

Retarded Potentials and Fields

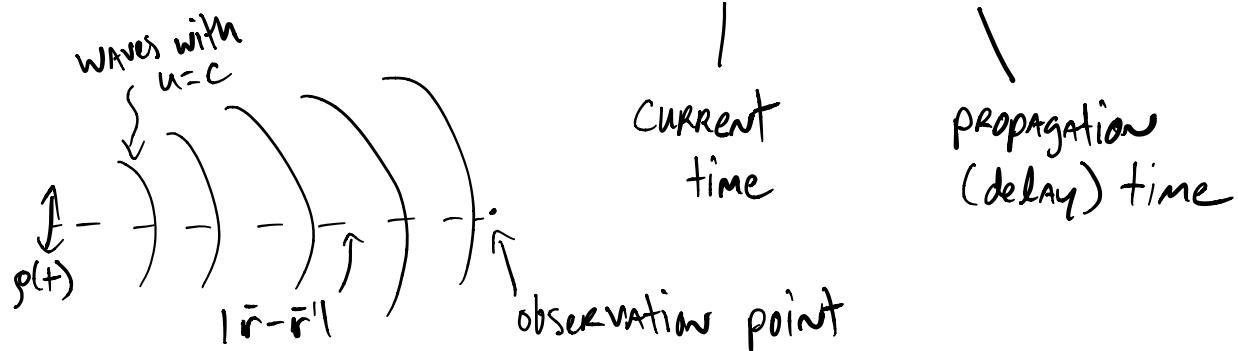
For static fields, our potentials are

$$\Phi(\bar{r}) = \int \frac{\rho(\bar{r}')}{|\bar{r} - \bar{r}'|} d\bar{r}'$$

$$\bar{A}(\bar{r}) = \frac{1}{c} \int \frac{\bar{s}(\bar{r}')}{|\bar{r} - \bar{r}'|} d\bar{r}'$$

If \bar{s} and ρ are functions at time,
then we have to account for the
finite propagation speed, c

$$\text{Retarded time} = t - \frac{|\bar{r} - \bar{r}'|}{c}$$



The retarded potentials are

$$\Phi(\bar{r}, t) = \int \frac{\rho(\bar{r}', t - |\bar{r} - \bar{r}'|/c)}{|\bar{r} - \bar{r}'|} d\bar{r}'$$

$$\bar{A}(\bar{r}, t) = \frac{1}{c} \int \frac{\bar{\jmath}(\bar{r}', t - |\bar{r} - \bar{r}'|/c)}{|\bar{r} - \bar{r}'|} d\bar{r}'$$

Substituting $R \equiv |\bar{r} - \bar{r}'|$

$$\rho(\bar{r}', t - R/c) \equiv [\rho(\bar{r}')]$$

↳ brackets mean evaluated
at retarded time

$$\Rightarrow \Phi(\bar{r}, t) = \int \frac{[\rho(\bar{r}')] }{R} d\bar{r}'$$

$$\bar{A}(\bar{r}, t) = \frac{1}{c} \int \frac{[\bar{\jmath}(\bar{r}')] }{R} d\bar{r}'$$

Retarded fields

From $\bar{E} = -\bar{\nabla}\bar{\Phi} - \frac{1}{c} \frac{\partial \bar{A}}{\partial t}$

we can calculate \bar{E} . However,

$$\bar{\nabla}\bar{\Phi} = \bar{\nabla} \left\{ \int \frac{g(\bar{r}')}{R} d\bar{r}' \right\} = \int \bar{\nabla} \frac{g(\bar{r}', + - \frac{R}{c})}{R} d\bar{r}'$$

acts on \bar{r} (unprimed, position of interest)

$R = |\bar{r} - \bar{r}'|$ is a function of \bar{r}

$$\bar{\nabla} \frac{1}{R} = -\frac{\hat{e}_R}{R^2}$$

$$\bar{\nabla}[g] = \left[\frac{\partial g}{\partial t} \right] \bar{\nabla} \left(+ - \frac{|\bar{r} - \bar{r}'|}{c} \right) = -\frac{1}{c} \left[\frac{\partial g}{\partial t} \right] \hat{e}_R$$

$$\Rightarrow \bar{E}(\bar{r}, t) = \int_v \left(\frac{[\rho] \hat{e}_R}{R^2} + \frac{[\partial \rho / \partial t] \hat{e}_R}{cR} - \frac{[\partial \bar{\rho} / \partial t]}{c^2 R} \right) dv'$$

