

Radiation Heat Transfer

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Lecture 1

Introduction

Reference books

- Incropera, F. P., Dewitt, D. P., Bergman, T., & Lavine, A. (2002). Fundamentals of mass and heat transfer. *Australia: John Wiley and Sons*.
- Holman, J. P. (1986). Heat transfer, 1986. *Mc Gran–Hill Book Company, Soythern Methodist University*.
- Modest, M. F. (2013). Radiative heat transfer. Academic press.

Various Electromagnetic Radiations

$$c = \lambda \nu$$

$$E = h\nu$$

Planck's constant

$$h = 6.625 \times 10^{-34} \text{ J.s}$$

Using relativistic relation and quantum statistical thermodynamics

$$E_b = \sigma T^4$$

Stefan-Boltzmann's constant

$$\sigma = 5.67 \times 10^{-8} \text{ W/m}^2 \cdot \text{K}^4$$

Microwave



Small [microwave oven](#) on a kitchen counter

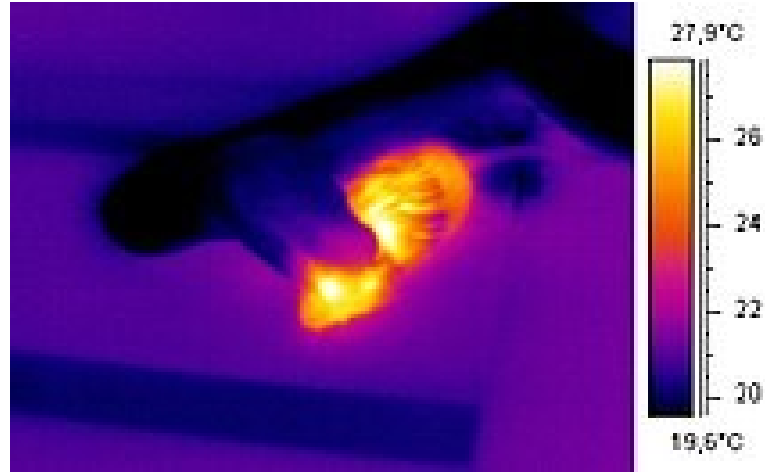


The [parabolic antenna](#) (lower curved surface) of an ASR-9 [airport surveillance radar](#) which radiates a narrow vertical fan-shaped beam of 2.7–2.9 GHz ([S band](#)) microwaves to locate aircraft in the airspace surrounding an airport



Some of the dish antennas of the [Atacama Large Millimeter Array](#) (ALMA) a radio telescope located in northern Chile. It receives microwaves in the [millimeter wave](#) range, 31 – 1000 GHz

Infrared radiation (IR)



Thermographic image of a snake eating a mouse



A comparison of a thermal image (top) and an ordinary photograph (bottom) shows that a trash bag is transparent but glass (the man's spectacles) is opaque in long-wavelength infrared.



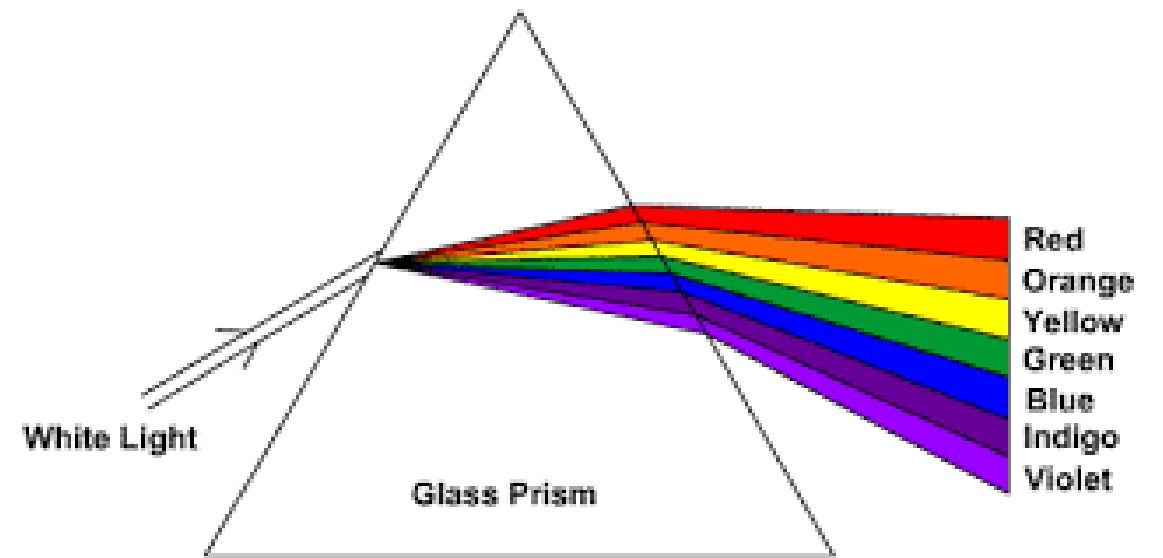
R Satellite picture taken 1315 Z on 15th October 2006. A [frontal](#) system can be seen in the [Gulf of Mexico](#) with embedded Cumulonimbus cloud.



