

Lunar Meteoroid Ejecta Engineering Model

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1 Executive Summary

2 Lunar Regolith Properties

2.1 Porosity

The lunar regolith porosity is related to the amount of free space between individual grains. The greater the porosity, the more void space is present. Table 3.4.2.3.4-1 of the DSNE gives values of the porosity as a function of depth down to 60 cm derived from Apollo core measurements (copied from Table 9.5 of the Lunar Sourcebook) and shown here in Table 1.

Table 1: Porosity for various depths.

Depth Range (cm)	Average Porosity, n (%)
0 – 15	52 ± 2
0 – 30	49 ± 2
30 – 60	44 ± 2
0 – 60	46 ± 2

2.2 Density

The bulk density (ρ) of the lunar regolith is defined as the mass of material in a given volume, which relates the particle density (ρ_p) and porosity (n) to the bulk density as (see Section 3.4.2.3.1 of the DSNE or Chapter 9 of the Lunar Sourcebook)

$$\rho = \rho_p(1 - n). \quad (2.1)$$

The DSNE suggests using $\rho_p = 3.1 \text{ g/cm}^3$ for the average particle density over the entire Moon. Otherwise, the typical highlands particle density is $\rho_p = 2.75 \pm 0.1 \text{ g/cm}^3$ whereas the typical mare particle density is $\rho_p = 3.35 \pm 0.1 \text{ g/cm}^3$.

The bulk density¹ as a function of depth, fit to Apollo data, is given by

$$\rho(z) = 1.92 \frac{z + 12.2}{z + 18}, \quad (2.2)$$

where z is the depth in cm and ρ is in units of g/cm^3 . At the surface ($z = 0$), the density is 1.30 g/cm^3 , and increases to 1.92 g/cm^3 for large depths. This expression is fairly reasonable down to 3 m (the limit reached by Apollo drill core samples). In order to get an up-to-depth average of the bulk density, take

$$\rho_{avg}(z) = \frac{1}{z} \int_0^z dz' \rho(z'), \quad (2.3)$$

¹ Found to follow the average particle density of 3.1 g/cm^3 for all depths with a porosity depth dependence following Table 1.

which gives

$$\rho_{avg}(z) = 1.92 \left[1 + \frac{5.8 \ln \left(\frac{18}{z+18} \right)}{z} \right]. \quad (2.4)$$

For example, the average bulk density of the regolith with a depth range of 0 – 60 cm would be $\rho_{avg}(60) = 1.65 \text{ g/cm}^3$. This expression is useful for computing the ejected mass from a crater, given a crater depth z .

2.3 Strength

2.4 Particle Size Distribution

2.5 Scaling Law Properties

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3.1 Space-Time Dependence of Environment

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5.3 Maximum Ejected Particle Mass

5.4 Mass Ejected from Crater

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