Generalized Coulomb - Faraday Law

Similarly, from $\vec{B} = \vec{\nabla} \times \vec{A}$

$$\bar{B}(\bar{r}, t) = \int \left(\frac{[\bar{s}] \times \hat{e}_R}{cR^2} + \frac{[\partial \bar{s}/\partial t] \times \hat{e}_R}{c^2 R} \right) d\nu'$$

Generalized Biot-Savart Law

Instead of analyzing distributions ρ and $\overline{\rho}$, we can focus on a point charge

$$\Phi(\bar{r}, t) = e^{\int_{-\infty}^{\infty} \frac{\delta(t' - t + |\bar{r} - \bar{r}'|/c)}{|\bar{r} - \bar{r}'|} dt'}$$

charge of e^-

location is a delta
fxn in time

Defining

$$\bar{R} \equiv R \hat{e}_R$$

$$\bar{\beta} \equiv \frac{\bar{u}}{c} \quad \text{--- } \bar{u} = \text{particle's velocity}$$

$$\bar{K} \equiv 1 - \frac{\bar{\beta} \cdot \bar{R}}{R}$$

We can rewrite Φ as (see book for details)

$$\Phi(\bar{r}, t) = \frac{e}{[KR]}$$

from $\bar{J} = \rho_e \bar{u}$ we obtain

$$\bar{A}(\bar{r}, t) = e \left[\frac{\bar{\beta}}{KR} \right]$$

Liénard-Wiechert potentials

From the potentials, "After some heroic Algebra"
 we obtain the Liénard-Wiechert fields

$$\bar{E} = e \left[\frac{(\hat{n} - \bar{\beta})(1 - \beta^2)}{K^3 R^2} + \frac{\hat{n} \times ((\hat{n} - \bar{\beta}) \times \bar{a})}{c^2 K^3 R} \right]$$

$$\bar{B} = [\hat{n}] \times \bar{E}$$

where $\hat{n} = \hat{r}_R = \frac{\bar{R}}{R}$

\bar{a} = particle's acceleration

Radiation From Accelerated Charged Particle at low velocities

For $u \ll c \Rightarrow \beta \ll 1, K \approx 1$

the acceleration fields become

$$\bar{E}_a = \frac{e}{c^2 R^3} \left\{ (\bar{R} \cdot \bar{a}) \bar{R} - R^2 \bar{a} \right\}$$

$$\bar{B}_a = \hat{n} \times \bar{E}_a$$

We have dropped the [retarded time] because we're now focused on radiation leaving rather than approaching.

The Poynting Vector is

$$\bar{S}_a = \frac{c}{4\pi} (\bar{E}_a \times \bar{B}_a) = \frac{c}{4\pi} \left\{ \frac{\bar{E}_a \times (\bar{R} \times \bar{E}_a)}{R} \right\}$$

$$= \frac{c}{4\pi} \left\{ \frac{E_a^2 \bar{R} - (\bar{E}_a \cdot \bar{R}) \bar{E}_a}{R} \right\}$$

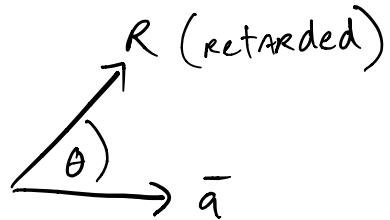
$$\text{But, } \bar{E}_a \perp \bar{R} \Rightarrow \bar{E}_a \cdot \bar{R} = 0$$

$$\Rightarrow \bar{s}_a = \frac{c}{4\pi} E_a^2 \frac{\bar{R}}{R} = \frac{c}{4\pi} E_a^2 \hat{n}$$

And

$$E_a^2 = \bar{E}_a \cdot \bar{E}_a = \frac{e^2}{c^4 R^4} \left\{ R^2 \dot{a}^2 - (\bar{R} \cdot \bar{a})^2 \right\}$$