Thermography helped to determine the temperature profile of the [Space Shuttle thermal protection system](#) during re-entry

Visible spectrum

<u>Color</u>	<u>Wavelength</u>	<u>Frequency</u>	<u>Photon energy</u>
<u>Violet</u>	380–450 nm	668–789 THz	2.75–3.26 eV
<u>Blue</u>	450–495 nm	606–668 THz	2.50–2.75 eV
<u>Green</u>	495–570 nm	526–606 THz	2.17–2.50 eV
<u>Yellow</u>	570–590 nm	508–526 THz	2.10–2.17 eV
<u>Orange</u>	590–620 nm	484–508 THz	2.00–2.10 eV
<u>Red</u>	620–750 nm	400–484 THz	1.65–2.00 eV

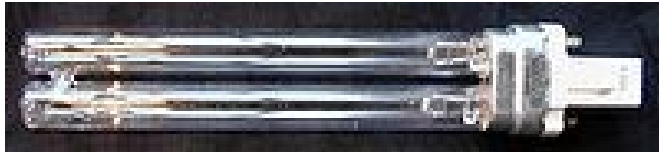


Rainbow

Ultraviolet (UV)



UV damaged [polypropylene](#) rope (left)
and new rope (right)

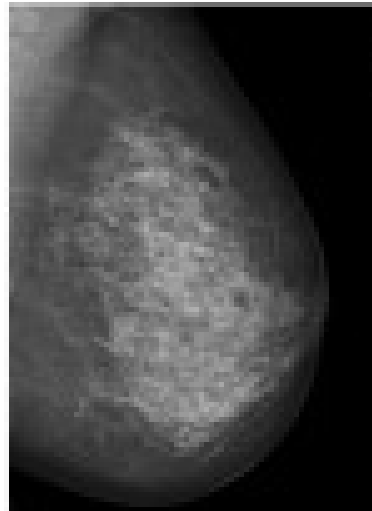
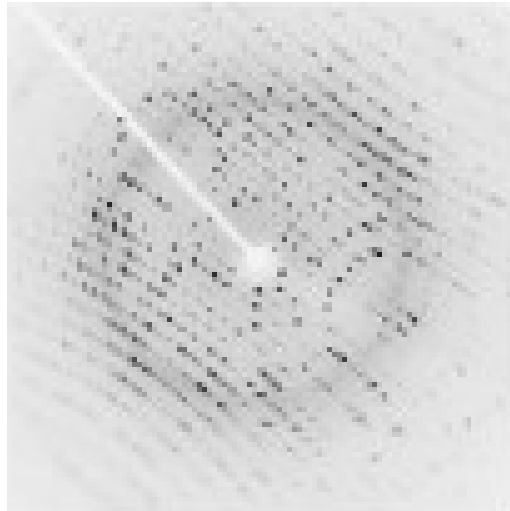
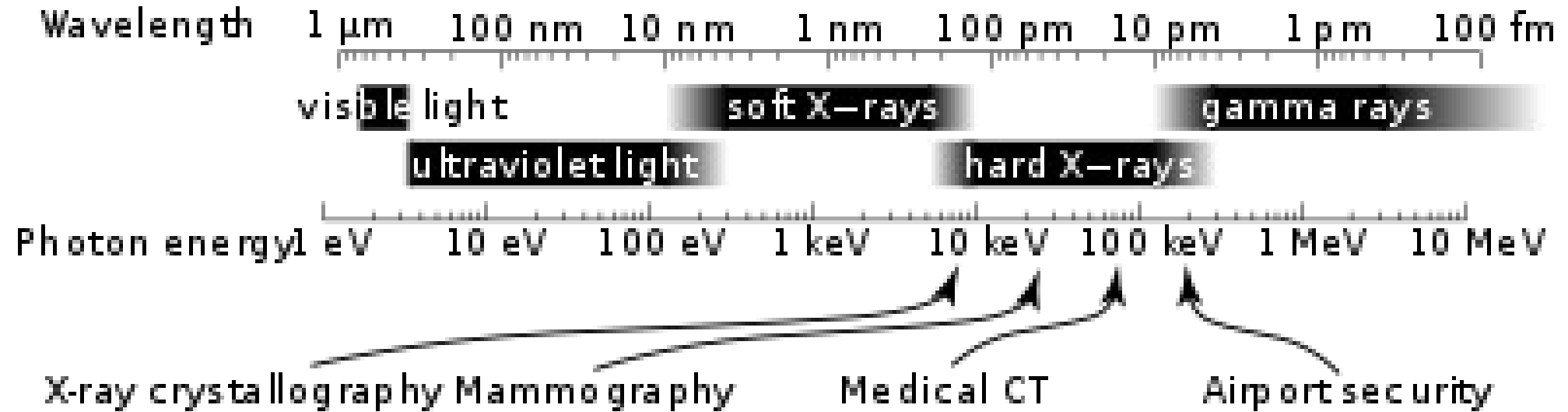


9-watt germicidal UV lamp, in compact fluorescent (CF)
form factor



UV radiation is also produced by [electric arcs](#).
[Arc welders](#) must wear [eye protection](#) and
cover their skin to prevent [photokeratitis](#) and
serious [sunburn](#)

