

# Lunar Meteoroid Ejecta Engineering Model

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## 1 Executive Summary

## 2 Lunar Regolith Properties

### 2.1 Porosity

The lunar regolith porosity is related to the amount of free space between individual grains. The greater the porosity, the more void space is present. Table 3.4.2.3.4-1 of the DSNE gives values of the porosity as a function of depth down to 60 cm derived from Apollo core measurements (copied from Table 9.5 of the Lunar Sourcebook) and shown here in Table 1.

Table 1: Porosity for various depths.

Depth Range (cm)	Average Porosity, $n$ (%)
0 – 15	$52 \pm 2$
0 – 30	$49 \pm 2$
30 – 60	$44 \pm 2$
0 – 60	$46 \pm 2$

### 2.2 Density

The bulk density ( $\rho$ ) of the lunar regolith is defined as the mass of material in a given volume, which relates the particle density ( $\rho_p$ ) and porosity ( $n$ ) to the bulk density as (see Section 3.4.2.3.1 of the DSNE or Chapter 9 of the Lunar Sourcebook)

$$\rho = \rho_p(1 - n). \quad (2.1)$$

The DSNE suggests using  $\rho_p = 3.1 \text{ g/cm}^3$  for the average particle density over the entire Moon. Otherwise, the typical highlands particle density is  $\rho_p = 2.75 \pm 0.1 \text{ g/cm}^3$  whereas the typical mare particle density is  $\rho_p = 3.35 \pm 0.1 \text{ g/cm}^3$ .

The bulk density<sup>1</sup> as a function of depth, fit to Apollo data, is given by

$$\rho(z) = 1.92 \frac{z + 12.2}{z + 18}, \quad (2.2)$$

where  $z$  is the depth in cm and  $\rho$  is in units of  $\text{g/cm}^3$ . At the surface ( $z = 0$ ), the density is  $1.30 \text{ g/cm}^3$ , and increases to  $1.92 \text{ g/cm}^3$  for large depths. This expression is fairly reasonable down to 3 m (the limit reached by Apollo drill core samples). In order to get an up-to-depth average of the bulk density, take

$$\rho_{avg,depth}(z) = \frac{1}{z} \int_0^z dz' \rho(z'), \quad (2.3)$$

<sup>1</sup>Follows the average particle density of  $3.1 \text{ g/cm}^3$  for all depths with a porosity depth dependence following Table 1, see the *porosity of lunar soil* paragraph on page 492 in the Lunar Sourcebook.

which gives (compare with the equation for  $d_m$  on page 494 of the Lunar Sourcebook)

$$\rho_{avg,depth}(z) = 1.92 \left[ 1 - \frac{5.8 \ln \left( \frac{z+18}{18} \right)}{z} \right]. \quad (2.4)$$

For example, the average bulk density of the regolith with a depth range of 0 – 60 cm would be  $\rho_{avg,depth}(60) = 1.65 \text{ g/cm}^3$ .

For a higher-fidelity estimate of the average bulk density sampled by the crater, a volume-average can be used instead of a depth-average, given by

$$\rho_{avg,volume}(z) = \frac{\int dV \rho(z')}{\int dV}. \quad (2.5)$$

Expanding the integral in a cylindrical coordinate system, Equation (2.5) becomes

$$\begin{aligned} \rho_{avg,volume}(z) &= \frac{\int_0^z \int_0^R \sqrt{1-z'^2/z^2} \int_0^{2\pi} d\phi r dr dz' \rho(z')}{\int_0^z \int_0^R \sqrt{1-z'^2/z^2} \int_0^{2\pi} d\phi r dr dz'} \\ &= \frac{1.92}{4z^3} \left[ z(6ab - 6b^2 - 3az + 3bz + 4z^2) + 6(a-b)(b^2 - z^2) \ln \left( \frac{b}{z+b} \right) \right], \end{aligned} \quad (2.6)$$

$$(2.7)$$

for the volume-averaged density in  $\text{g/cm}^3$  with  $z$  in cm, where  $a = 12.2$  and  $b = 18$ . Following the example from earlier, the average bulk density of the regolith with a depth range of 0 – 60 cm would be  $\rho_{avg,volume}(60) = 1.60 \text{ g/cm}^3$ , which is  $\sim 3\%$  less than  $\rho_{avg,depth}(60) = 1.65 \text{ g/cm}^3$ . The expression given in Equation (2.7) is useful for computing the ejected mass from a crater<sup>2</sup>, given a crater depth  $z$ .

The expressions for the regolith density at a certain depth  $z$ , weighted by depth, and weighted by crater volume are given by Equations (2.2), (2.4), and (2.7), respectively, are compared in Figure 1. The crater volume is approximated as a half-ellipsoid with two of the dimensions scaled by the crater radius  $R$  and one dimension scaled by the crater depth  $z$ , sliced such that the half-ellipsoid is symmetric about the surface normal. For a given crater, more of the volume is near the surface so that more weight is given by bulk densities that originate near the surface. In contrast, the depth-averaged bulk density takes the bulk density at each depth equally. This results in the volume-averaged bulk density to be slightly less than the depth-averaged bulk density, as shown in Figure 1.

<sup>2</sup>In an iterative fashion, since the crater radius depends on the regolith density.

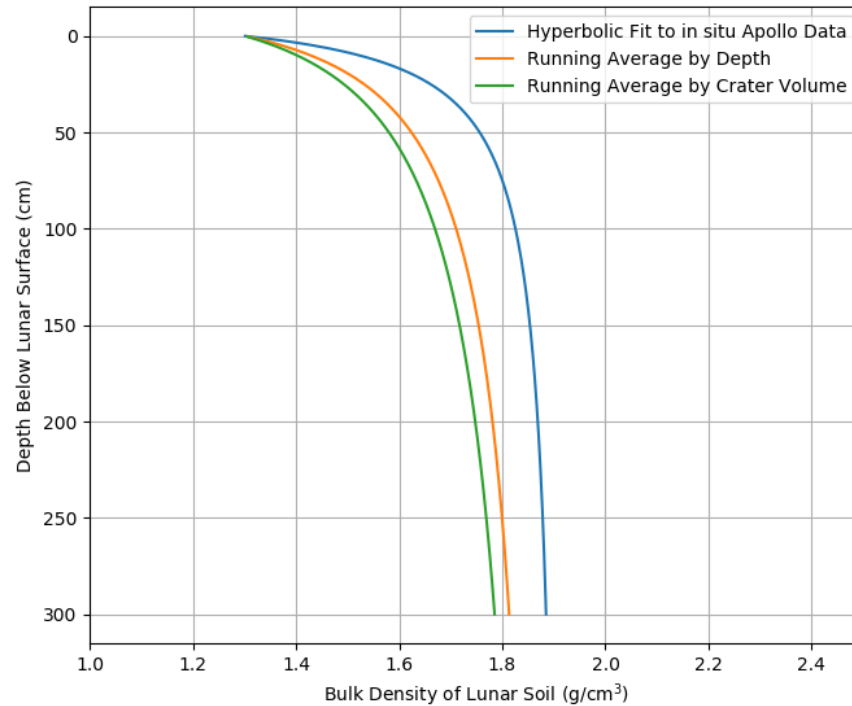


Figure 1: A comparison of the regolith bulk density for a certain depth depth (blue), the depth-averaged bulk density (orange), and the volume-averaged bulk density (green). See also, Figure 9.16 of the Lunar Sourcebook.

## **2.3 Strength**

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## **5.3 Maximum Ejected Particle Mass**

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## References