Expansion of Idempotency Constraint in Perturbation Strength $\lambda$ to first order

In Time-Dependent Density Functional Theory (TDDFT), the \*\*one-particle density matrix\*\*  $\gamma(t)$  satisfies the \*\*idempotency condition\*\*:

$$\gamma(t) = \gamma(t)\gamma(t) \tag{5.10}$$

When external perturbations characterized by coupling constants  $\lambda_{\alpha}$  are applied, the density matrix becomes dependent on these perturbations:

$$\gamma_{\lambda}(t) = \gamma(t; \{\lambda_{\alpha}\}) \tag{5.11}$$

Our goal is to expand  $\gamma_{\lambda}(t)$  in powers of  $\lambda_{\alpha}$  and derive the equation of motion for the first-order response  $\gamma^{(\alpha)}$ .

Step 1: Taylor Expansion of the Density Matrix and Hamiltonian Assume that the density matrix  $\gamma_{\lambda}(t)$  and the Hamiltonian  $H_{\lambda}(t)$  can be expanded in a Taylor series around  $\lambda = 0$ :

$$\gamma_{\lambda}(t) = \gamma^{(0)}(t) + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)}(t) + \sum_{\alpha \leq \beta} \lambda_{\alpha} \lambda_{\beta} \gamma^{(\alpha\beta)}(t) + \cdots$$
 (5.12)

$$H_{\lambda}(t) = H^{(0)} + \sum_{\alpha} \lambda_{\alpha} H^{(\alpha)}(t) + \sum_{\alpha < \beta} \lambda_{\alpha} \lambda_{\beta} H^{(\alpha\beta)}(t) + \cdots$$
 (5.13)

where:

•  $\gamma^{(0)}(t)$  is the \*\*zeroth-order\*\* (unperturbed) density matrix.

- $\gamma^{(\alpha)}(t)$  is the \*\*first-order\*\* response to the perturbation  $\lambda_{\alpha}$ .
- $\gamma^{(\alpha\beta)}(t)$  is the \*\*second-order\*\* response involving perturbations  $\lambda_{\alpha}$  and  $\lambda_{\beta}$ .
- Similarly for the Hamiltonian terms.

Step 2: Applying the Idempotency Condition Substitute the expansion from Equation (A) into both sides of the idempotency condition (8):

$$\gamma_{\lambda}(t) = \gamma_{\lambda}(t)\gamma_{\lambda}(t)$$

$$\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \sum_{\alpha \leq \beta} \lambda_{\alpha} \lambda_{\beta} \gamma^{(\alpha\beta)} + \dots = \left(\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \sum_{\alpha \leq \beta} \lambda_{\alpha} \lambda_{\beta} \gamma^{(\alpha\beta)} + \dots\right) \left(\gamma^{(0)} + \sum_{\gamma} \lambda_{\gamma} \gamma^{(\gamma)} + \sum_{\alpha \leq \beta} \lambda_{\alpha} \lambda_{\beta} \gamma^{(\alpha\beta)} + \dots\right) \left(\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \sum_{\alpha \leq \beta} \lambda_{\alpha} \lambda_{\beta} \gamma^{(\alpha\beta)} + \dots\right) \left(\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \sum_{\alpha \leq \beta} \lambda_{\alpha} \lambda_{\beta} \gamma^{(\alpha\beta)} + \dots\right) \left(\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \sum_{\alpha \leq \beta} \lambda_{\alpha} \lambda_{\beta} \gamma^{(\alpha\beta)} + \dots\right) \left(\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \sum_{\alpha \leq \beta} \lambda_{\alpha} \lambda_{\beta} \gamma^{(\alpha\beta)} + \dots\right) \left(\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \sum_{\alpha \leq \beta} \lambda_{\alpha} \lambda_{\beta} \gamma^{(\alpha\beta)} + \dots\right) \left(\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \sum_{\alpha \leq \beta} \lambda_{\alpha} \lambda_{\beta} \gamma^{(\alpha\beta)} + \dots\right) \left(\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \sum_{\alpha \leq \beta} \lambda_{\alpha} \lambda_{\beta} \gamma^{(\alpha\beta)} + \dots\right) \left(\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \sum_{\alpha \leq \beta} \lambda_{\alpha} \lambda_{\beta} \gamma^{(\alpha\beta)} + \dots\right) \left(\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \sum_{\alpha \leq \beta} \lambda_{\alpha} \lambda_{\beta} \gamma^{(\alpha\beta)} + \dots\right) \left(\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \dots\right) \left(\gamma^{(0)} + \dots\right) \left(\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \dots\right) \left(\gamma^{(0)} + \dots\right) \left(\gamma^$$