Defining



$$\Rightarrow E_a^2 = \frac{e^2 \dot{a}^2}{c^4 R^2} (1 - \cos^2 \theta) = \frac{e^2 \dot{a}^2}{c^4 R^2} \sin^2 \theta$$

$$\Rightarrow \bar{s}_a = \frac{e^2 \dot{a}^2 \sin^2 \theta}{4\pi c^3 R^2} \hat{n}$$

$$S_a = \frac{\text{POWER}}{\Delta A} = \frac{\text{POWER}}{R^2 \Delta \Omega}$$

↳ unit area ↳ solid angle

$$\Rightarrow (\bar{S}_a \cdot \hat{n}) R^2 = \frac{dP}{d\Omega} = \frac{e^2 a^2}{4\pi c^3} \sin^2 \theta$$

Radiated power per unit solid angle

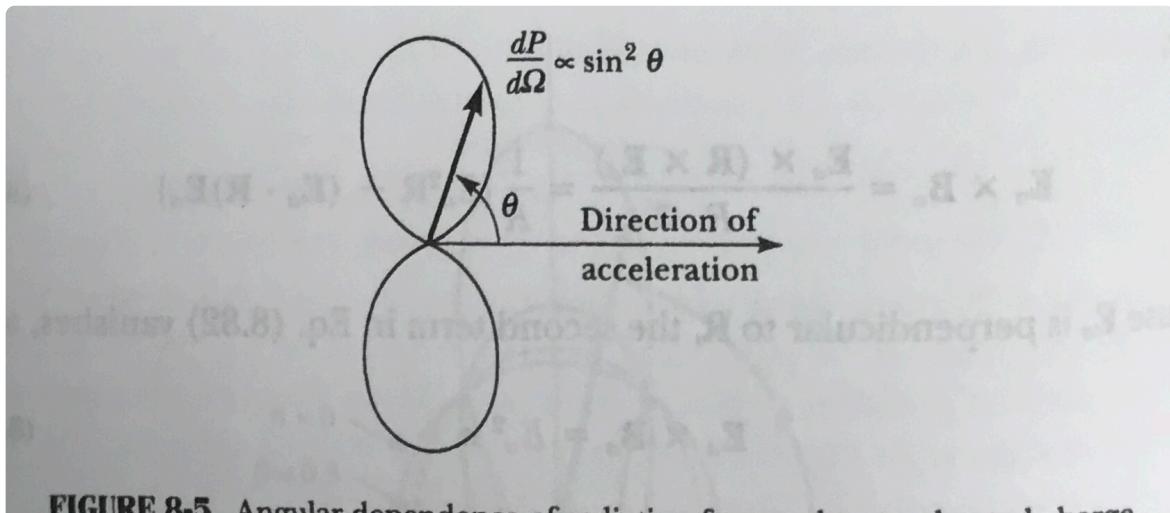


FIGURE 8-5. Angular dependence of radiation from a slow accelerated charge.

3-D extension = doughnut

[H+M]

→ This radiation pattern is characteristic
of dipole radiation And simple antennas

The total radiated power is

$$P = \int_{4\pi} \frac{dP}{d\Omega} d\Omega = \frac{2e^2 a^2}{3c^3}$$

Larmor Formulas

Collinear Velocity And Acceleration Radiation

For $\bar{u} \parallel \bar{a} \Rightarrow \bar{\beta} \times \bar{a} = 0$

$$\Rightarrow \bar{E}_a = \frac{e}{c^2 K^3 R^3} \left\{ (\bar{R} \cdot \bar{a}) \bar{R} - R^2 \bar{a} \right\}$$

↳ now has retardation factor

$$K = 1 - \hat{n} \cdot \bar{\beta}$$

$$\Rightarrow \bar{S}_a = \frac{c}{4\pi} E_a^2 \hat{n} = \frac{e^2 a^2 \sin^2 \theta}{4\pi c^3 K^6 R^2} \hat{n}$$

The incremental amount of energy lost by the particle and radiated to solid Angle θ measured during time dt :

$$-dW(\theta) = (\bar{S}_a \cdot \hat{n}) R^2 dt$$

$$\Rightarrow \frac{dP}{d\Omega} = - \frac{dW(\theta)}{dt'} = (\bar{S}_a \cdot \hat{n}) R^2 \frac{dt}{dt'}$$