X-Rays



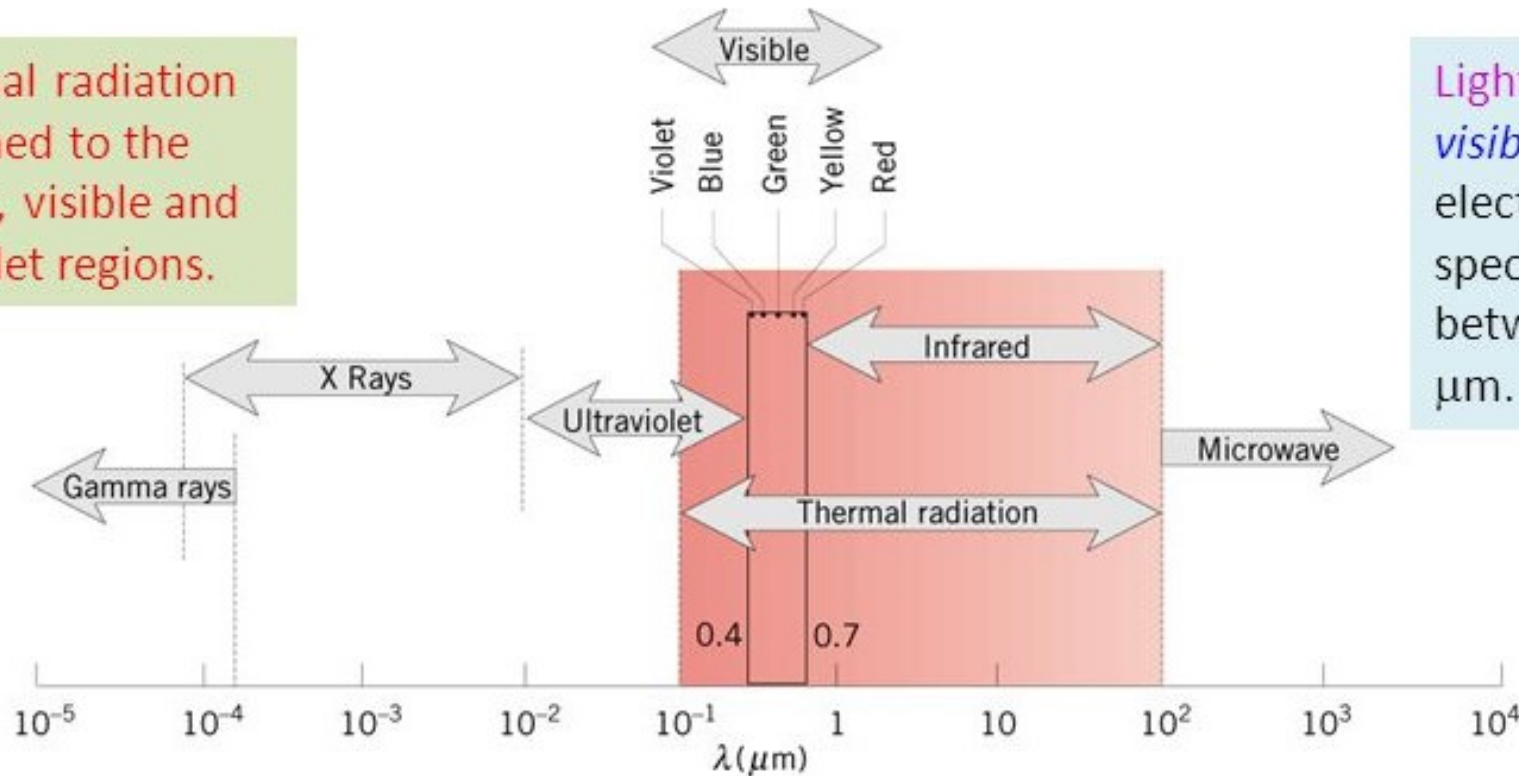
γ -Rays



Gamma rays are emitted during [nuclear fission](#) in nuclear explosions

The Electromagnetic Spectrum

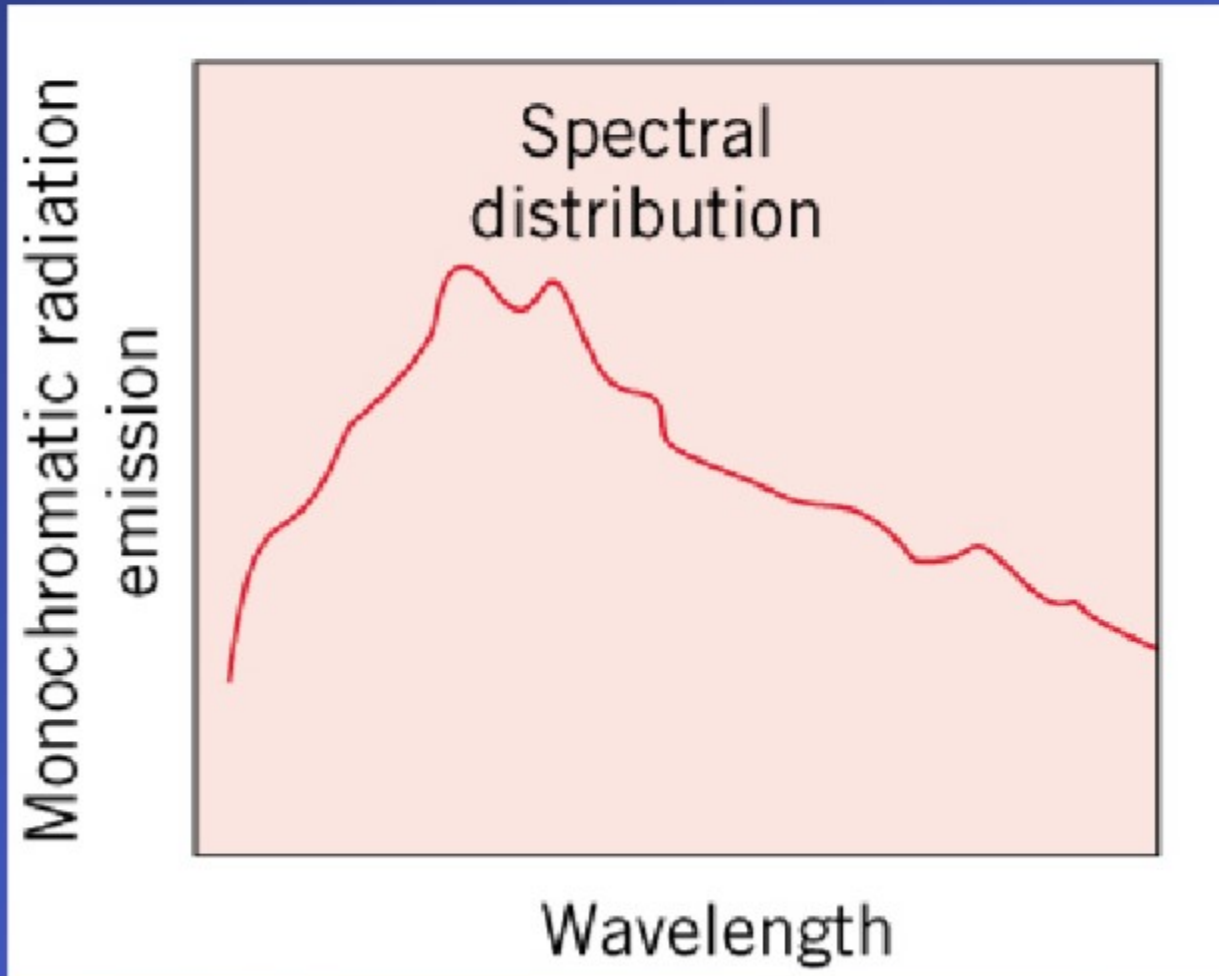
