

# Correlation function of a clipped Gaussian field

S. Plaszczynski

April 13, 2018

Let  $X$  follow a central normal distribution. We wish to avoid the unphysical region  $X < -1$  by applying to the random variable some local transform

$$Y = L(X) \quad (1)$$

Since  $X(\vec{r})$  actually represents a Gaussian field with a given covariance matrix ( $\Sigma$ ), a pair of values follows the (central) bivariate Gaussian distribution

$$(X_1, X_2) \sim \mathcal{N}_2(0, \Sigma), \quad (2)$$

where

$$\Sigma = \begin{pmatrix} \sigma_G^2 & \xi_G \\ \xi_G & \sigma_G^2 \end{pmatrix}. \quad (3)$$

The covariance of the transformed  $Y$  variable

$$\xi_Y = \mathbb{E}[y_1 y_2] - \mathbb{E}^2[y] \quad (4)$$

can be computed using probability conservation

$$\begin{aligned} \mathbb{E}[y_1 y_2] &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} y_1 y_2 f_Y(y_1, y_2) dy_1 dy_2 \\ &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} L(x_1) L(x_2) \mathcal{N}_2(x_1, x_2) dx_1 dx_2. \end{aligned} \quad (5)$$

Eq.(5) is rarely analytical (but for the log-normal case why it is often used).

Bel et al. (2016) introduces a very interesting method to compute this integral numerically. Let us first assume that  $\sigma = 1$ . Then one can use a formula from an obscure German mathematician named Mehler, which, quoted in our notations read

$$\mathcal{N}_2(x_1, x_2) = \mathcal{N}(x_1) \mathcal{N}(x_2) \sum_{n=0}^{\infty} H_n(x_1) H_n(x_2) \frac{\xi_G^n}{n!}, \quad (6)$$

$\mathcal{N}$  representing here the Normal distribution, and  $H_n$  the (probabilistic) Hermite polynomials which are orthogonal wrt to the Gaussian measure:

$$\int_{-\infty}^{+\infty} H_n(x) H_m(x) \mathcal{N}(x) dx = n! \delta_{nm} \quad (7)$$

Then if we decompose the local transform onto them

$$L(x) = \sum_{n=0}^{\infty} h_n H_n(x), \quad (8)$$

we obtain the series expansion

$$\xi_Y = \sum_{n=0}^{\infty} \alpha_n \xi_G^n, \quad (9)$$

where (here)  $\alpha_n = n!h_n^2$ .

The  $h_n$ 's are obtained from the orthogonality relation through:

$$h_n = \frac{1}{n!} \int_{-\infty}^{+\infty} L(x) H_n(x) \mathcal{N}(x) dx \quad (10)$$

Now if  $\sigma^2 \neq 1$  one rescales the variables  $x'_1 = x_1/\sigma, x'_2 = x_2/\sigma$  so that Eq.(5) can be written as

$$\mathbb{E}[y_1 y_2] = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} L(\sigma x'_1) L(\sigma x'_2) \mathcal{N}_2(x'_1, x'_2) dx'_1 dx'_2 \quad (11)$$

but now

$$(x'_1, x'_2) \sim \mathcal{N}_2 \left[ 0, \begin{pmatrix} 1 & \frac{\xi_G}{\sigma^2} \\ \frac{\xi_G}{\sigma^2} & 1 \end{pmatrix} \right]. \quad (12)$$

We can now safely apply Mehler's formula and the generalized expansion reads

$$\xi_Y(r) = \sum_{n=0}^{\infty} \alpha_n \left[ \frac{\xi_G(r)}{\xi_G(0)} \right]^n \quad (13)$$

where

$$\alpha_n = \frac{h_n^2}{n!} \quad (14)$$

$$h_n(\sigma) = \int_{-\infty}^{+\infty} L(\sigma x) H_n(x) \mathcal{N}(x) dx, \quad (15)$$

and note that in cosmology  $\frac{\xi_G(r)}{\xi_G(0)} \leq 1$ .

The  $n = 0$  coefficient is

$$h_0 = \int_{-\infty}^{+\infty} L(\sigma x) H_0(x) \mathcal{N}(x) dx = \frac{1}{\sigma} \int_{-\infty}^{+\infty} L(x) \mathcal{N}(x/\sigma) dx \quad (16)$$

since  $H_0 = 1$ . The 1D conservation of probability being

$$p(y) dy = \mathcal{N}(x) dx = \sigma \mathcal{N}(x/\sigma) dx, \quad (17)$$

and since  $L(x) = y$ , we find that

$$h_0 = \mathbb{E}[y]. \quad (18)$$

Recalling Eq.(4) we see that the squared mean value cancels with  $h_0^2$  so that Eq.(13) with  $\alpha_0 = 0$  is actually valid for *any local transform*. This makes sense since we introduced no stochasticity that could add some extra DC power.