Expanding the right-hand side (RHS) and collecting terms up to second order:

$$\gamma_{\lambda}(t)\gamma_{\lambda}(t) = \gamma^{(0)}\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \left( \gamma^{(0)}\gamma^{(\alpha)} + \gamma^{(\alpha)}\gamma^{(0)} \right) + \sum_{\alpha \leq \beta} \lambda_{\alpha}\lambda_{\beta} \left( \gamma^{(0)}\gamma^{(\alpha\beta)} + \gamma^{(\alpha)}\gamma^{(\beta)} + \gamma^{(\beta)}\gamma^{(\alpha)} + \gamma^{(\alpha\beta)}\gamma^{(0)} \right) + \cdots$$

$$(5.14)$$

Step 3: Equating Terms Order by Order To satisfy the idempotency condition at each order of  $\lambda$ , equate the coefficients of corresponding powers of  $\lambda$  on both sides (Left-Hand Side and Right-Hand Side).

1. Zeroth Order (
$$\lambda^0$$
): 
$$\gamma^{(0)} = \gamma^{(0)} \gamma^{(0)}$$
 (5.15)

2. First Order (
$$\lambda^1$$
): 
$$\gamma^{(\alpha)} = \gamma^{(0)} \gamma^{(\alpha)} + \gamma^{(\alpha)} \gamma^{(0)}$$
 (5.16)

3. Second Order  $(\lambda^2)$ :

$$\gamma^{(\alpha\beta)} = \gamma^{(0)}\gamma^{(\alpha\beta)} + \gamma^{(\alpha)}\gamma^{(\beta)} + \gamma^{(\beta)}\gamma^{(\alpha)} + \gamma^{(\alpha\beta)}\gamma^{(0)}$$
(5.17)

These equations ensure that the density matrix remains idempotent at each order of perturbation.

Step 4: Expanding the Equation of Motion Up to First Order Consider the \*\*equation of motion\*\* for the density matrix:

$$i\frac{\partial}{\partial t}\gamma_{\lambda}(t) = [H_{\lambda}(t), \gamma_{\lambda}(t)]$$
 (5.18)

Substitute the expansions from Equations (A) and (B) into this equation:

$$i\frac{\partial}{\partial t}\left(\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \cdots\right) = \left(H^{(0)} + \sum_{\alpha} \lambda_{\alpha} H^{(\alpha)} + \cdots\right) \left(\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \cdots\right) - \left(\gamma^{(0)} + \sum_{\alpha} \lambda_{\alpha} \gamma^{(\alpha)} + \cdots\right) \left(H^{(0)} + \sum_{\alpha} \lambda_{\alpha} H^{(\alpha)} + \cdots\right) = \left[H^{(0)}, \gamma^{(0)}\right] + \sum_{\alpha} \lambda_{\alpha} \left(\left[H^{(0)}, \gamma^{(\alpha)}\right] + \left[H^{(\alpha)}, \gamma^{(0)}\right]\right) + \cdots$$

$$(5.19)$$

Step 5: Equating First-Order Terms Collecting first-order terms in  $\lambda_{\alpha}$ , we obtain:

$$i\frac{\partial}{\partial t}\gamma^{(\alpha)}(t) = [H^{(0)}, \gamma^{(\alpha)}(t)] + [H^{(\alpha)}(t), \gamma^{(0)}]$$

$$(5.20)$$

Step 6: Frequency Domain Representation Assume that the first-order response  $\gamma^{(\alpha)}(t)$  oscillates harmonically at frequency  $\omega_{\alpha}$ :

$$\gamma^{(\alpha)}(t) = \gamma^{(\alpha)}(\omega_{\alpha})e^{-i\omega_{\alpha}t} + \gamma^{(\alpha)}(-\omega_{\alpha})e^{i\omega_{\alpha}t}$$
(5.21)

Taking the time derivative:

$$i\frac{\partial}{\partial t}\gamma^{(\alpha)}(t) = \omega_{\alpha}\gamma^{(\alpha)}(\omega_{\alpha})e^{-i\omega_{\alpha}t} - \omega_{\alpha}\gamma^{(\alpha)}(-\omega_{\alpha})e^{i\omega_{\alpha}t}$$
(5.22)