* Thermal radiation is confined to the infrared, visible and ultraviolet regions.



Light is simply the *visible* portion of the electromagnetic spectrum that lies between 0.4 and 0.7 μm .

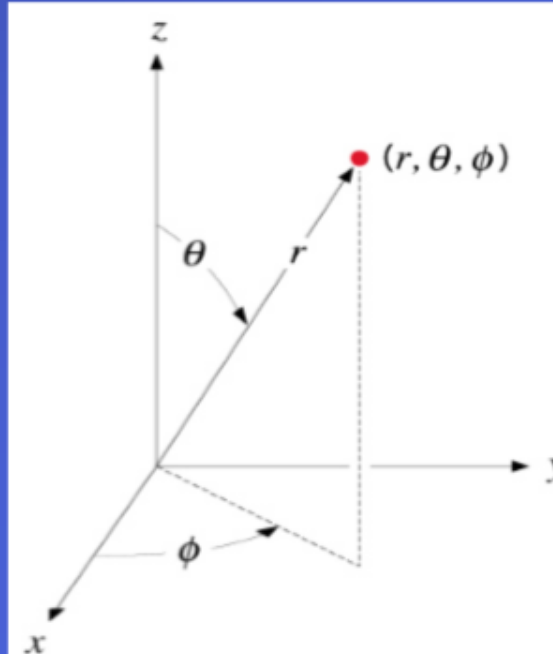
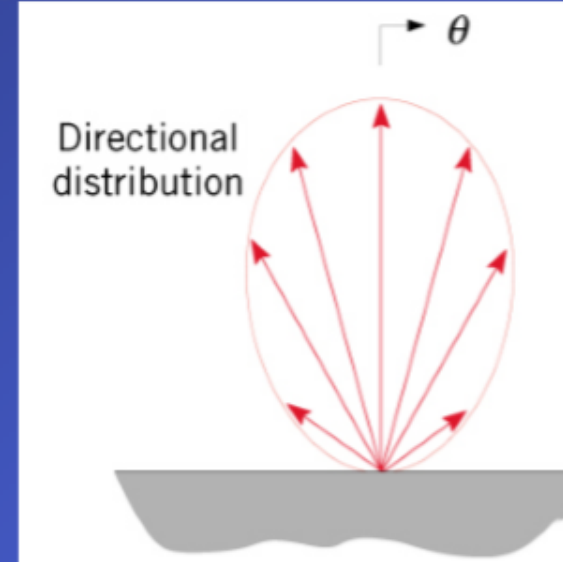
- Thermal radiation is confined to the **infrared**, **visible** and **ultraviolet** regions of the spectrum $0.1 < \lambda < 100 \mu\text{m}$