## Log-normal case

$$L(x) = e^{x - \frac{\sigma^2}{2}} - 1 \quad (19)$$

Eq.(5) can be computed directly (using for instance the multi-normal moments generating function) which gives

$$\xi_Y = e^{\xi_G} - 1 \quad (20)$$

Let's see how the previous approach works:

$$\begin{aligned} h_n &= \frac{1}{n!} \int_{-\infty}^{+\infty} (e^{\sigma x - \frac{\sigma^2}{2}} - 1) \mathcal{N}(x) H_n(x) dx \\ &= \frac{1}{n!} \left[ \int_{-\infty}^{+\infty} \mathcal{N}(x - \sigma) H_n(x) dx - \int_{-\infty}^{+\infty} \mathcal{N}(x) H_n(x) dx \right] \\ &= \frac{1}{n!} [\sigma^n - \delta_{n0}]. \end{aligned} \quad (21)$$

where we used the fact that  $H_0(x) = 1$  and  $H_n(x + \sigma) = \sum_{k=0}^n \binom{n}{k} \sigma^k H_{n-k}(x)$ .

The series expansion then reads (Eq.(14))

$$\alpha_n = \left[ \frac{\sigma^n}{n!} \right]^2 \frac{n!}{\sigma^{2n}} - \delta_{n0} \quad (22)$$

$$\rightarrow \xi_Y = \sum_{n=1}^{\infty} \frac{\xi_G}{n!} = e^{\xi_G} - 1 \quad (23)$$

## Clipping

$$L(x) = \begin{cases} x & x \geq -1 \\ -1 & x < -1 \end{cases} \quad (24)$$

Eq.(5) is not analytical.

The Hermite coefficients are

$$h_n = \frac{1}{n!} \left[ \int_{-1/\sigma}^{+\infty} \sigma x \mathcal{N}(x) H_n(x) dx - \int_{1/\sigma}^{+\infty} \mathcal{N}(x) H_n(-x) dx \right] \quad (25)$$

which has no general analytical form. However only the very first coefficients (but 0) contribute significantly as shown on Fig 1 for  $\sigma = 1$ .

- $n = 1$  is the linear component that is majority,
- the non-linear coefficients  $n \geq 2$  drop rapidly.

This is observed for  $\sigma = 1$ . When  $\sigma$  is lower ( which is in practice always the case in LSS) the transformed variable should be even more linear (less clipping) so that keeping only a few terms is sufficient. The first 10 ones are given in Table 1

## References

Bel, J., Branchini, E., Di Porto, C., et al. 2016, A&A, 588, A51

Kibble, W. F. 1945, Mathematical Proceedings of the Cambridge Philosophical Society, 41

n	$\alpha_n$
1	$\frac{\sigma^2}{4} \left( \operatorname{erf} \left( 1 + \frac{1}{\sqrt{2}\sigma} \right) \right)^2$
2	$\frac{\sigma^2}{4\pi} e^{-\frac{1}{\sigma^2}}$
3	$\frac{1}{12\pi} e^{-\frac{1}{\sigma^2}}$
4	$\frac{(\sigma^2 - 1)^2}{48\pi\sigma^2} e^{-\frac{1}{\sigma^2}}$
5	$\frac{(1 - 3\sigma^2)^2}{240\pi\sigma^4} e^{-\frac{1}{\sigma^2}}$
6	$\frac{(3\sigma^4 - 6\sigma^2 + 1)^2}{1440\pi\sigma^6} e^{-\frac{1}{\sigma^2}}$
7	$\frac{(15\sigma^4 - 10\sigma^2 + 1)^2}{10080\pi\sigma^8} e^{-\frac{1}{\sigma^2}}$
8	$\frac{(1 - 15(\sigma^6 - 3\sigma^4 + \sigma^2))^2}{80640\pi\sigma^{10}} e^{-\frac{1}{\sigma^2}}$
9	$\frac{(1 - 21(5\sigma^6 - 5\sigma^4 + \sigma^2))^2}{725760\pi\sigma^{12}} e^{-\frac{1}{\sigma^2}}$
10	$\frac{(7(15\sigma^2(\sigma^4 - 4\sigma^2 + 2) - 4)\sigma^2 + 1)^2}{7257600\pi\sigma^{14}} e^{-\frac{1}{\sigma^2}}$

Table 1: Explicit coefficients of Eq.(13) expansion

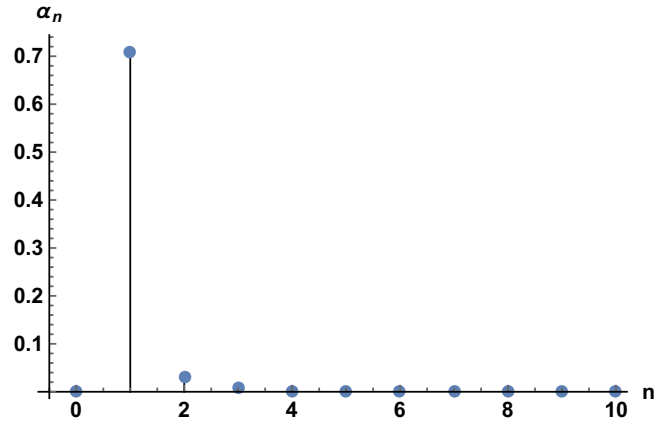


Figure 1: Coefficients entering the series expansion Eq.(13) for the -1 clipping transform ( $\sigma = 1$ )