↑
Radiated during t' to be measured at t

From $t' = t - R(t')/c$

$$\Rightarrow \frac{dt}{dt'} = 1 - \frac{\bar{R} \cdot \bar{R}}{R} = k = 1 - \beta \cos \theta$$

$$\Rightarrow \boxed{\frac{dP}{d\Omega} = \frac{e^2 a^2 \sin^2 \theta}{4\pi c^3 (1 - \beta \cos \theta)^5}}$$

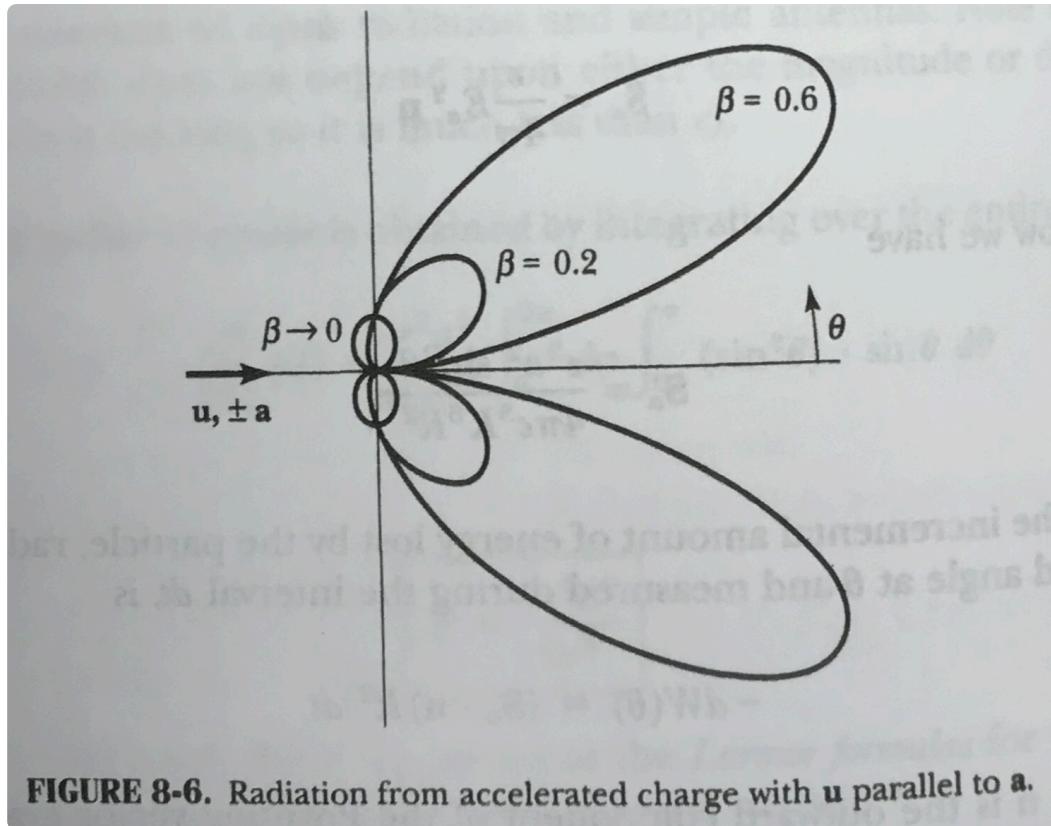
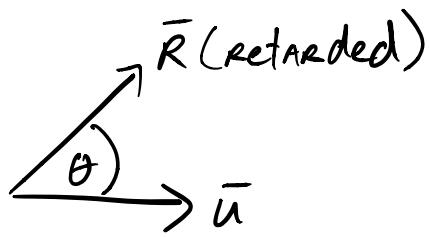


FIGURE 8-6. Radiation from accelerated charge with \mathbf{u} parallel to \mathbf{a} .

- For $\beta \ll 1$, we get the Larmor formula
- As $\beta \rightarrow 1$, the radiation intensity increases in the forward direction
 - ↳ bremsstrahlung (breaking radiation)
 - ↳ produce x-rays by "breaking" fast e^-