- The amount of radiation emitted by an opaque surface varies with wavelength, and we may speak of the **spectral distribution** over all wavelengths or of **monochromatic/spectral components** associated with particular wavelengths.



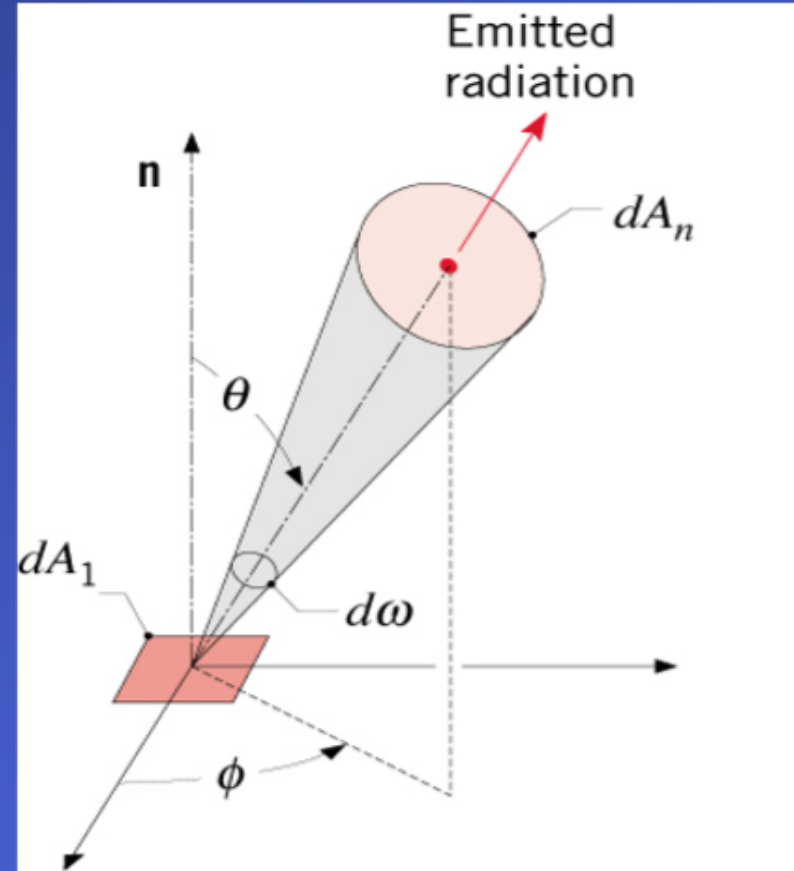
Directional Considerations and the Concept of Radiation Intensity

- Radiation emitted by a surface will be in all directions associated with a hypothetical hemisphere about the surface and is characterized by a **directional distribution**.
- Direction may be represented in a spherical coordinate system characterized by the zenith or **polar angle** θ and the **azimuthal angle** ϕ .