**Step 7: Substituting into the Equation of Motion** Substitute Equations (G) and (H) into Equation (F):

$$\begin{split} \omega_{\alpha}\gamma^{(\alpha)}(\omega_{\alpha})e^{-i\omega_{\alpha}t} - \omega_{\alpha}\gamma^{(\alpha)}(-\omega_{\alpha})e^{i\omega_{\alpha}t} &= [H^{(0)},\gamma^{(\alpha)}(\omega_{\alpha})e^{-i\omega_{\alpha}t} + \gamma^{(\alpha)}(-\omega_{\alpha})e^{i\omega_{\alpha}t}] \\ &+ [H^{(\alpha)}(\omega_{\alpha})e^{-i\omega_{\alpha}t} + H^{(\alpha)}(-\omega_{\alpha})e^{i\omega_{\alpha}t},\gamma^{(0)}] \\ &= [H^{(0)},\gamma^{(\alpha)}(\omega_{\alpha})]e^{-i\omega_{\alpha}t} + [H^{(0)},\gamma^{(\alpha)}(-\omega_{\alpha})]e^{i\omega_{\alpha}t} \\ &+ [H^{(\alpha)}(\omega_{\alpha}),\gamma^{(0)}]e^{-i\omega_{\alpha}t} + [H^{(\alpha)}(-\omega_{\alpha}),\gamma^{(0)}]e^{i\omega_{\alpha}t} \\ &= \left( [H^{(0)},\gamma^{(\alpha)}(\omega_{\alpha})] + [H^{(\alpha)}(\omega_{\alpha}),\gamma^{(0)}] \right)e^{-i\omega_{\alpha}t} \\ &+ \left( [H^{(0)},\gamma^{(\alpha)}(-\omega_{\alpha})] + [H^{(\alpha)}(-\omega_{\alpha}),\gamma^{(0)}] \right)e^{i\omega_{\alpha}t} \end{split}$$

Step 8: Matching Frequency Components By matching the coefficients of  $e^{-i\omega_{\alpha}t}$  and  $e^{i\omega_{\alpha}t}$  on both sides, we obtain two separate equations:

$$\omega_{\alpha} \gamma^{(\alpha)}(\omega_{\alpha}) = [H^{(0)}, \gamma^{(\alpha)}(\omega_{\alpha})] + [H^{(\alpha)}(\omega_{\alpha}), \gamma^{(0)}]$$
(5.23)

$$-\omega_{\alpha}\gamma^{(\alpha)}(-\omega_{\alpha}) = [H^{(0)}, \gamma^{(\alpha)}(-\omega_{\alpha})] + [H^{(\alpha)}(-\omega_{\alpha}), \gamma^{(0)}]$$

$$(5.24)$$

Step 9: Rearranging to Obtain Equation (20) Focusing on Equation (H1), which corresponds to the positive frequency component:

$$\omega_{\alpha} \gamma^{(\alpha)}(\omega_{\alpha}) = [H^{(0)}, \gamma^{(\alpha)}(\omega_{\alpha})] + [H^{(\alpha)}(\omega_{\alpha}), \gamma^{(0)}]$$
(5.25)

This matches \*\*Equation (20)\*\* from your paper:

$$\omega_{\alpha} \gamma^{(\alpha)} = [H^{(0)}, \gamma^{(\alpha)}] + [H^{(\alpha)}, \gamma^{(0)}]$$
(5.26)

\*\*Note:\*\* In Equation (20), the frequency dependence is implicit. Here,  $\gamma^{(\alpha)}$  is understood to be associated with the frequency  $\omega_{\alpha}$ .

Summary of the Derivation 1. \*\*Expansion:\*\* The density matrix  $\gamma_{\lambda}(t)$  and Hamiltonian  $H_{\lambda}(t)$  are expanded in powers of the perturbation parameters  $\lambda_{\alpha}$ .

- 2. \*\*Idempotency Condition:\*\* The idempotency condition  $\gamma_{\lambda}(t) = \gamma_{\lambda}(t)\gamma_{\lambda}(t)$  is applied, leading to Equations (16)-(18).
- 3. \*\*Equation of Motion:\*\* The time-dependent equation of motion is expanded and matched order by order in  $\lambda$ .
- 4. \*\*Frequency Components:\*\* By assuming harmonic time dependence for the first-order response, the equation of motion yields a linear equation for  $\gamma^{(\alpha)}(\omega_{\alpha})$ , resulting in Equation (20).

**Implications** - \*\*Linear Equation for  $\gamma^{(\alpha)}$ :\*\* Equation (20) is a linear equation that can be solved to find the first-order response of the density matrix to the perturbation  $\lambda_{\alpha}$ .

- \*\*Frequency Dependence:\*\* The factor  $\omega_{\alpha}$  arises from the time derivative, indicating that the response oscillates at the frequency  $\omega_{\alpha}$ .
- \*\*Commutators:\*\* The commutators involve the unperturbed Hamiltonian  $H^{(0)}$  and the first-order perturbation  $H^{(\alpha)}$ , showing how these operators influence the response.

Conclusion Through this systematic expansion and application of the idempotency condition, we derive Equation (20), which governs the first-order response of the density matrix in the presence of external perturbations. This equation is fundamental for calculating how observables respond to external fields within the density matrix-based response theory framework.

If you have any further questions or need additional clarification on specific steps, feel free to ask!