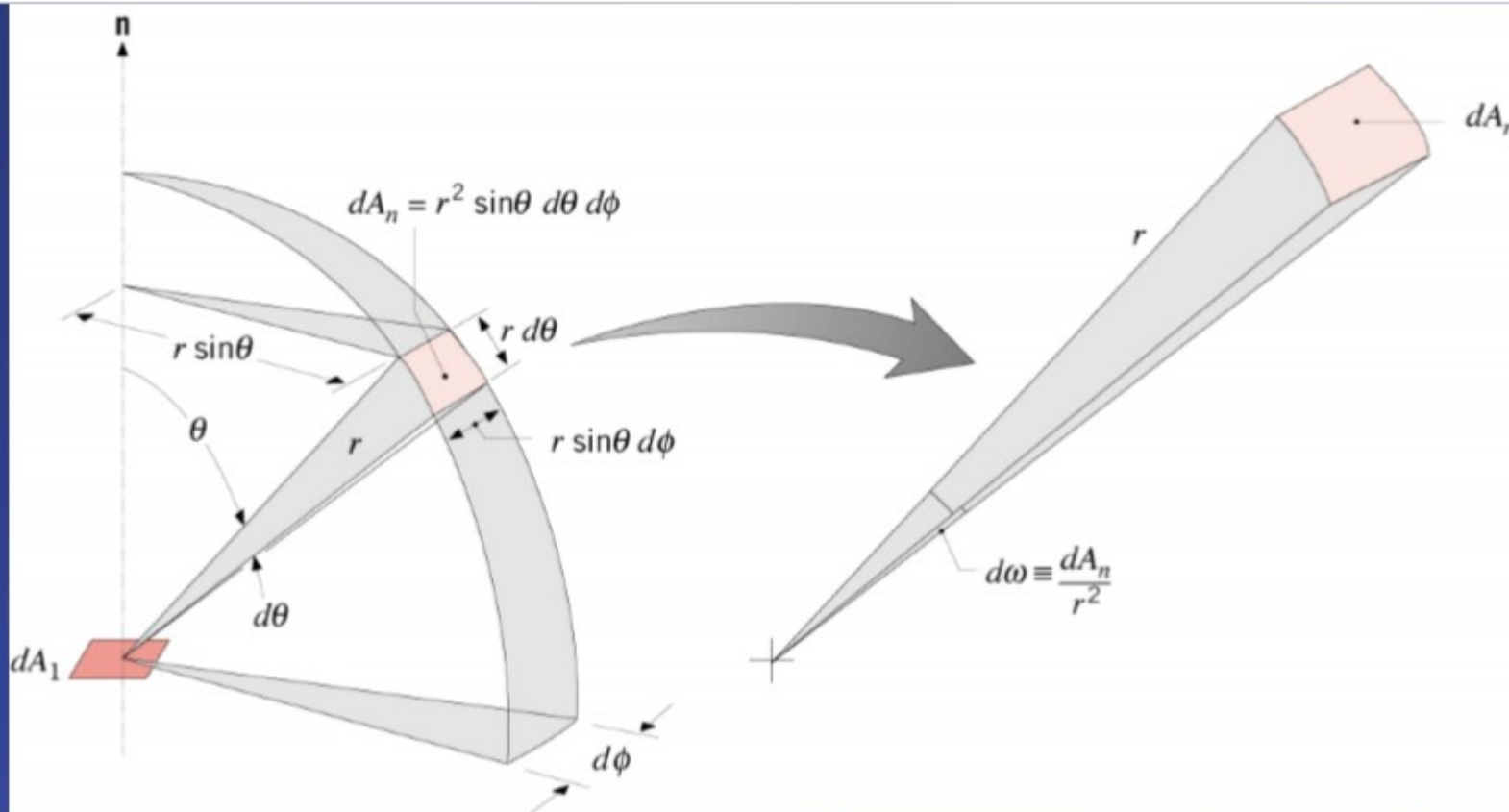


- The amount of radiation emitted from a surface, dA_1 , and propagating in a particular direction, θ, ϕ , is quantified in terms of a **differential solid angle** ($d\omega$) associated with the direction.

$$d\omega \equiv \frac{dA_n}{r^2}$$



$dA_n \rightarrow$ unit element of surface on a hypothetical sphere and normal to the θ, ϕ direction.



$$dA_n = r^2 \sin\theta d\theta d\phi$$

$$d\omega \equiv \frac{dA_n}{r^2} = \sin\theta d\theta d\phi$$

- The solid angle ω has units of **steradians (sr)**.
- The solid angle associated with a complete hemisphere is is

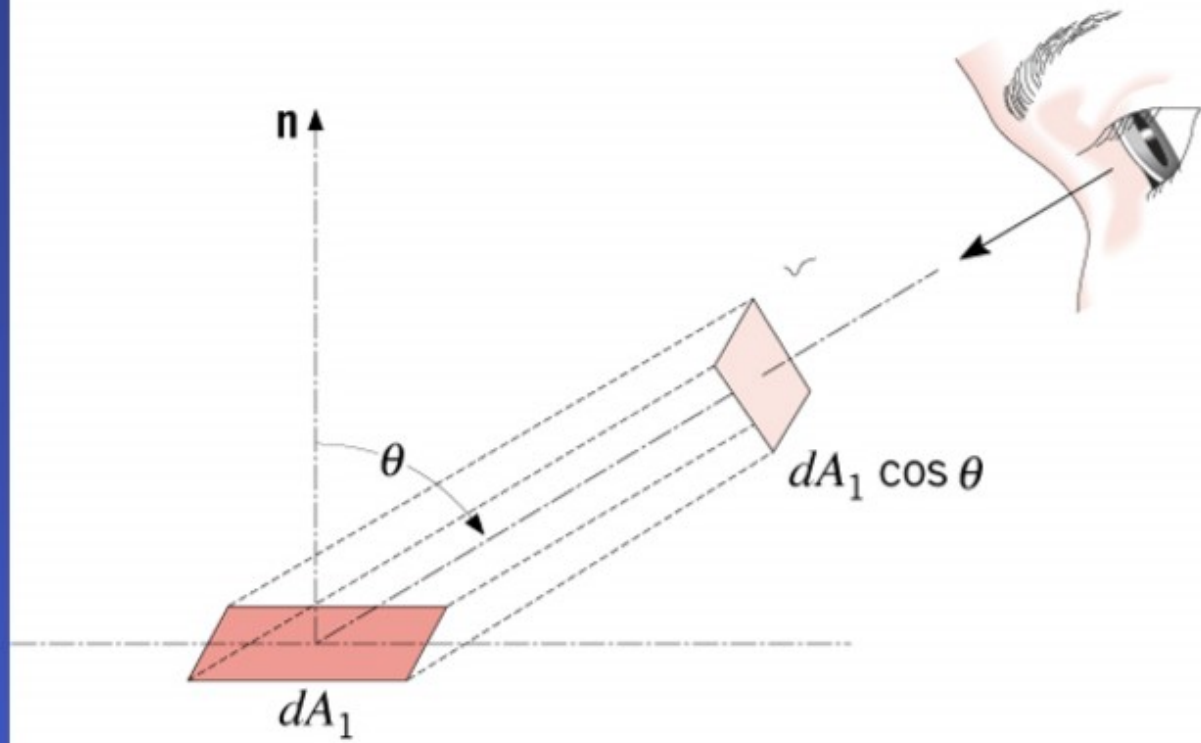
$$\omega_{\text{hemi}} = \int_0^{2\pi} \int_0^{\pi/2} \sin\theta d\theta d\phi = 2\pi \quad (\text{unit: sr})$$

- **Spectral Intensity:**

A quantity used to specify the radiant **heat flux** (W/m^2) **within a unit solid angle** about a prescribed direction ($\text{W/m}^2 \cdot \text{sr}$) and **within a unit wavelength interval** ($\text{W/m}^2 \cdot \text{sr} \cdot \mu\text{m}$). about a prescribed wavelength

- The **spectral intensity** $I_{\lambda,e}$ associated with emission from a dA_1 surface element in the solid angle $d\omega$ about θ, ϕ and the wavelength interval $d\lambda$ about λ is defined as:

$$I_{\lambda,e}(\lambda, \theta, \phi) \equiv \frac{dq}{(dA_1 \cos \theta) \cdot d\omega \cdot d\lambda}$$



Formal definition of $I_{\lambda,e}$

The rate at which radiant energy is emitted at the wavelength λ in the (θ, ϕ) direction, per unit area of the emitting surface normal to this direction, per unit solid angle about this direction, and per unit wavelength interval $d\lambda$ about λ .

Thank you